

INTRODUCTION

This handbook is intended to be a collection of circuit diagrams and other relevant information about the locking servos of the 40 meter prototype Gravitational Wave Detector at Caltech.

The information contained in this handbook reflects the state of the servos as of July 1989. The handbook should be updated continuously; otherwise it will be obsolete very soon.

In order to update a circuit diagram in the handbook; please make a copy of the old diagram and mark your changes on the copy. Then put the upgraded copy in front of the old diagram in the handbook. A short explanation of the changes on the diagram is also very desirable.

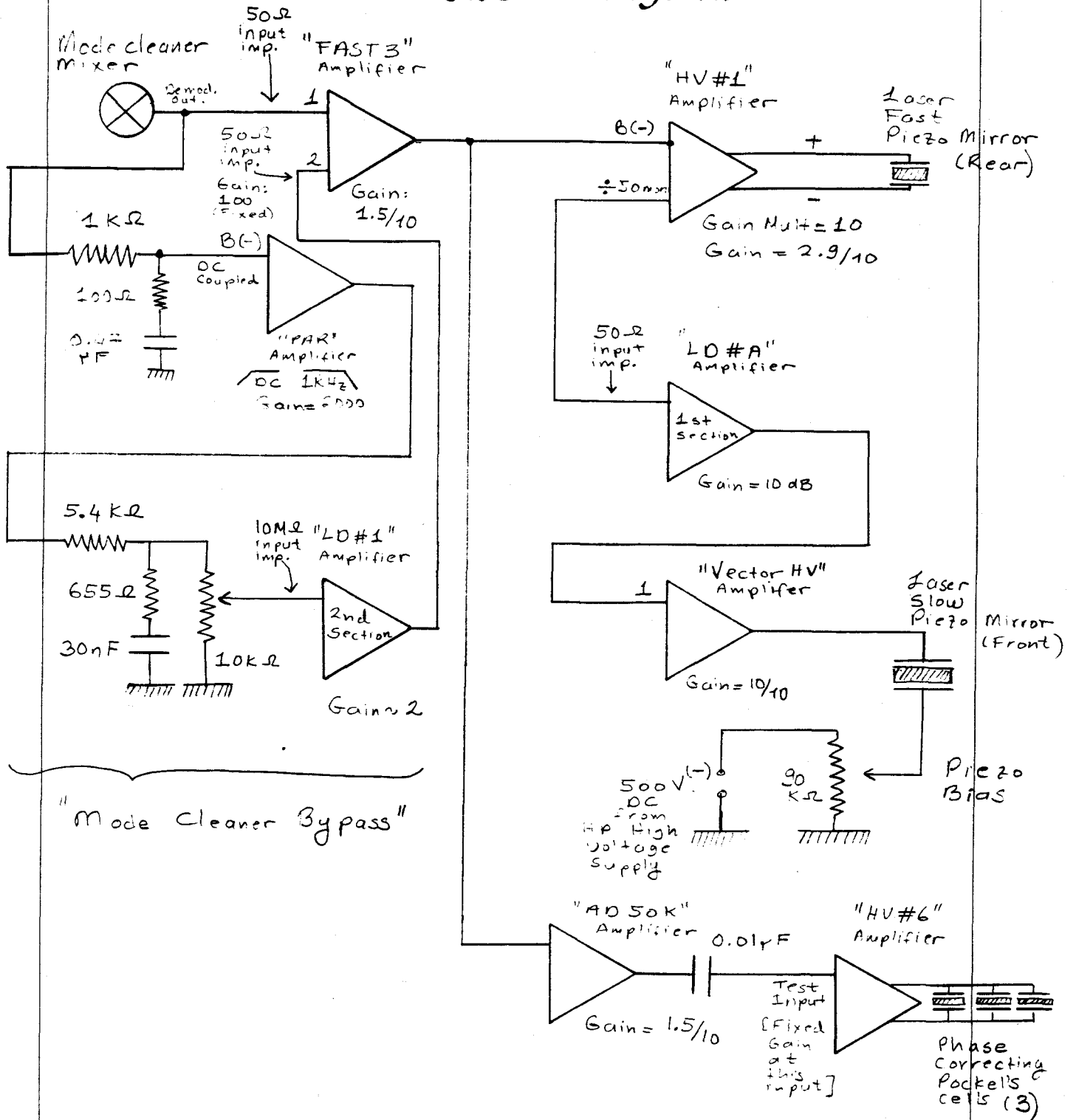
Thank you very much for your cooperation.

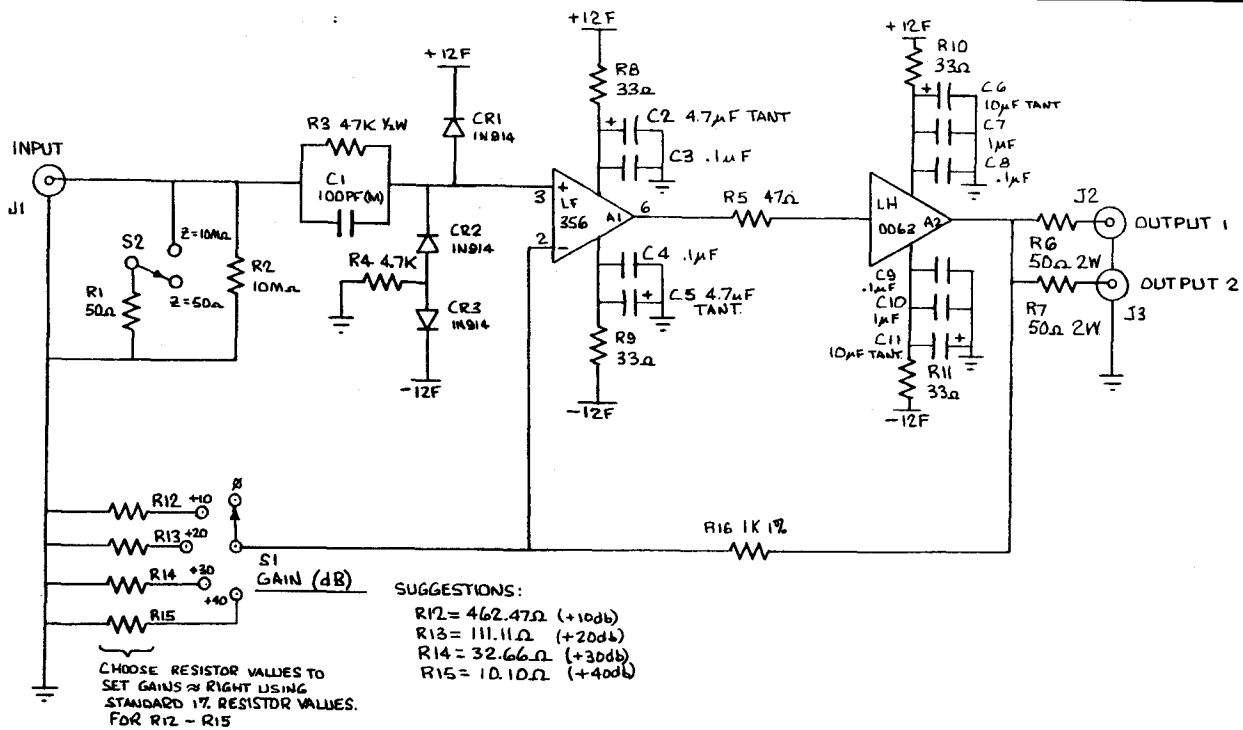
Jeha Cui
July 24, 1989

BATCH
START

STAPLE
OR
DIVIDER

Mode Cleaner Servo Loop Block Diagram





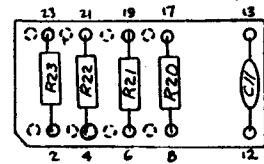
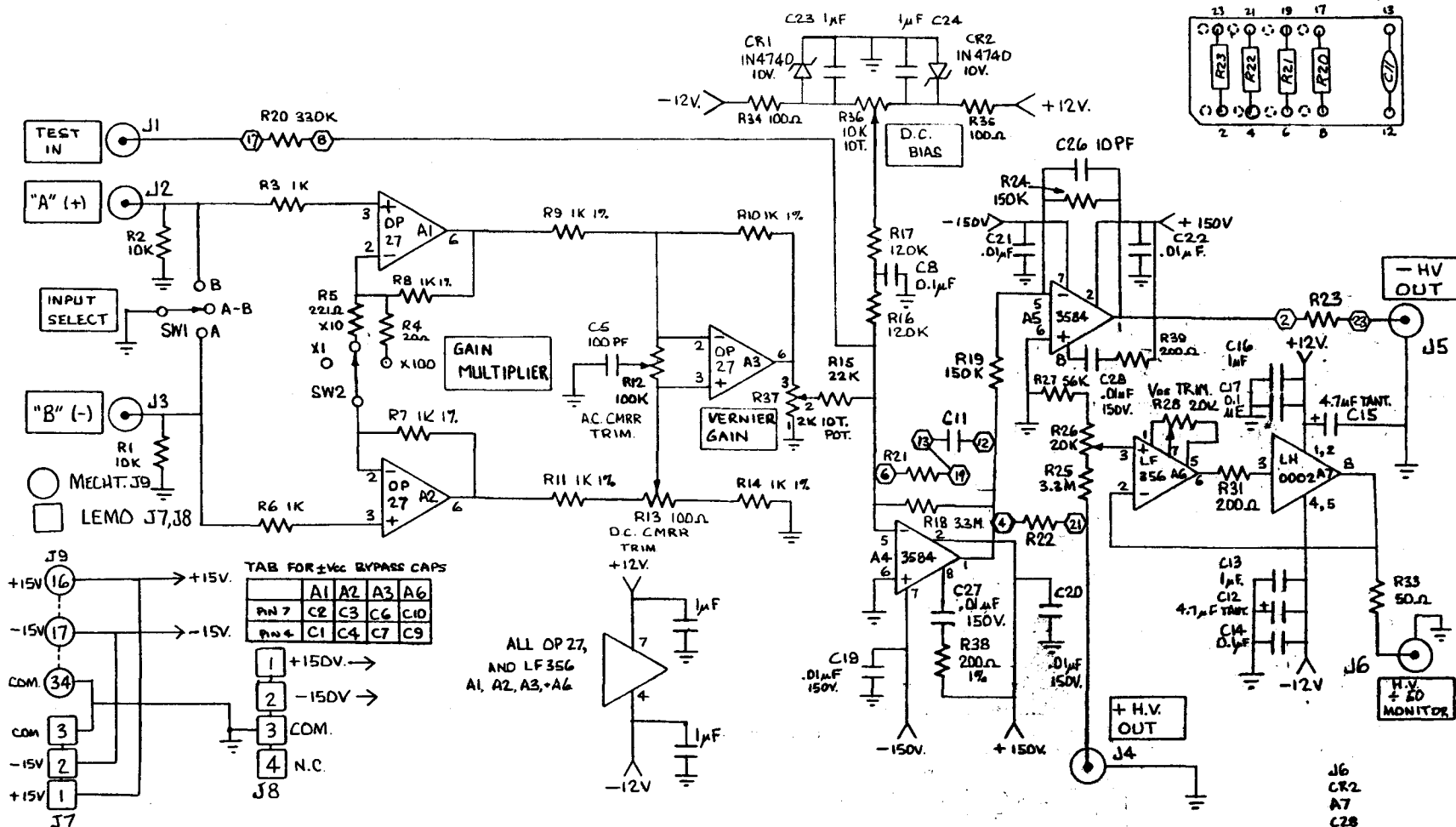
CHOOSE RESISTOR VALUES TO SET GAINS AS RIGHT USING STANDARD 1% RESISTOR VALUES. FOR R12 - R15

- SUGGESTIONS:
- R12 = 462.47Ω (+10db)
 - R13 = 111.11Ω (+20db)
 - R14 = 32.66Ω (+30db)
 - R15 = 10.10Ω (+40db)

A 2
CR 3
S 2
J 3
R 16

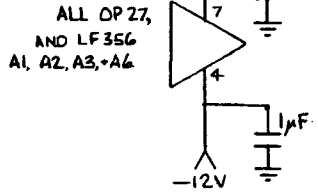
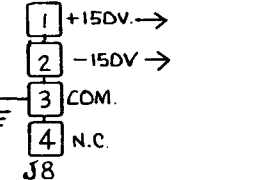
L&T C

CALIFORNIA INSTITUTE OF TECHNOLOGY GRAVITATIONAL PHYSICS		
LINE DRIVER		
DRAWN BY B.T.	DATE 6-13-88	DRAWING NO.
CHECKED BY	SCALE	
APPROVED BY	W.D.	



TAB FOR $\pm V_{cc}$ BYPASS CAPS

	A1	A2	A3	A6
Pin 7	C2	C3	C6	C10
Pin 4	C1	C4	C7	C9



2. J7, J8, and J9 ARE ALL MOUNTED ON BACK PANEL.
 NOTES 1. INDICATES JUMPER TO SHORT.

SMARTWRK FILE C:\SMARTWRK HVAMP.PCB
 DRAWN FROM E.LINDELF OF 9-1-87 P20 HVAMP.PL
 ADDED COMP CARRIED PIN NO'S TO DWG.

REPLACES 87-0901-1

J6 CR2
 A7
 C28
 LAST R39

R32
 C25 R30
 C18 R29
 NO'S SKIPPED

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GRAVITATIONAL PHYSICS

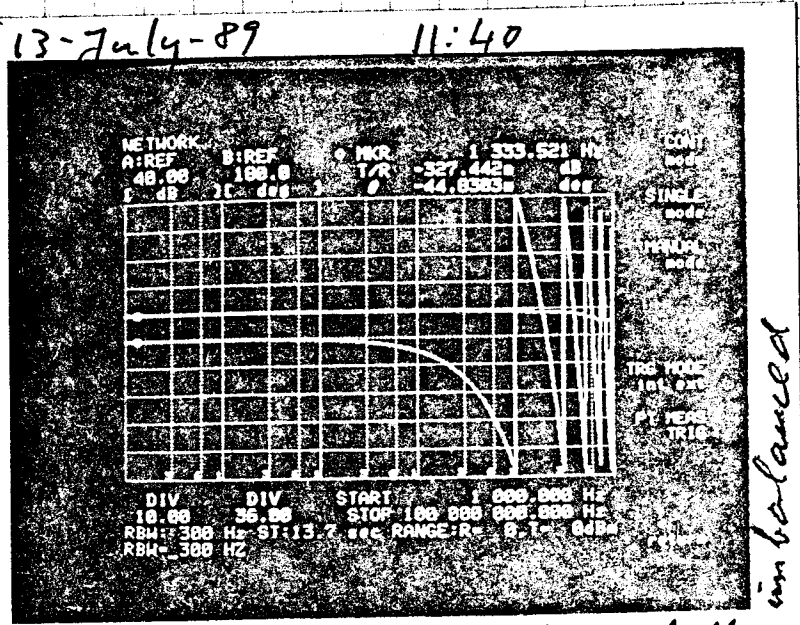
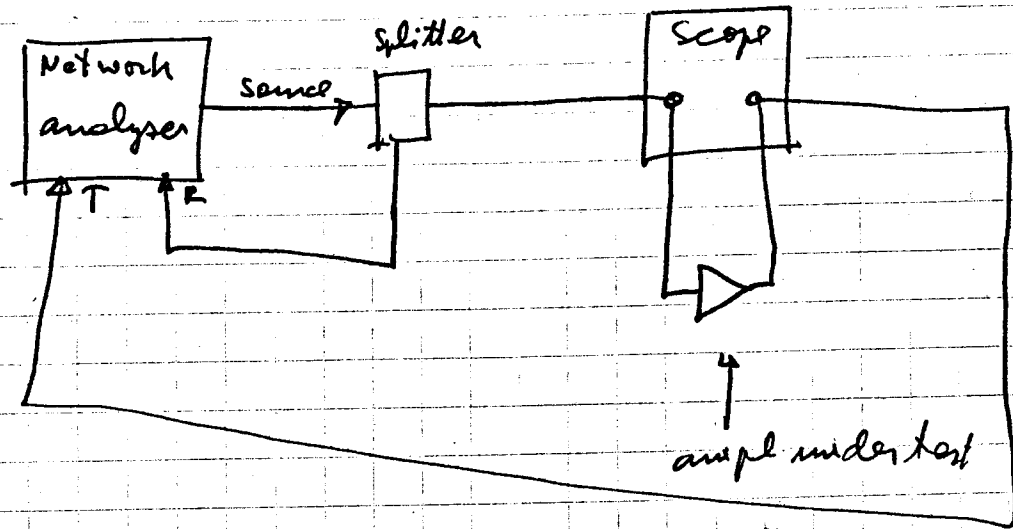
HV AMPLIFIER $\pm 150V$.

DRAWN BY B.TWIKER	DATE 3-4-88	DRAWING NO.
CHECKED BY	SCALE	
APPROVED BY	W.D.	

13-July-89

078

Since ~30' of cables were used to connect the amplifier under test (see diagram below), we took the response of the cables themselves (trace at 11:40)

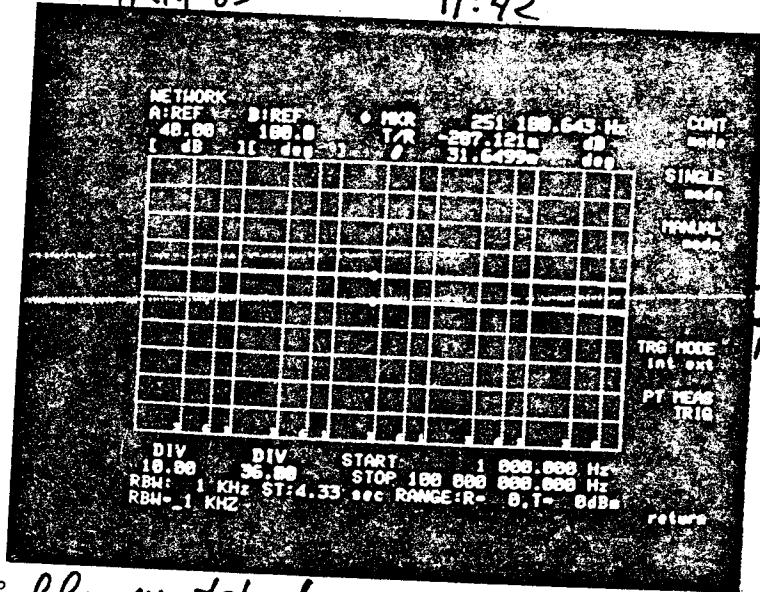


TR. FUNCTION OF CABLES - they are badly

- It turns out that at 250 kHz the cables alone show a phase shift of 5°.
- Therefore, the ~3' cable from splitter to R input was replaced with a long cable, matched to the one in the test path. See traces overlaid

13-July-89

11:42

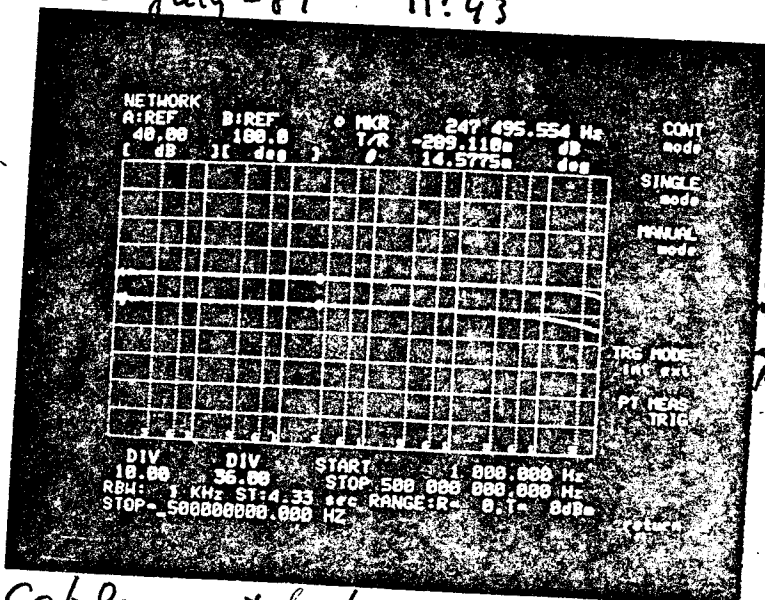


15
ph

Cables matched

13-July-89

11:43

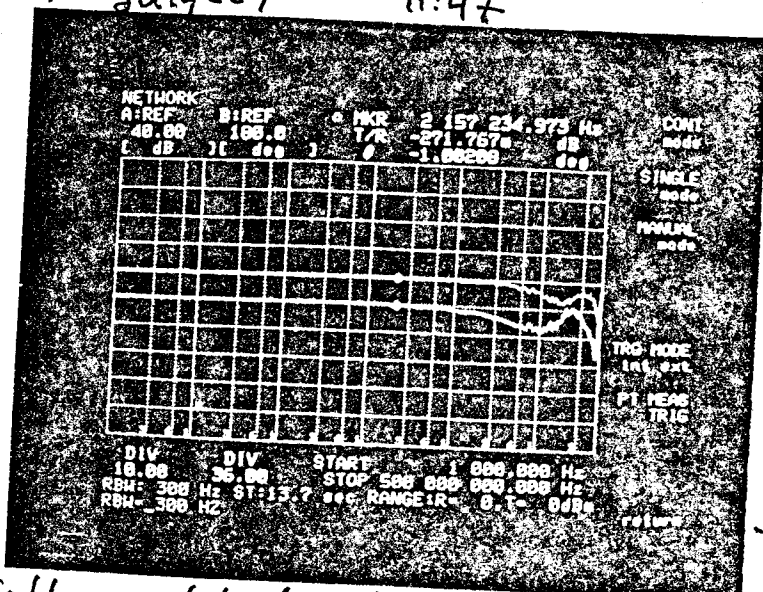


15
ph

Cables matched

13-July-89

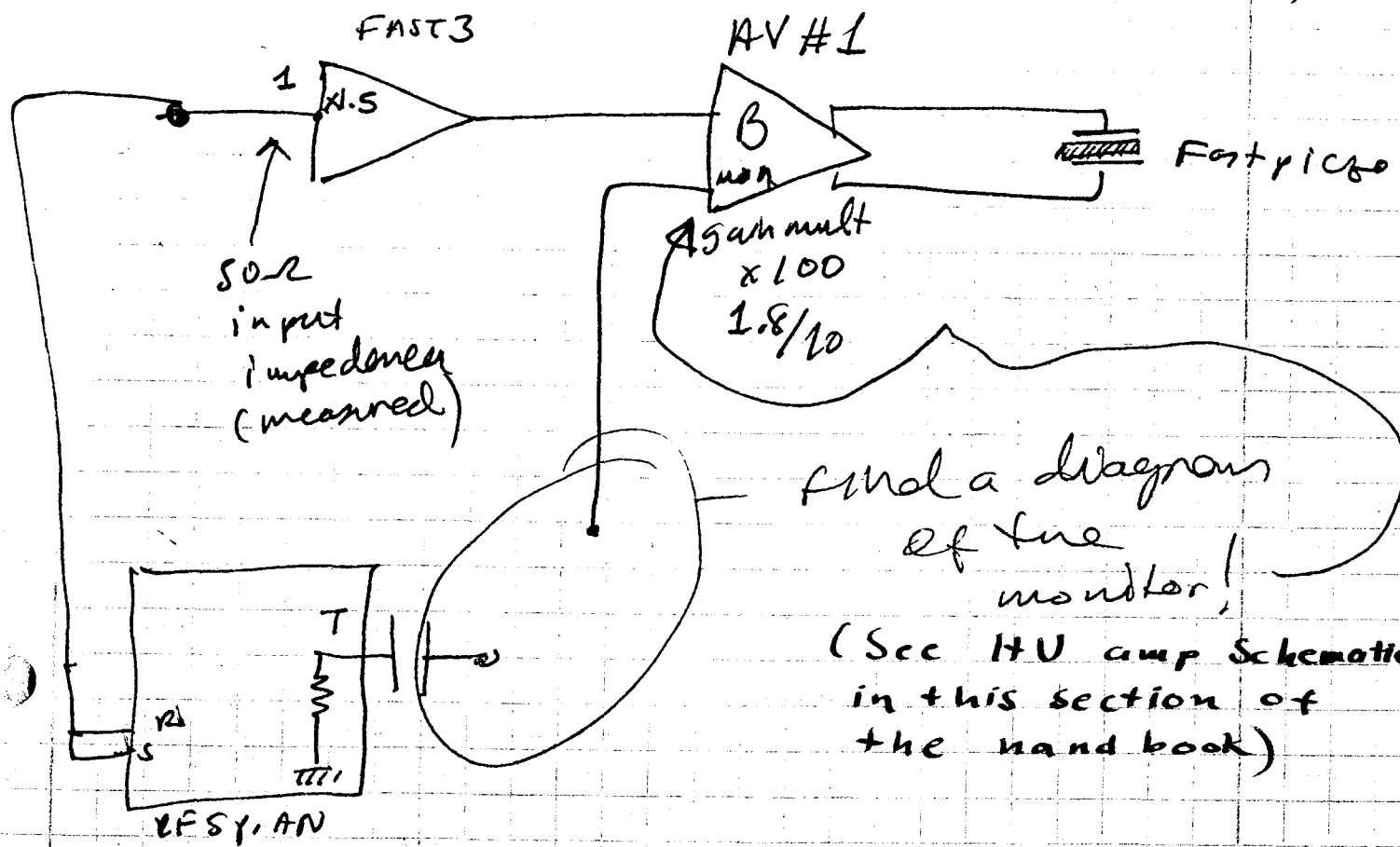
11:47



to scope

Cables matched, going through T's connected

Fast Piezo branch of mode cleaner loops (Transfer Function)

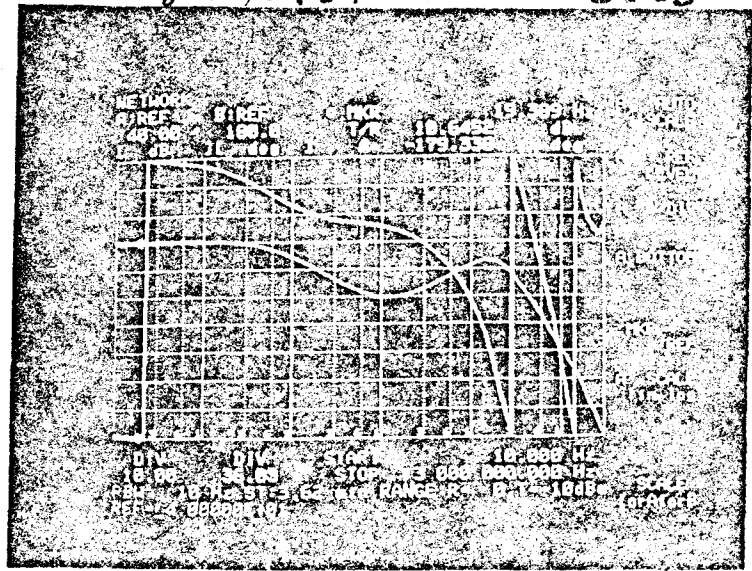


find a diagram of the monitor!

(See HV amp Schematic in this section of the handbook)

July 18, 1989

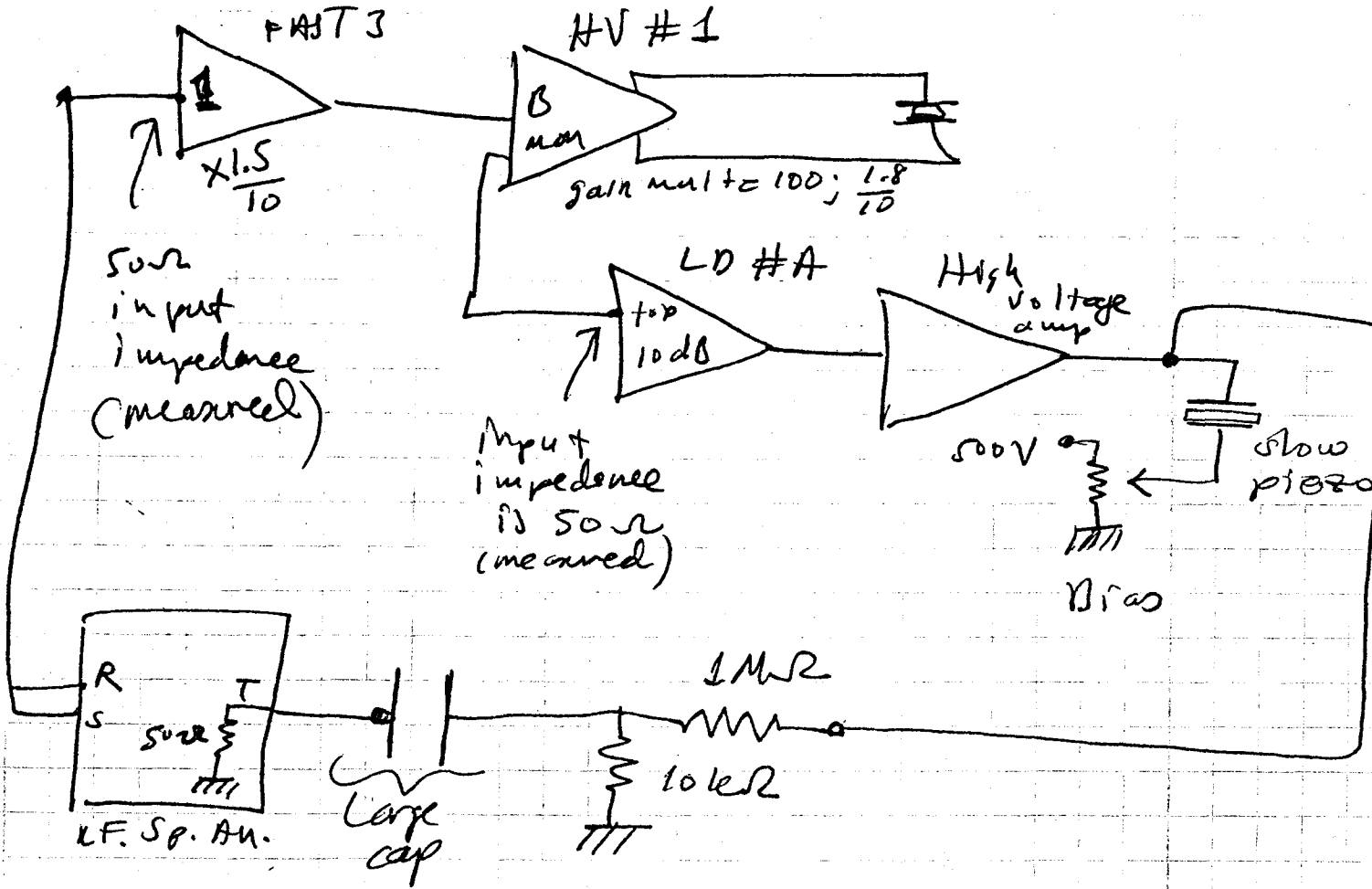
15:55



AF 672014

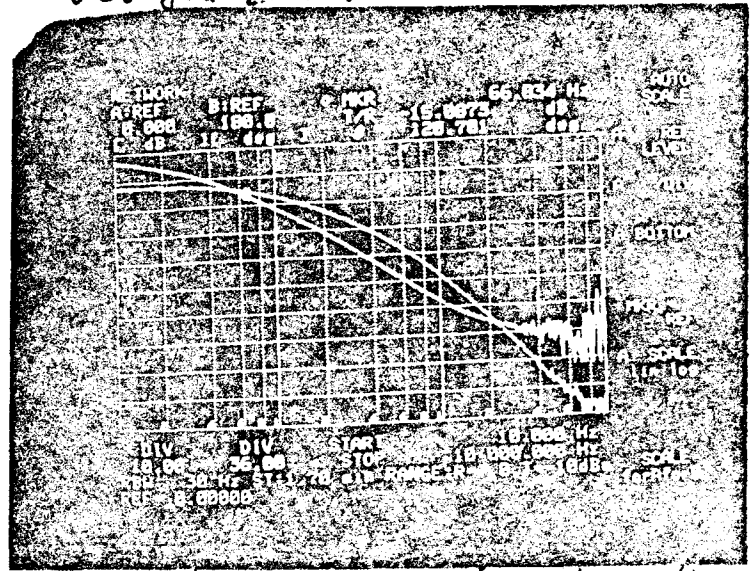
Fast piezo branch of mode cleaner Loop

Mode cleaner slow piezo branch ~~piezo~~ transfer function:



July 18, 1984

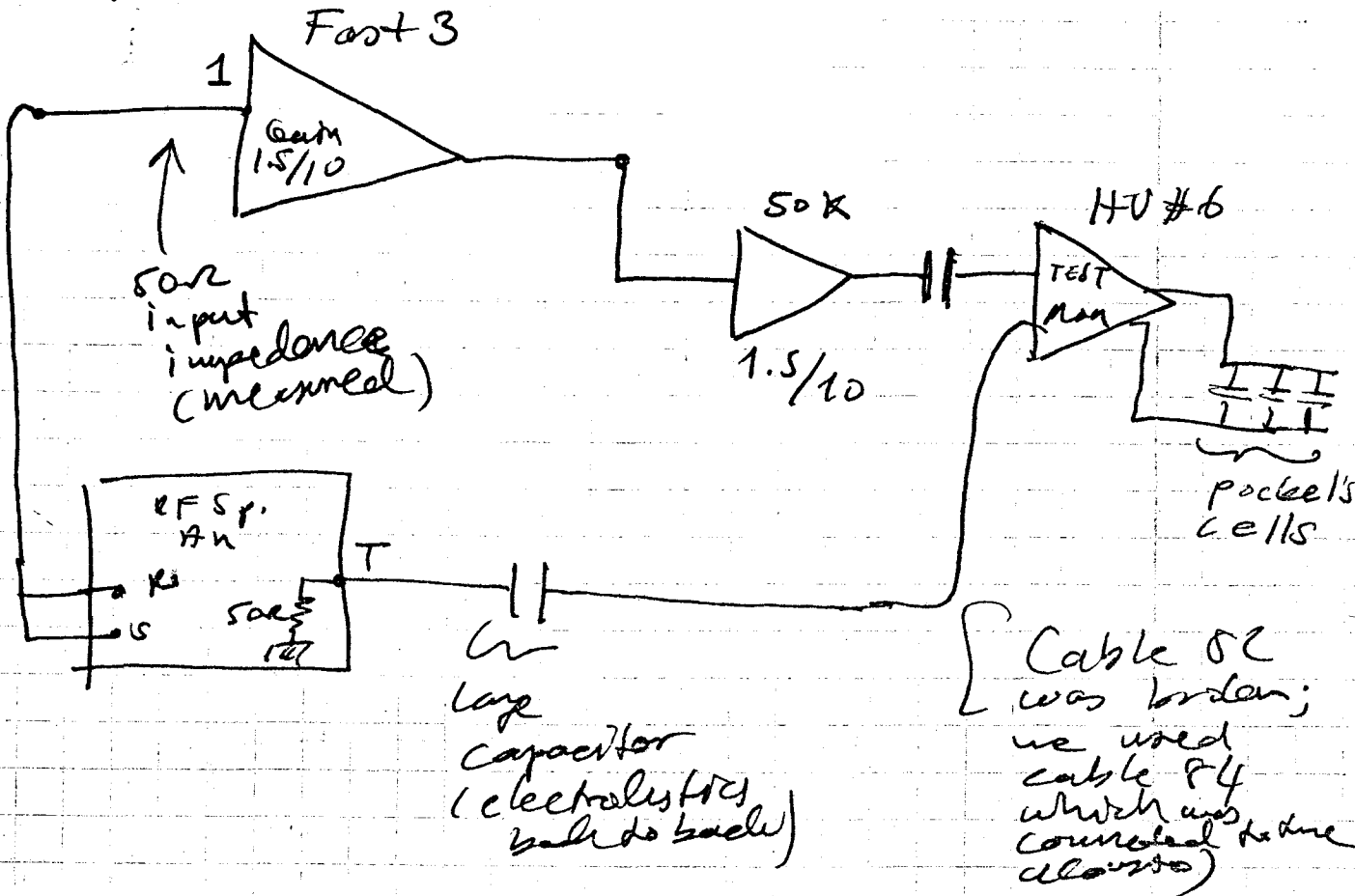
16:25



MDCLP718

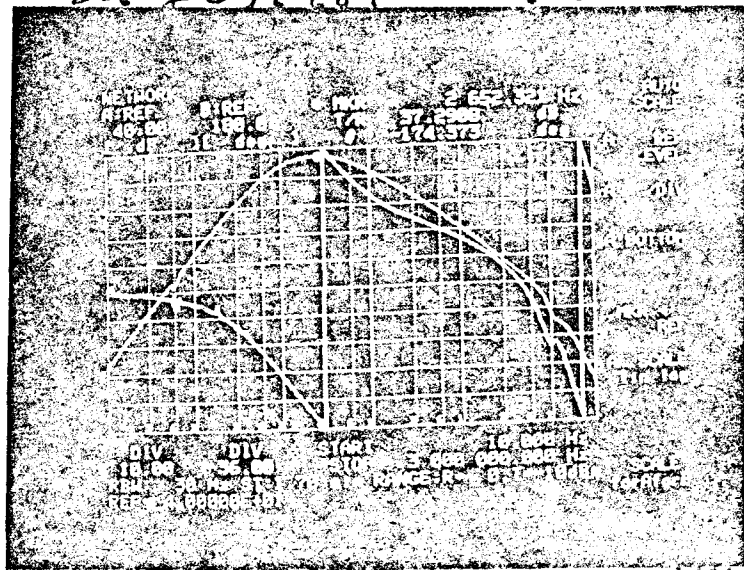
mode cleaner slow piezo branch transfer function

Mode cleaner phase correctly
pocket's cells branch transfer
function:



Cable 8C was broken; we used cable 84 which was connected to the cleanroom

July 25, 1979 11:50

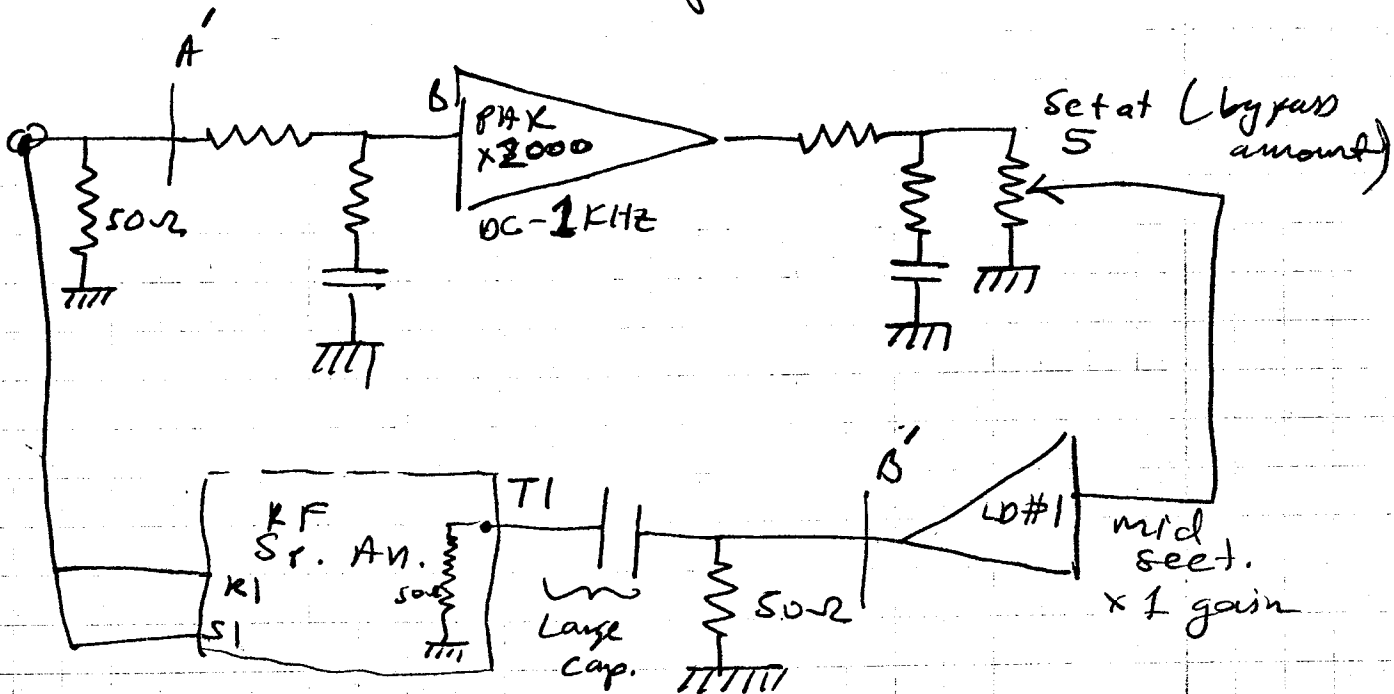


MDCLPC 7/8

mode cleaner phase correctly
pocket's cell branch tr. function.

