Conceptual Design of the Hannover 10m Interferometer for sub SQL measurement

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for Hannover-10m team

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Hannover 10m



- Multiple experiments
- Sub-SQL measurement



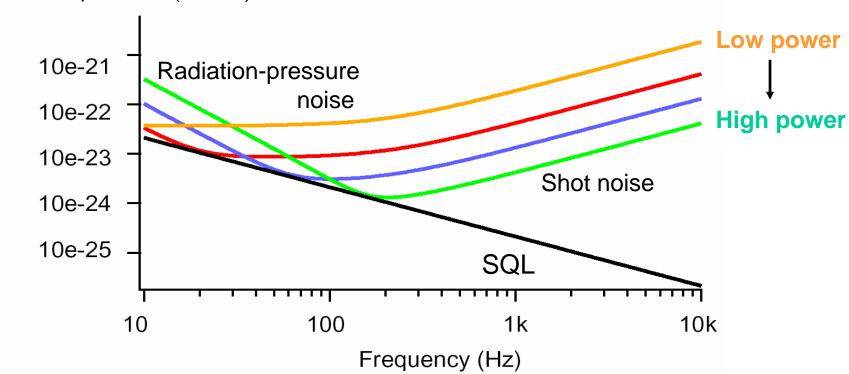
vacuum chamber φ=3m, h=3.2m

Today's talks for the 10m

- 1. Noise analysis and IFO design (K.Somiya)
 - SQL
 - Reduction of thermal noise
 - Parameter optimization
- 2. Discussion of various configurations (F.Kawazoe)
- 3. Overview and status of the experiment (S.Gossler)

Standard Quantum Limit (SQL)

Noise Spectrum (1/rtHz)

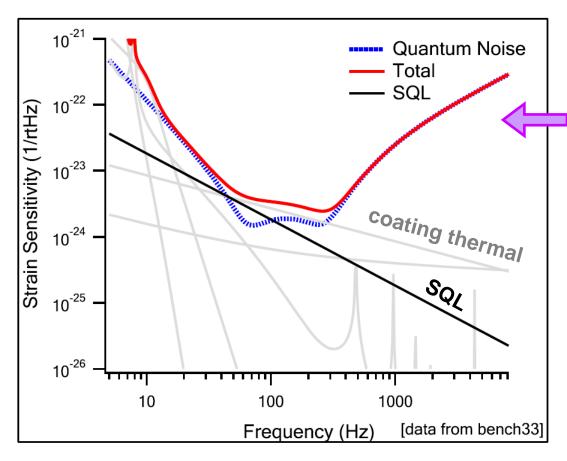


- Quantum limit of conventional measurement
- Ideas to overcome the limit

(squeezing, homodyne detection, speed-meter, opt-spring, ... etc.)

Essential to test quantum mechanics

Can we reach the SQL?



AdLIGO's sensitivity will be close to the SQL

40kg masses are almost in the quantum superposition!!

$$h_{SQL} = \sqrt{\frac{8\hbar}{m\Omega^2 L^2}}$$

- What if we reduce the mass?
- Can we reduce thermal noise?

Mass-dependence of thermal noise

1. SQL and coating Brownian TN

$$SQL \propto 1/\sqrt{m}$$

coating TN
$$\propto 1/w_0 \propto 1/m^{1/3}$$

[fixed aspect-ratio]



Gain is only by m^1/6

2. SQL and suspension TN

$$\operatorname{suspTN} \propto \begin{cases} 1/m^{3/4}(LF) \\ 1/m^{5/4}(HF) \end{cases}$$

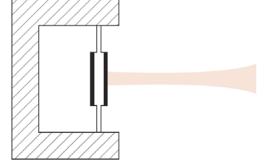
[fixed wire-thickness]



Susp TN increases faster than SQL

Low-mass regime and high-mass regime

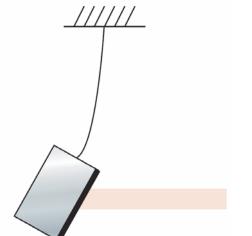




- high SQL
- high frequency
- low Q, but easy to cool down

 $300K \rightarrow 4K$



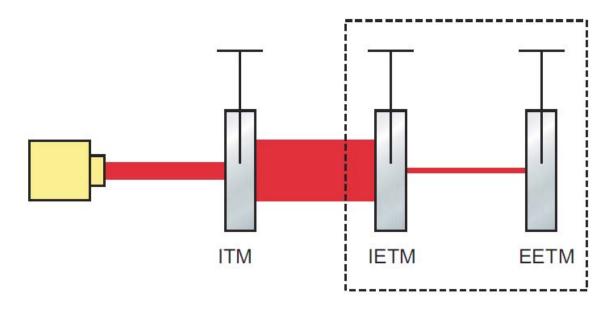


high-Q pendulum with dilution by gravity

- knowledge from GW detectors
- need to do something to reduce coating TN

Reduction of coating TN with end-mirror cavity

[Khalili PLA 05]