July 18, 1989
 YB

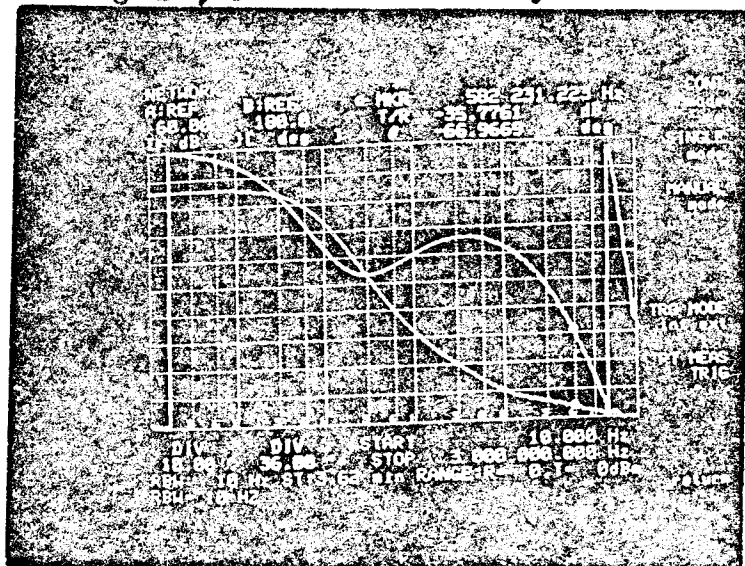
Mode cleaner loop servo
 branch transfer functions



A' to B' : Mode cleaner by pass circuit.

July 18, 1989

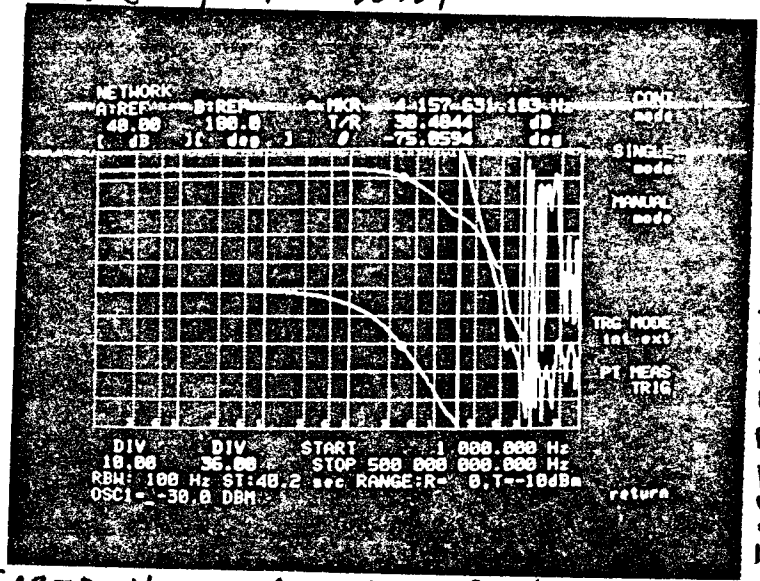
14:30



MDC LBP 718

Mode cleaner by pass transfer function.

14-July 89 10:59



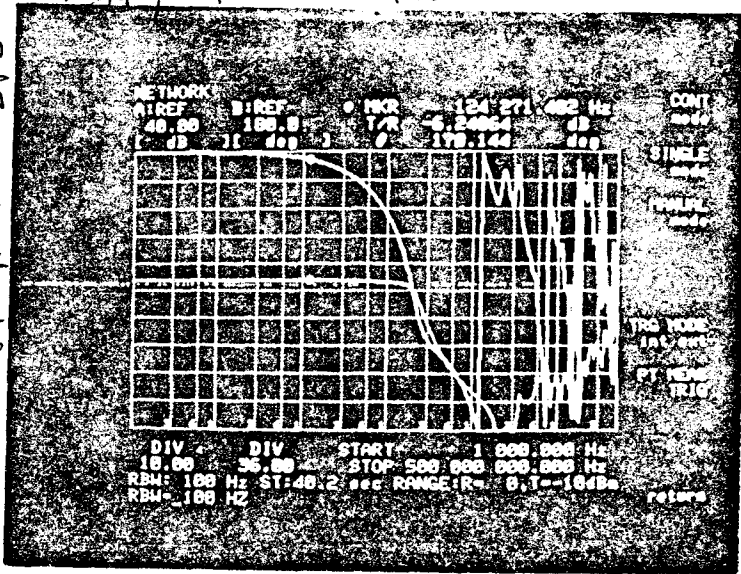
FAST 3711V



FAST3, Var. in (max) 10°: 461.6 kHz

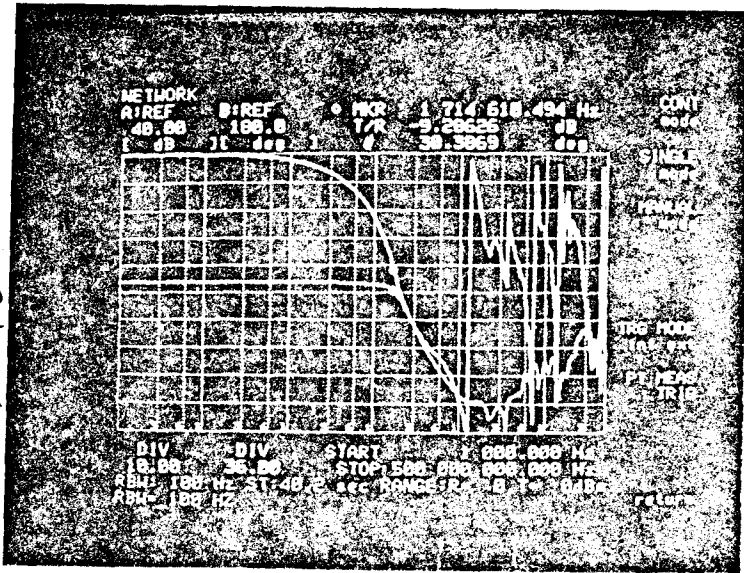
17 July 89 13:35

3 dB point at: 1.95 MHz



Line driver A, mid, 0 dB, 50 ohm out.

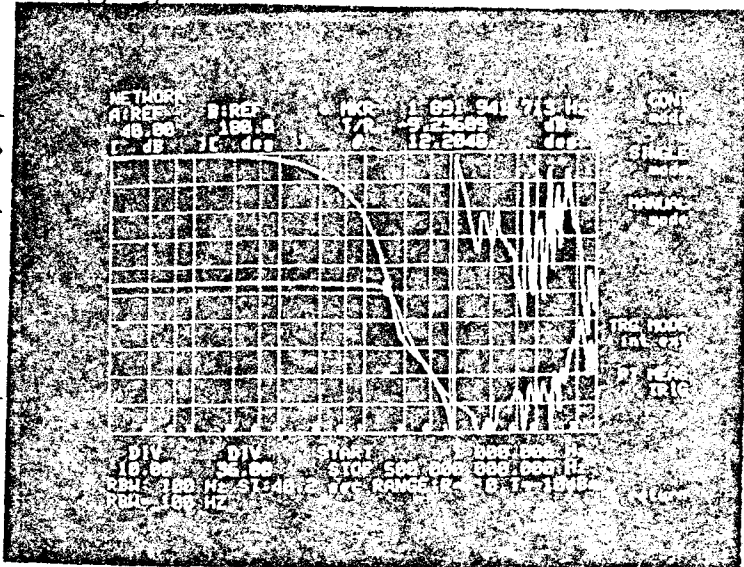
100 : 128.4 kHz



Line driver A lower, 50 ohm out, 0 dB

17 July 89 13:30

100 at: 124 kHz



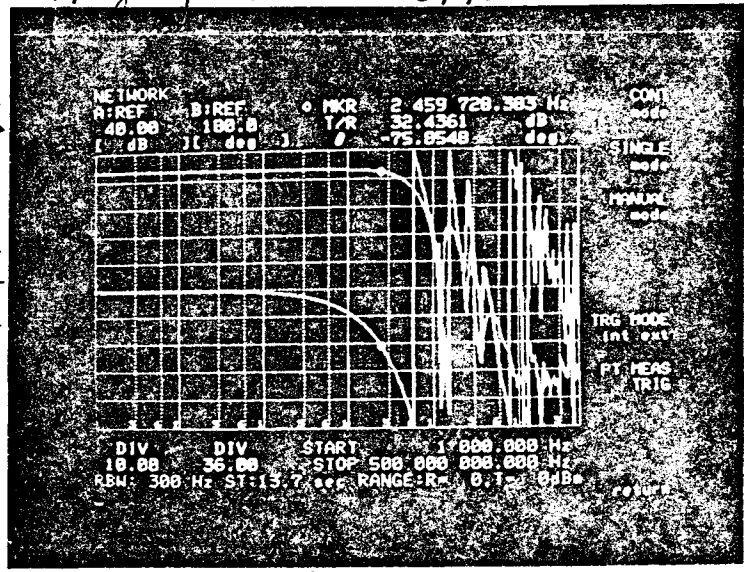
Line driver A, upper, 0 dB, 50 ohm at outp.



17 July 1989

17 July 89 9:40

10³ : 321 kHz

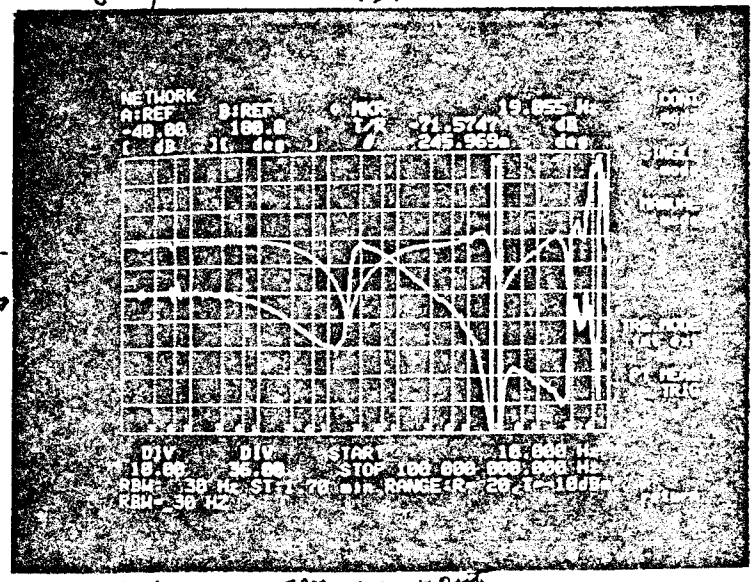


FAST 1 V AXES: 100 MHz/40dB

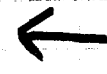
FAST 1 50 Ω out, v in, full gain (33.5 dB)

17 July 89 13:43

2
dB



CROSS 1



Crosstalk between v_T and lower ^v out on LD# A

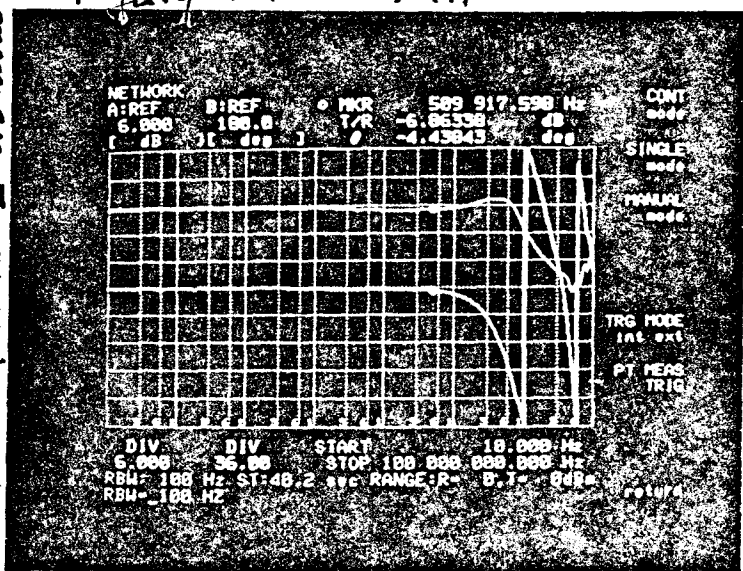
14-July-89

How transfer functions taken

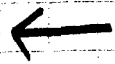
GAIN = 0 dB nominal

14-July-89

10:17



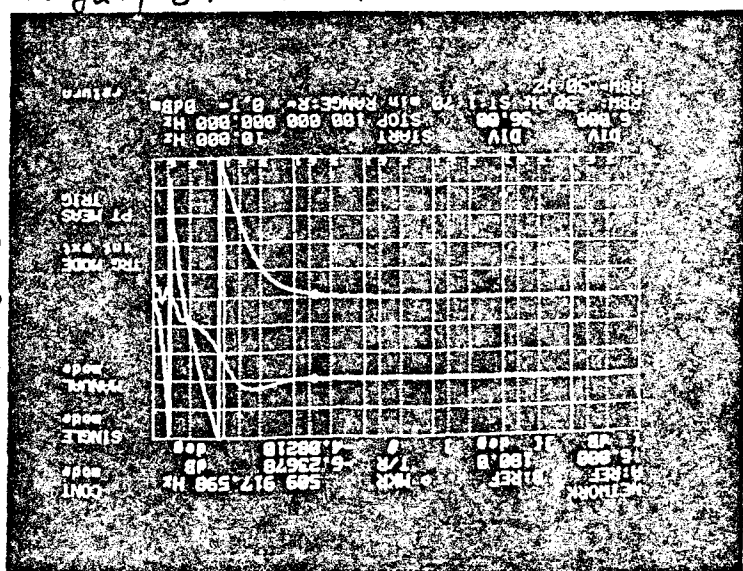
LD1 MID



LINE DRIVER 1, Middle, 50 Ohm in, out

14-July-89 10:25

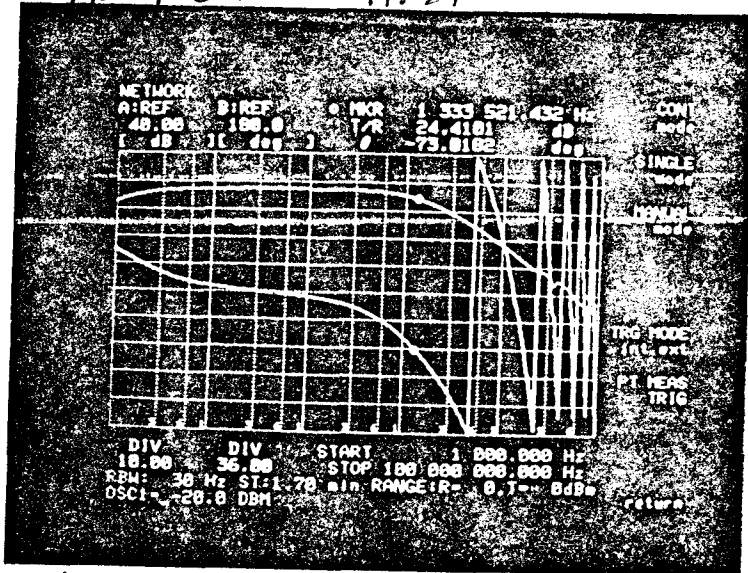
GAIN = 0 dB nominal



LD1 Low

Line driver 1, Lower, 50 Ohm in, out

13-July-89 11:21



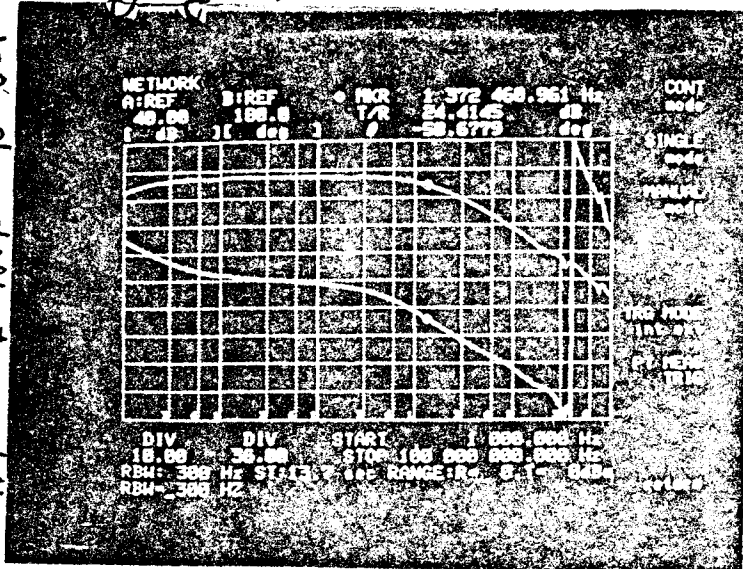
AXES: 100 MHz 40 dB
50k 7/3

← unmatched cables



50k 50Ω in and out

13-July 89 12:02



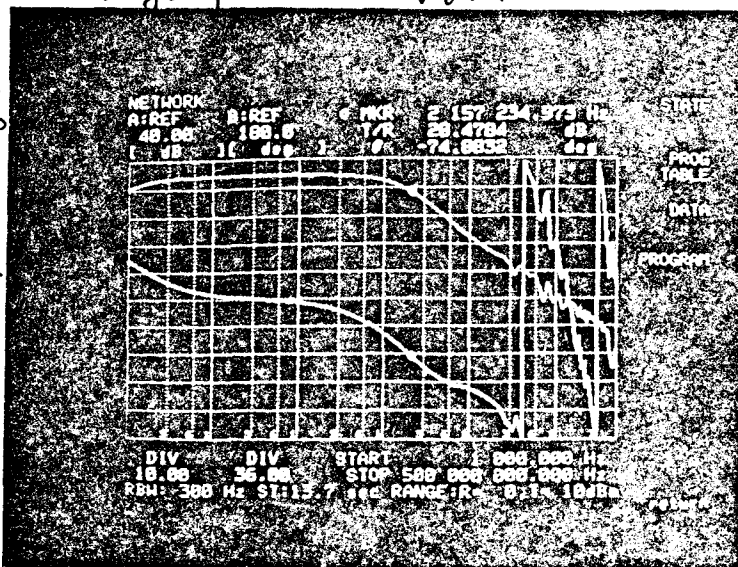
50k in and out 50k 7/3

← matched cables



50k 50Ω in and out

14-July 89 17:40



50k 7/3



50k 50Ω in and out

100° phase shift at 230 kHz

100° 203 MHz

CALIFORNIA INSTITUTE OF TECHNOLOGY

GRAVITATIONAL PHYSICS 130-33

November 16, 1988

Jeff Livas
MIT Gravitational Physics Group

Howdy, Jeff—

I've finally come up with the much-promised description and evaluation of our current (reasonably successful) fast piezo laser stabilization servo. The numbers are very crude but give all the qualitative features we actually observe in their respective places. Some of the numbers I quoted to you on the phone weren't quite right; one that I remember misquoting is the responsivity of our "slow" pzt stack. It is really more like $40 \text{ \AA} / \text{volt}$.

The procedure I use for working out the crossover points between the various parallel signal paths in these "bypass topology" loops is to break them apart into subunits at those points where they share a common input or output. In this example, I've taken the output of the "Fast # 3" video preamplifier as one common signal path, and the laser output frequency (after all the contributions from fast and slow piezos and the phase-correcting Pockels cells have been added together) as another. The various electronic transfer functions of the parallel subunits are computed, multiplied by the electro-optic transfer functions of their respective actuators, and added together. It's easiest to just do this graphically on a log-log plot, since your input data is rarely more accurate than your pencil and (for a good robust design) the major stability criteria should be extremely insensitive to small changes in gain.

You should be warned of a few glitches that we'd rather not have. Since our PA-85 amp blew up we've been using the "HV # 6" in its place for the pockels cells. It unfortunately is not fast enough

and has this weird wrinkle around 100 kHz where the inverting half poops out. With the PA-85 in place we got a unity-gain frequency of about 1 MHz rather than 350 kHz and no wrinkle. Also, you will note that the slow pzt/fast pzt crossover violates the canonical rule of "no crossovers at greater than 6 db/octave relative slope." We just barely get away with this one because this crossover is not far enough away from the "extra" pole to have the full effect of its phase shift. This is a marginal situation which we hope to correct in the next version.

In our conversation you asked about dealing with the resonance of the PZ80 piezo mirror at 5 kHz and I'm afraid I wasn't too clear about my recommendation. Just in case, let me state the problem more formally; one wants to attenuate the signal reaching the nasty element (the PZ80) by a large factor to prevent its resonant behavior from dominating over the righteous and properly-phased feedback going to, say, your fast piezo at that resonant frequency f_r . At the same time, the crossover rule prohibits you from adding too much phase shift at some lower frequency $f_{\text{crossover}}$. The optimization therefore is for maximum attenuation in the stop band with minimum phase shift in the passband, and the solution we tend to use is a Butterworth filter with the 3dB point chosen at the geometric mean $f_{3\text{dB}} = (f_r f_{\text{crossover}})^{1/2}$ with whatever number of poles it takes to control the resonance.

Give me a ring if you have a chance to look at this stuff and we can discuss it. Happy frequency stabilization !!

Closing the loop,

M. E. Zucker

cc: Alex Abramovici
Bob Spero
Jeff Harman

11/14/78 MEG

MODE CLEANER SERVO BLOCK DIAGRAM

NOTE: THIS IS A WORKING DOCUMENT ONLY. ALL DATA ARE PRELIMINARY.

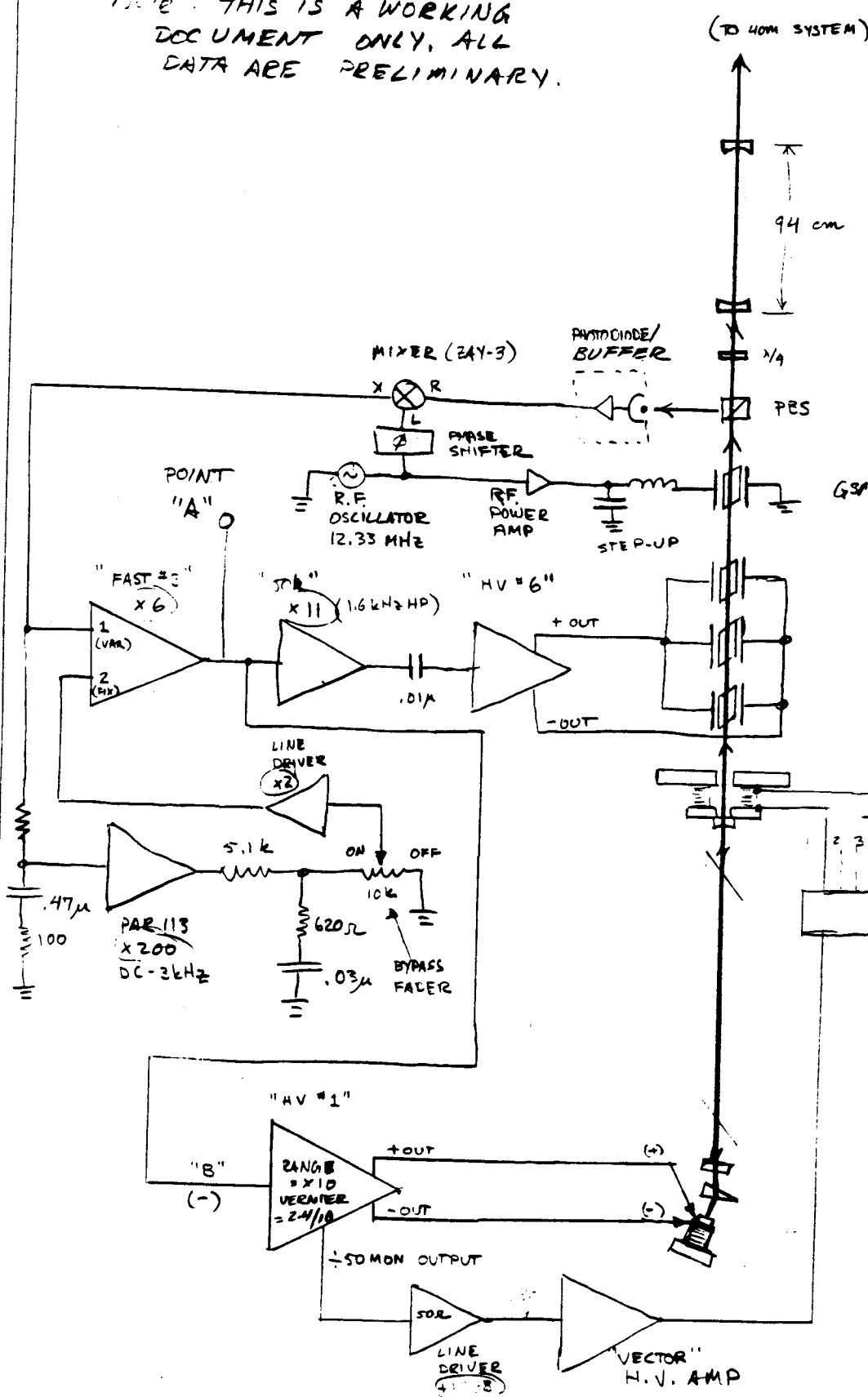
(TO 40M SYSTEM)

DC KEYS
TELL
DRIVE

PASS

FAST
PZT

SLOW
PZT



MODE CLEANER
FSR = 160 MHz
 $T_c = 2.4 \mu\text{sec}$

(ISOLATOR & MODE MATCHING LENSES OMITTED)

94 cm

1/4

PES

DM-25

DM-25

DM-25

DM-25

11. $C \approx 30 \text{ pF}$ EACH CELL

SLOW PZT (3 STACKS)
 $\sim 90 \text{ V/ORDER}$ $C \approx .187 \mu\text{F}$

3-STACK BALANCING NETWORK

FAST PZT
 550 V/ORDER $C \approx 500 \text{ pF}$

11/14/88 MEG

MODE CLEANER SERVO TRANSDUCERS (IN TERMS OF EQUIVALENT LASER FREQUENCY)

1. POKKELS CELLS

$$2\pi \nu_{\text{mirr}} = \frac{\partial \phi_{\text{mirr}}}{\partial t} ; \quad \phi_{\text{mirr}} = 2\pi \nu_0 t + \phi_{\text{pc}} ; \quad \phi_{\text{pc}} = \frac{2\pi n_{\text{pc}} l_{\text{pc}}}{\lambda} ;$$

$$n_{\text{pc}} l_{\text{pc}} = k + \alpha V_{\text{pc}} \quad (\alpha \approx .22 \text{ nm/V for PM-25})$$

$$\phi_{\text{pc}} = 2\pi (k + \alpha V_{\text{pc}}) / \lambda$$

$$\Rightarrow \delta \nu_{\text{mirr}} = \frac{2\pi \alpha}{\lambda} \frac{dV_{\text{pc}}}{dt}$$

i.e., for $V_{\text{pc}} = V_0 \sin(2\pi f t)$ (single Fourier component at f)

$$\frac{\partial(\delta \nu_{\text{mirr}})}{\partial V_0} = 2.7 \times 10^{-3} \frac{f}{V_0} = 2.7 \text{ Hz} \left(\frac{f}{1 \text{ kHz}} \right) \left(\frac{1 \text{ volt}}{V_0} \right)$$

NOTE: We're using \approx phase-adjusting po cells in series, driven in parallel.

2. FAST PIEZO MIRROR

Since $\frac{dV_{\text{PZT}}}{dl} = 550 \text{ V/order}$ and the laser is ≈ 2.3 meters long

$$\Rightarrow \nu_{\text{c}} = 65 \text{ MHz} = 1 \text{ order}$$

$$\Rightarrow \frac{\partial \nu}{\partial V_{\text{PZT}}} = \frac{65 \text{ MHz}}{550 \text{ V}} \approx \frac{120 \text{ kHz}}{\text{volt}}$$

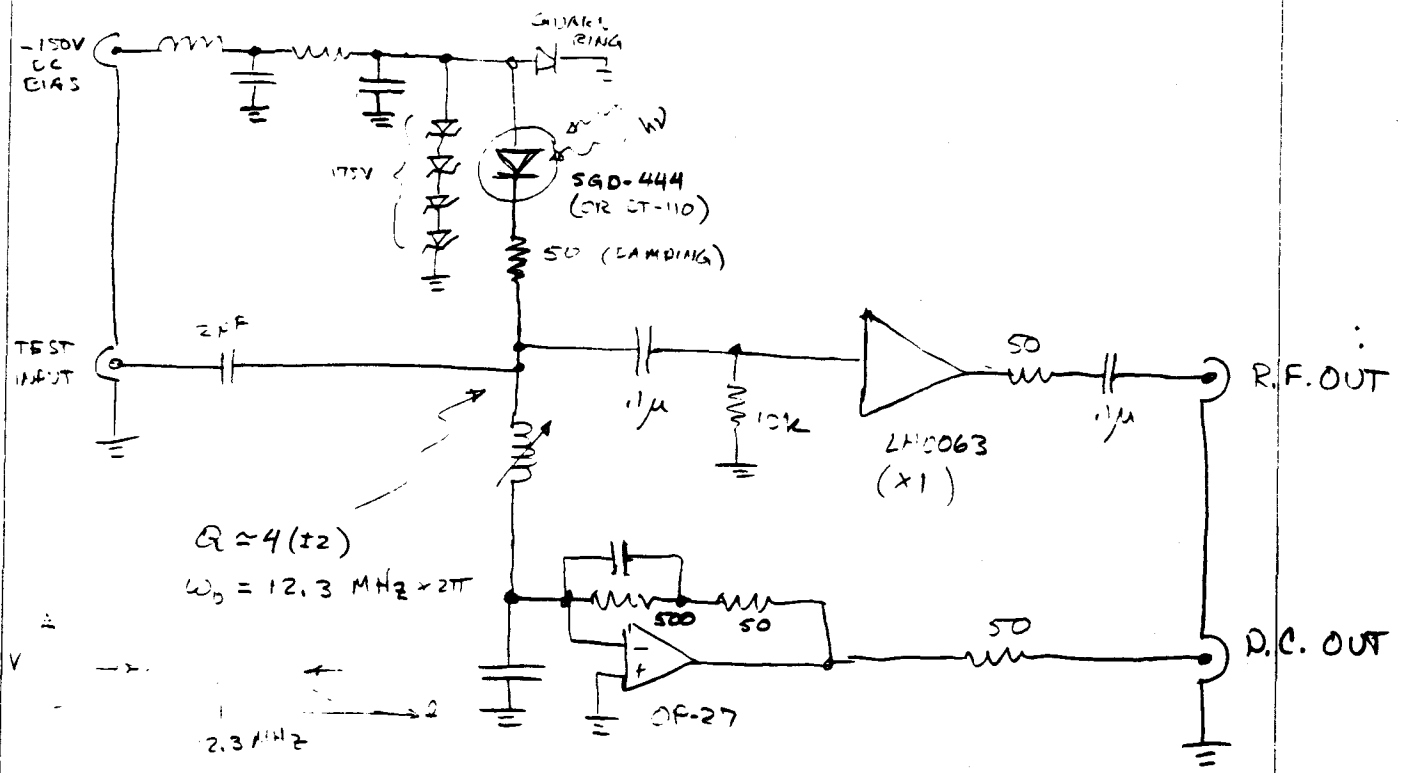
3. SLOW PIEZO MIRROR

Similarly, with $\frac{dV_{\text{SPZT}}}{dl} \approx 70 \text{ V/order}$, we get

$$\frac{\partial \nu}{\partial V_{\text{SPZT}}} = \frac{65 \text{ MHz}}{70 \text{ V}} \approx \frac{930 \text{ kHz}}{\text{volt}}$$

11/14/88 M.E.S.