Replacing ETM to an anti-resonant cavity



~20% TN reduction with rietm=ritm

→ can be better with lower rietm

Setup plan of the experiment ref. cav. **ITMs** BS **IETMs EETMs** input optics

- Folded FPMI with end-mirror cavities
- Ref. cavity with heavy mirrors for freq-stabilization
- Seismic isolation with SAS and SPI

Parameters to be optimized

1. Mass

Trade-off of susp-TN/SQL vs coat-TN/SQL

2. Aspect ratio of a mirror

Finite-size analysis

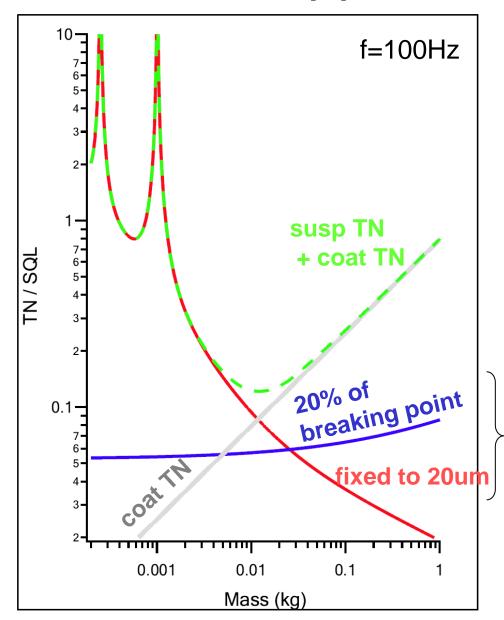
3. Reflectivity of each mirror

Trade-off of IETM TN vs EETM TN (+ IETM TR noise)

4. Arm length

Trade-off of freq noise vs cavity stability

(1) Mass optimization



- Manufacture limit ~ 20um
- Breaking point ~ 2GPa
- 2 fibers on each side
- 30cm length (1st violin mode = 500Hz w/ m=100g)



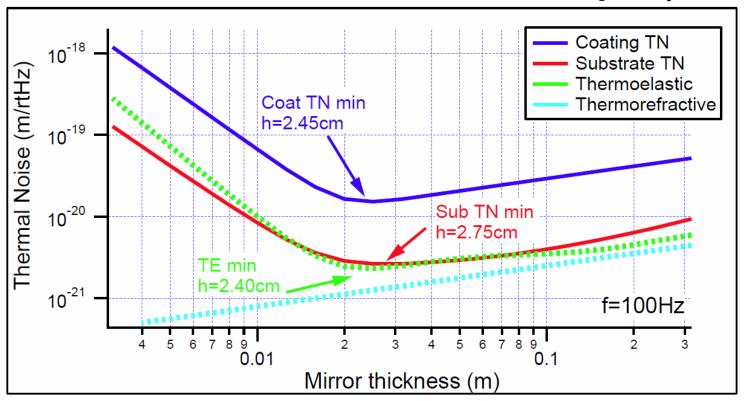
susp TN



Our choice: m=100g

(2) Aspect ratio

[Somiya PRD 09]

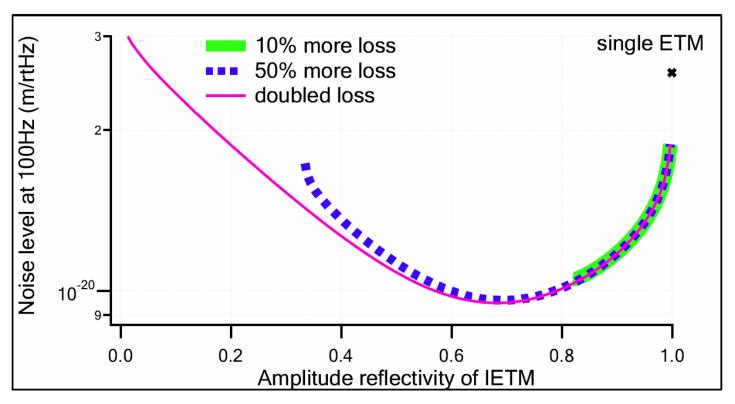


Mass is fixed to 100g. Beam radius is mirror radius/2.5.