MODE CLEANER SERVO PHOTODIODE / BUFFER (STANDARD CALTECH FRONT END)

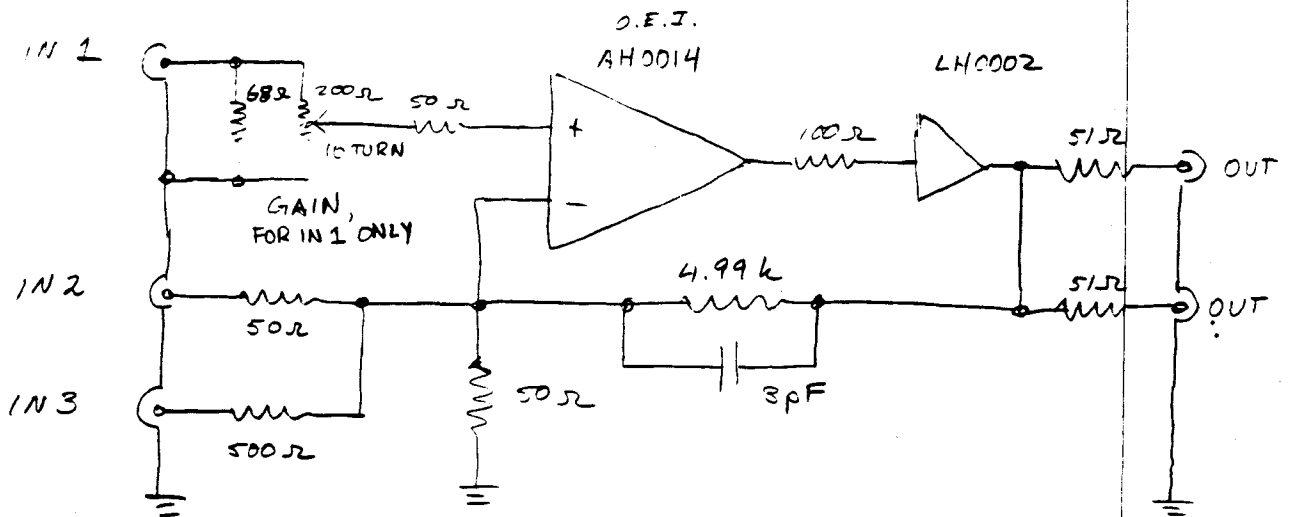


$Q \approx 4(\pm 2)$
 $\omega_0 = 12.3 MHz \times 2\pi$

R.F. DARK NOISE \approx SHOT NOISE AT $1mW$, 5145 \AA

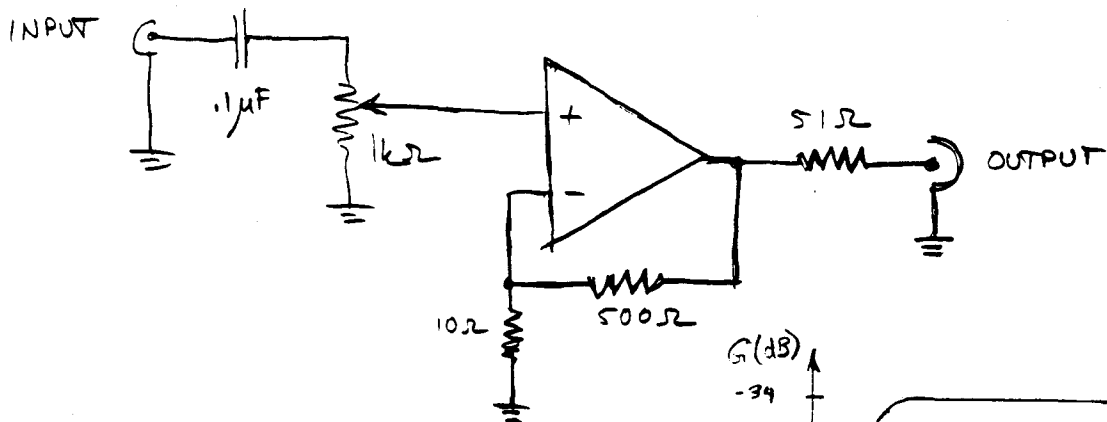
11/14/88 MEC

MODE CLEANER SERVO "FAST #3" AMPLIFIER

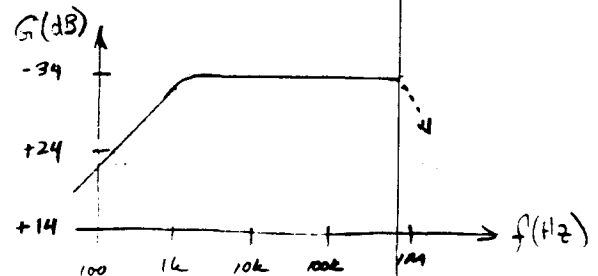


GAIN; IN 1 $\times 0 \rightarrow \times 100$, DC - 10 MHz
 IN 2 $\times -100$
 IN 3 $\times 10$

"50k" AMPLIFIER

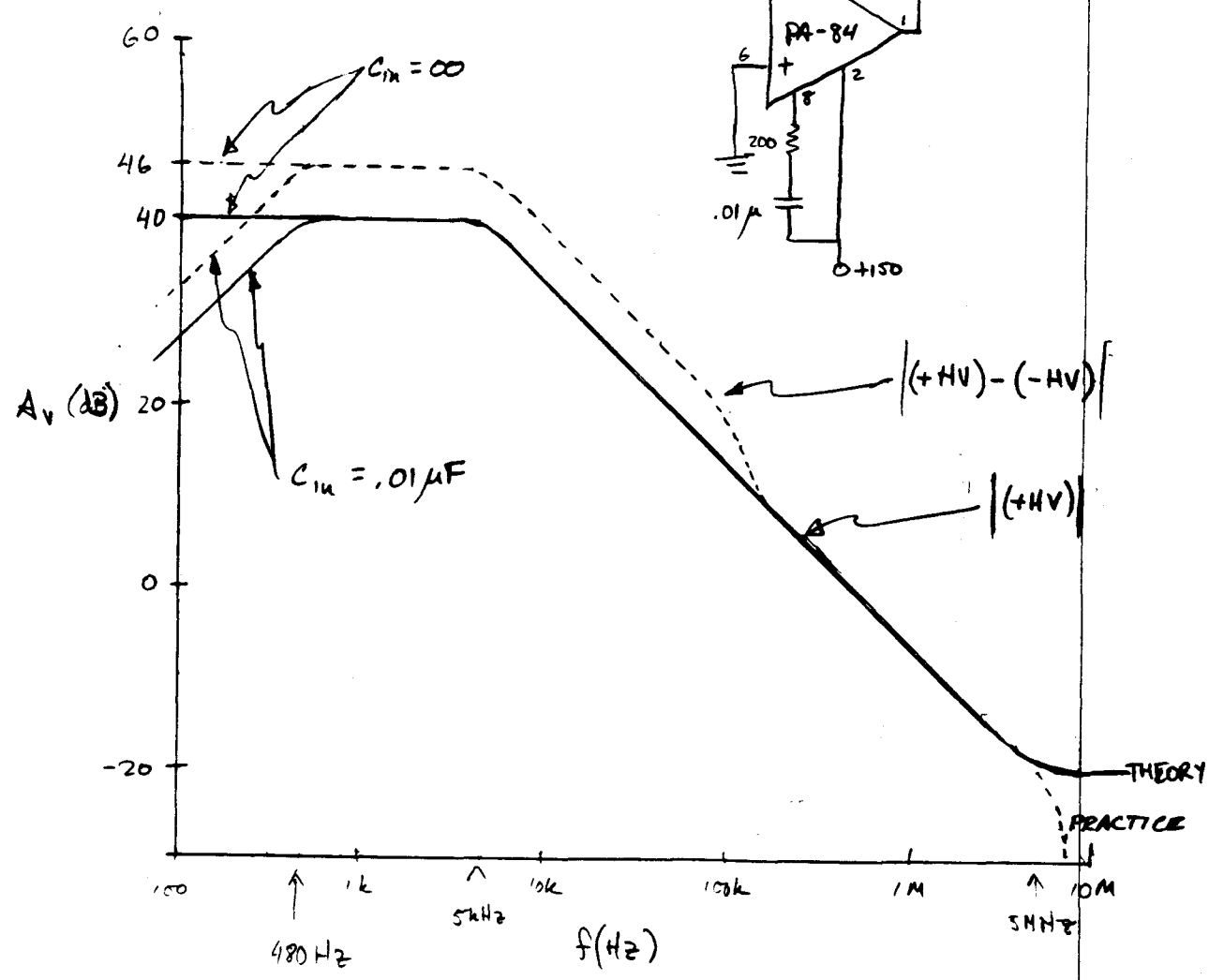
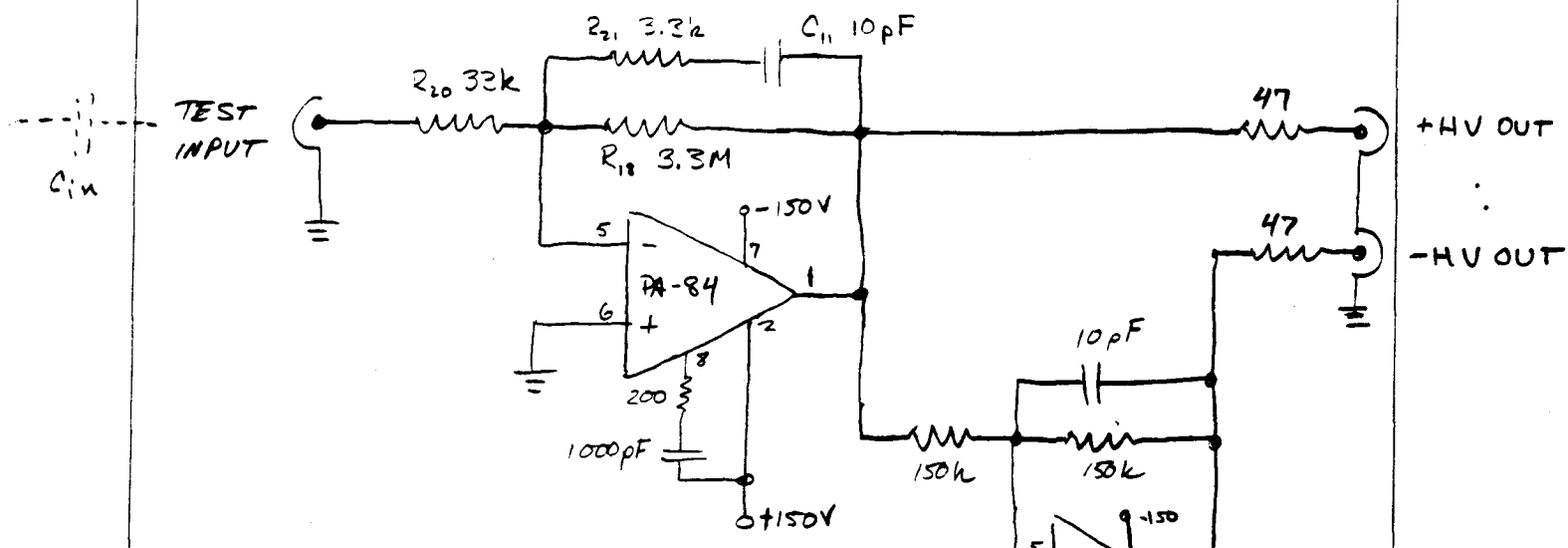


GAIN; $\times 50$ MAX., 1.5 kHz $\rightarrow \sim 700$ kHz



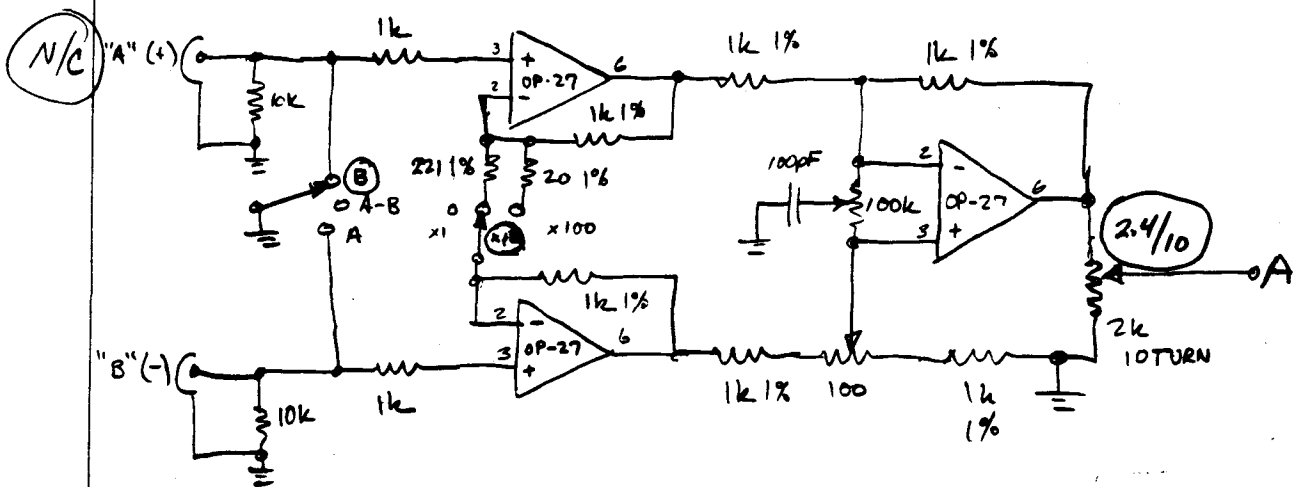
11/15/88 MES

MODE CLEANER SERVO LOOP "HV#6" AMP
 (IMPORTANT CIRCUITRY SHOWN) - SEE GENERAL
 "HV AMP" SCHEMATIC FOR FULL DETAILS)



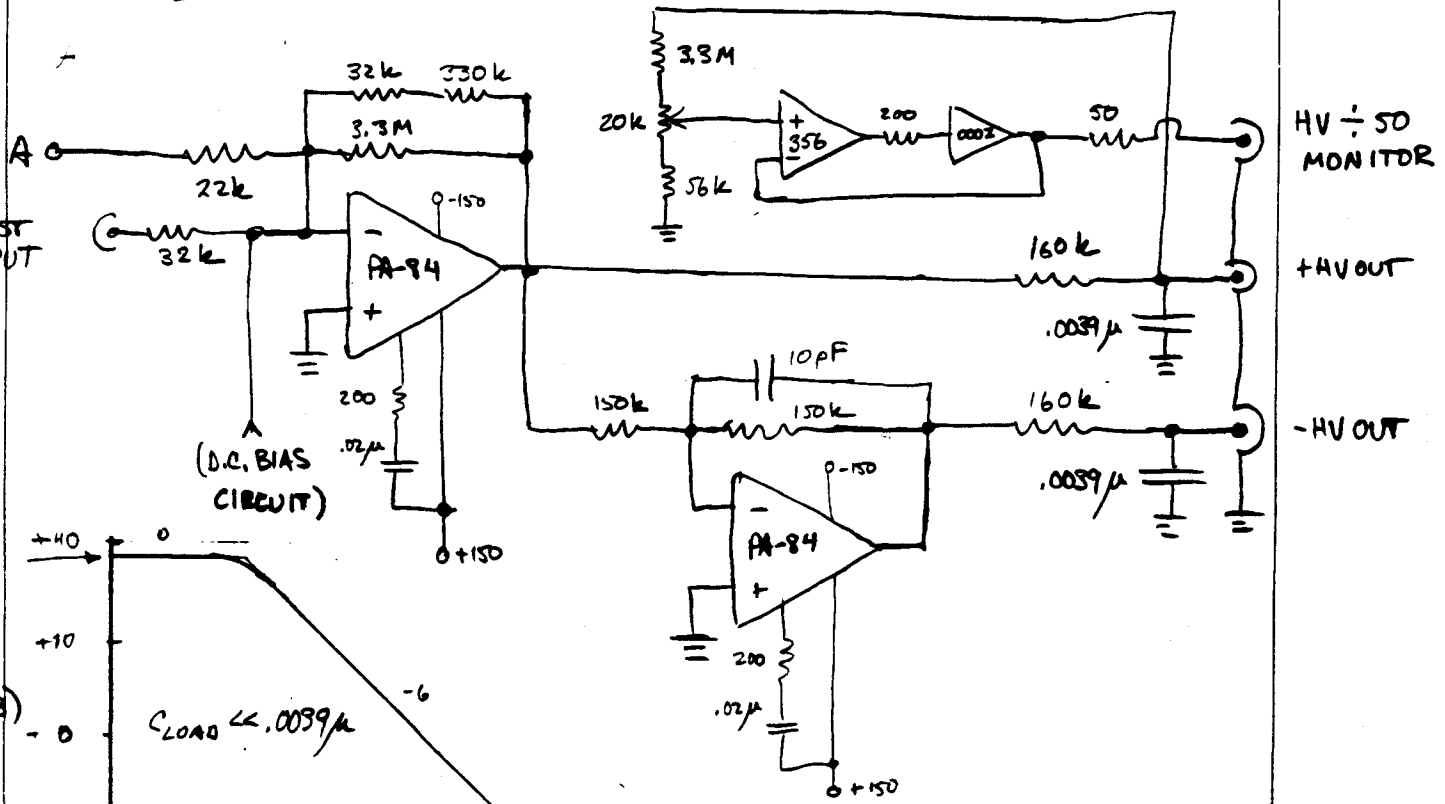
11/15/88 WSG

MODE CLEANER SERVO LOOP "HV ±1" AMP
(IMPORTANT CIRCUITRY SHOWN - SEE GENERAL
"HV AMP" SCHEMATIC FOR DETAILS).



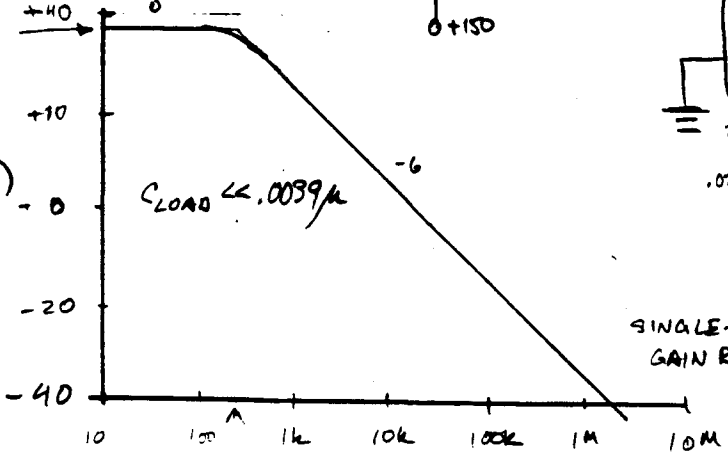
N/C

TEST INPUT



+37dB

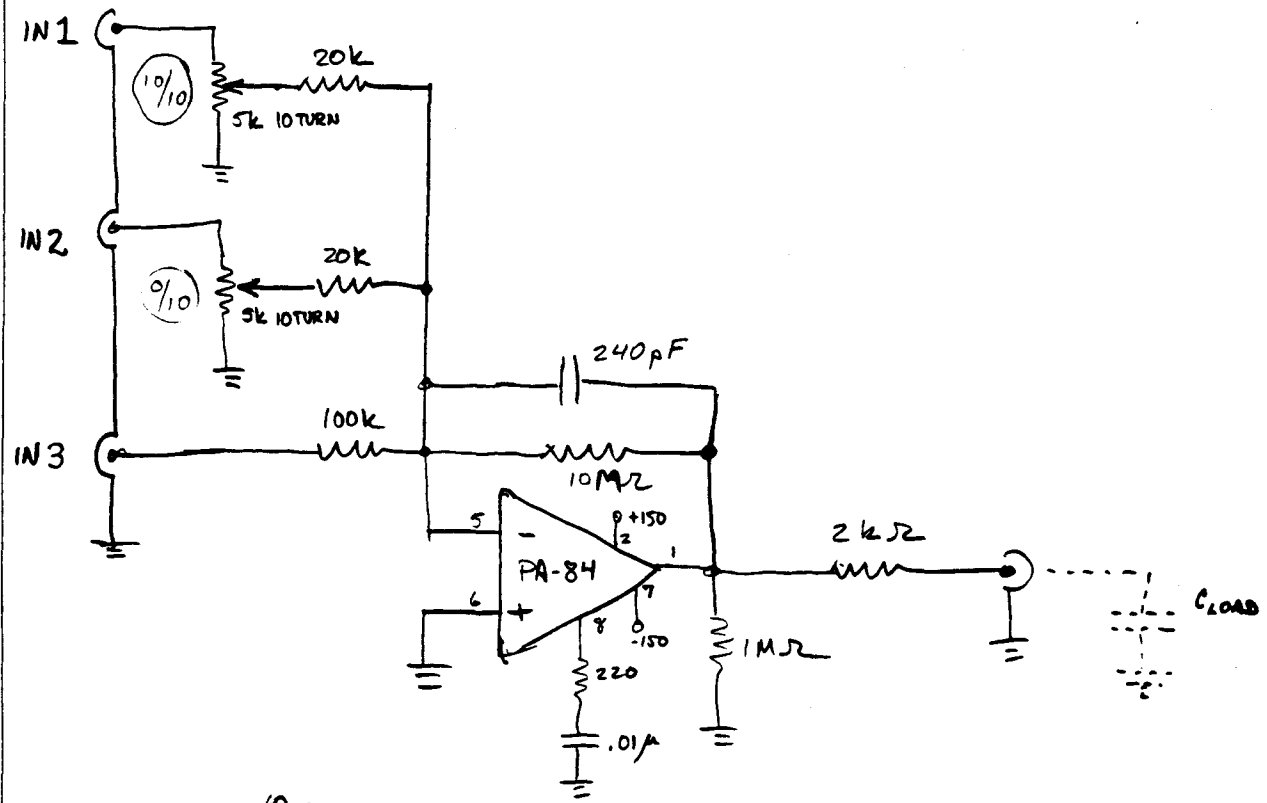
A_v (dB)



SINGLE-ENDED,
GAIN RANGE $\times 10$, VERNIER $\frac{2.4}{10}$

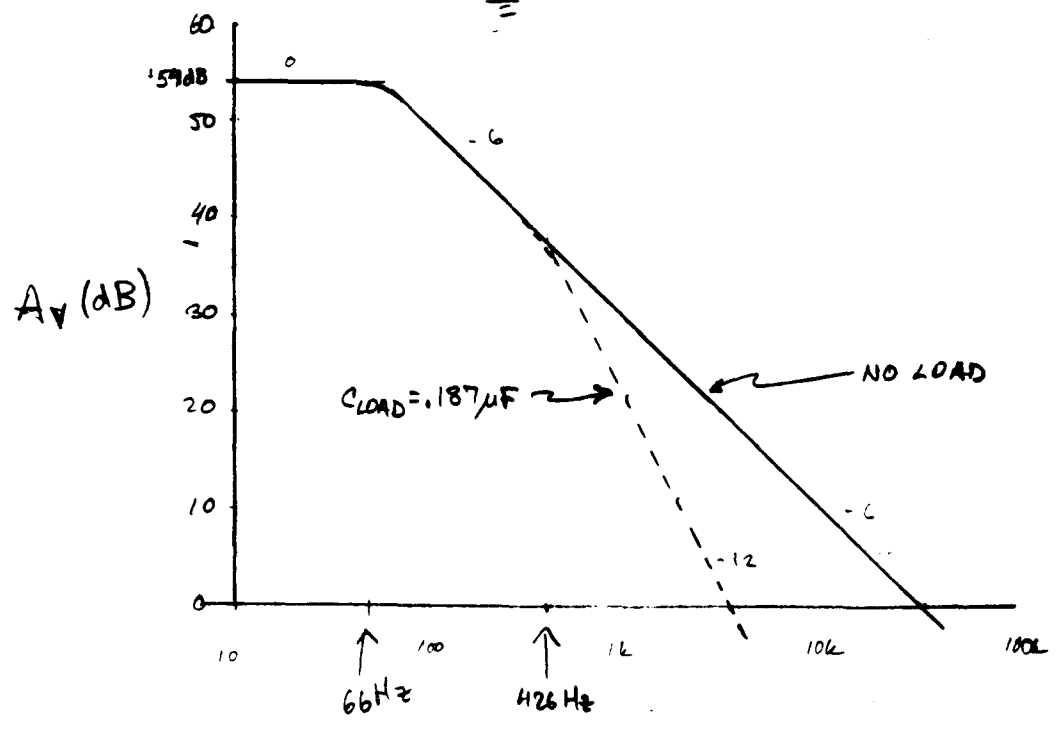
11/15/88 WEF

MODE CLEANER SERVO LOOP "VECTOR H.V. AMP"



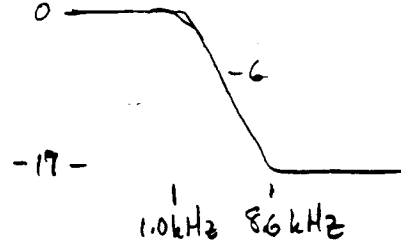
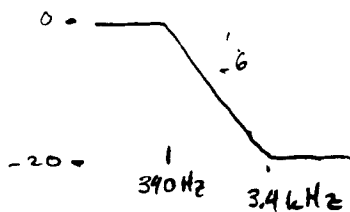
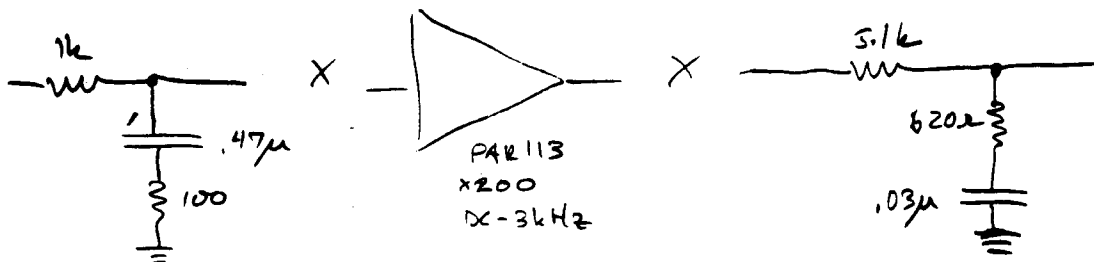
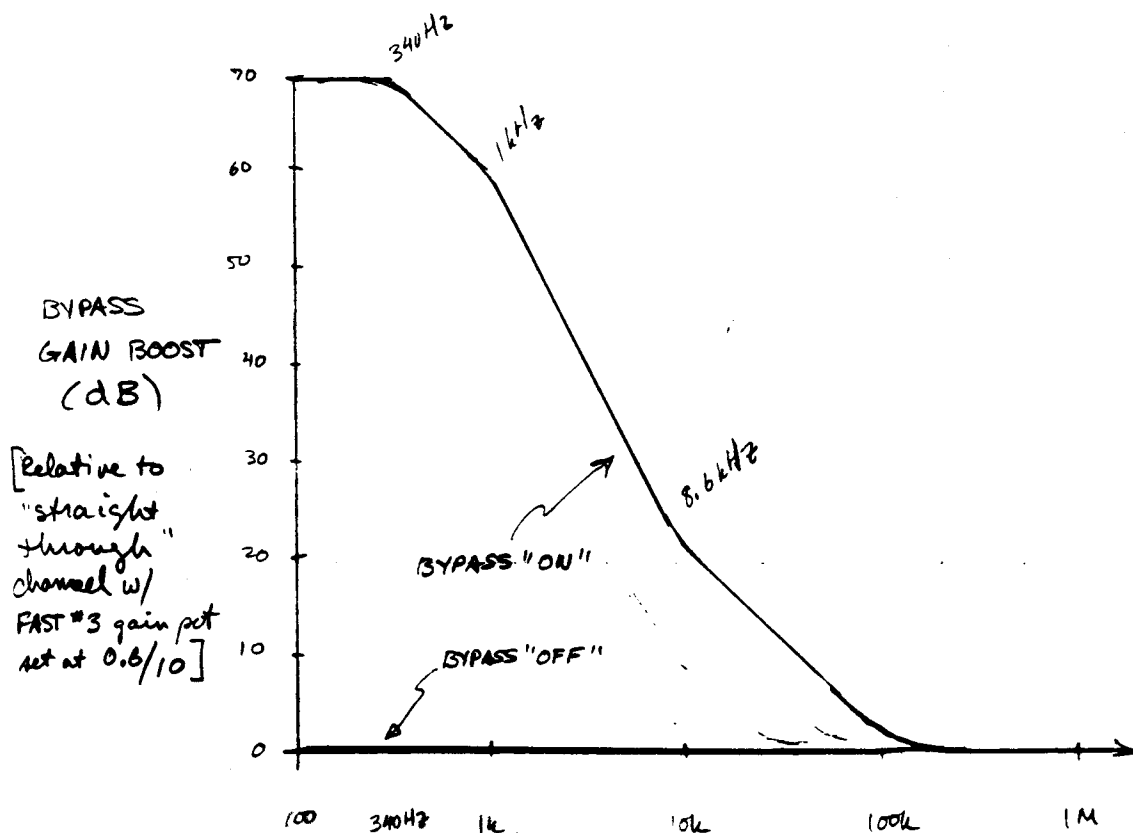
N/C

N/C



11/15/88 MEZ

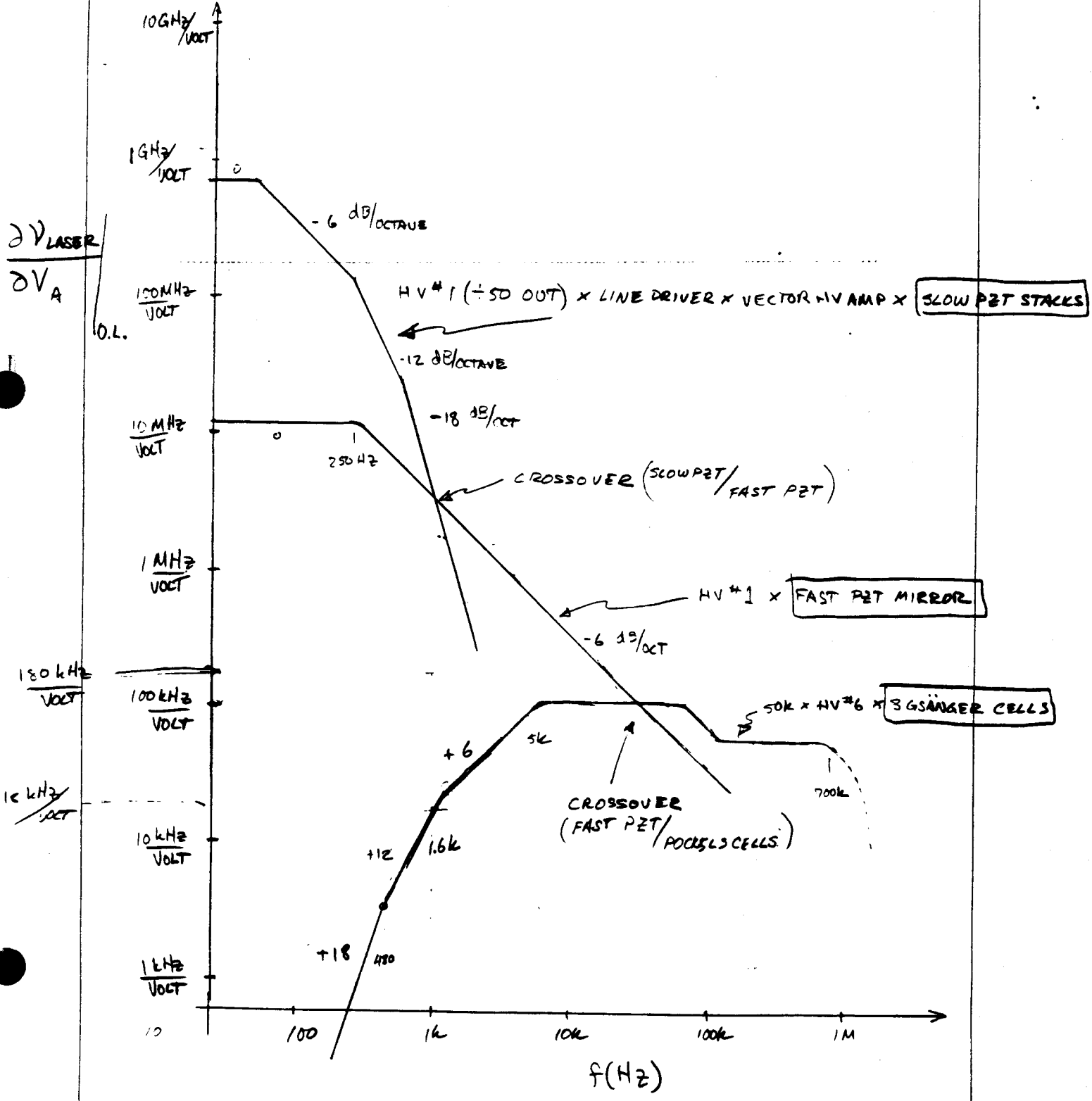
MODE CLEANER SERVO BYPASS GAIN ENHANCEMENT



11/15/88 MEZ

MODE CLEANER SERVO ELECTRONIC AND TRANSDUCER OPEN-LOOP TRANSFER FUNCTION

LASER FREQUENCY EXCURSION PER VOLT APPLIED AT POINT "A" (OPEN-LOOP)



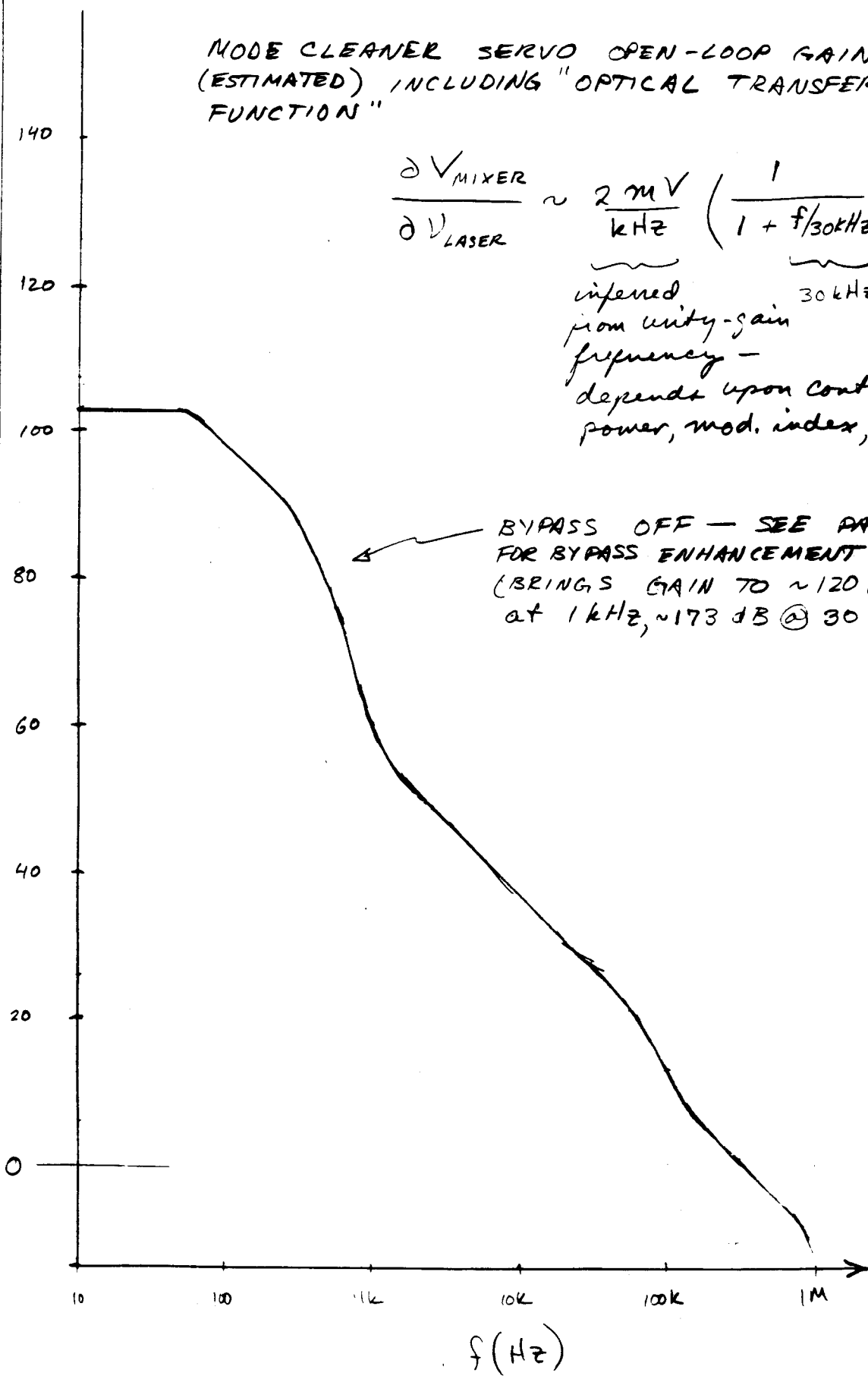
11/15/88 MEZ

MODE CLEANER SERVO OPEN-LOOP GAIN (ESTIMATED) INCLUDING "OPTICAL TRANSFER FUNCTION"

$$\frac{\partial V_{MIXER}}{\partial V_{LASER}} \sim \frac{2 \text{ mV}}{\text{kHz}} \left(\frac{1}{1 + f/30\text{kHz}} \right)$$

inferred from unity-gain frequency - depends upon contrast, power, mod. index,

$30 \text{ kHz} = \frac{1}{4\pi \tau_c}$



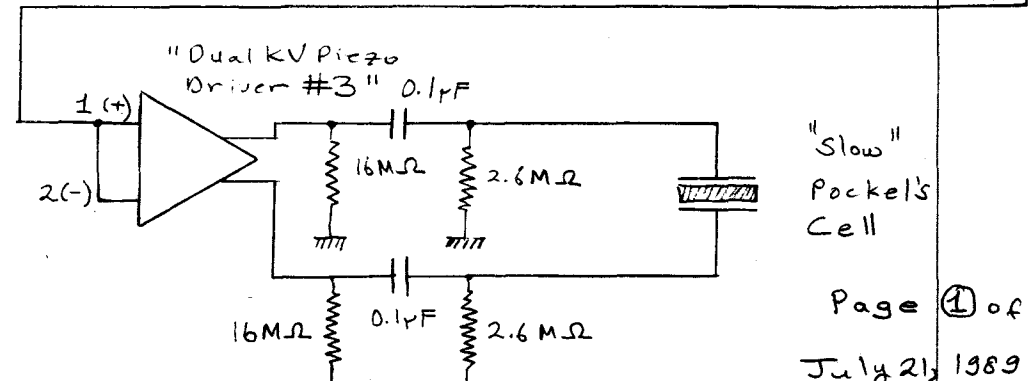
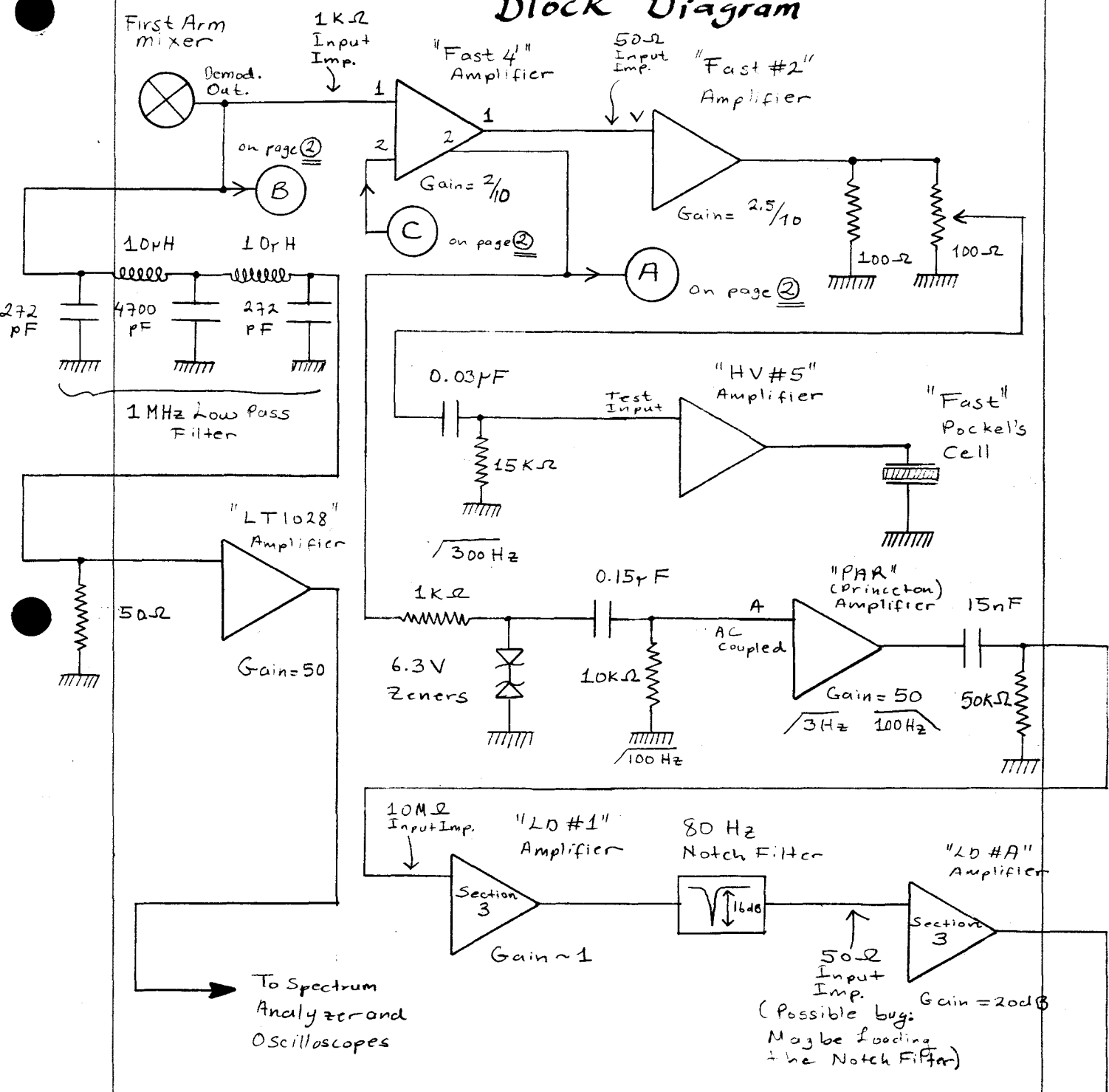
← BYPASS OFF - SEE PAGE 8 FOR BYPASS ENHANCEMENT (BRINGS GAIN TO ~120 dB at 1 kHz, ~173 dB @ 30 Hz)

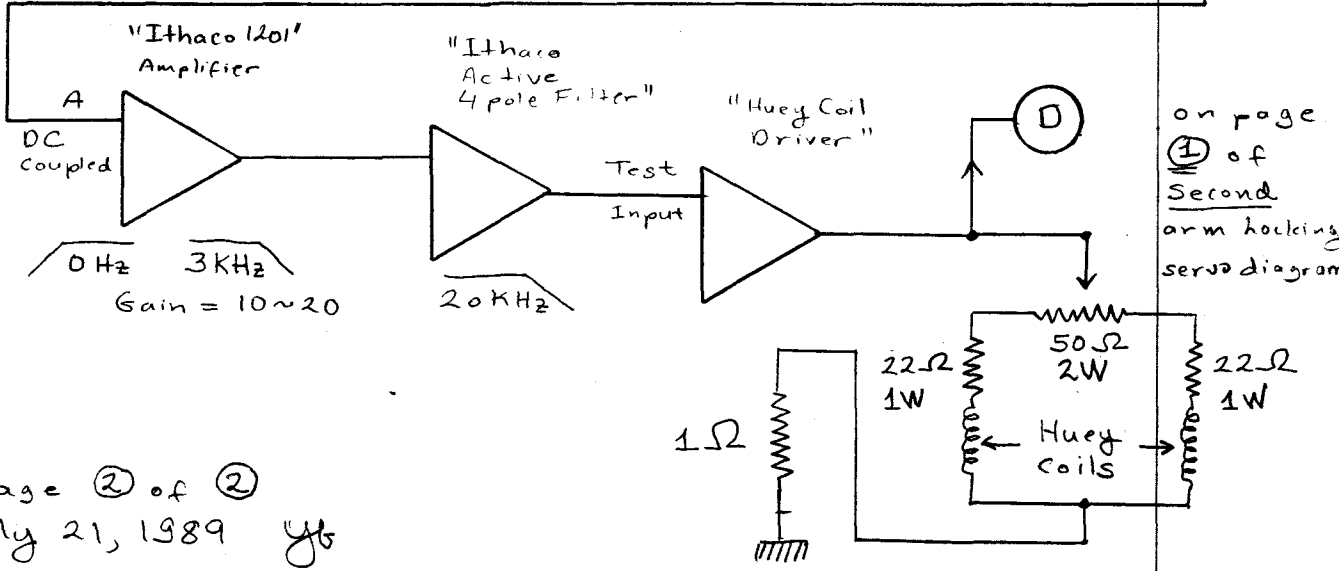
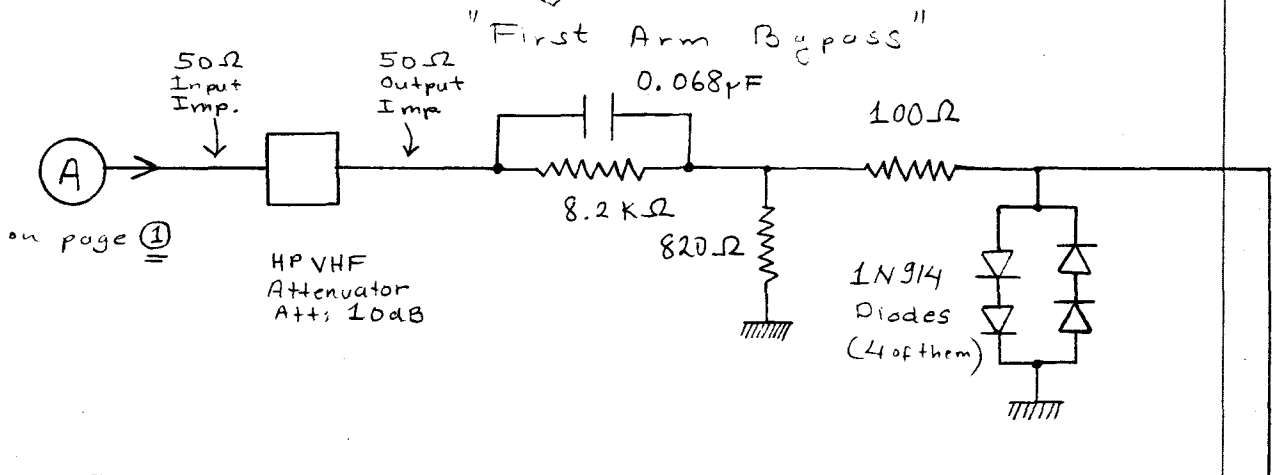
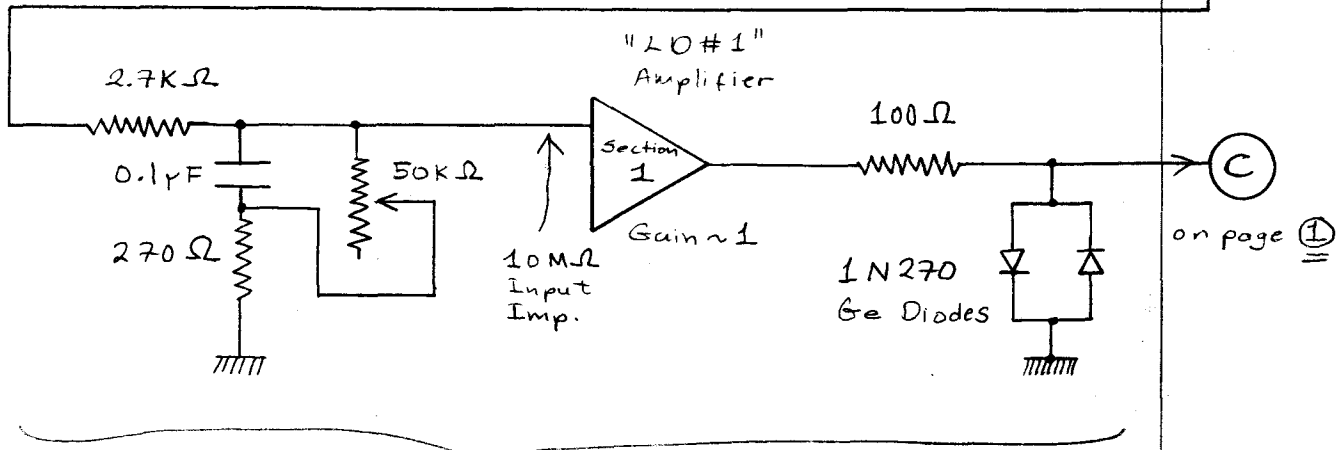
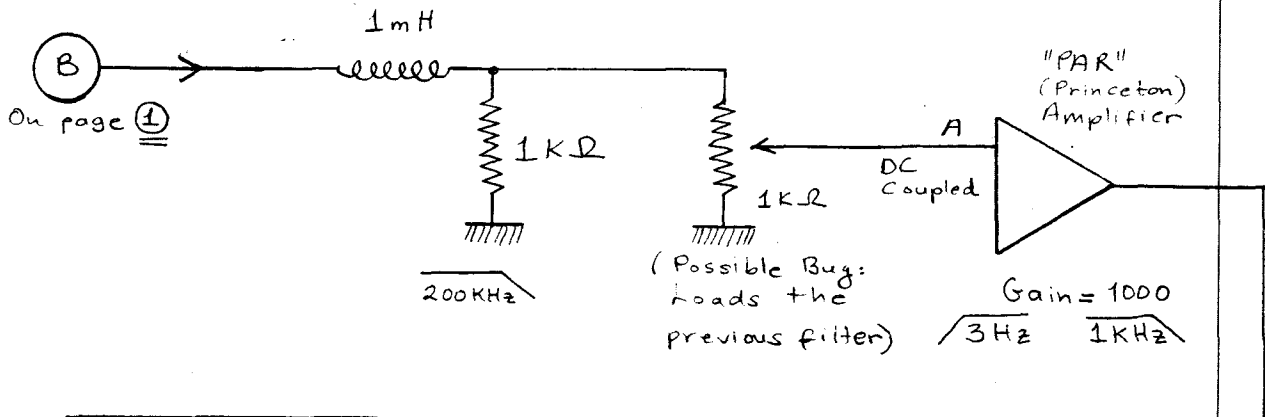
AOL
(5)

BATCH
START

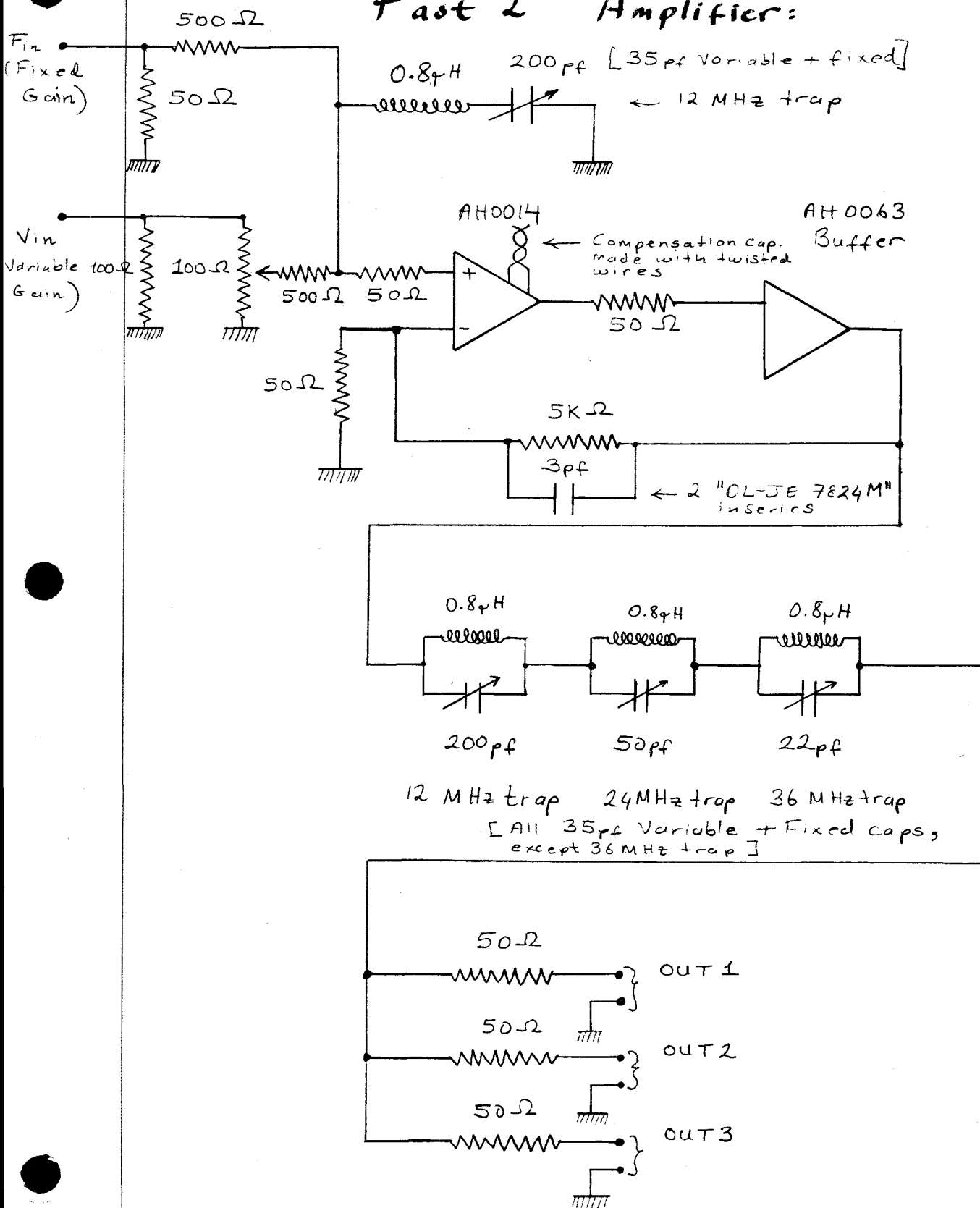
STAPLE
OR
DIVIDER

First Arm Servo Loop Block Diagram



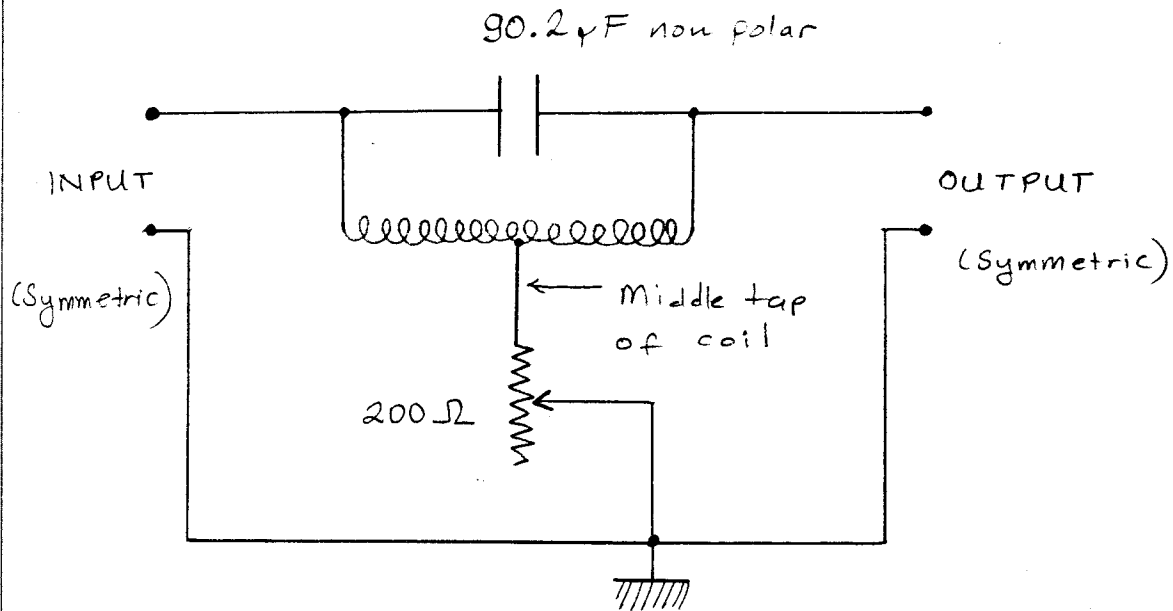


'Fast 2' Amplifier:



July 24, 1989
 YG

80 Hz Notch Filter:



The coil is chosen in such a way that the resonance of the LC combination is $80 \pm 1 \text{ Hz}$.

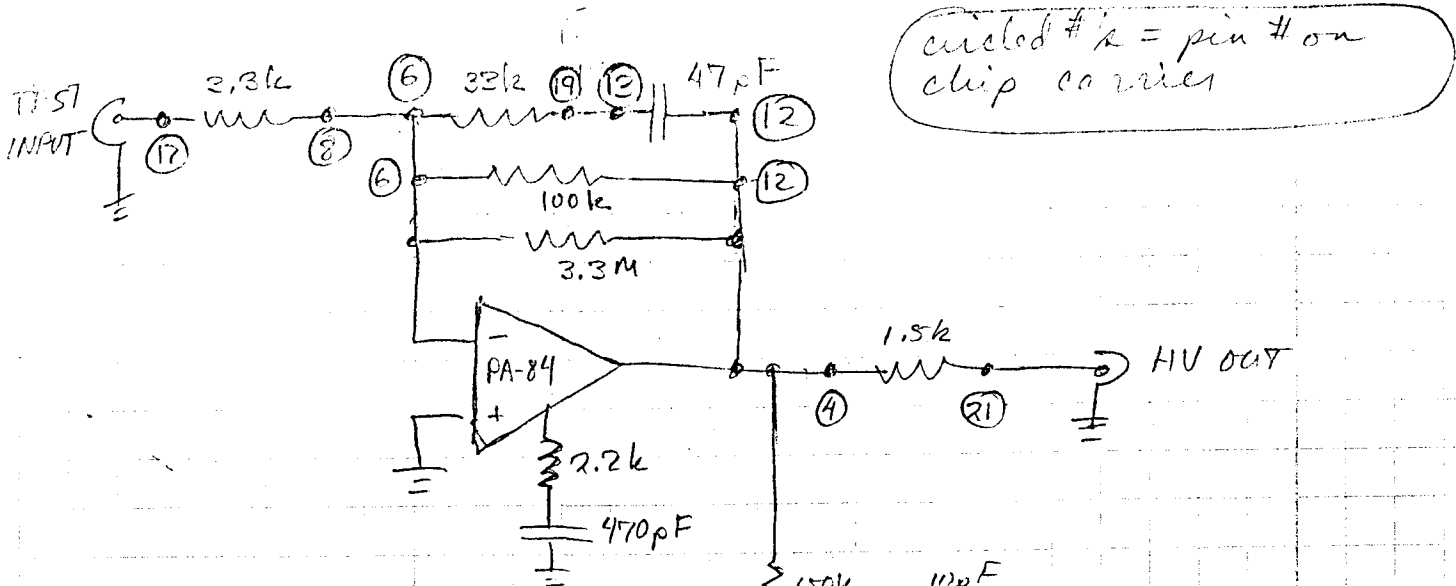
July 24, 1989 YB

2/24/88

LABELLED "Y"

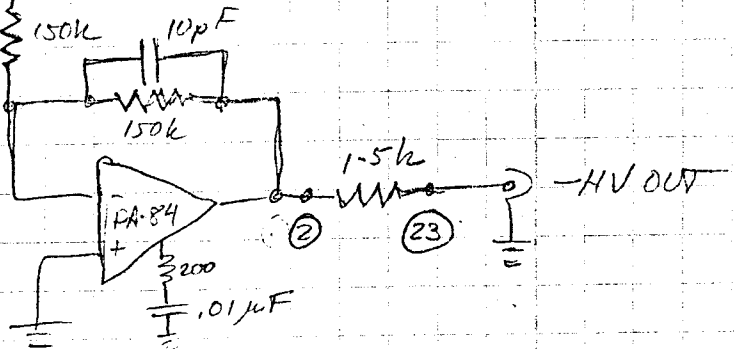
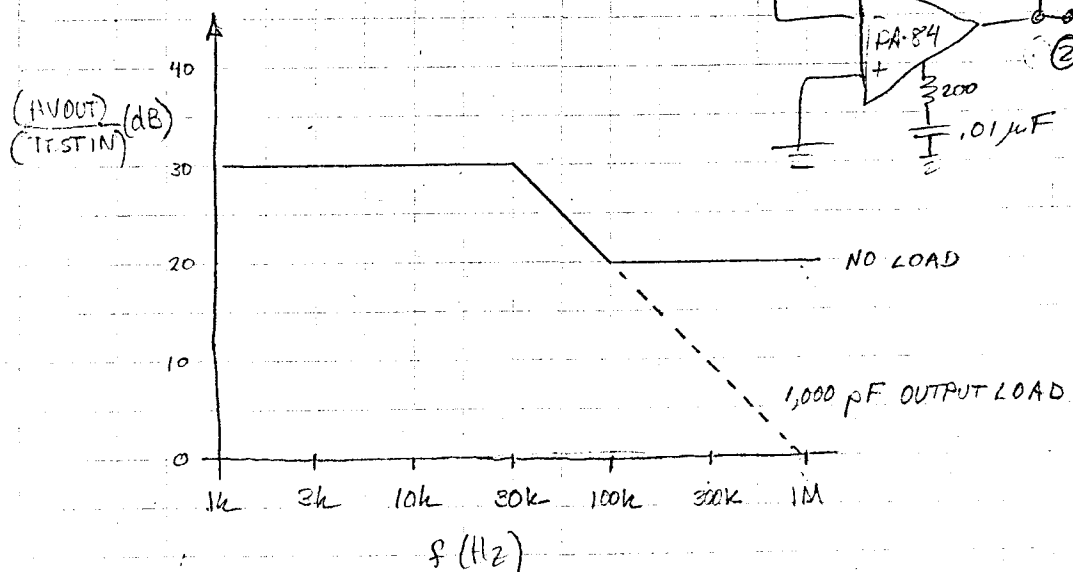
Quotient H.V. amp has been modified for use in phase connecting loop (to replace/supplement "push" amplifier used for part M.C. in previous setups).

Relevant parts of circuit;

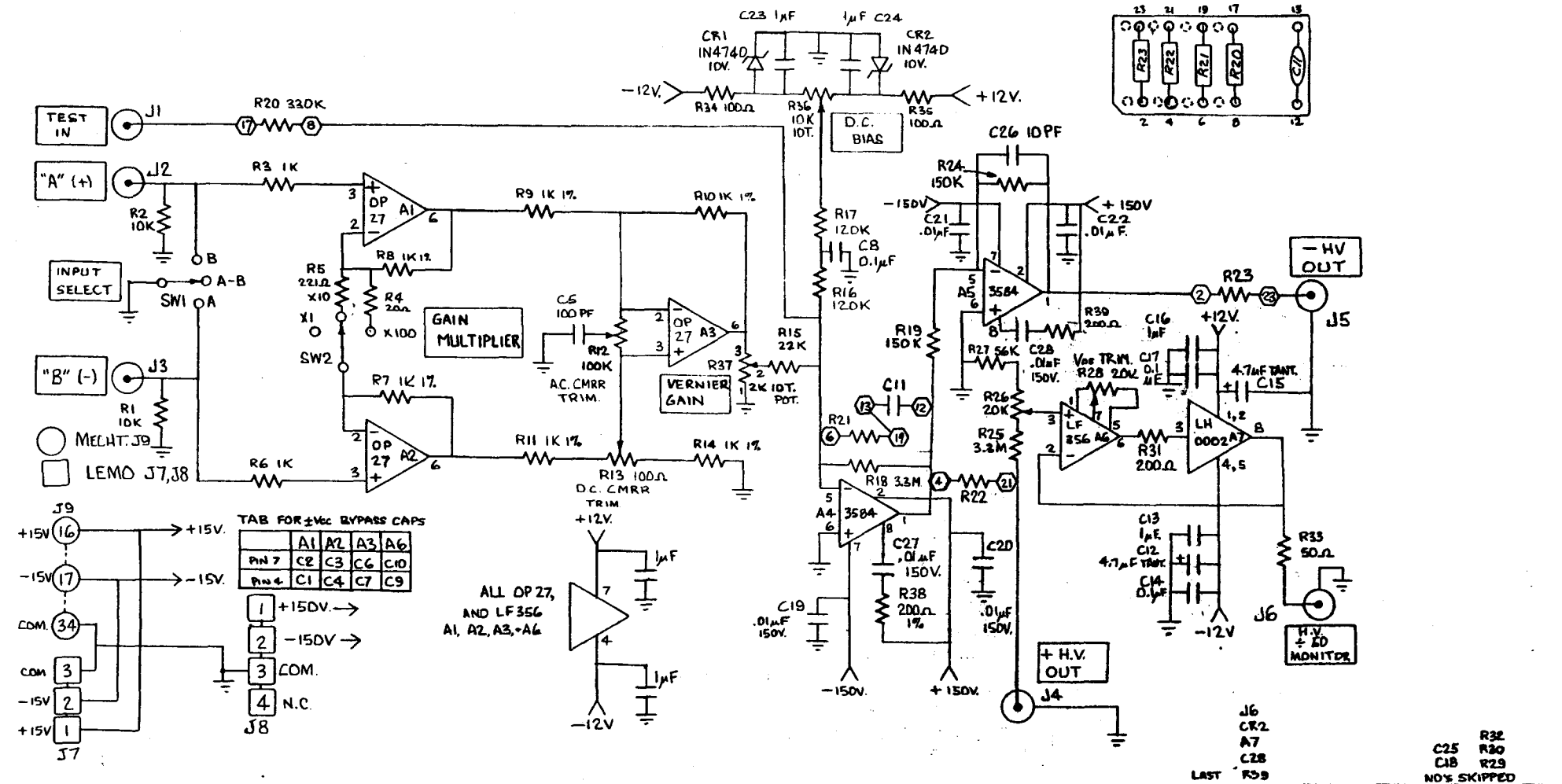


circled #'s = pin # on chip carrier

Transfer Function (theoretical);



Actual transfer function very close to theoretical, but peaky at ~2MHz

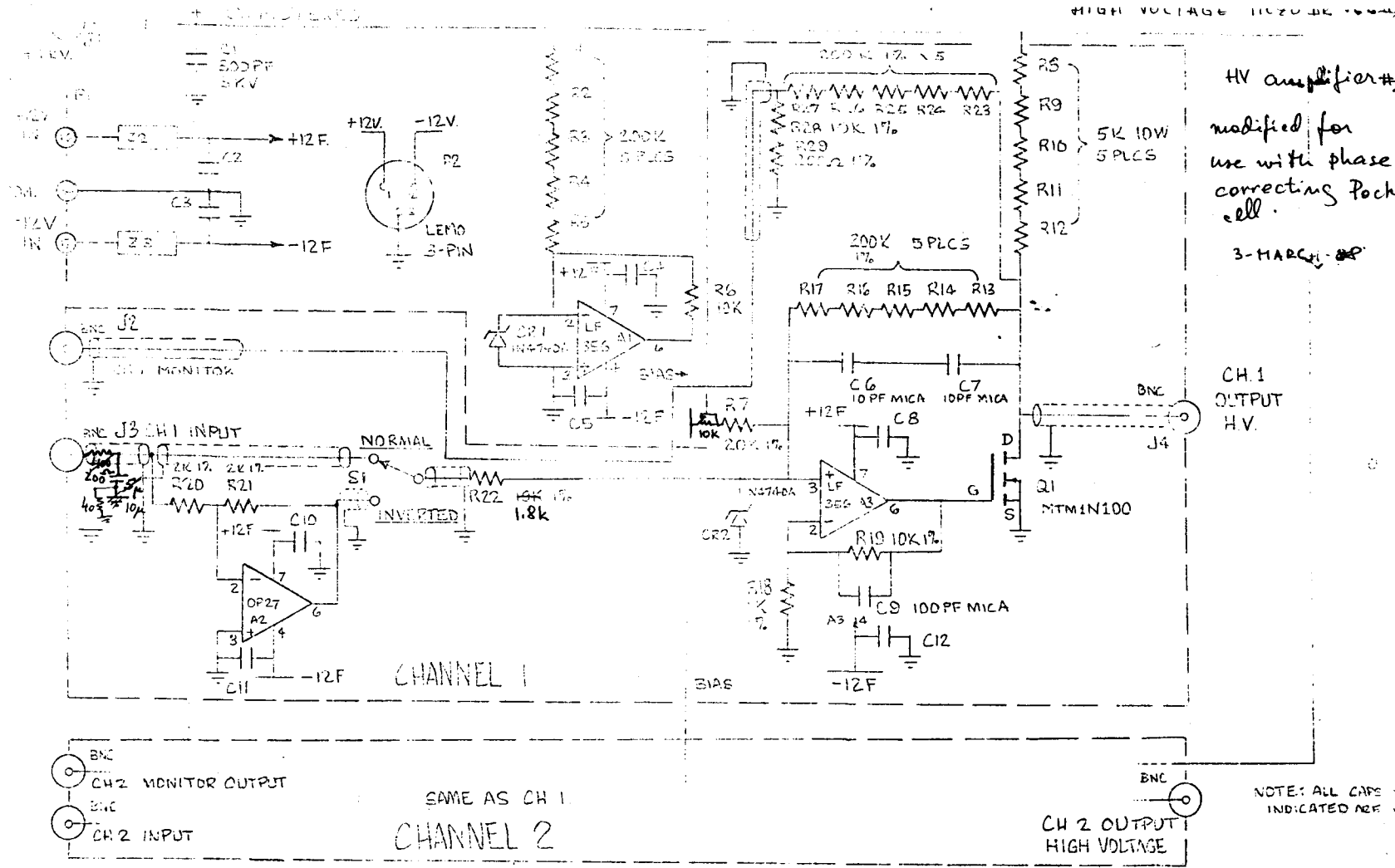


2. J7, J8, and J9 ARE ALL MOUNTED ON BACK PANEL.
 NOTES 1. J INDICATES JUMPER TO SHORT.

SMARTWRK FILE C:\SMARTWRK HVAMP.PCB
 DRAWN FROM E.LINDELEF OF 9-1-87 PRO HVAMP.PL
 ADDED COMP. CARRIED PIN NO'S TO DWG.

REPLACES 87-0901-1

CALIFORNIA INSTITUTE OF TECHNOLOGY GRAVITATIONAL PHYSICS		R32 R30 R29 R23
HV AMPLIFIER ± 150V.		C25 C18 R29 NO'S SKIPPED
DRAWN BY B.TINKER	DATE 3-4-88	DRAWING NO.
CHECKED BY	SCALE	
APPROVED BY	W.O.	



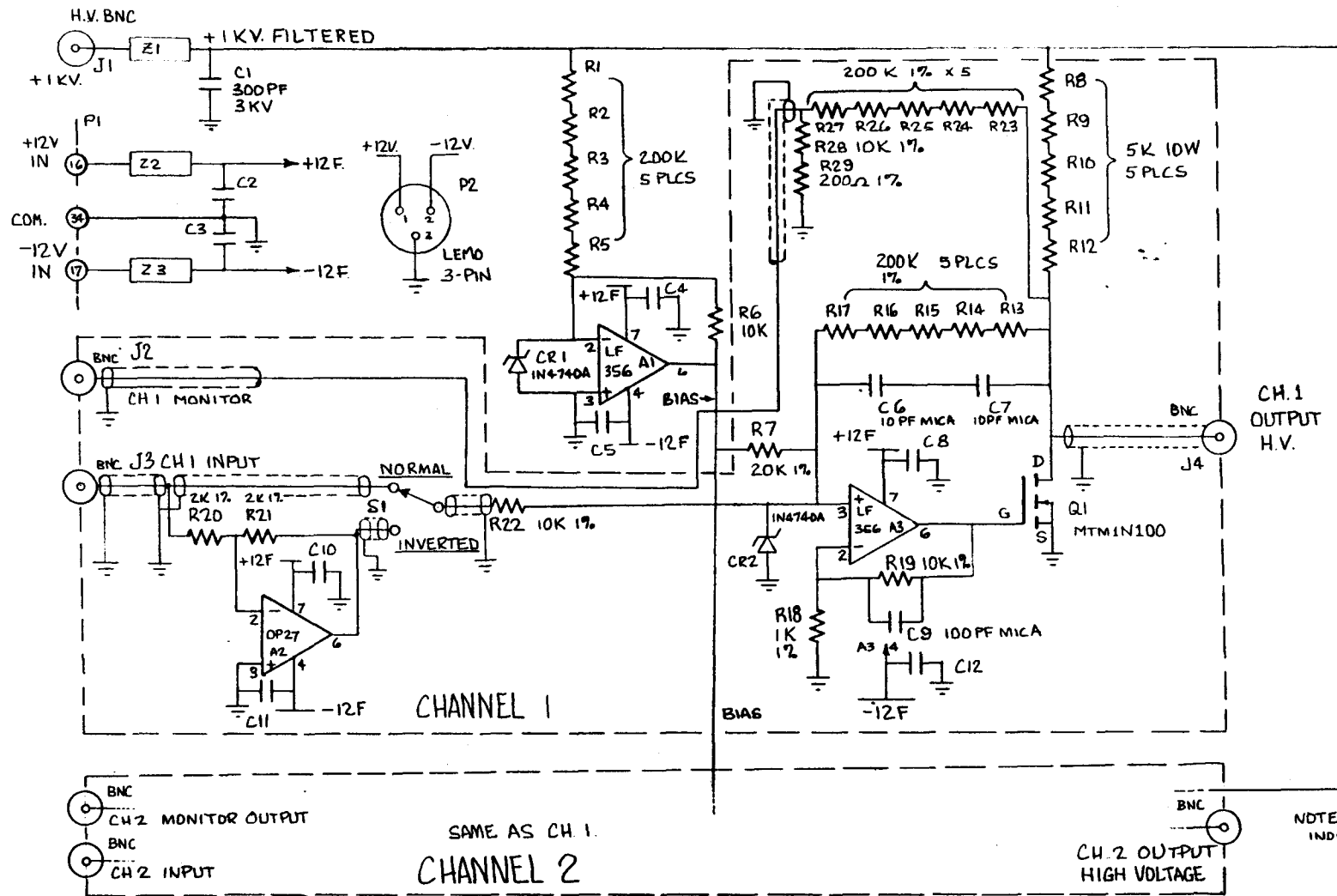
HV amplifier #
 modified for
 use with phase
 correcting Poche
 cell.

3-MARCH-88

CH.1
 OUTPUT
 H.V.

SAME AS CH.1.
 CHANNEL 2

CH.2 OUTPUT
 HIGH VOLTAGE

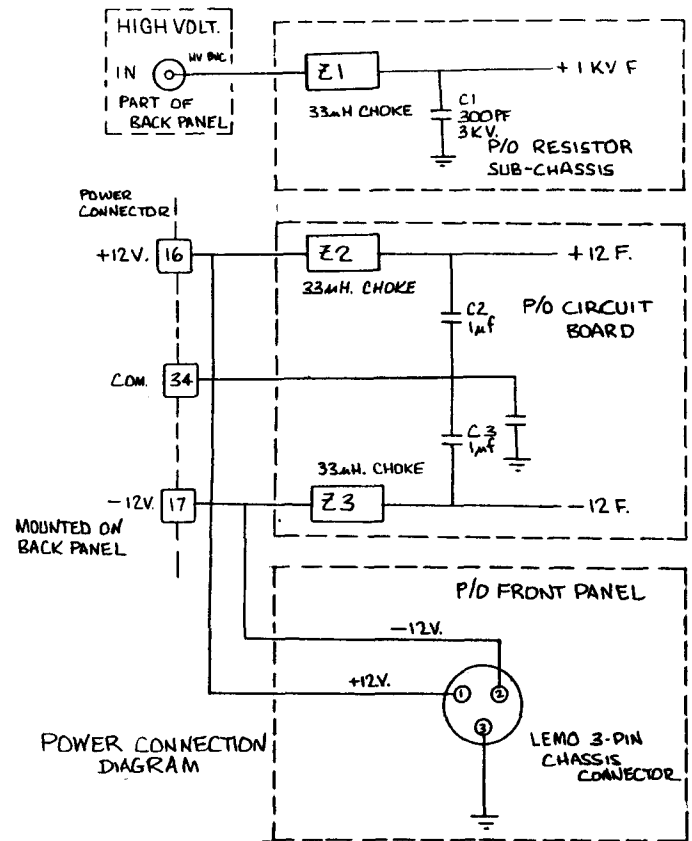


- R29
- C12
- CR2
- Q1
- A3
- Z3
- J4
- P2
- S1

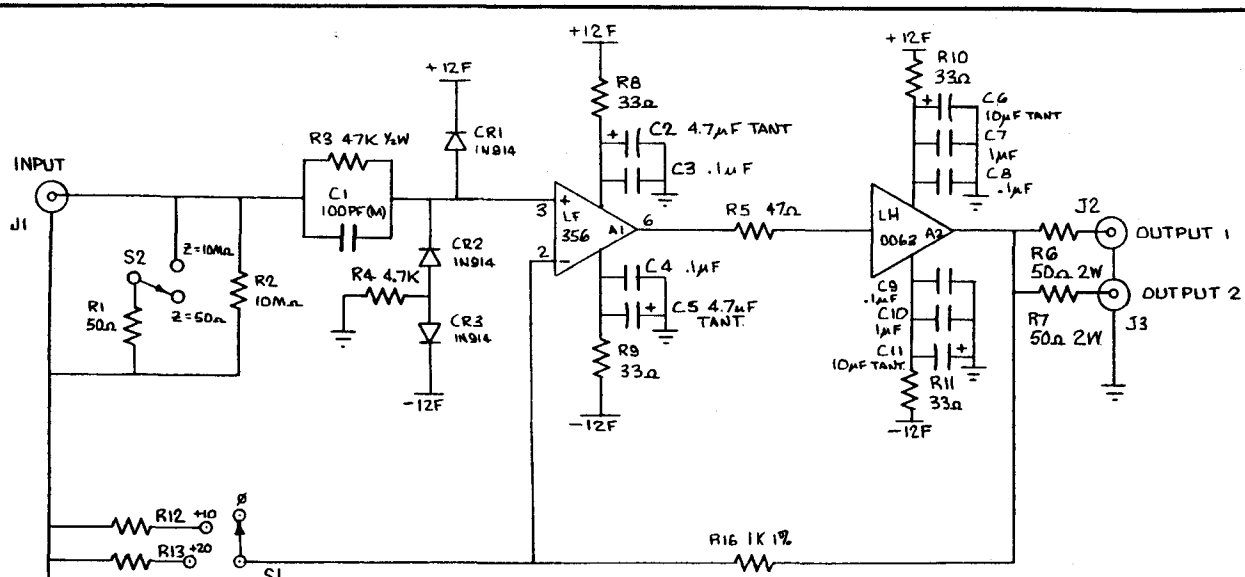
WARNING: CIRCUITS MAY CARRY UP TO 1KV.

CALIFORNIA INSTITUTE OF TECHNOLOGY GRAVITATIONAL PHYSICS		
DUAL KILOVOLT PIEZO DRIVER		
DRAWN BY B. TINKER	DATE 11/24/87	DRAWING NO.
CHECKED BY	SCALE	-1
APPROVED BY	W.D.	

87-1124-1



CALIFORNIA INSTITUTE OF TECHNOLOGY GRAVITATIONAL PHYSICS		
DUAL KILOVOLT PIEZO DRIVER		
DRAWN BY B. TINKER	DATE 2-2-88	DRAWING NO. -2
CHECKED BY	SCALE	
APPROVED BY	W.O.	