- big h → small beam radius → noise increases
- small h → high susceptibility → noise increases

Our choice: h=2.45cm, a=2.43cm

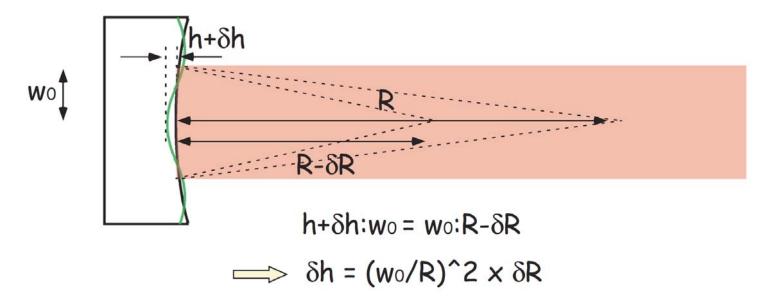
(3) Reflectivity optimization



- high IETM-ref → more coat TN on IETM
- low IETM-ref → more coat TN from EETM (also IETM's substrate thermorefractive)

Our choice: N=2 (r=0.72) for IETM
[N=15 for EETM]

(4) Arm length and cavity stability

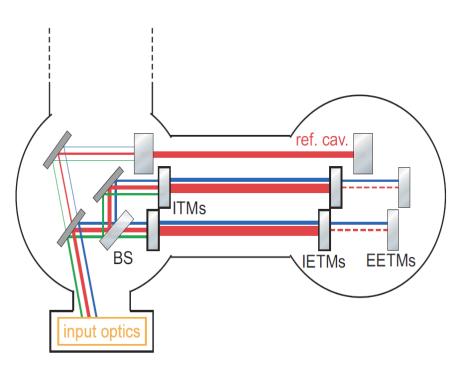


Cavity stability:
$$w_0 < \frac{\lambda}{\pi} \sqrt{\frac{L}{2\delta h}} \implies L > 2\delta h \left(\frac{w_0 \pi}{\lambda}\right)^2$$

- mirror radius / 2.5 = 9.7mm →
- $\delta h \sim 2nm$

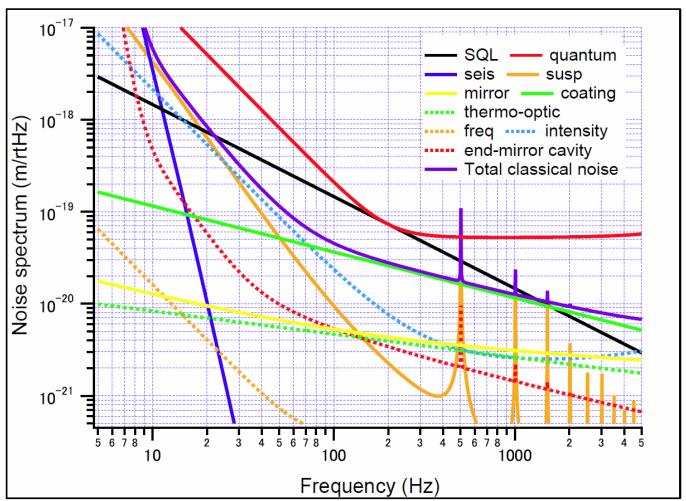
Our choice: L=10m (1.5m chamber + 10m pipe)

Parameters are determined



- number of coating layers
 N=8(ITM), 2(IETM), 15(EETM)
 finesse = 600
- L=10m
- m=100g (\$\psi_2.45cm x 2.43cm)
- input power = 5.5W (touching SQL at 200Hz)
- suspension l=30cm, d=28um, N=4
- mechanical losses (at 100Hz)
 susp:6e-7, subs:5e-10,
 coat (silica):1e-4, coat (tantala):4e-4
- ISS: 100mW PD
- FSS: reference cavity (F=6000, I=100mW, m=860g)

Estimated noise curve



- SQL is a factor ~3.2 above classical noise (100Hz)!
- Coat TN (ITM) and intensity noise limit the margin

Discussion to increase the margin

Coating TN reduction

- ITM cavity (or power-recycling)
- doped tantala
- non-quarter-wave-length coatings

Intensity noise reduction

- stabilization with L+
- stabilization with squeezing

Mass reduction

- ribbon suspension
- magnetic levitation