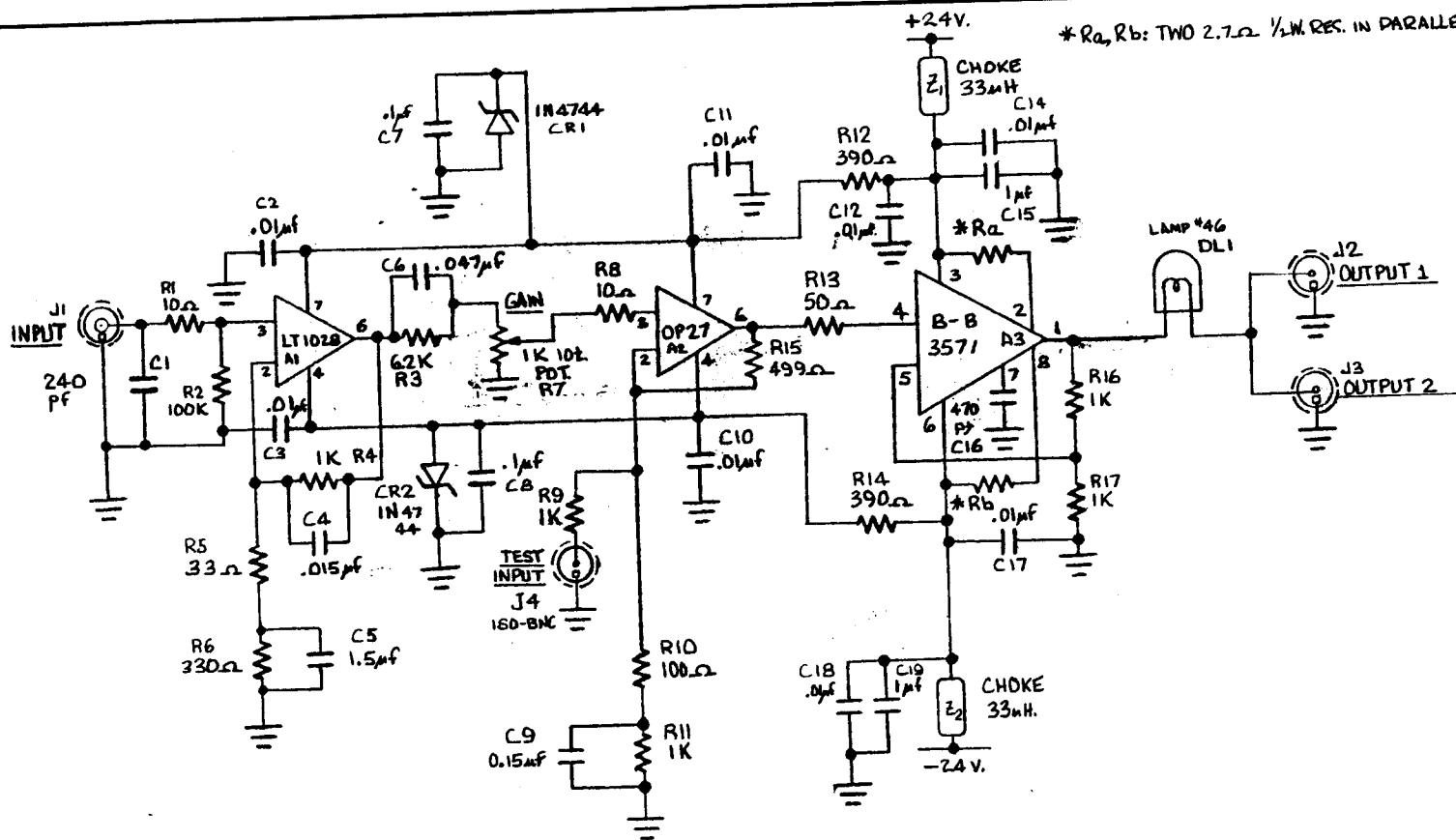
SUGGESTIONS:
 R12 = 462.47Ω (+10db)
 R13 = 111.11Ω (+20db)
 R14 = 32.66Ω (+30db)
 R15 = 10.10Ω (+40db)

CHOOSE RESISTOR VALUES TO SET GAINS AS RIGHT USING STANDARD 1% RESISTOR VALUES. FOR R12 - R15

A 2
 CR3
 S2
 J3
 R16

L.M.T. C

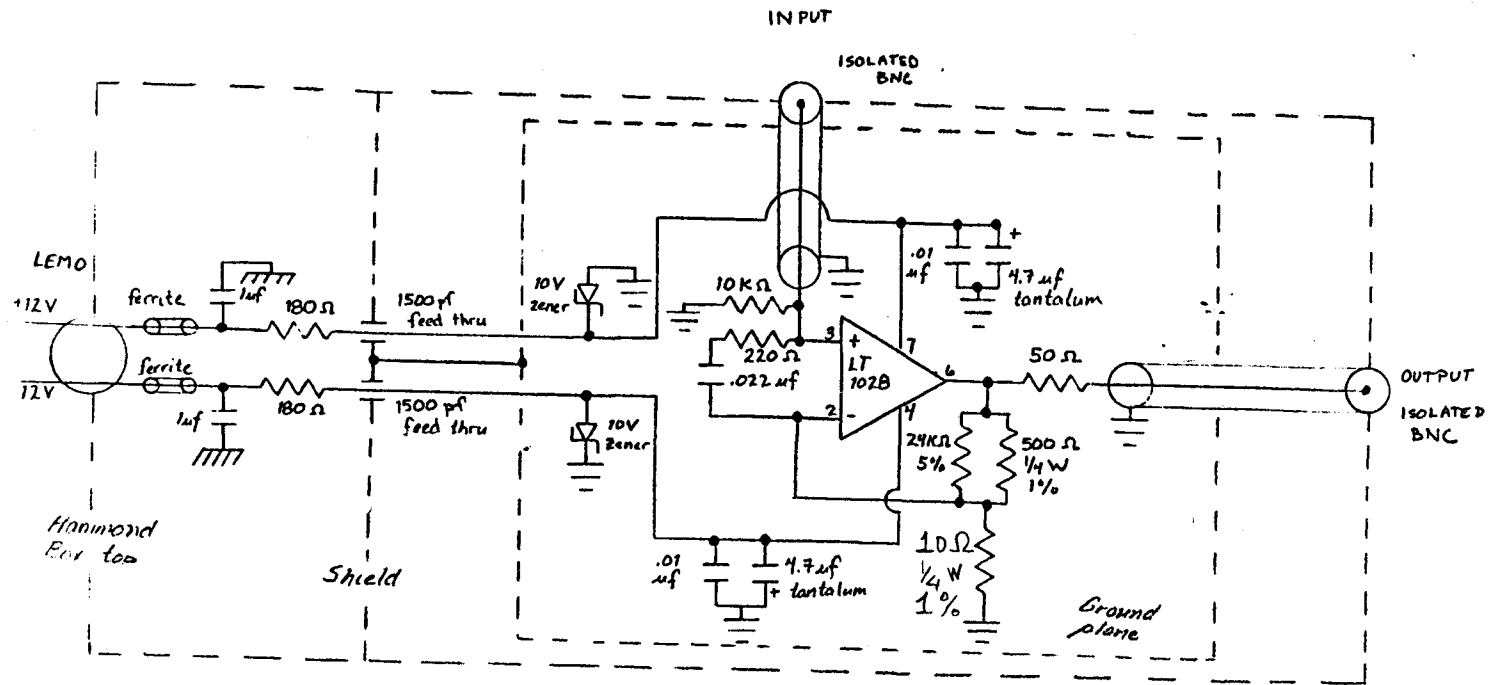
CALIFORNIA INSTITUTE OF TECHNOLOGY GRAVITATIONAL PHYSICS		
LINE DRIVER		
DRAWN BY B.T.	DATE 6-13-88	DRAWING NO.
CHECKED BY	SCALE	
APPROVED BY	W.O.	



3-1-88 R15 MOVED: 1K TO 500Ω
 1-28-88 C16 FROM 220pf TO 470pf
 UPDATE TO 9-8-78

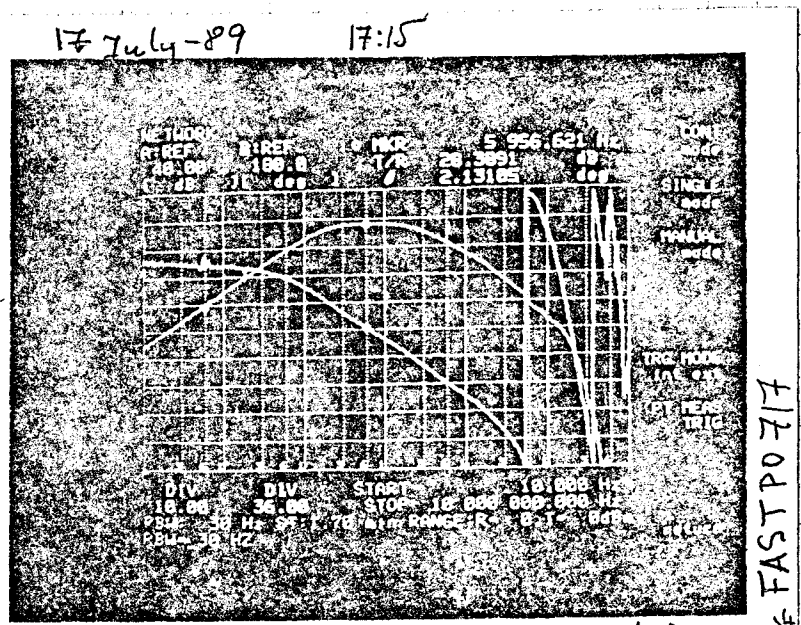
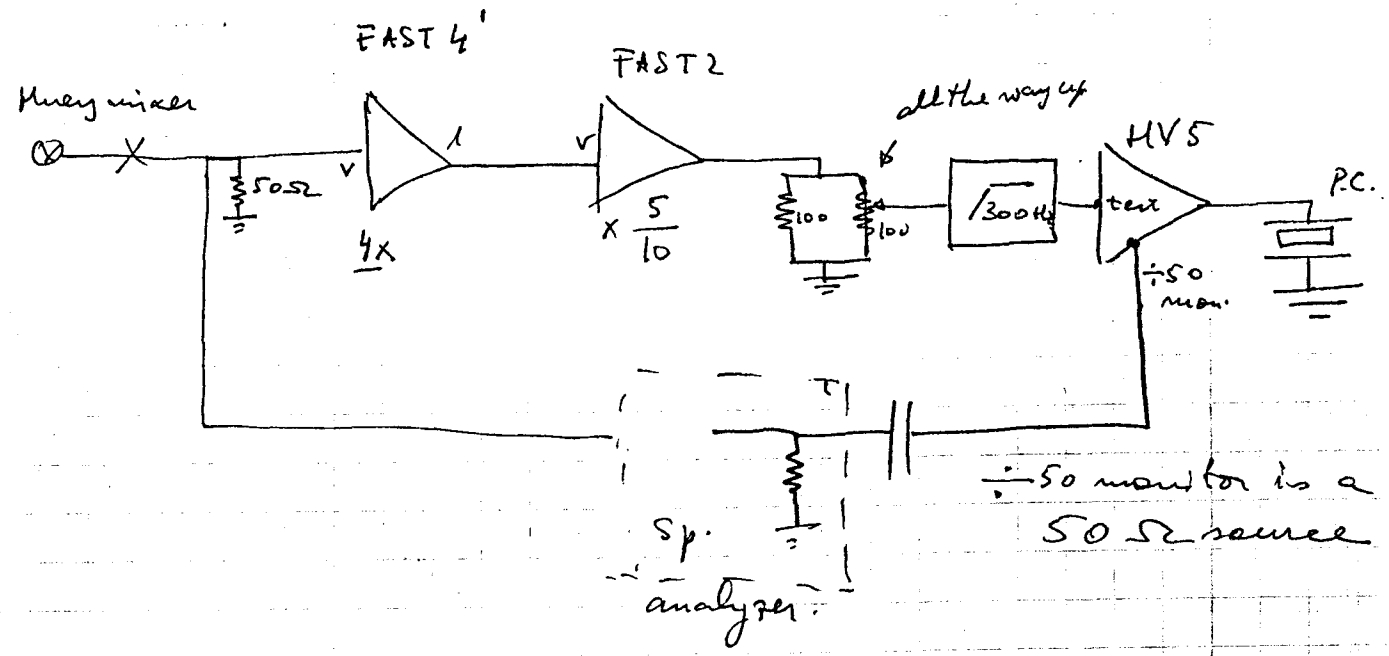
CALIFORNIA INSTITUTE OF TECHNOLOGY GRAVITATIONAL PHYSICS		
COIL DRIVER HIGH CURRENT		
DRAWN BY B.T.	DATE 9-2-87	DRAWING NO. -1
CHECKED BY	SCALE	
APPROVED BY	W.G.	

87-0902-1



LOW NOISE BUFFER
8-21-87

1st arm (phase correction) servo: fast rocket cell leg

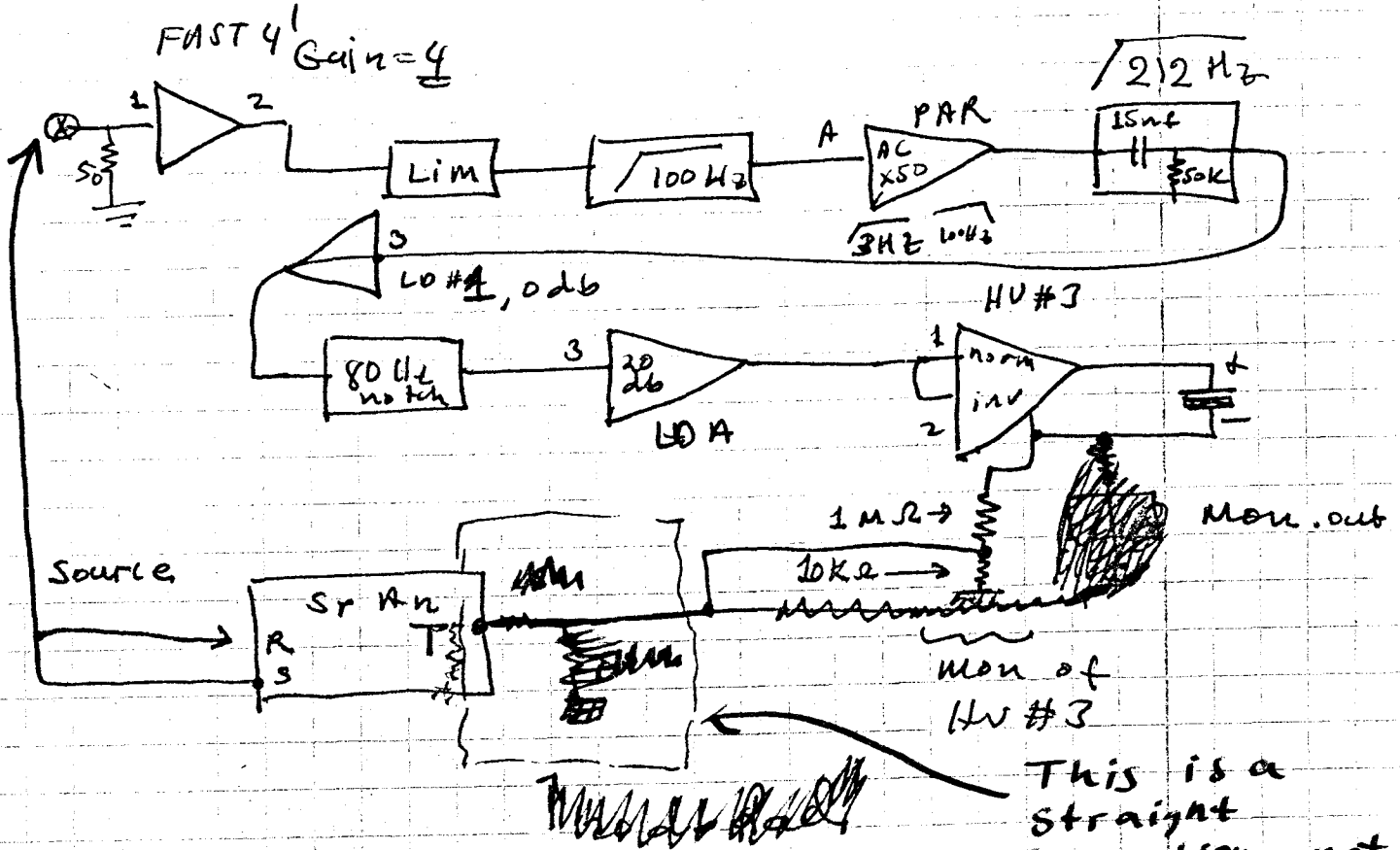


Fast PC leg transf. function, 50 ohm at inf of FAST

July 17, 1989
 JLS, ARJ, SH

Servo block frequency and phase response:

1st arm: slow pocket's cell loop:

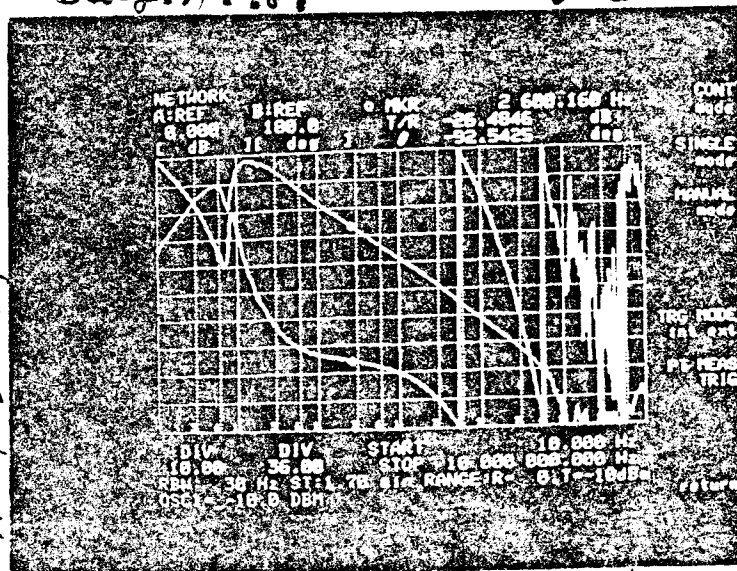


This is a straight connection, not a 20dB pad.

← real gain:
 add 92dB
 (re divider on sketch above)

July 17, 1989 16:20

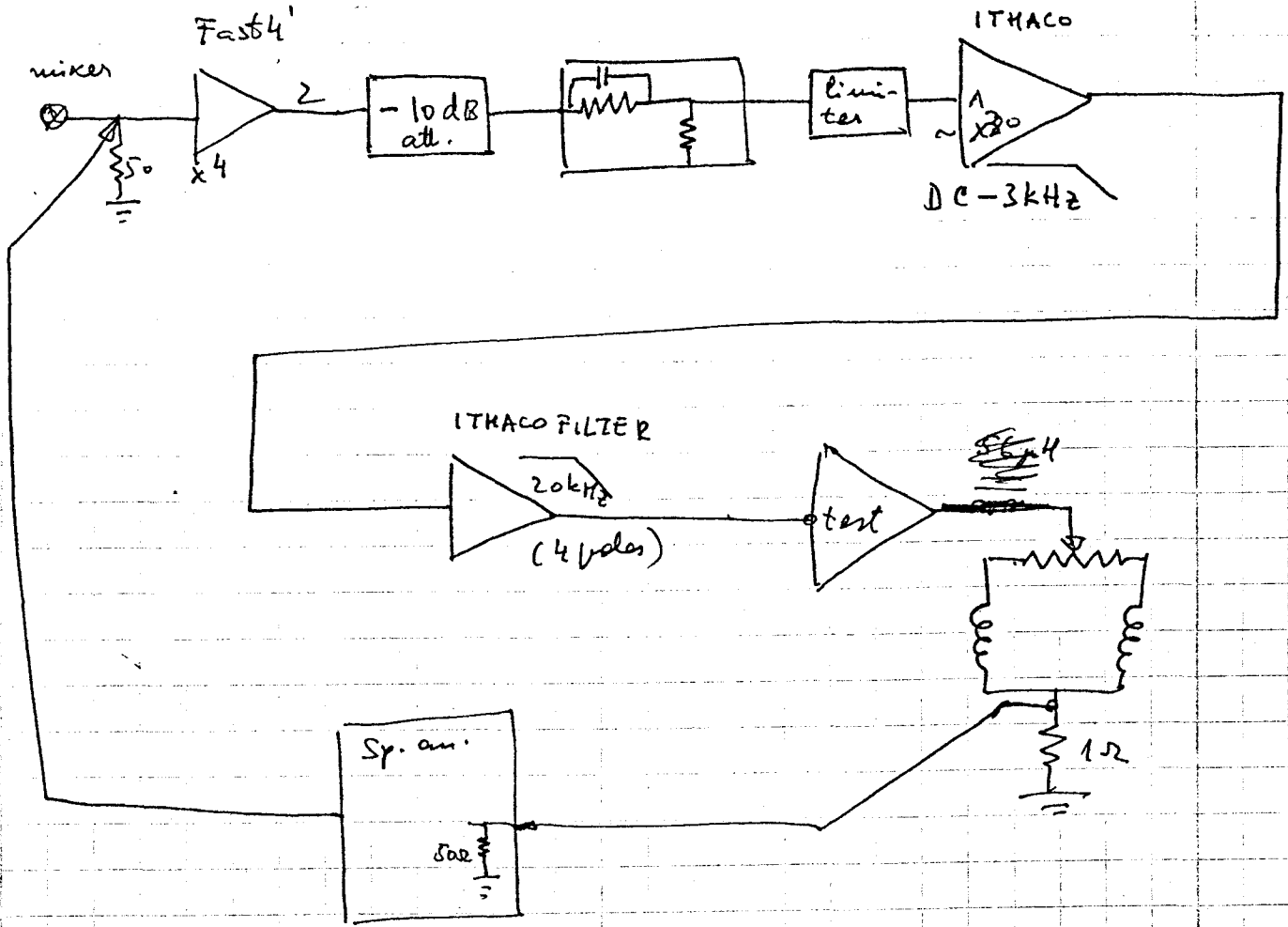
FAST 4 gain 4



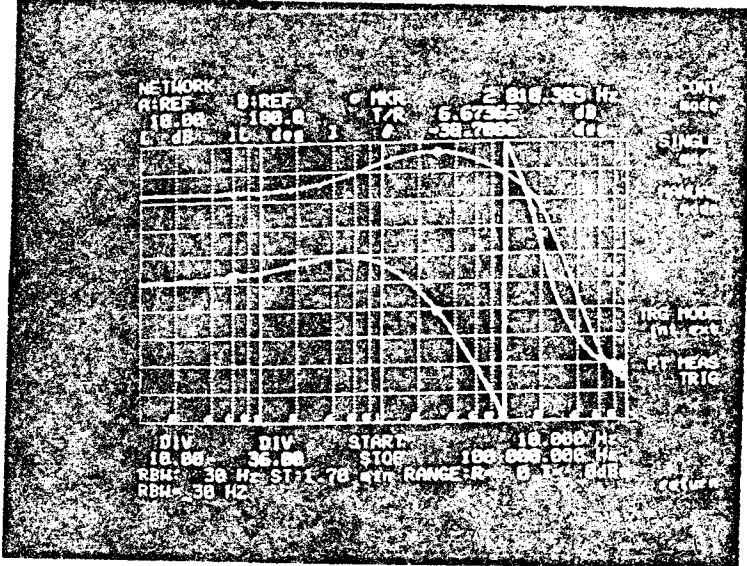
slow pocket's cell branch of 1st arm servo loop

50Ω term at Fast 4 ampl.

1st arm (please connecting) servo; coil driver leg



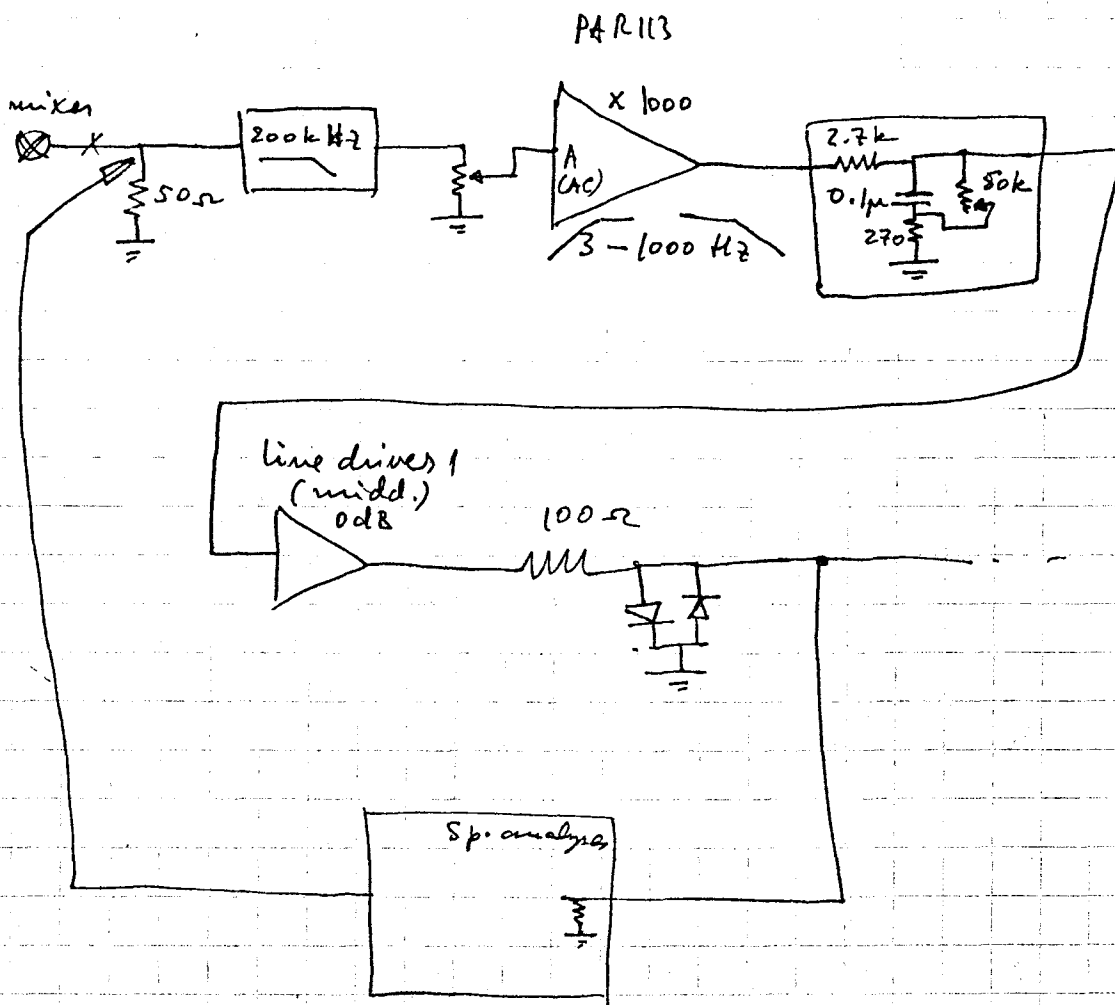
17 July 89 17:30



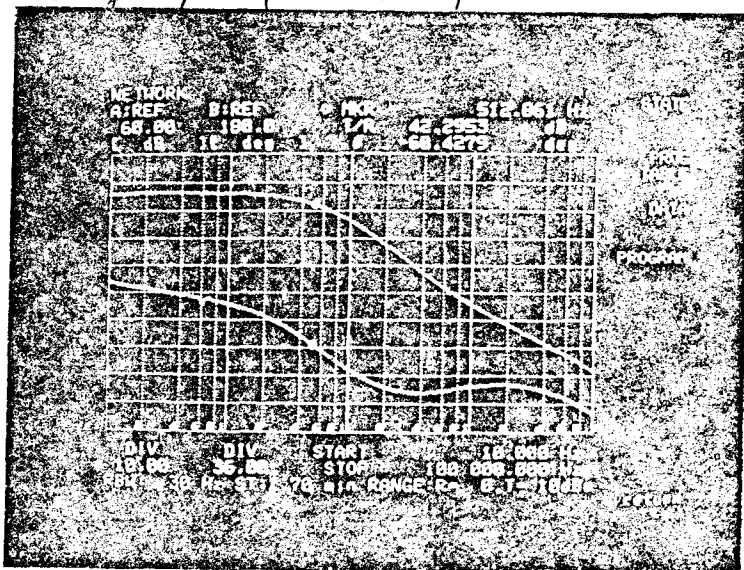
Coil leg response, 50 ohm at in of FAST4 1

COIL LEG

1st arm locking (phase correction) servo: by pas



17 - July - 89 18:09



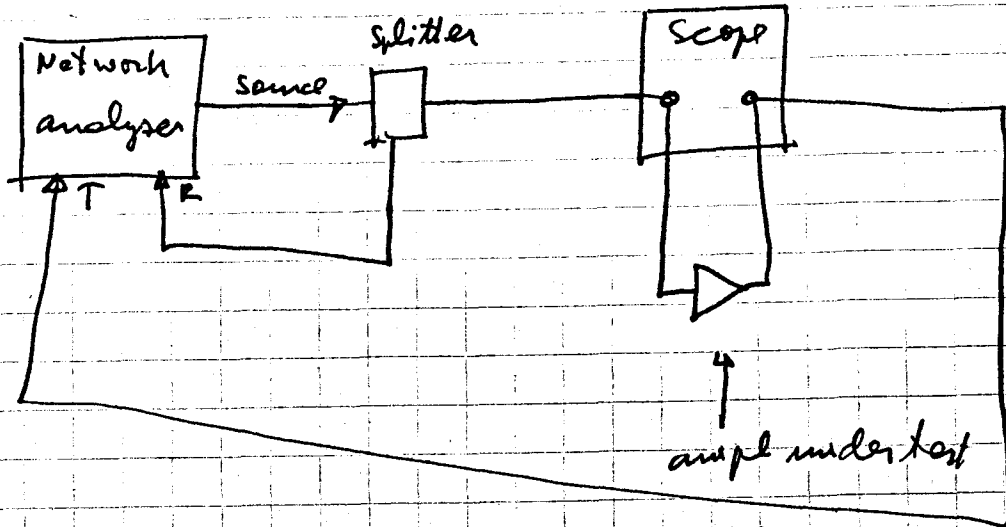
BYP 717

1st arm bypass response

13-July-89

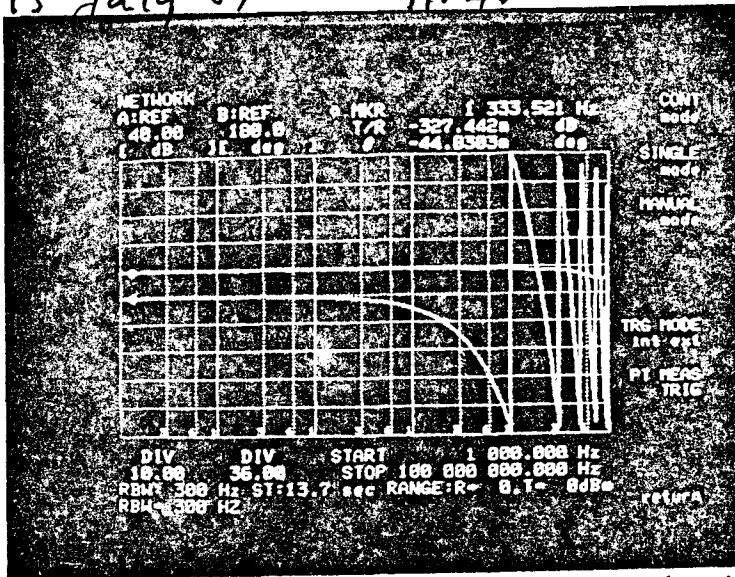
078

Since ~30' of cables were used to connect the amplifier under test (see diagram below), we took the response of the cables themselves (trace at 11:40)



13-July-89

11:40



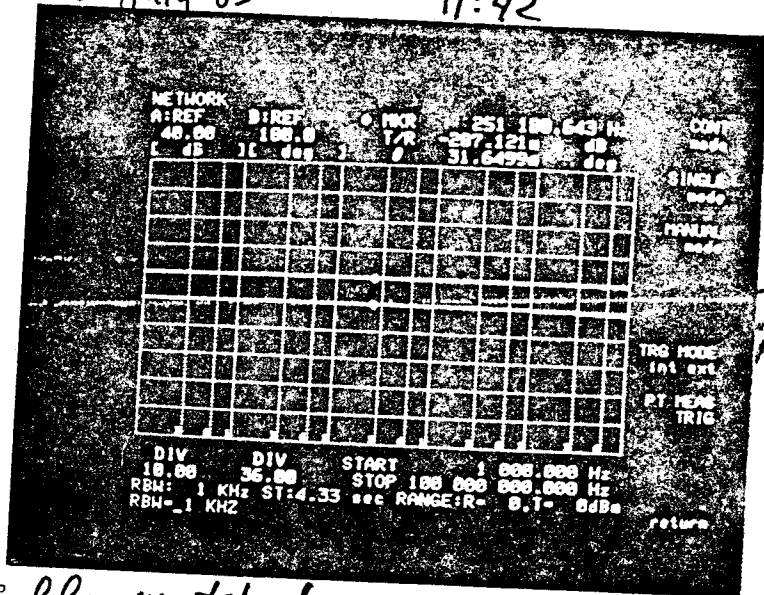
imbalanced

TR. FUNCTION OF CABLES - they are badly

- It turns out that at 250 kHz the cables alone show a phase shift of 5°.
- Therefore, the ~3' cable from splitter to R input was replaced with a long cable, matched to the one in the test path. See traces overlaid

13-July-89

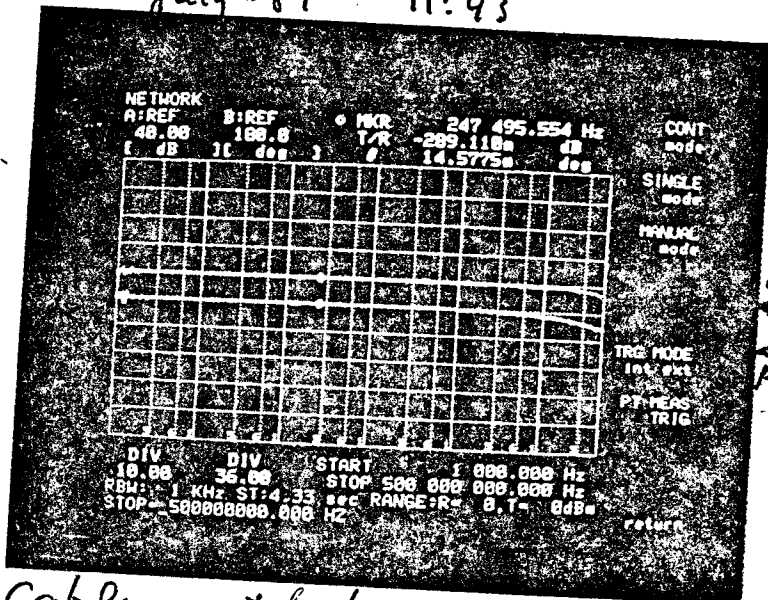
11:42



Cables matched

13-July-89

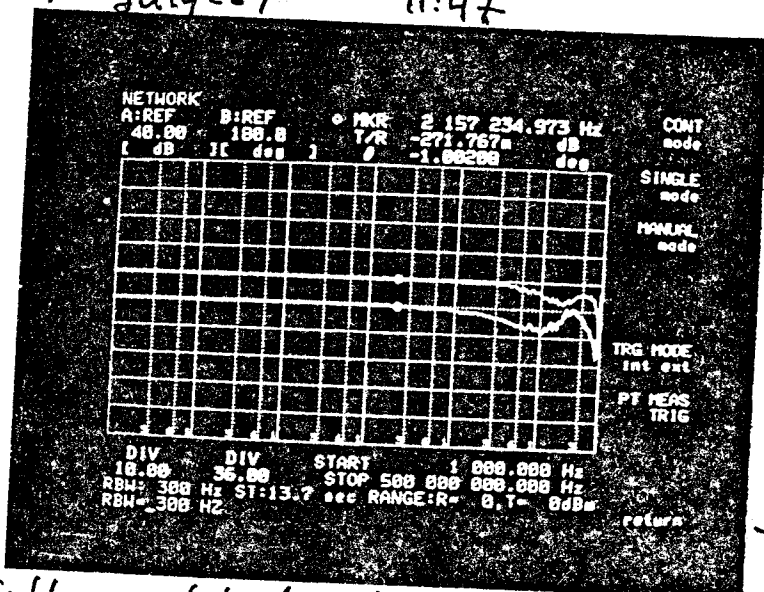
11:43



Cables matched

13-July-89

11:47



Cables matched, going through T's connected

to scope

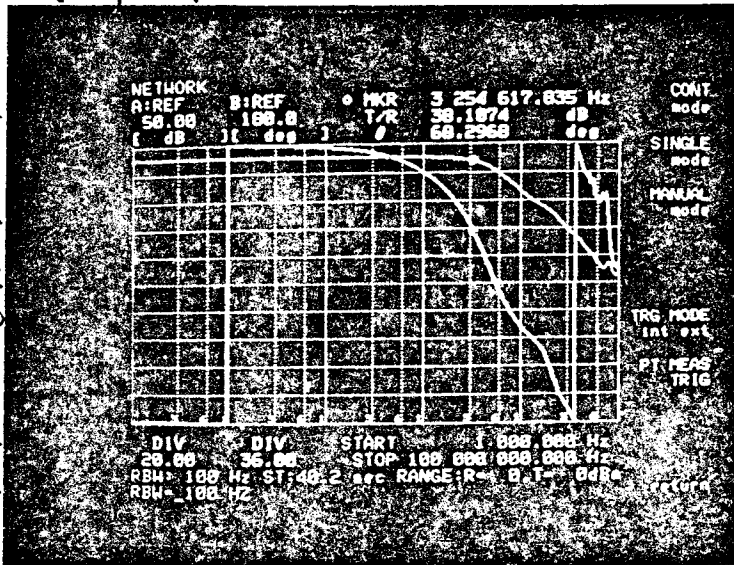
To obtain the correct phase response of the amplifiers in the 40 m system server, we decided to measure their response again, this time with the matched cable.

The two ~~transfer~~ transfer functions (opposite page) clearly illustrate the difference between using unmatched/matched cables. All the following transfer functions have been measured

with matched cables (see cable response on p. 78Y overleaf)

13-July-89 14:09

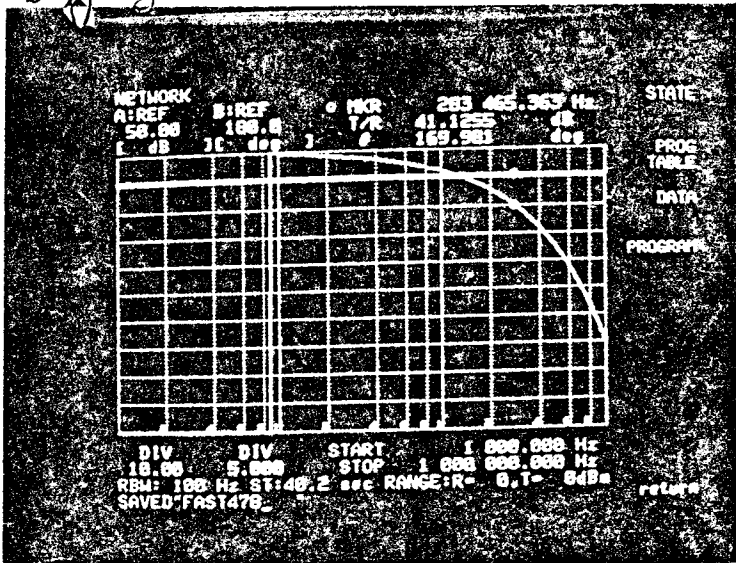
90° phase shift @ 282 KHz



FAST 477

FAST 4' 50 on out imp 1 (v) full gain

13-July-89 14:19

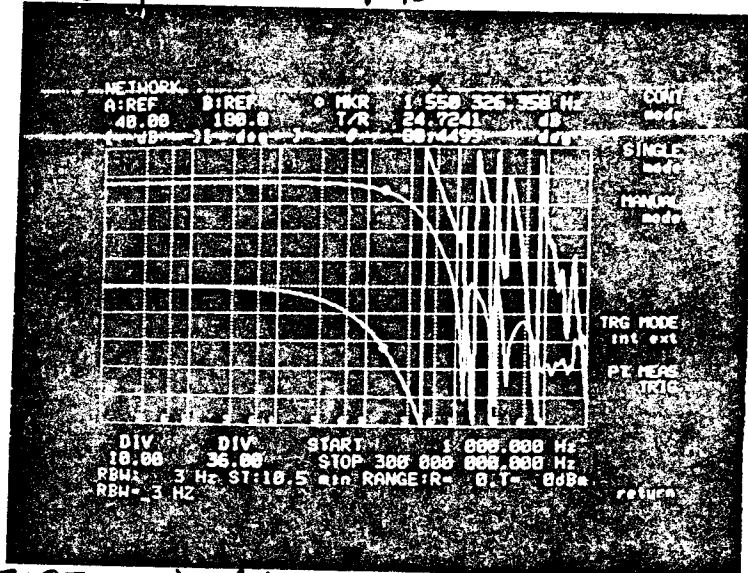


FAST 478

FAST 4' imp 1 (v)

13-July-89

14:45

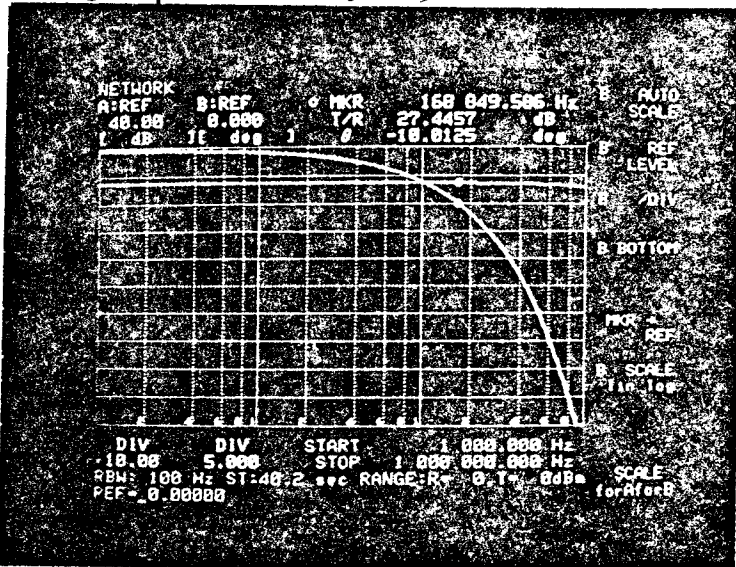


FAST 2 7101

FAST2 imp(V) full gain

13-July-89

14:35



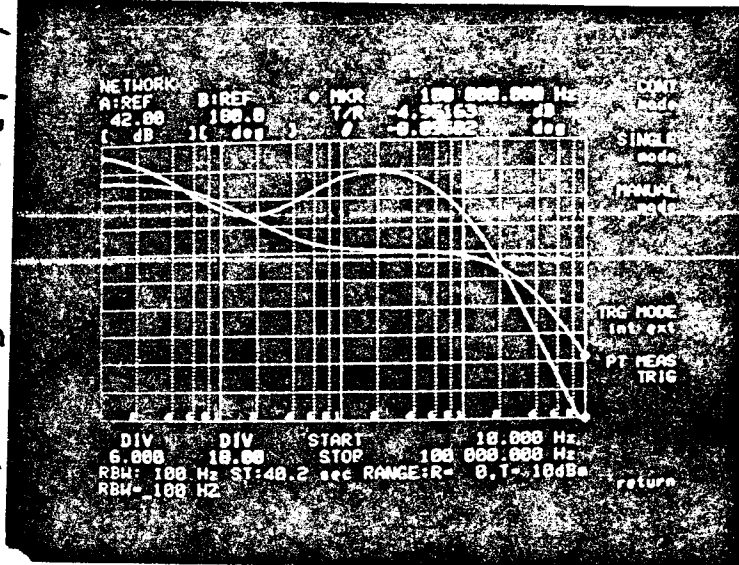
FAST 2 7102

FAST2, imp(V) full gain

13-July-89

15:30

Low 2 ITHACO, 50m out



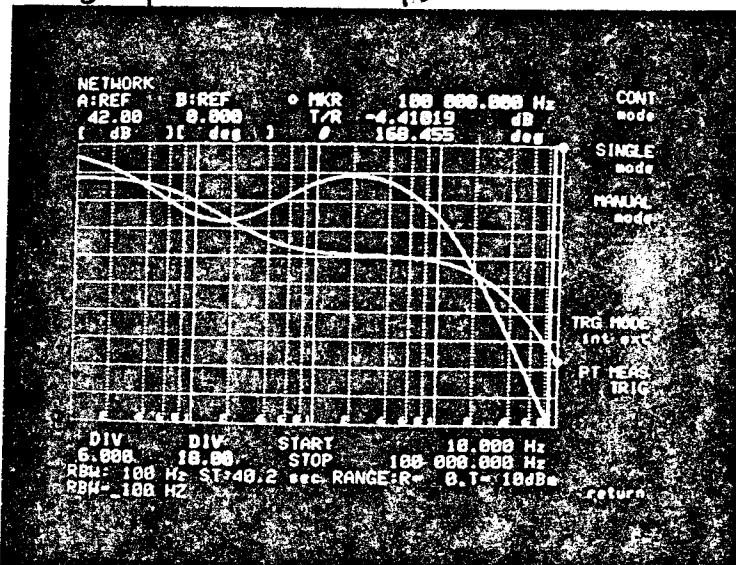
HV3710

wide open

HV3 monitor, left channel, through ITHACO X10

13-July-89

15:40



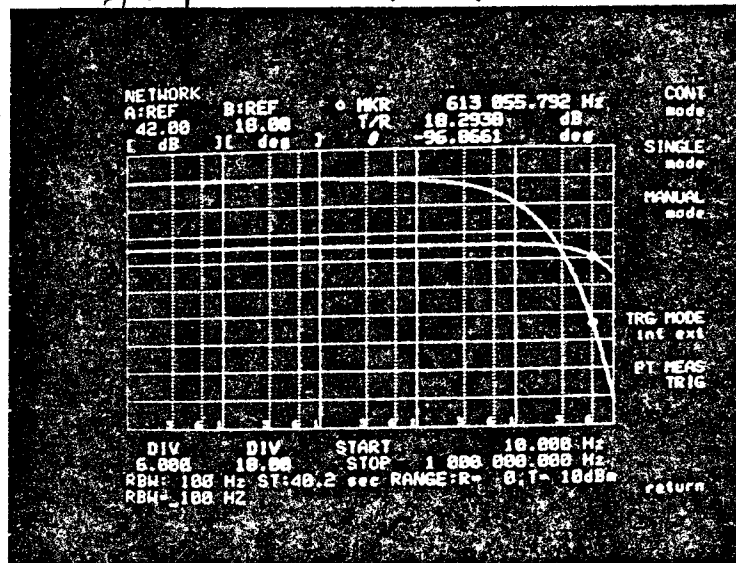
HV3710R

HV3 monitor, as 15:30, right channel

13-July-89

15:51

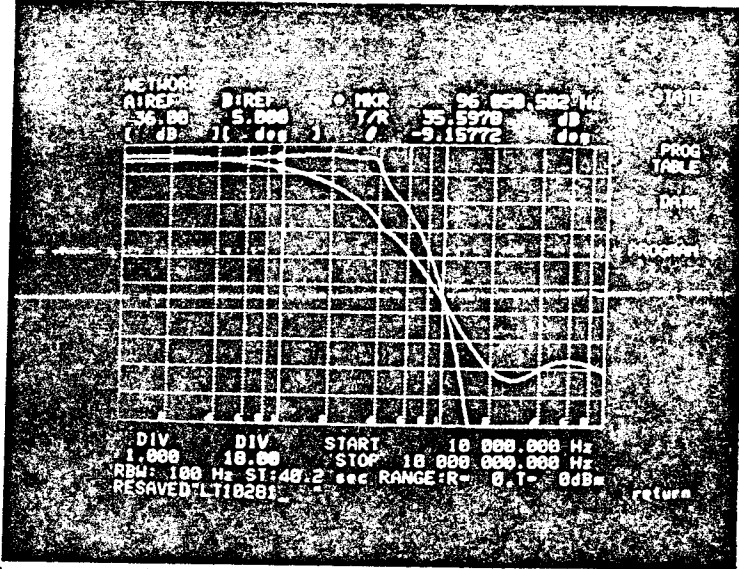
Alt Z in



ITHACO X10

ITHACO X10, LO Z out, wide open

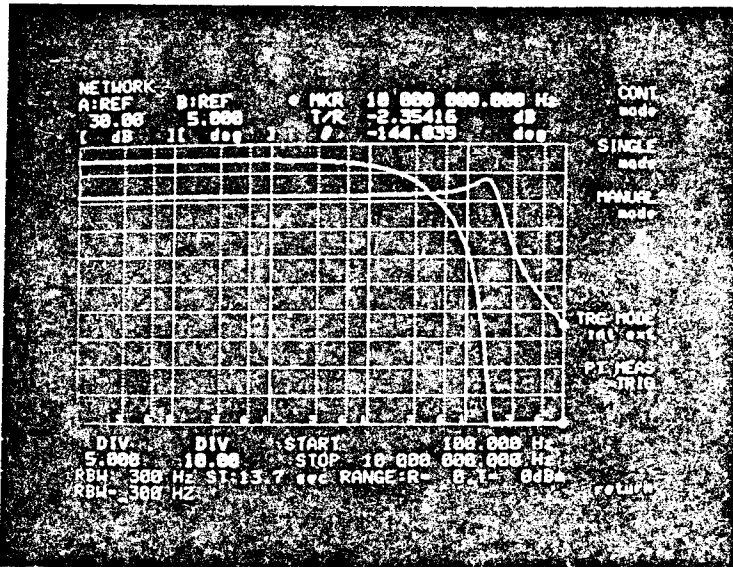
14-July-89 9:40



LT10281 AXES: 10MHz 36dB

LT 10281 High Z in, 50 out

14-July-89 10:00

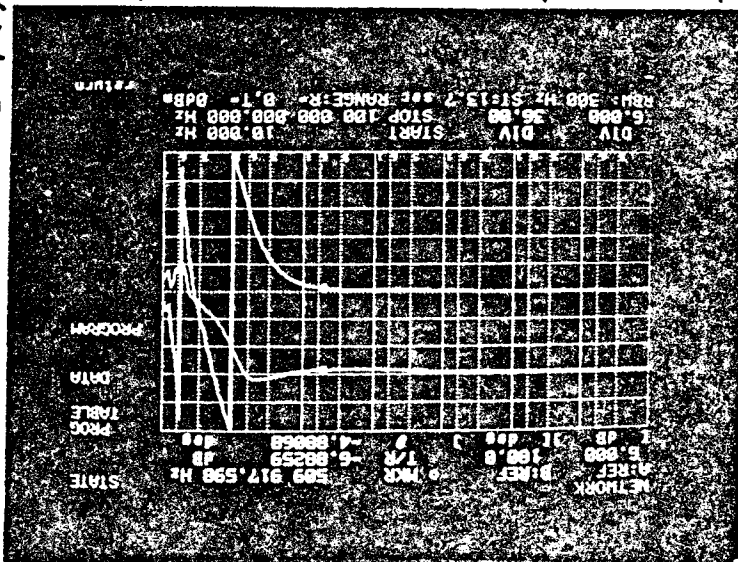


LT 10282

LT1028 #2 High Z in, 50 out

LINE DRIVER 1, water, 50 ohm out

LD4UP



GAIN = 0 dB nominal

1014

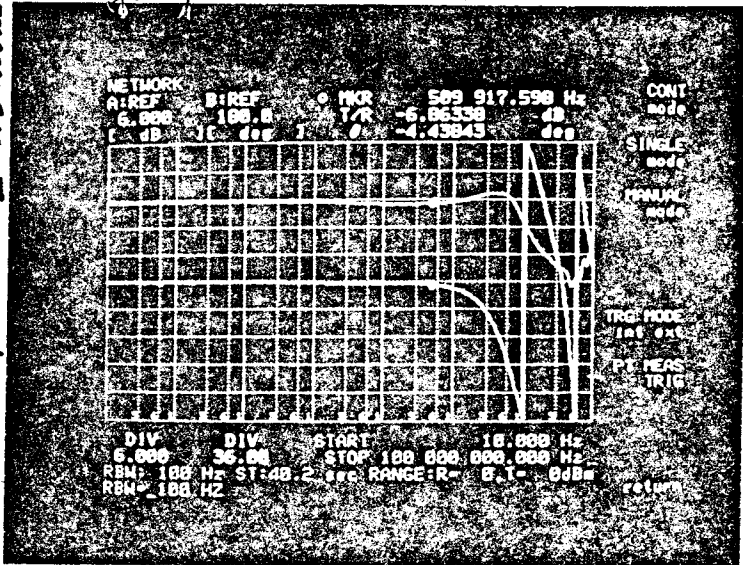
14-July-89

14-July-89

Nontransfer functions taken

GAIN = 0 dB nominal

14-July-89 10:17

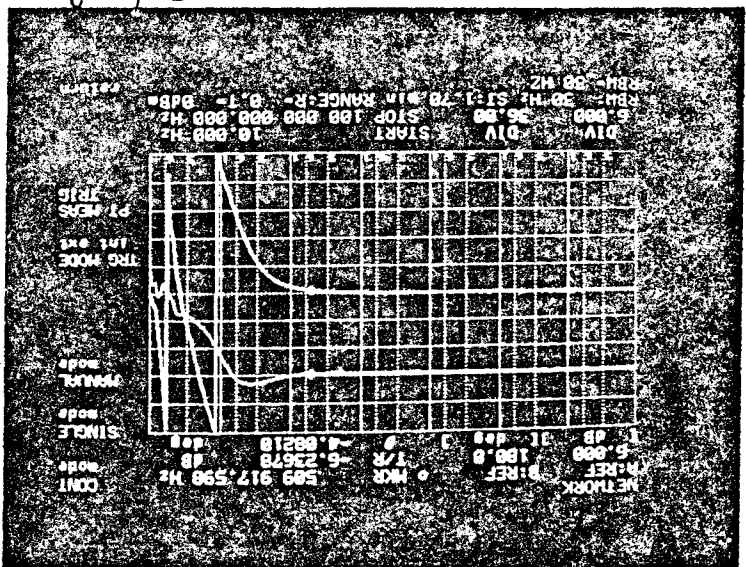


LD1 MID

LINE DRIVER 1, Middle, 50Ω in, out

14-July-89 10:25

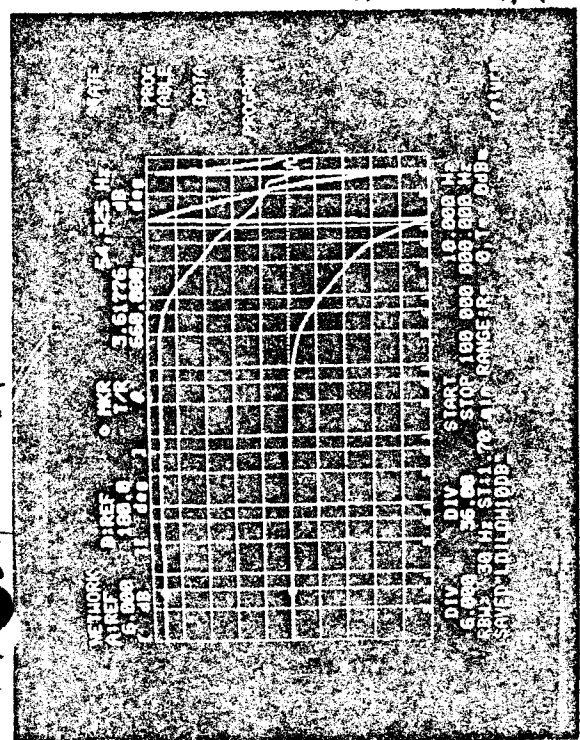
GAIN = 0 dB nominal



LD1 Low

Line driver 1, Lower, 50Ω in out

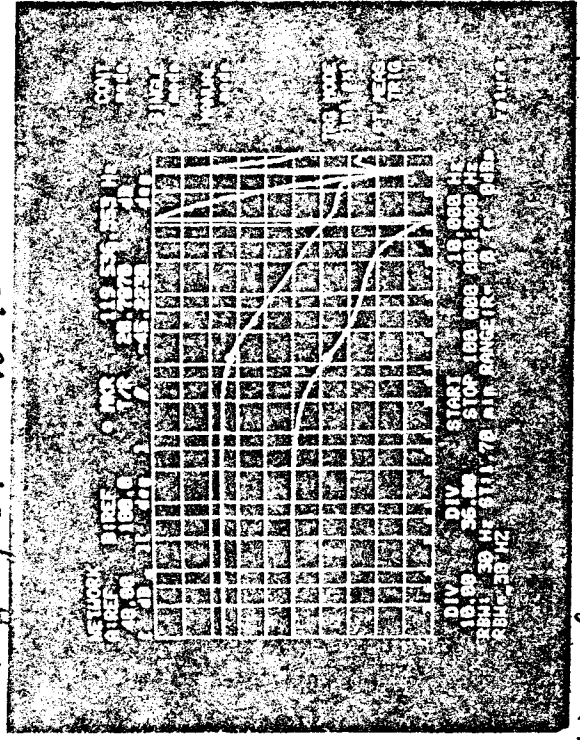
14 July 10:34



GAIN: 1x10dB nominal

Line driver 1 lower, 50-Ω in, out

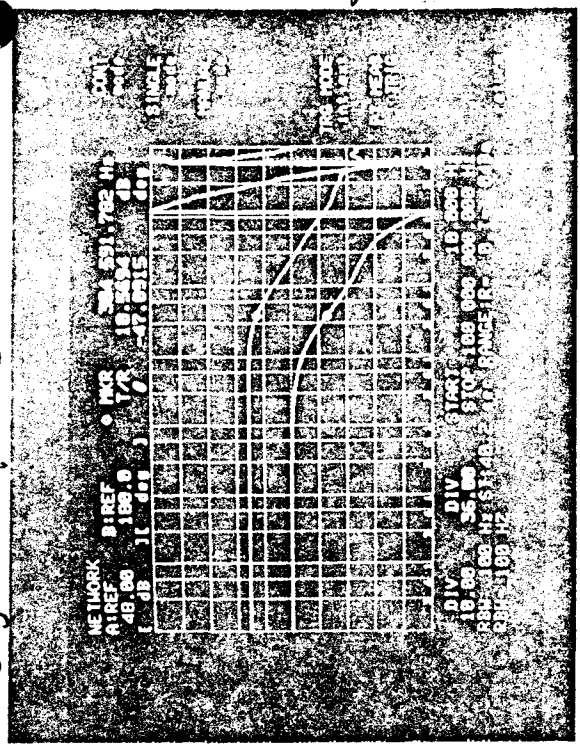
14 July-89 10:40



GAIN: 3x10 dB nominal

Line driver 1, lower, 50-Ω in, out

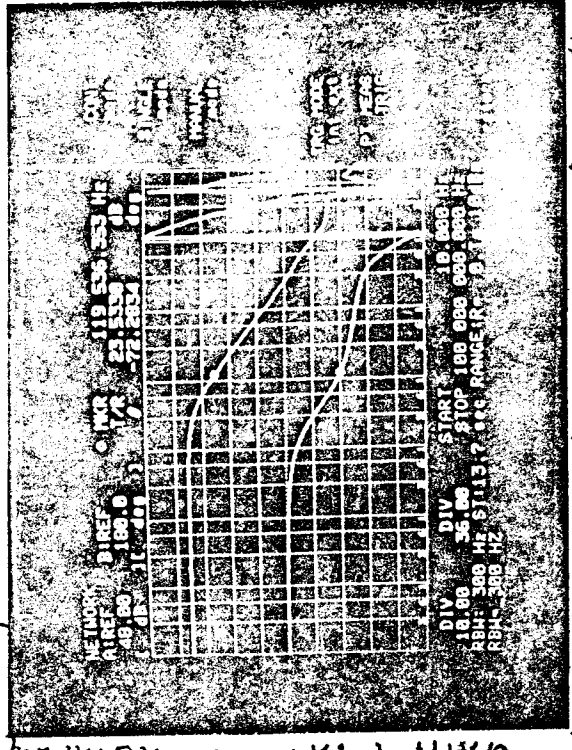
14-July-89 10:36



GAIN: 2x10dB nominal

Line driver 1, lower, 50-Ω in, out

14 July-89 10:45

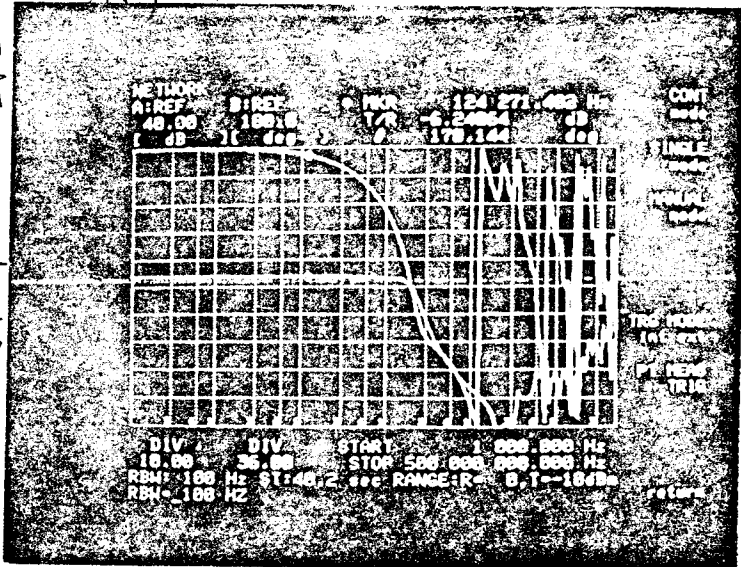


GAIN: 4x10 dB nominal

Line driver 1, lower, 50-Ω in, out

17 July 89 13:35

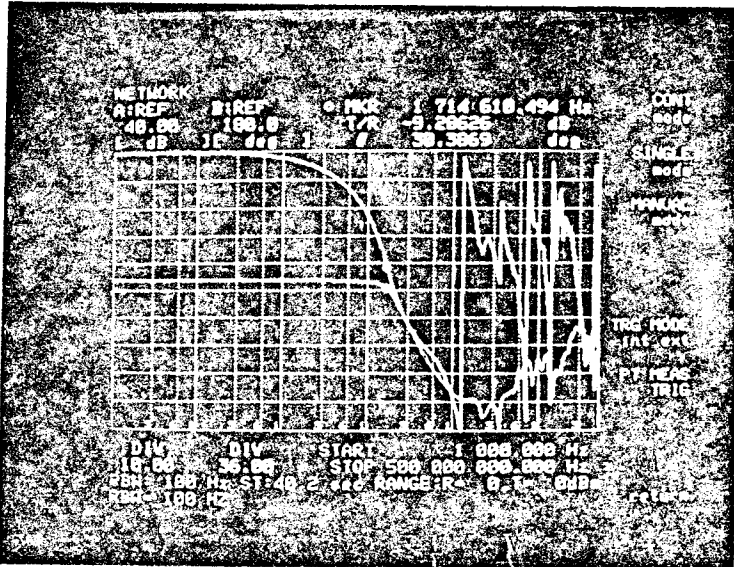
3 dB point at: 1.955 MHz



LDA MID 0dB

Line driver A, mid, 0dB, 50 ohm out.

100 dB at: 128.4 kHz

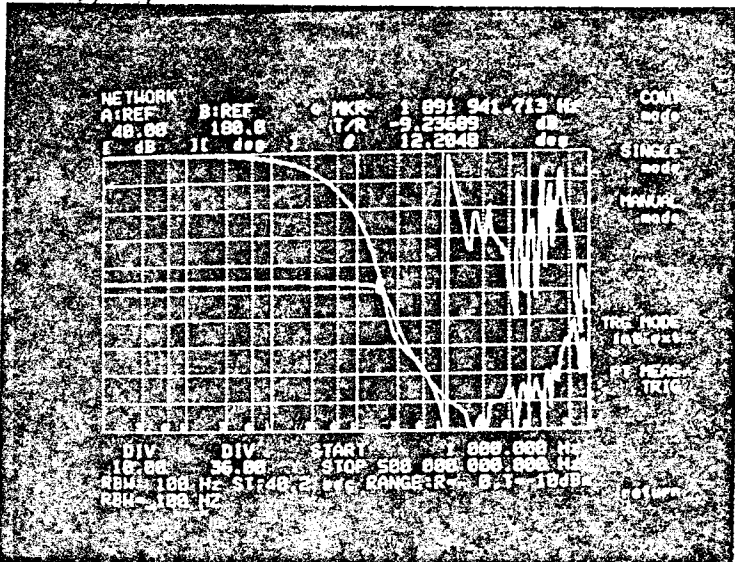


LDA LOW

Line driver A, lower, 50 ohm out, 0 dB

17 July 89 13:30

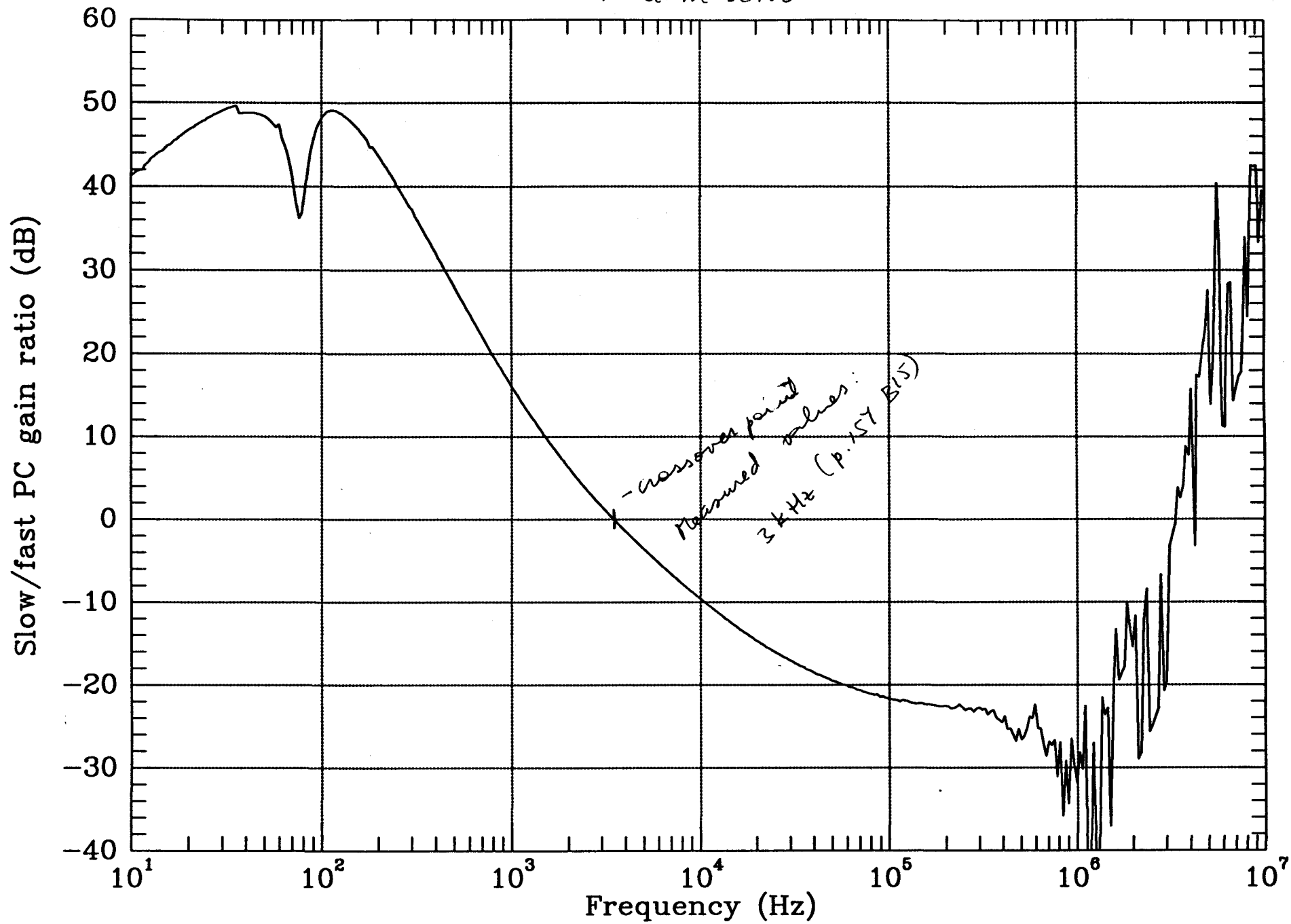
100 dB at: 124 kHz



LDA UP 0dB

Line driver A, upper, 0 dB, 50 ohm at outp.

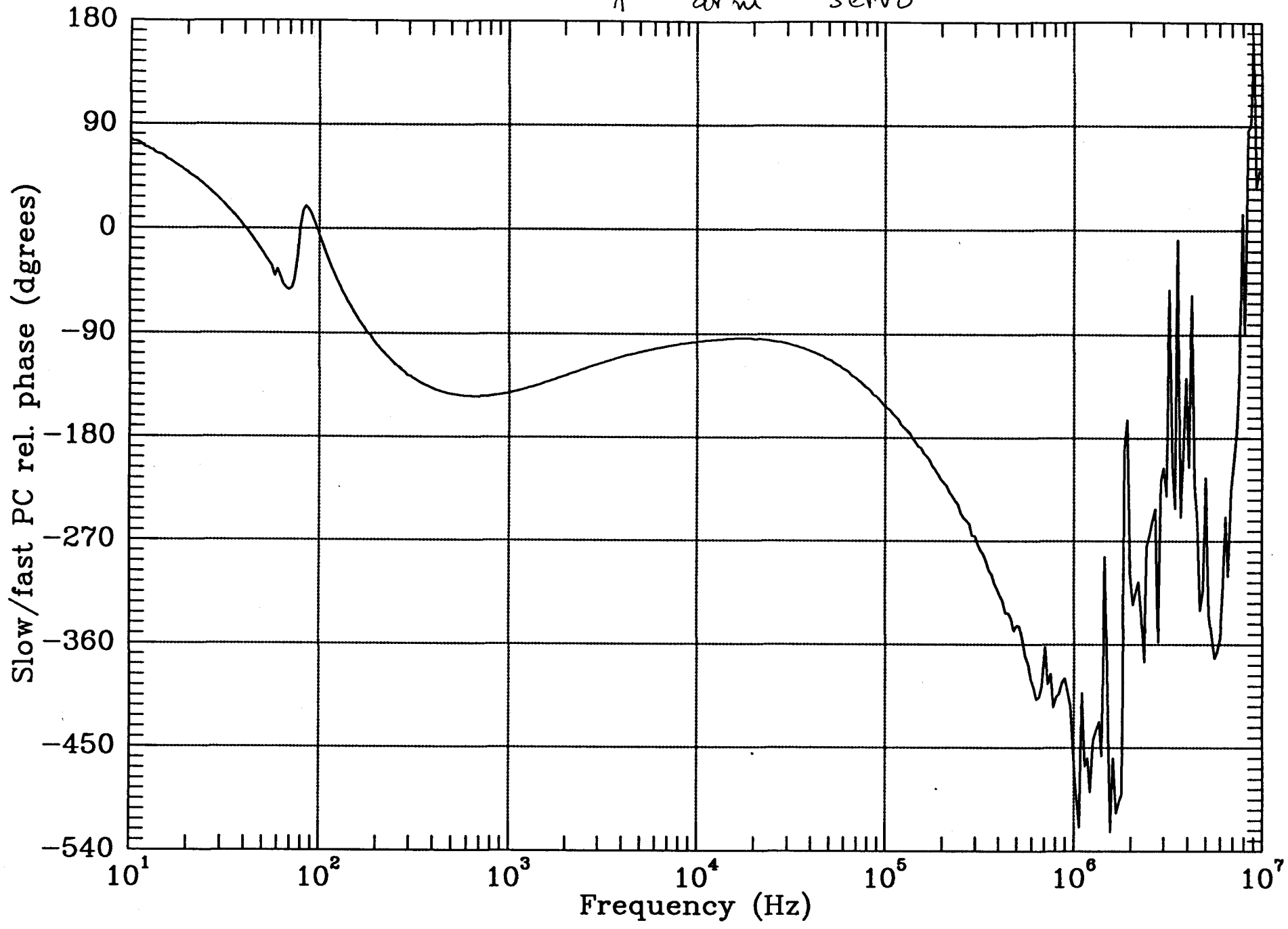
1st arm servo



21 July - 89

AA, Y6, JA
measured
processed
by AA

1st arm servo

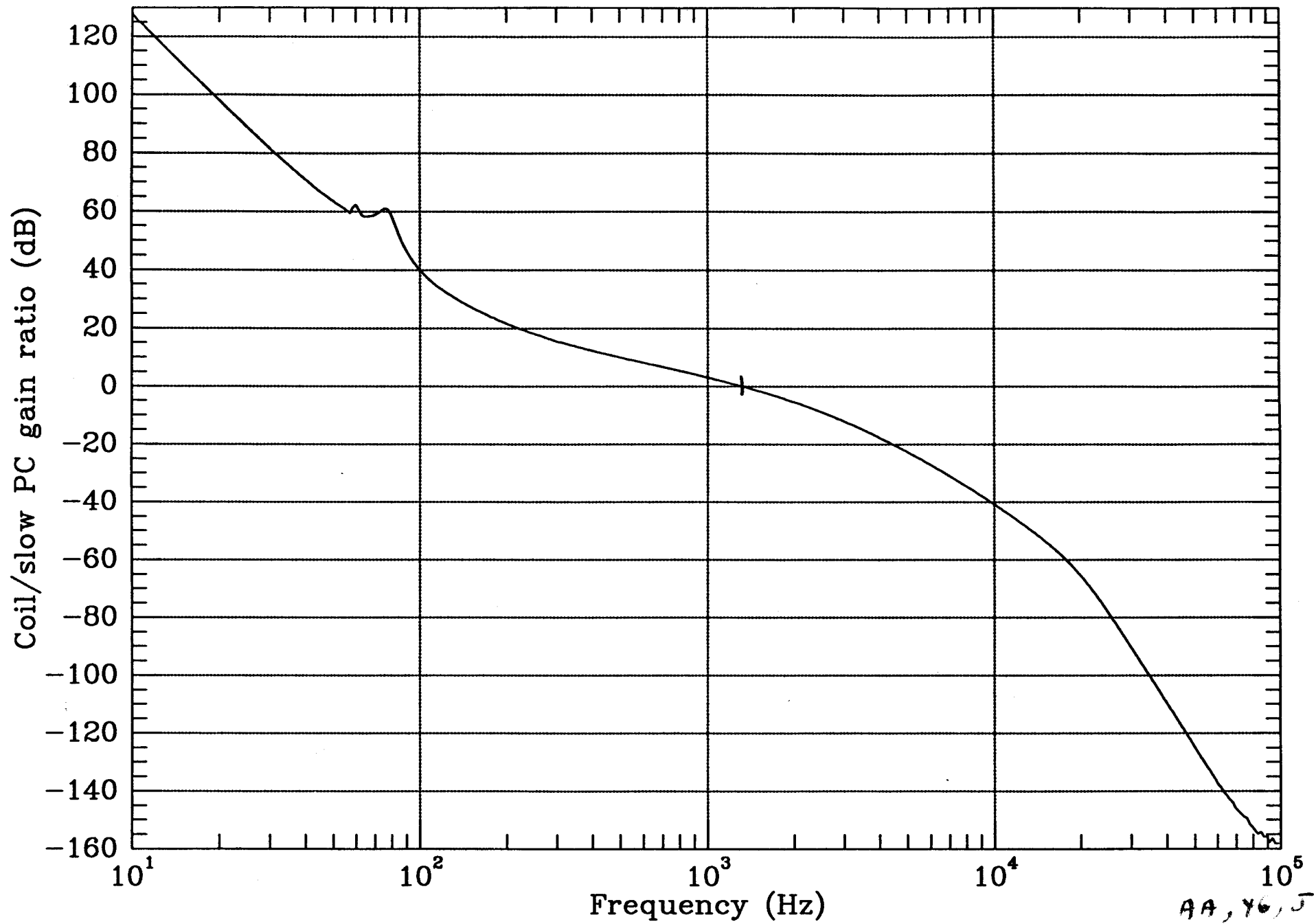


USES FILES: SLOPO721
FASTPO

21-July-89

AA, Y6, JH
measurements
processed
by AA

1st arm servo

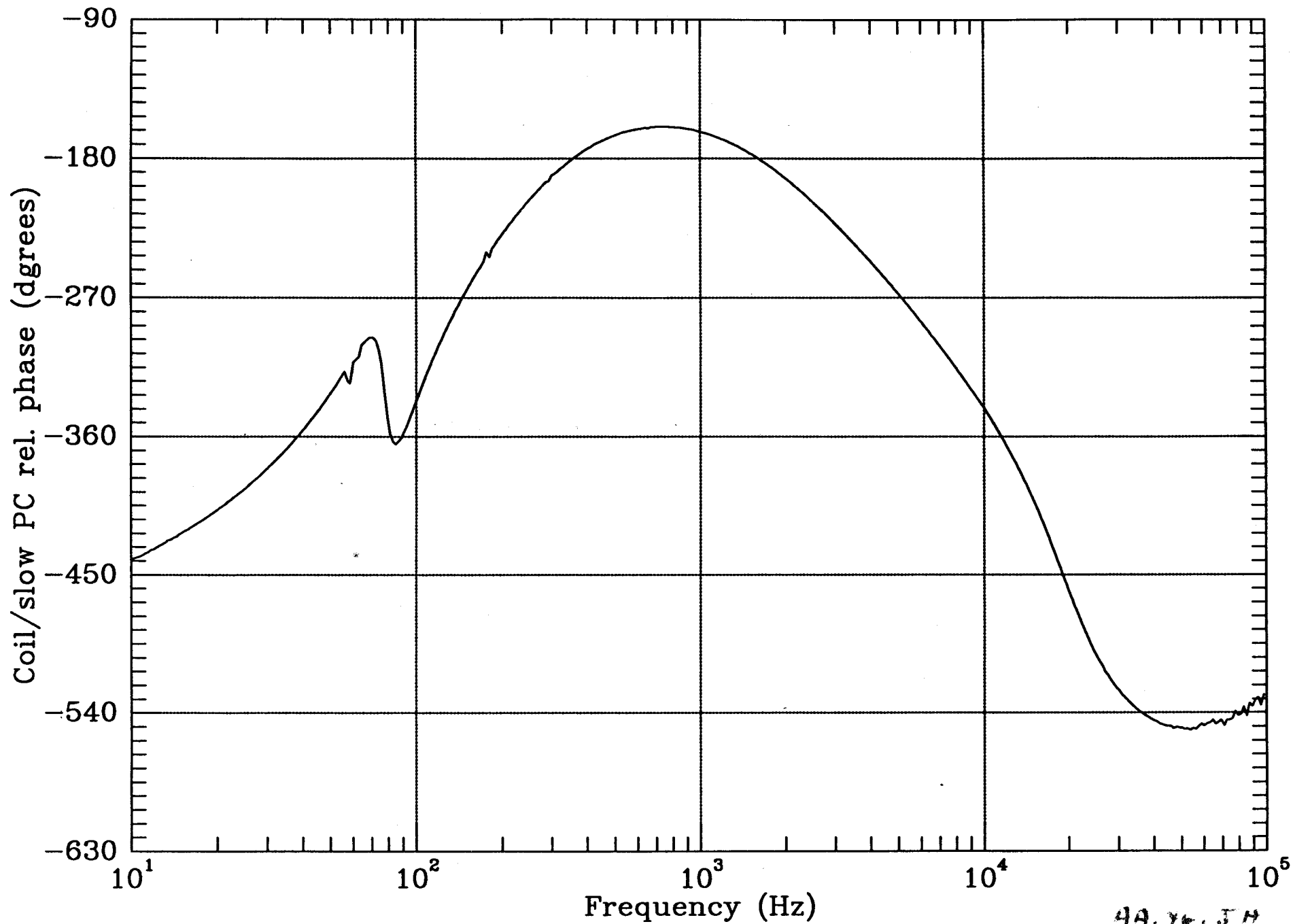


AA, Y6, JH
measurements
processed by AA

21-Jul-89

uses coil, c.f., coil pepl

1st arm servo



uses coil p●, coil pepl1

AA, Y6, JH
measurements
processed by AA

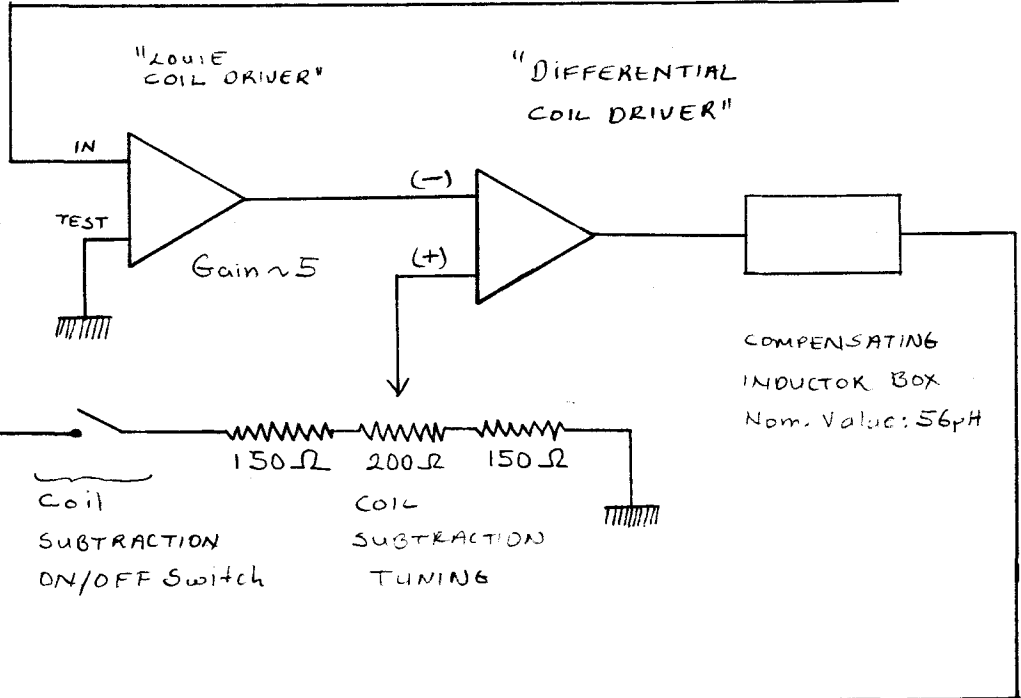
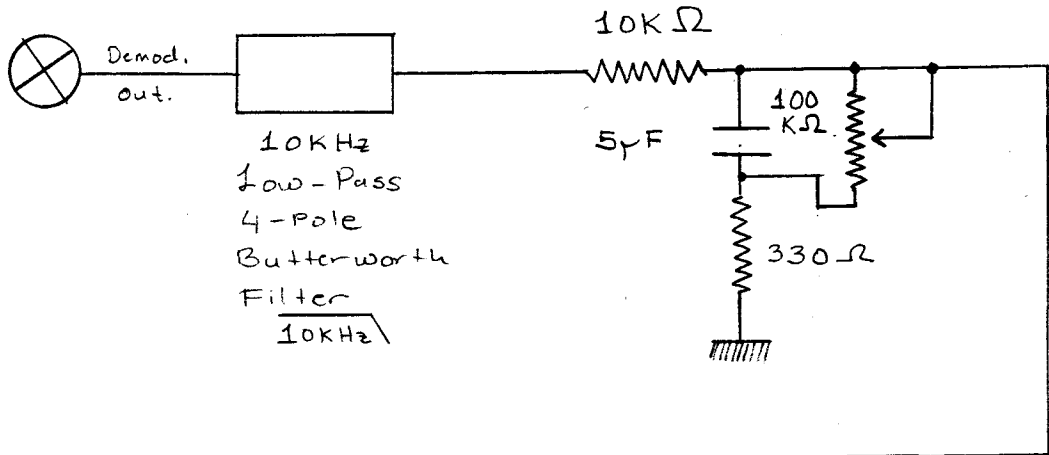
24-● ly-89

BATCH
START

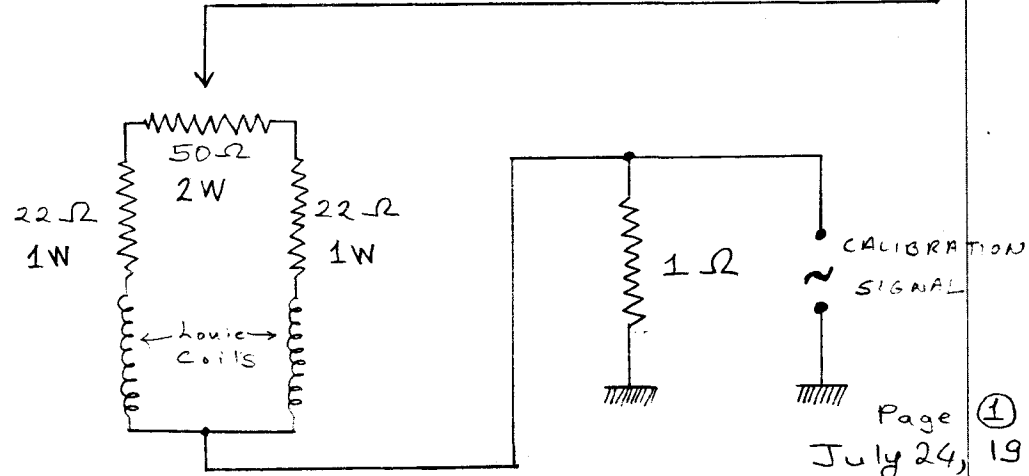
STAPLE
OR
DIVIDER

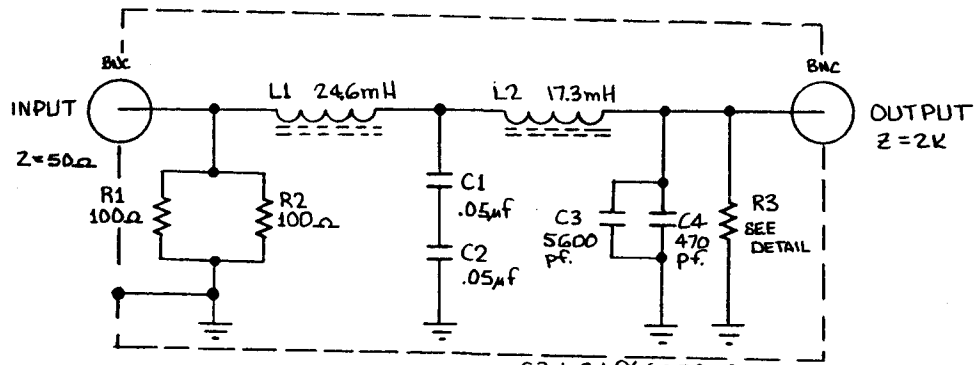
Second Arm Servo Loop Block Diagram:

Second Arm
Mixer



On page ②
of First arm
Servo
diagram.



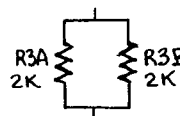


$C3 + C4 \approx 6070 \text{ pf}$
 $C1 \text{ AND } C2 = .025 \mu\text{f}$

- L1— WIND 248 T. ON BOBBIN. USE CORE G-42213-40 AND MTG. CLIP. 24.6mH.
- L2— WIND 208 T. ON BOBBIN. SAME CORE AS L1. 17.3mH.
 (WIRE IS #32 INSULATED COPPER.)

NOTE: ALL CAPS. ARE MYLAR OR MICA.

R3 DETAIL



1. FOR 1K OUTPUT Z 2K. RES IN PARALLEL
2. CUT ONE RESISTOR FOR 2K OUTPUT Z.

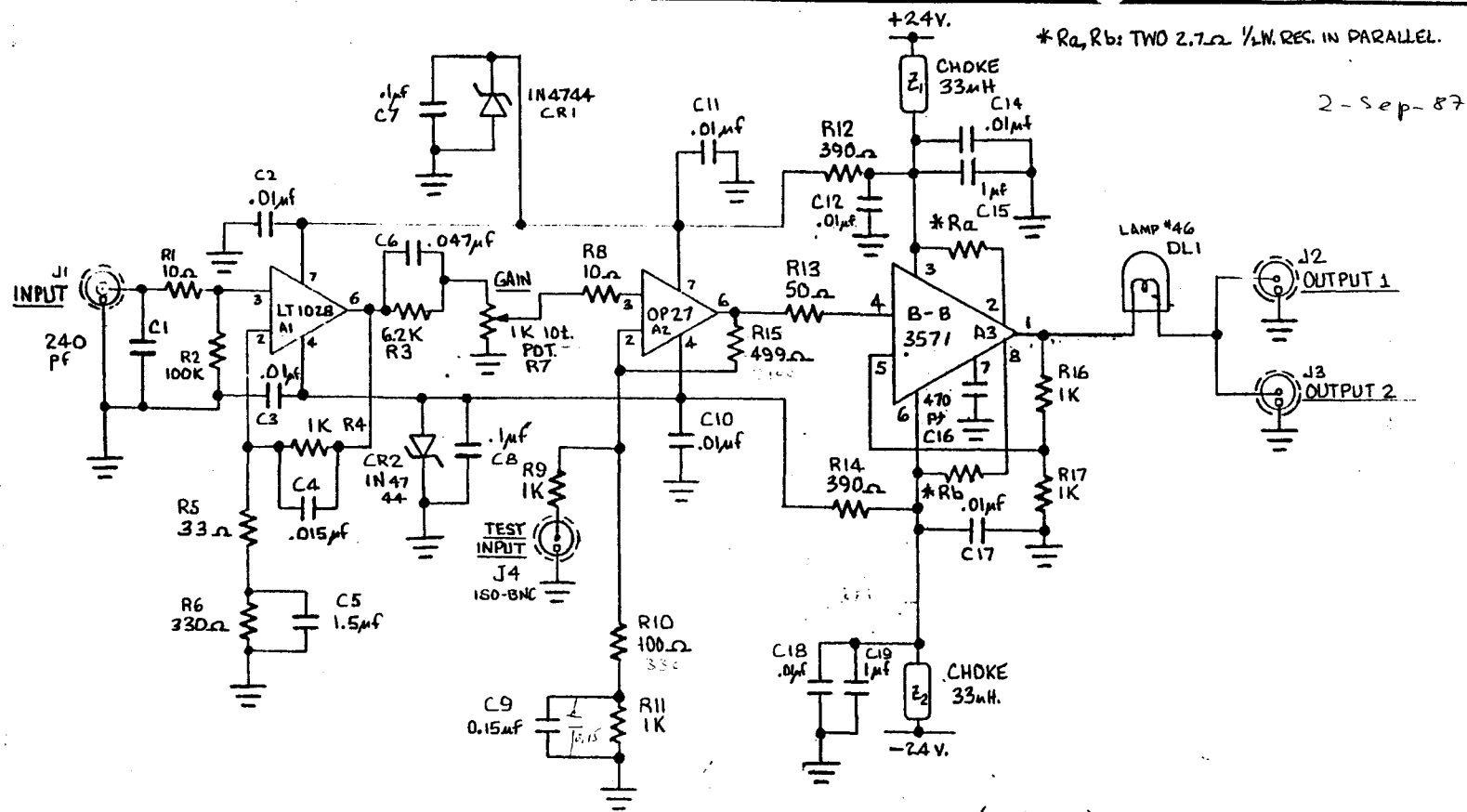
CUT-OFF=10 KHz at 3db POINT

2-29-88 ADD R3 DETAIL

CALIFORNIA INSTITUTE OF TECHNOLOGY
 GRAVITATIONAL PHYSICS

FOUR-POLE BUTTERWORTH FILTER

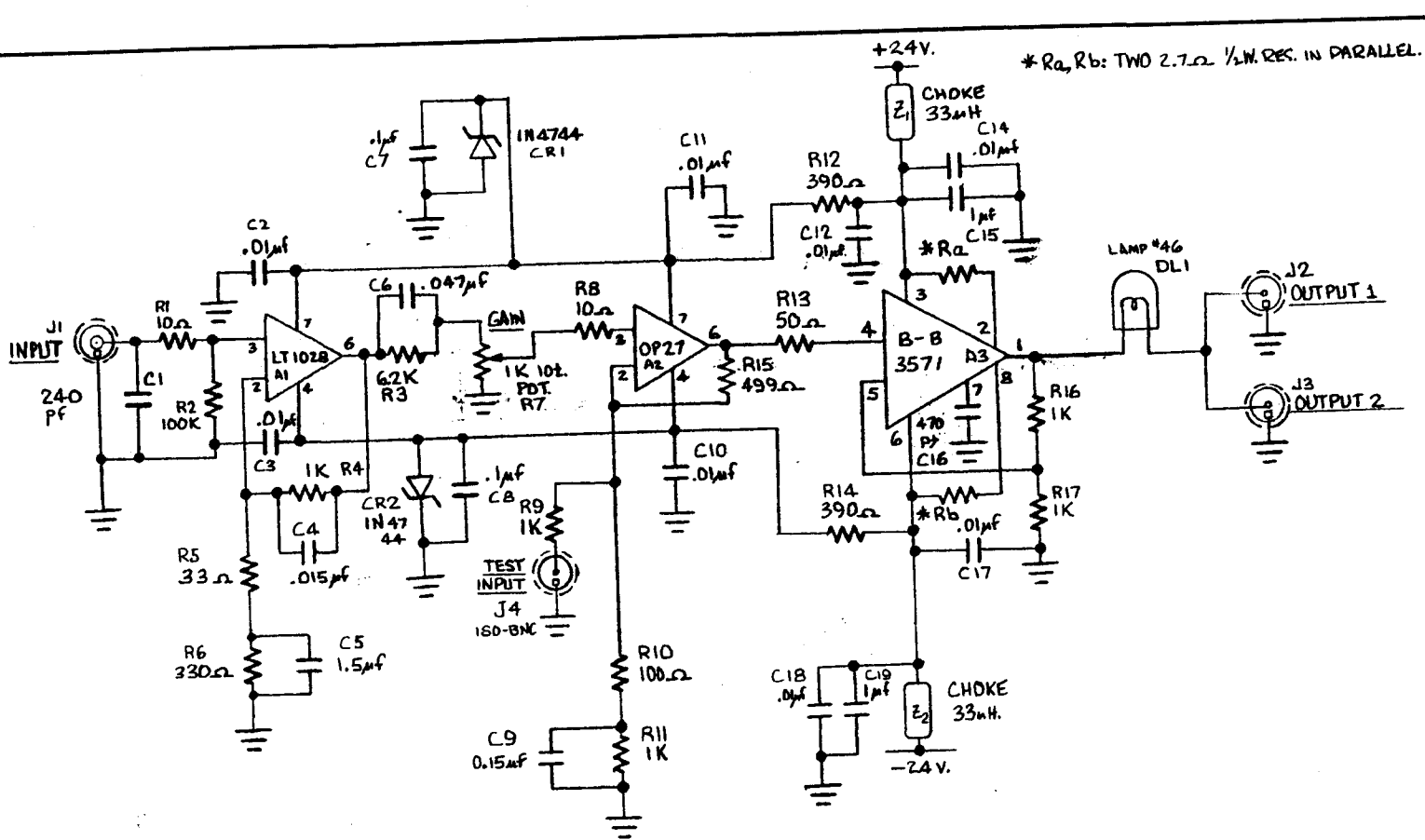
DRAWN BY B. Tinker	DATE 08-11-87	DRAWING NO.
CHECKED BY	SCALE	-1
APPROVED BY	W.O.	



* R_a, R_b: TWO 2.7Ω 1/2W RES. IN PARALLEL.

2-Sep-87

Red: modification on 6 June-89 - on basic coil driver only

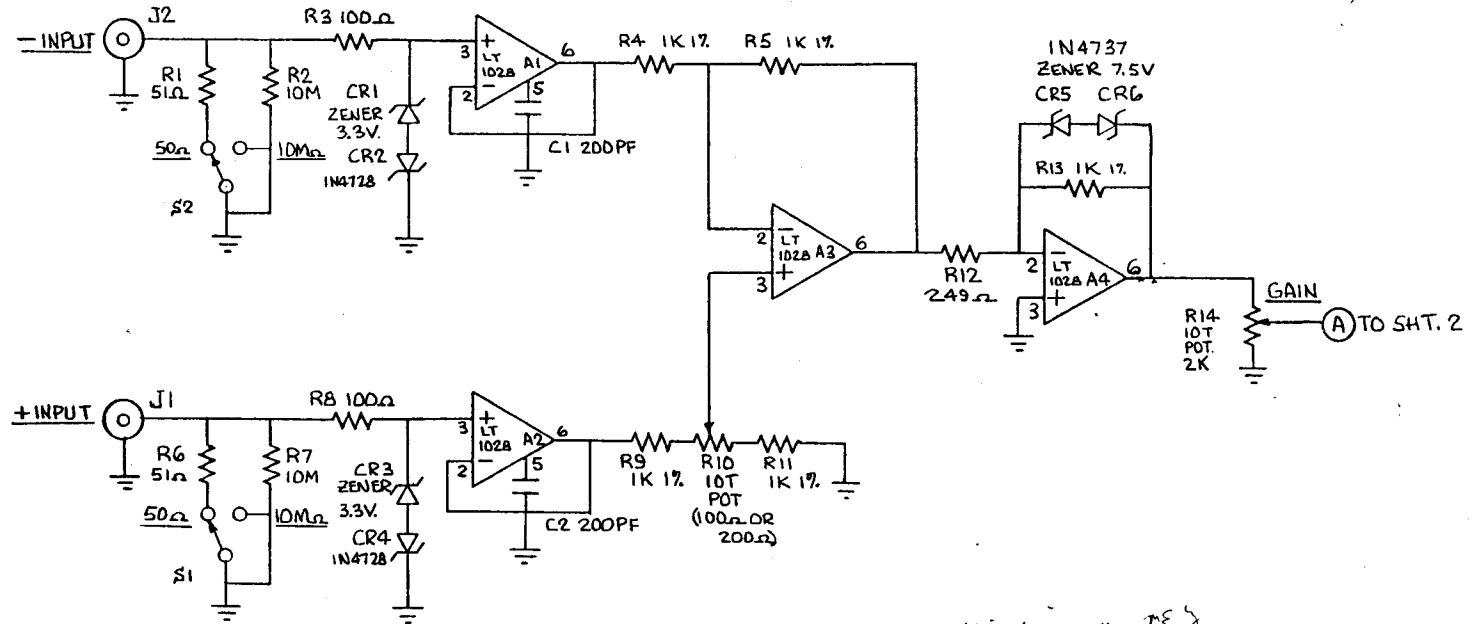


* Ra, Rb: TWO 2.7Ω 1/2W RES. IN PARALLEL.

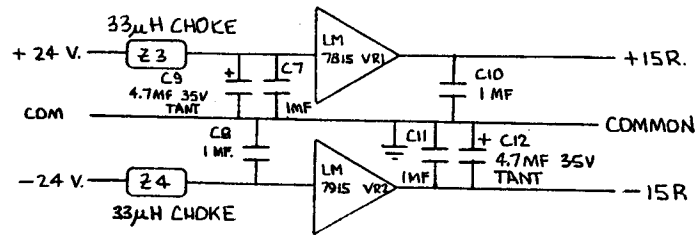
3-1-88 R15 MOVED: 1K TO 500Ω
 1-28-88 C16 FROM 220pf TO 470pf
 UPDATE TO 9-8-78

CALIFORNIA INSTITUTE OF TECHNOLOGY GRAVITATIONAL PHYSICS		
COIL DRIVER HIGH CURRENT		
DRAWN BY B.T.	DATE 9-2-87	DRAWING NO. -1
CHECKED BY	SCALE	
APPROVED BY	W.Q.	

High current, programmable, differential coil drive (MK II)



NEJ MF = μ F in this diagram MEJ



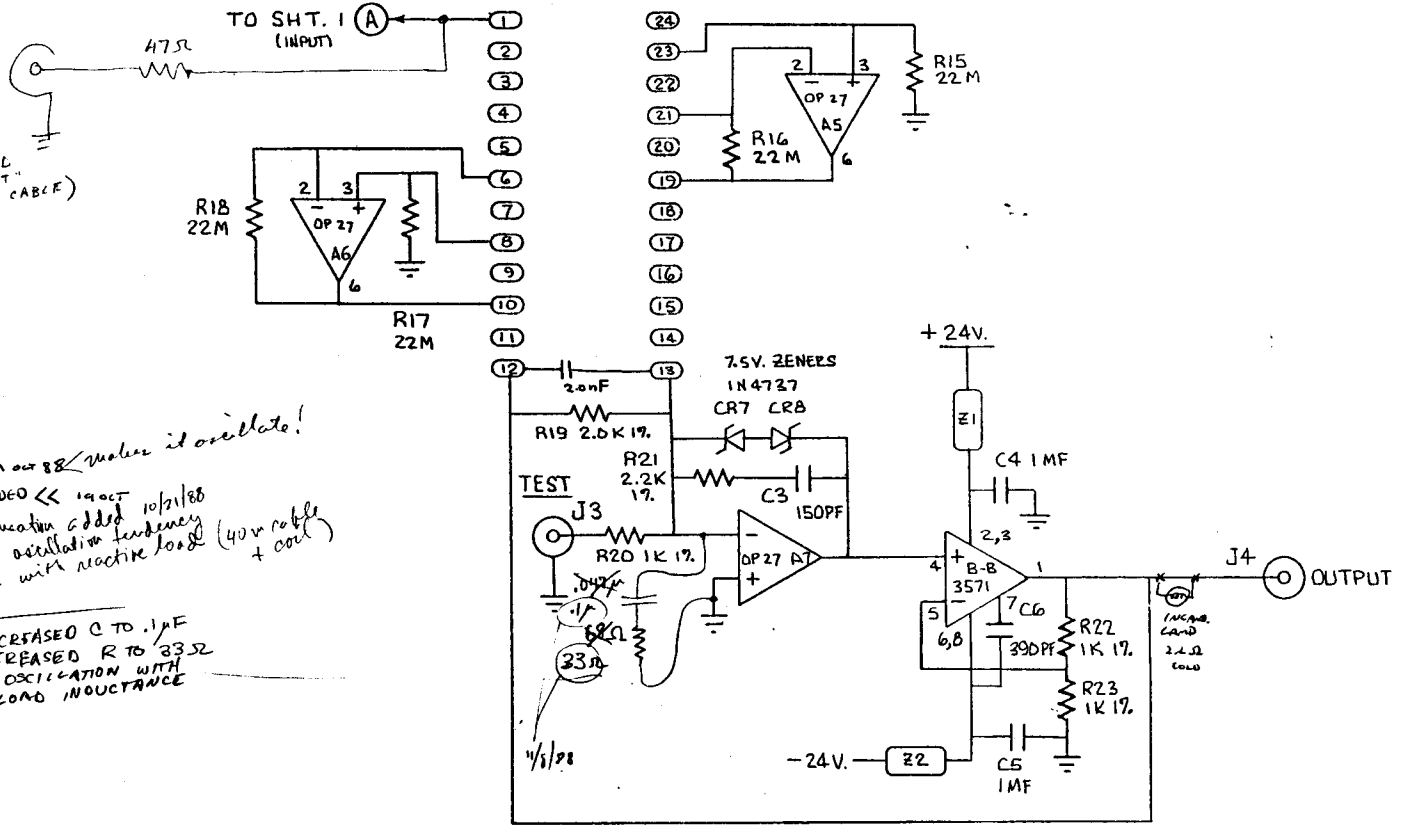
DIFF COIL DRIVER SHT 1

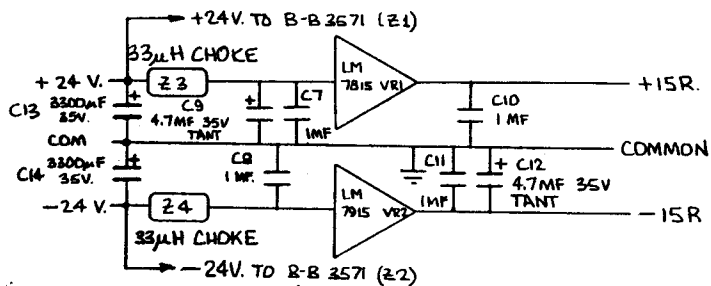
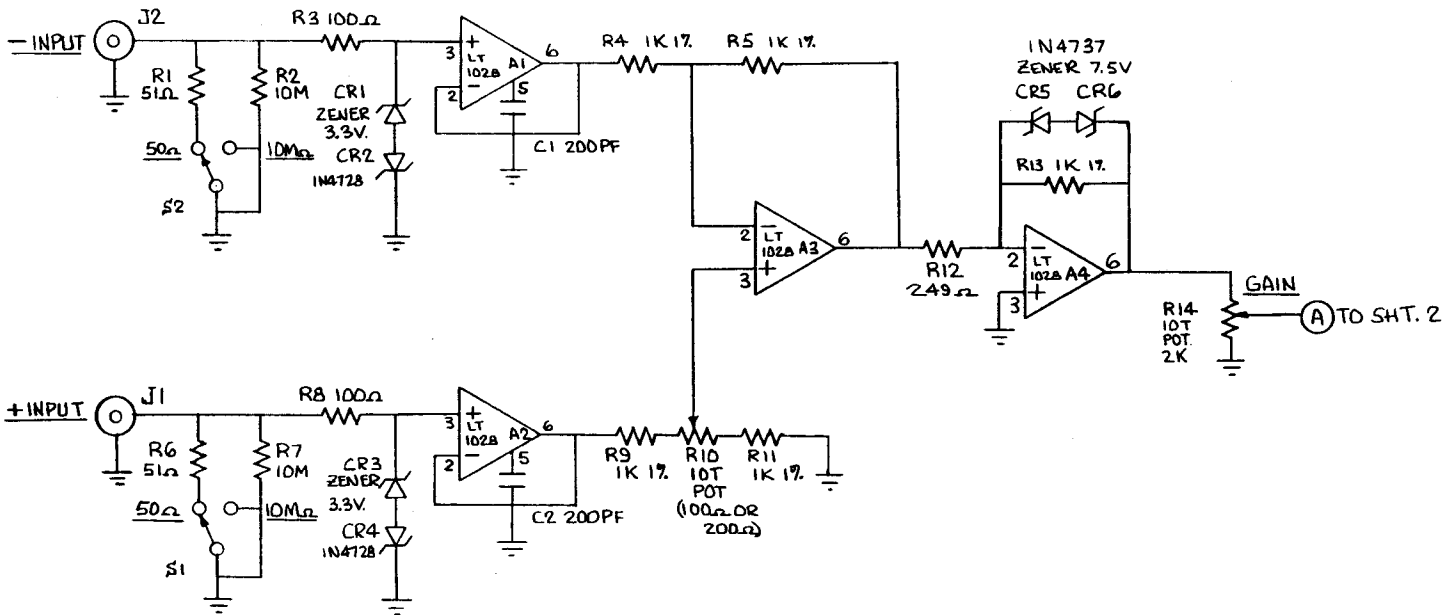
PERSONALITY MODULE

"POT WIPER
OUT"
ON BACK
PANEL
ALSED 10/88
(CAN BE CONNECTED
TO "TEST INPUT"
BY SHORT BNC CABLE)

20nF ADDED 19 OCT 88 *makes it oscillate!*
INLAND LAMP ADDED << 19 OCT
input compensation added
to reduce oscillation tendency
@ 500 kHz with reactive load (40v cable
+ coil)

11/8/88 INCREASED C TO .1uF
DECREASED R TO 33.5k
TO CURE OSCILLATION WITH
HIGHER LOAD INDUCTANCE

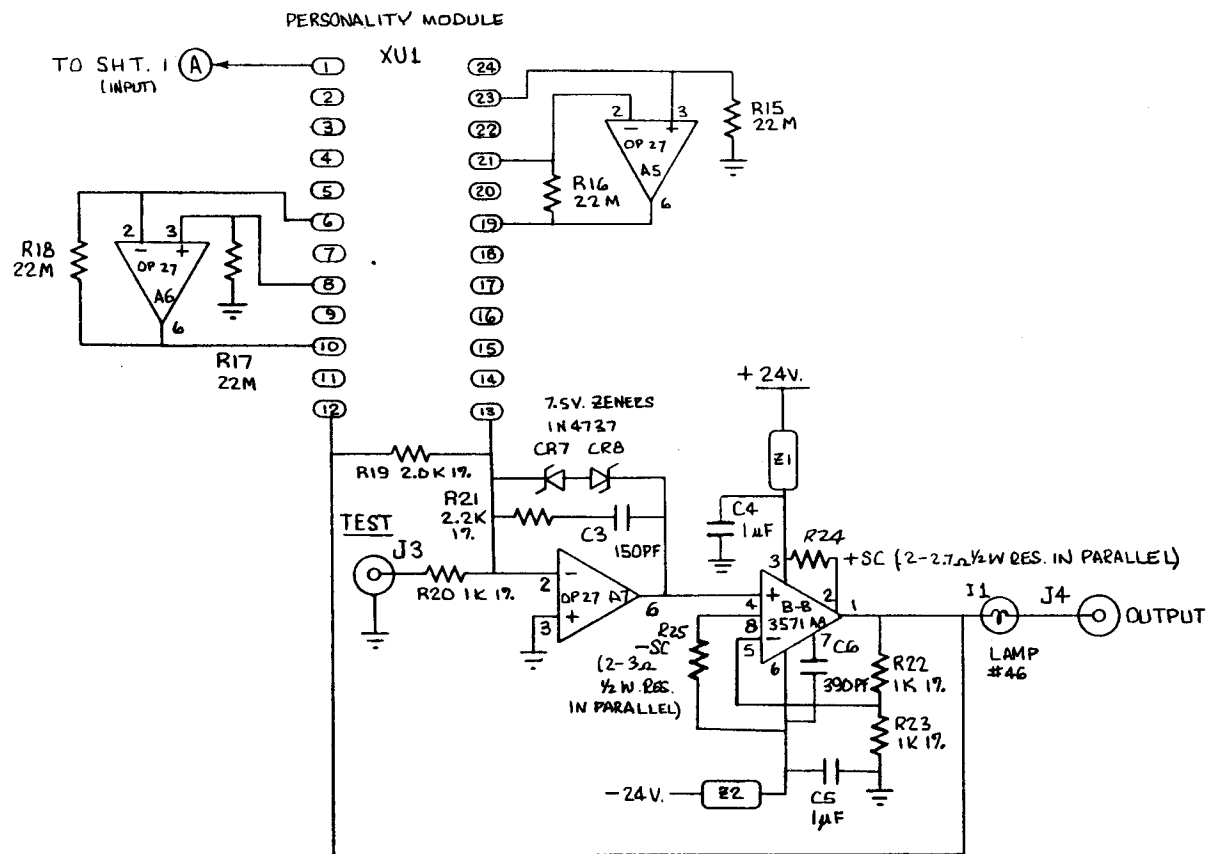




DIFF COIL DRIVER SHT 1

UPDATE 10/4/88 B.

CALIFORNIA INSTITUTE OF TECHNOLOGY GRAVITATIONAL PHYSICS		
PRG. DIFF. COIL DRIVER		SHT. 1
DRAWN BY B.T.	DATE 4-29-88	DRAWING NO.
CHECKED BY	SCALE	
APPROVED BY	W.G.	



DIFF COIL DRIVER SHT 2

UPDATE 10/4/88

CALIFORNIA INSTITUTE OF TECHNOLOGY
GRAVITATIONAL PHYSICS

PRG. DIFF. COIL DRIVER

SHT. 2

DRAWN BY B.T.

DATE 4-29-88

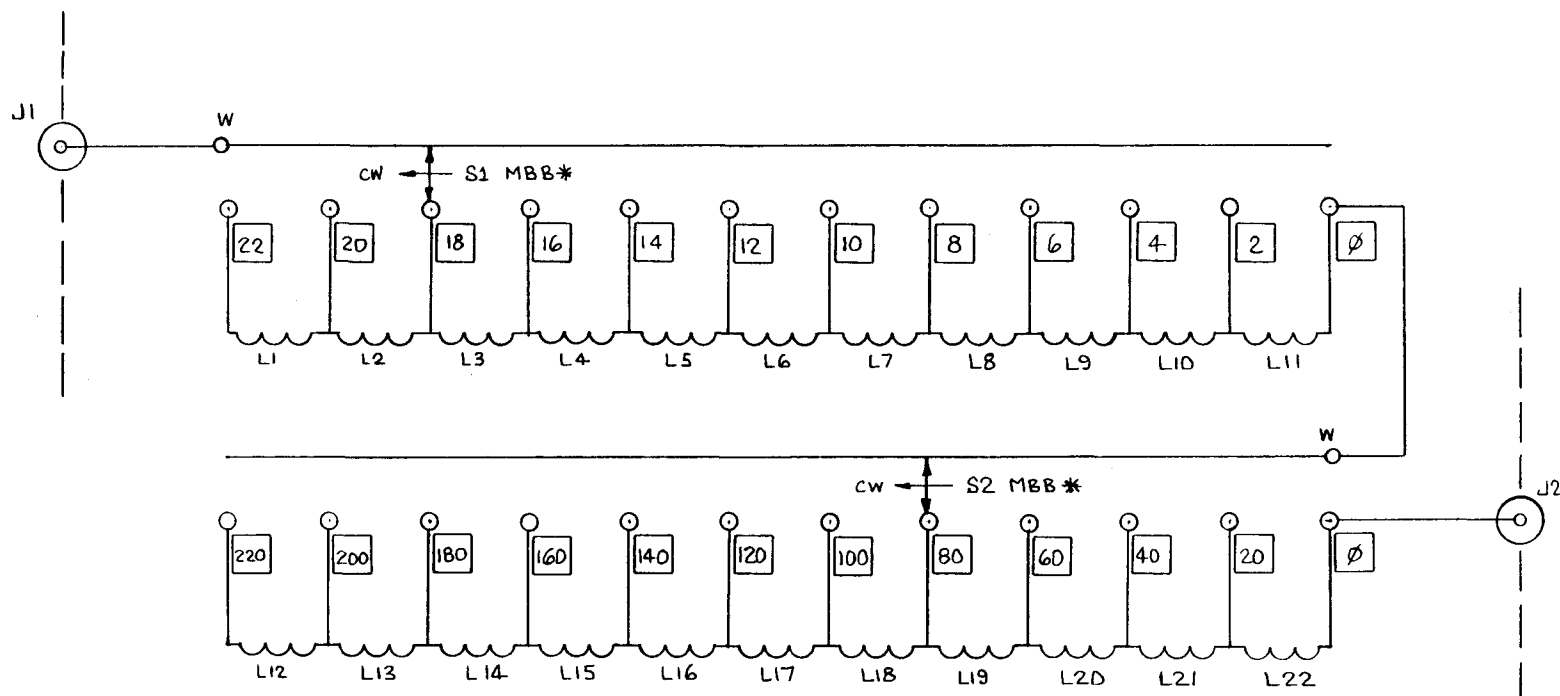
DRAWING NO.

CHECKED BY

SCALE

APPROVED BY

W.G.



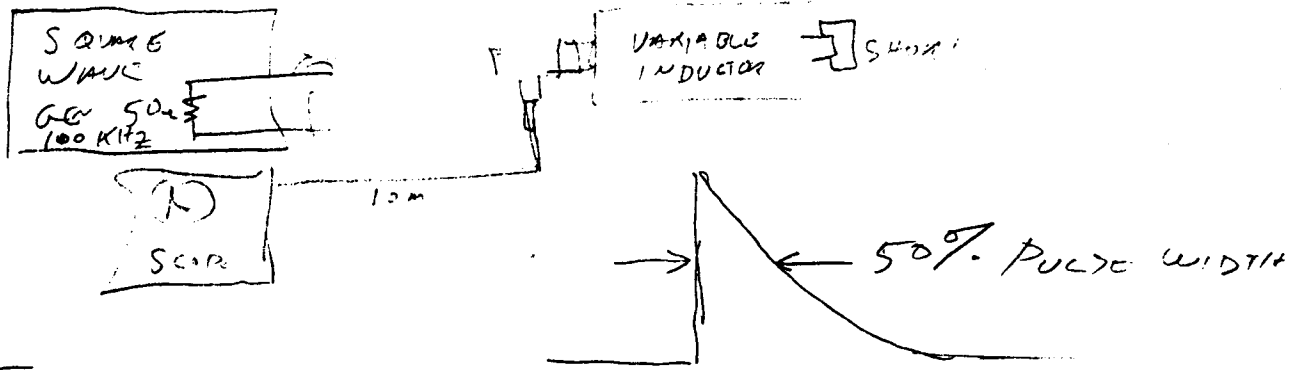
4. *S1 & S2 ARE MAKE-BEFORE-BREAK, 12-POSITION SWITCHES.
 3. VALUES IN BOXES ARE IN μH .
 2. L12-L22 = $20\mu\text{H}$
 1. L1-L11 = $2\mu\text{H}$
 NOTES:

CALIFORNIA INSTITUTE OF TECHNOLOGY
 GRAVITATIONAL PHYSICS

INDUCTANCE BOX: GRAVITY PHYSICS
 RANGE: 0 TO $242\mu\text{H}$ IN $2\mu\text{H}$ STEPS

DRAWN BY R.T.	DATE 5-8-89	DRAWING NO.
CHECKED BY	SCALE	
APPROVED BY	W.G.	

UNAKI 10/15/72 INDUCTANCE BOX TEST



TEST SET UP

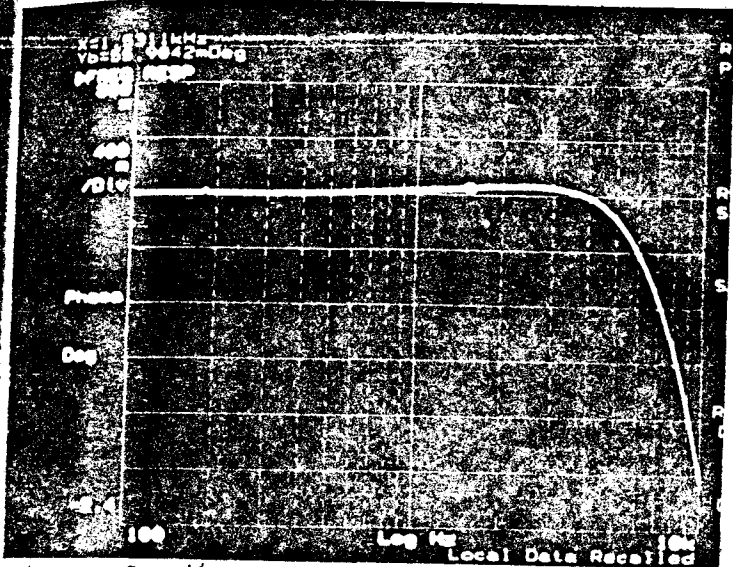
	INDICATED	INDUCTANCE	50% PULSE WIDTH	INDUCTANCE
		MHY	MS	MHY
	0	"	10	2.7
	2	"	46	3.3
	4	"	86	6.2
	6	"	112	8.1
	8	"	160	11.5
VEEPEK	10	"	180	13.0
	12	"	220	15.9
	14	"	250	18.0
	16	"	270	19.5
	18	"	290	20.9
	20	"	315	22.7
	22	"	350	25.2
	20	"	330	23.8
	40	"	660	47.6
	60	"	960	69.2
	80	"	1240	89.4
	100	"	1520	110.0
	120	"	1840	133.0
	140	"	2150	155.0
COARSE	160	"	2450	177.0
	180	"	2700	195.0
	200	"	3000	216.0
	220	"	3250	234.0

INDUCTANCE BOX TEST

9-MAY-89

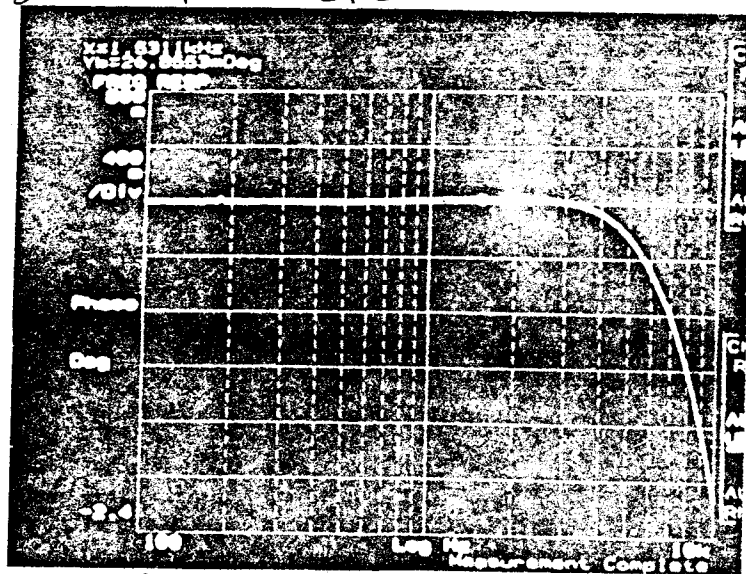
9-MAY-89

9:20

 $L = 52 \mu\text{H}$

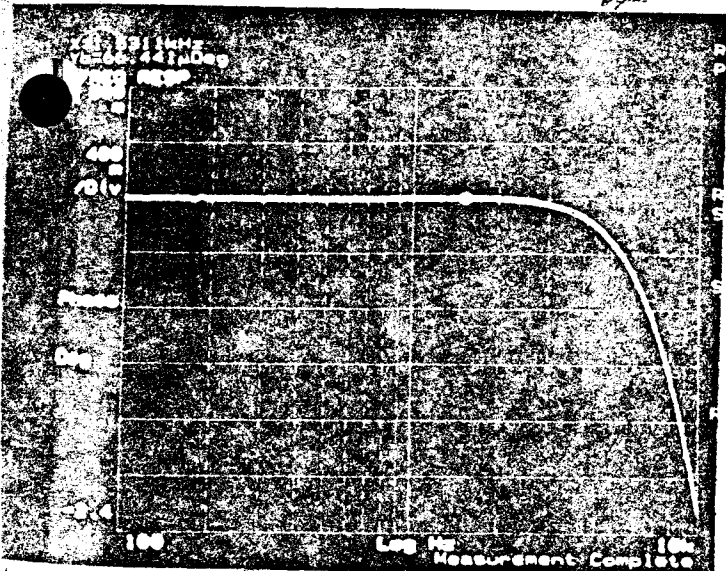
9-MAY-89

9:25

 $L = 54 \mu\text{H}$

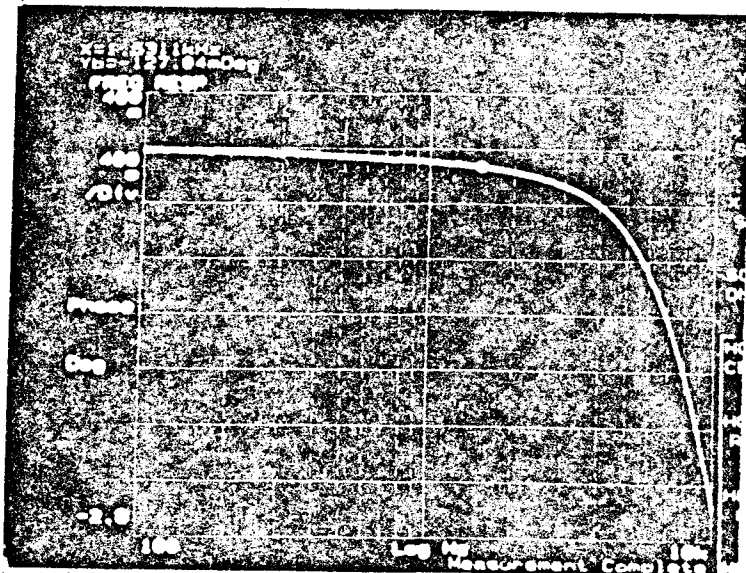
9-MAY-89

9:35

 $L = 56 \mu\text{H}$

9-MAY-89

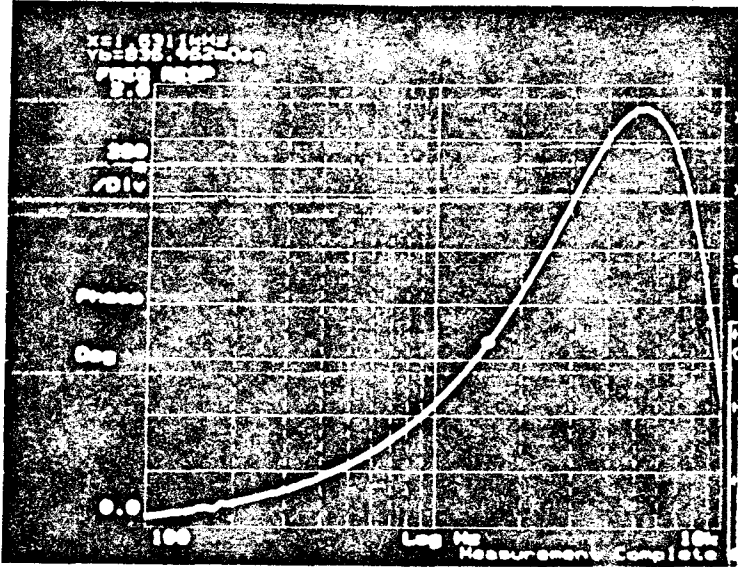
9:00

 $L = 60 \mu\text{H}$

Using the new adjustable inductor in series w. Lorie coils (instead of the fixed $80 \mu\text{H}$ one), phase matching of currents in the two coils for they, Lorie (respectively), was checked w. the set-up on p. 77. Results are shown in order of increasing L . Best matching: $L = 56 \mu\text{H} \rightarrow$ if the resulting $66 \mu\text{deg}$ were the only limit, a coil subtraction of $\sim 10^5$ would be possible. Of, course, at this level differences in the man/pendulum (etc) response become the real limits.

9-MAY-89

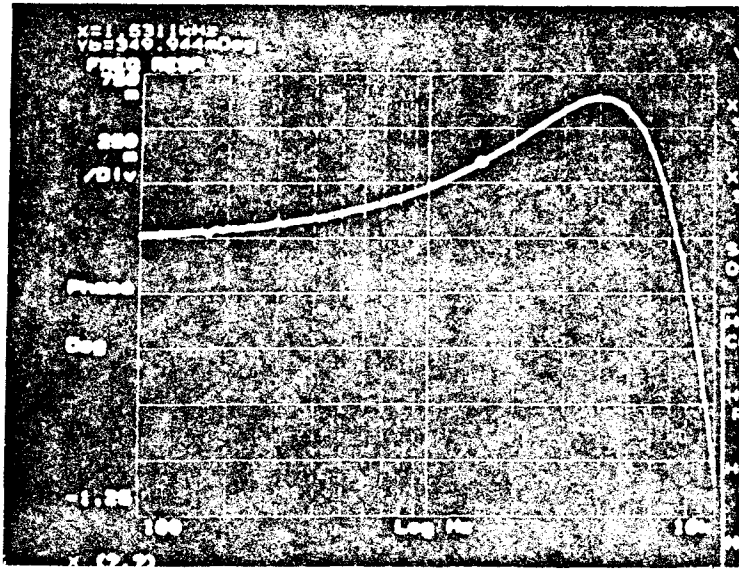
8:50



$L = 0 \mu H$

9-MAY-89

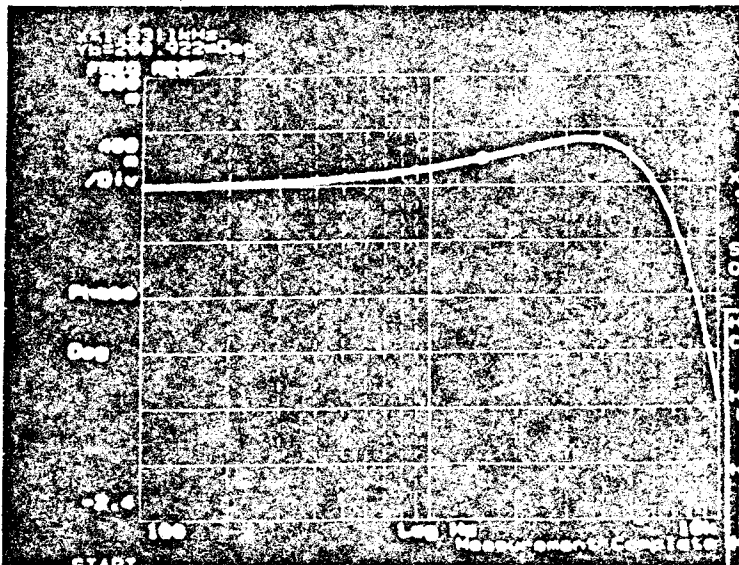
8:45



$L = 34 \mu H$

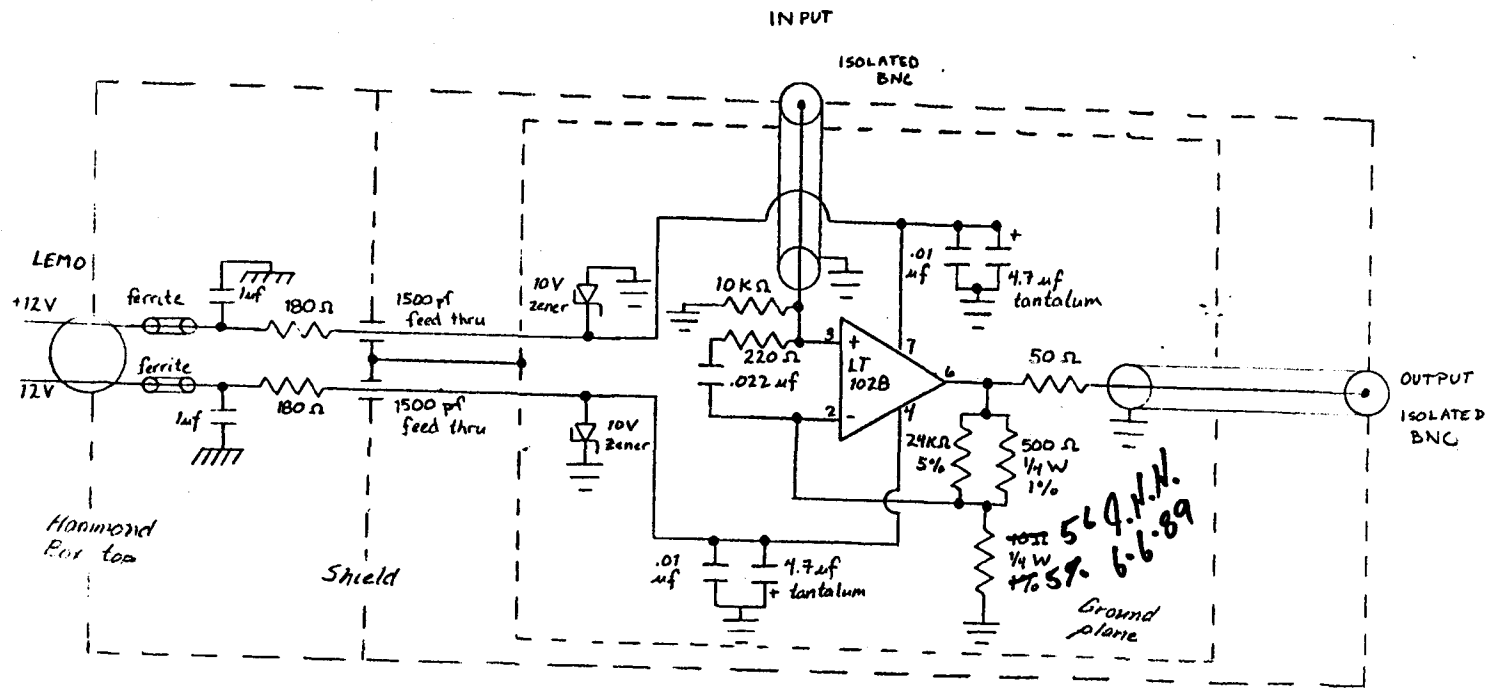
9-MAY-89

9:15



$L = 40 \mu H$

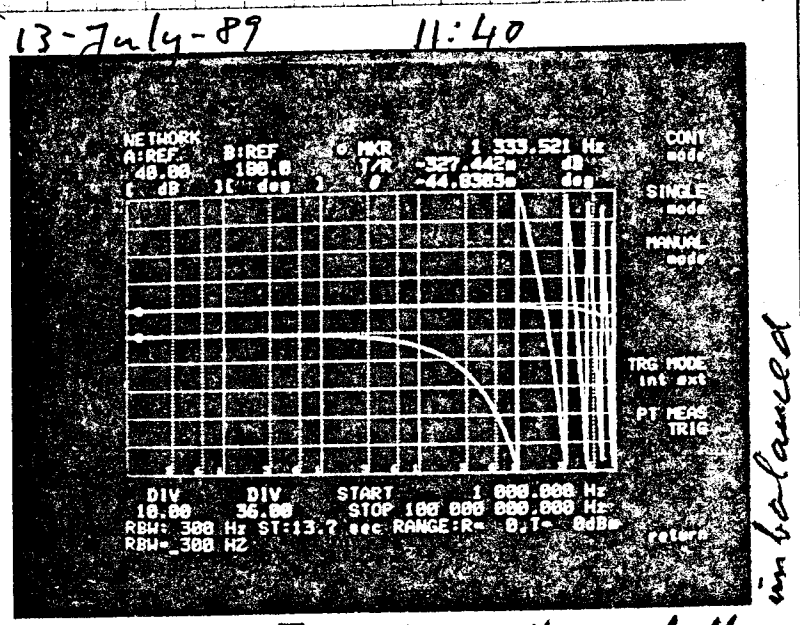
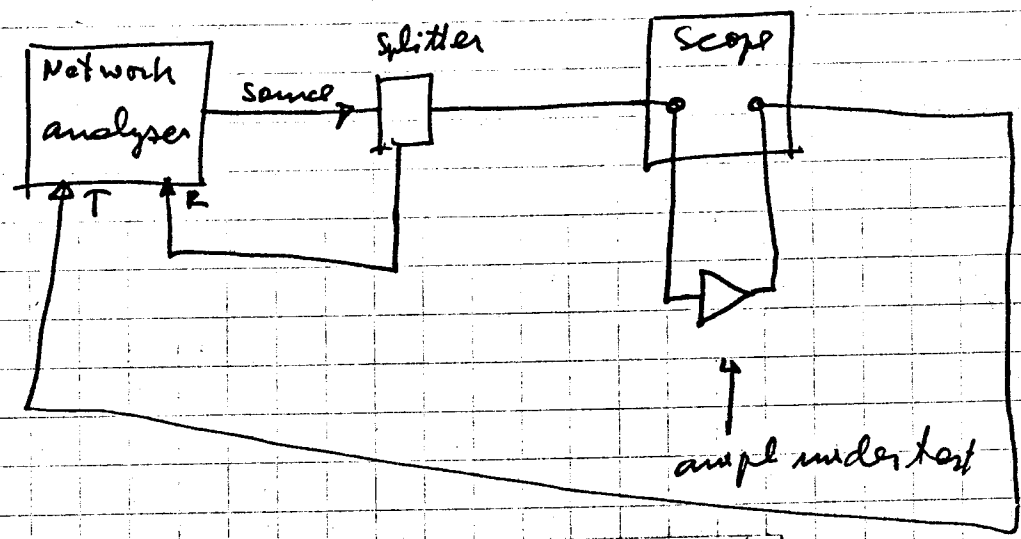
$\frac{dF}{di} =$



LOW NOISE BUFFER
 8-21-87

13-July-89

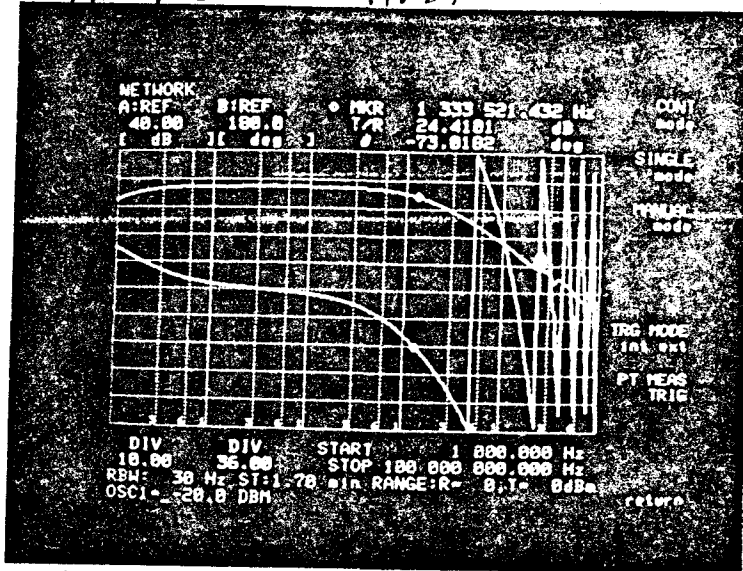
Since ~30' of cables were used to connect the amplifiers under test (see diagram below), we took the response of the cables themselves (trace at 11:40)



TR. FUNCTION OF CABLES - they are badly

- It turns out that at 250 kHz the cables alone show a phase shift of 5°.
- Therefore, the ~3' cable from splitter to R input was replaced with a long cable, matched to the one in the test path. See traces overleaf

13-July-89 11:21



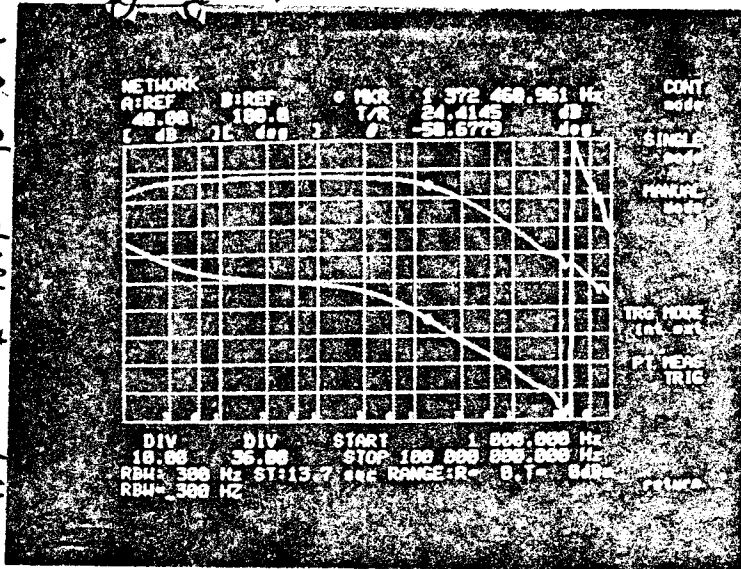
AXES: 100 MHz 40dB
50K 713

← unbalanced cables

50K 50 Ohm in and out

13-July 89 12:02

100 phase shift at 230 KHz



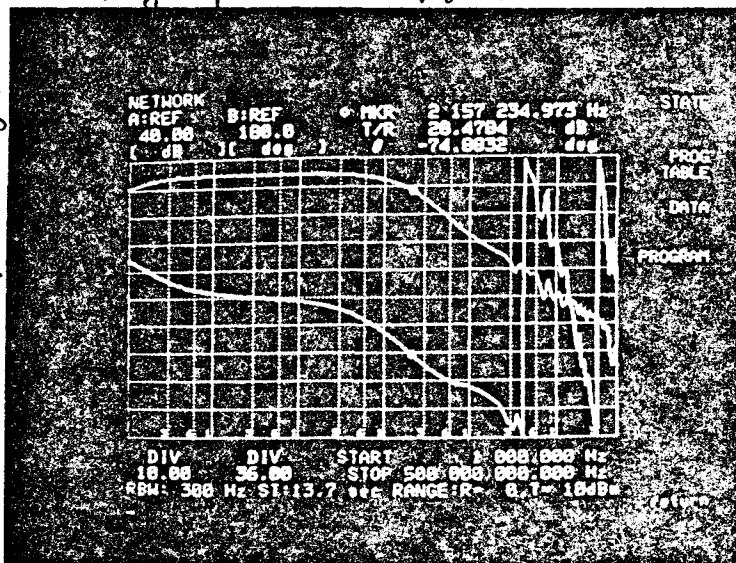
50K in and out 50K 713

← matched cables

50K 50 Ohm in and out

14-July 89 17:40

100: 203 12MHz

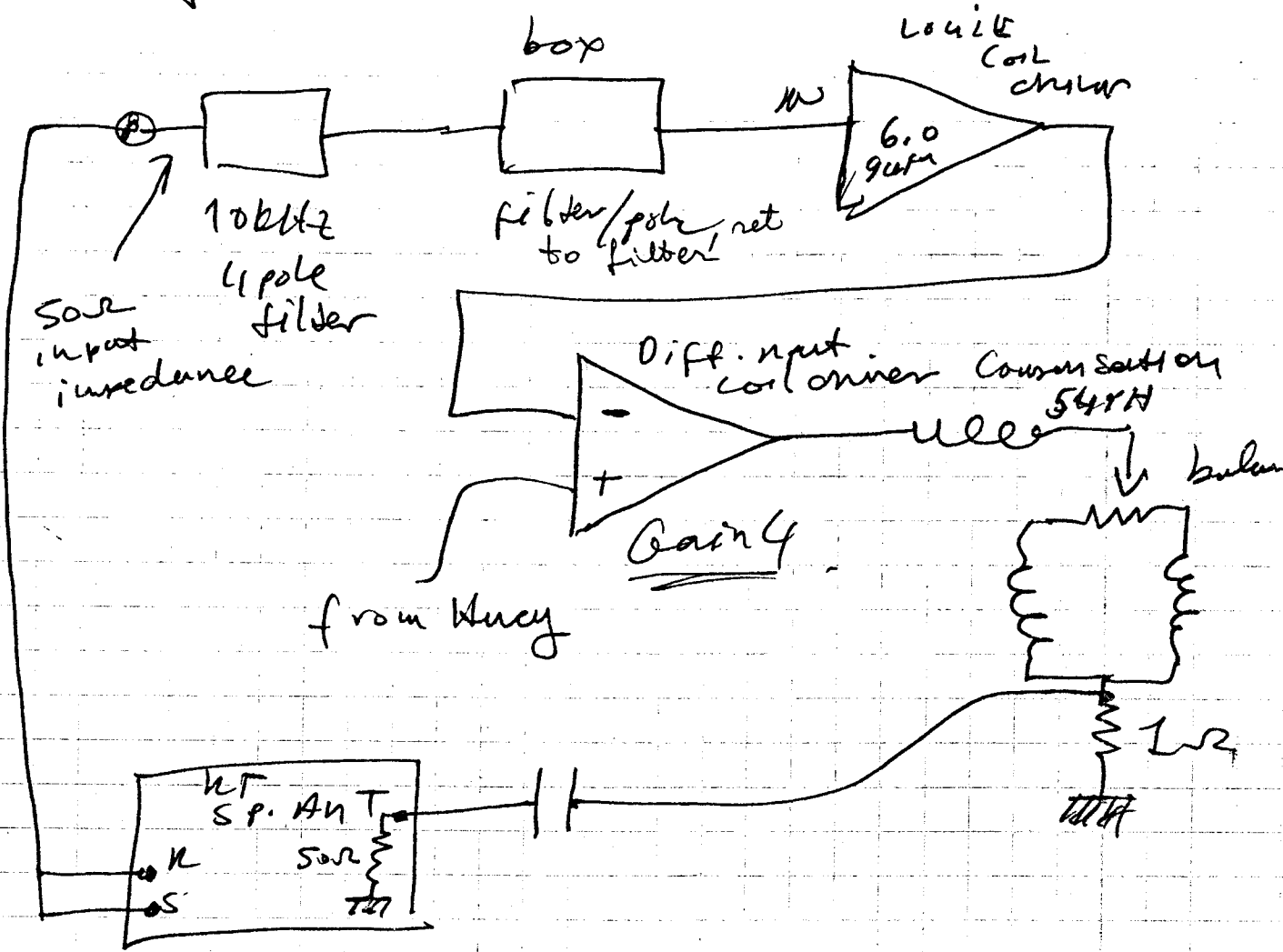


50K 713

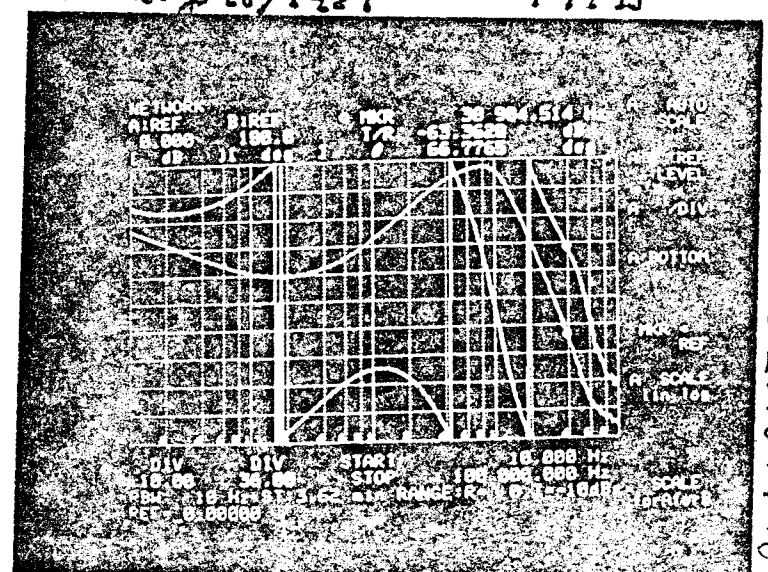
50K 50 Ohm in and out

(1)

Second arm chain transfer functions



July 18, 1989 17:15

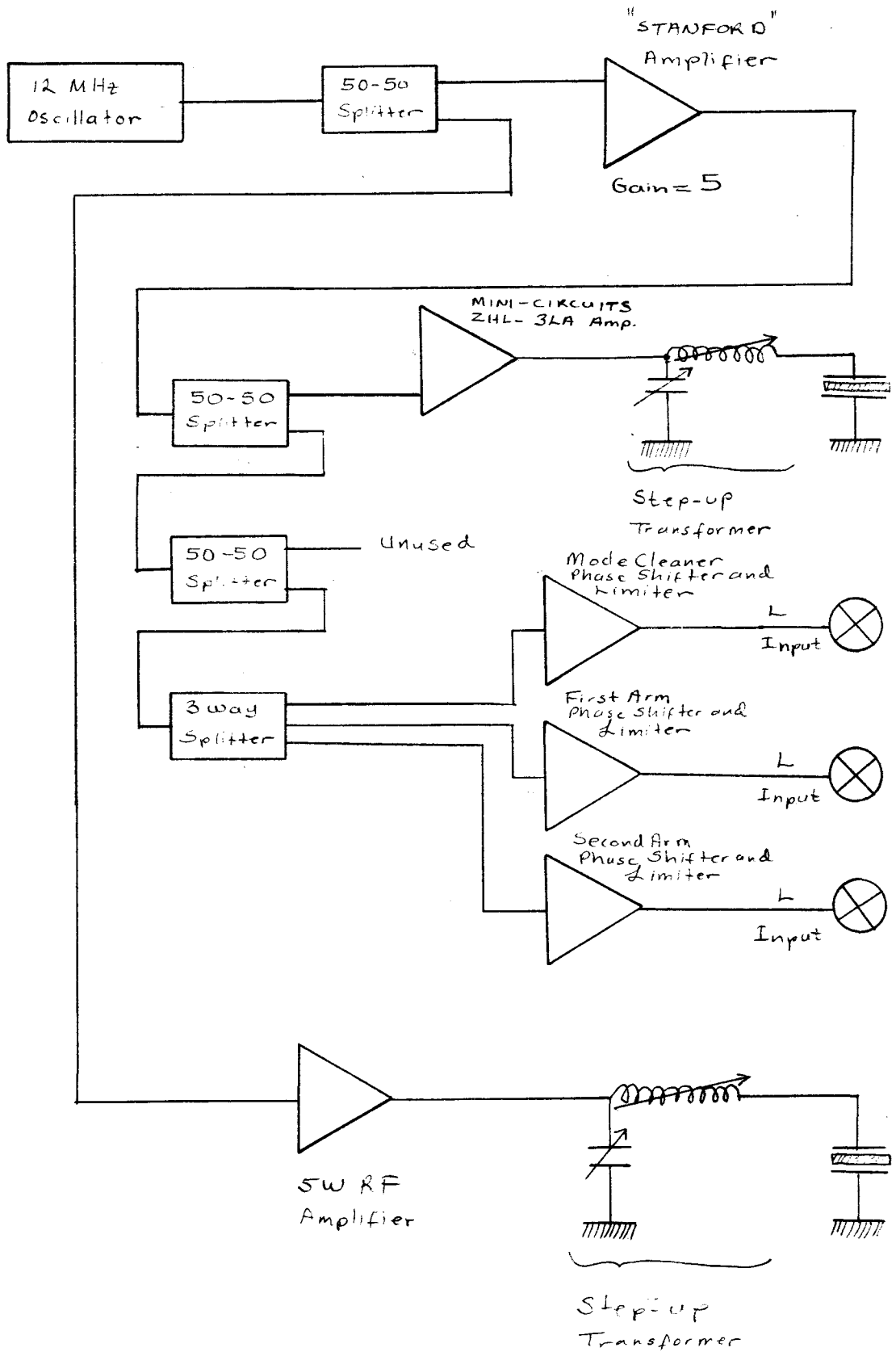


second arm chain transfer function

BATCH
START

STAPLE
OR
DIVIDER

12 MHz RF distribution for Modulation:



Mode cleaner
Modulation
Pockels
Cell

Mode
Cleaner
Mixer

First Arm
Mixer

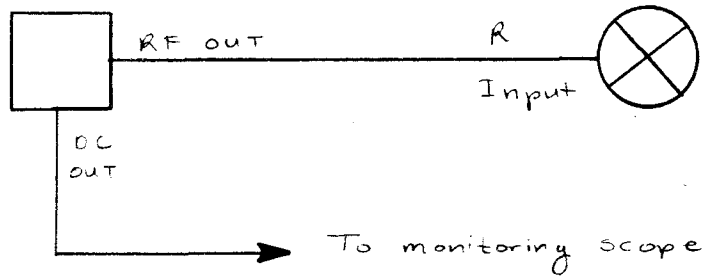
Second
Arm
Mixer

Main
System
Modulation
Pockels
Cell

July 24, 1984
YB

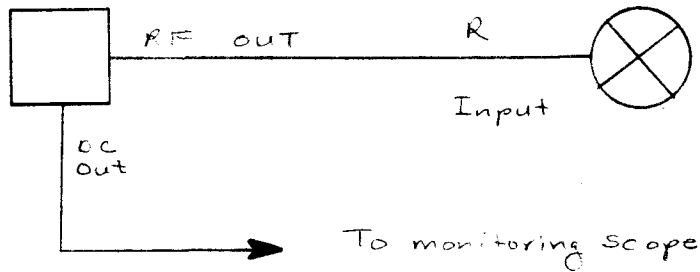
Photo Detector Electrical Wiring Diagrams:

Mode Cleaner
Photodiode and Amplifier



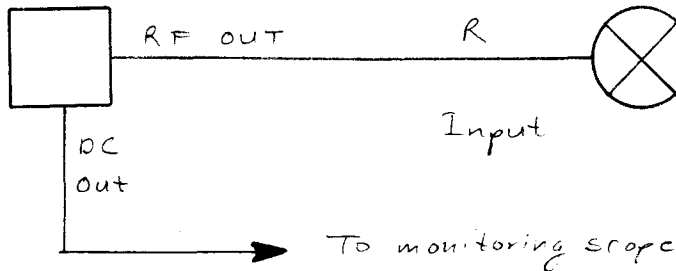
Mode Cleaner
Mixer

First Arm Photodiode
and Amplifier



First Arm
Mixer

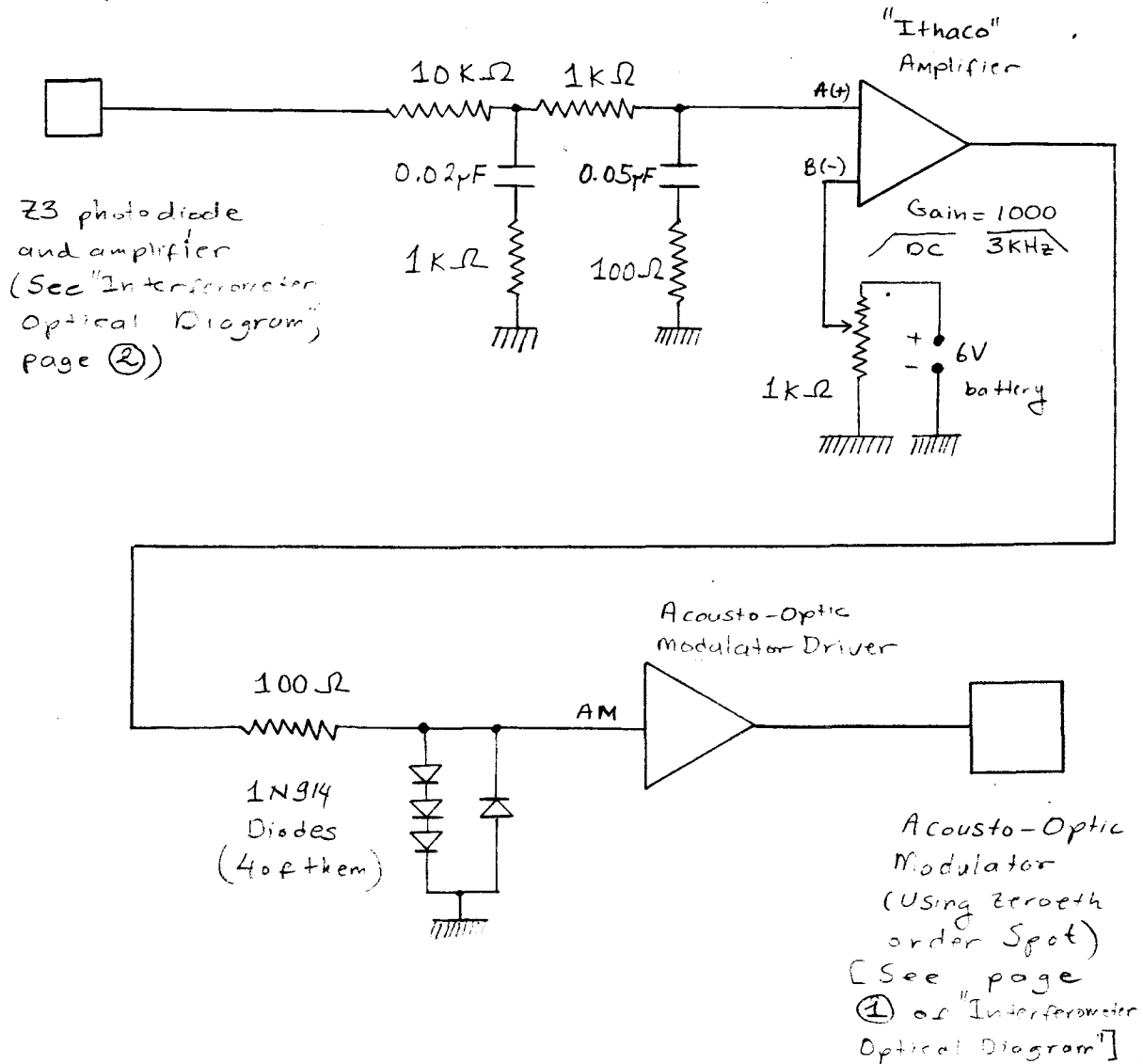
Second Arm Photodiode
and Amplifier



Second Arm
Mixer

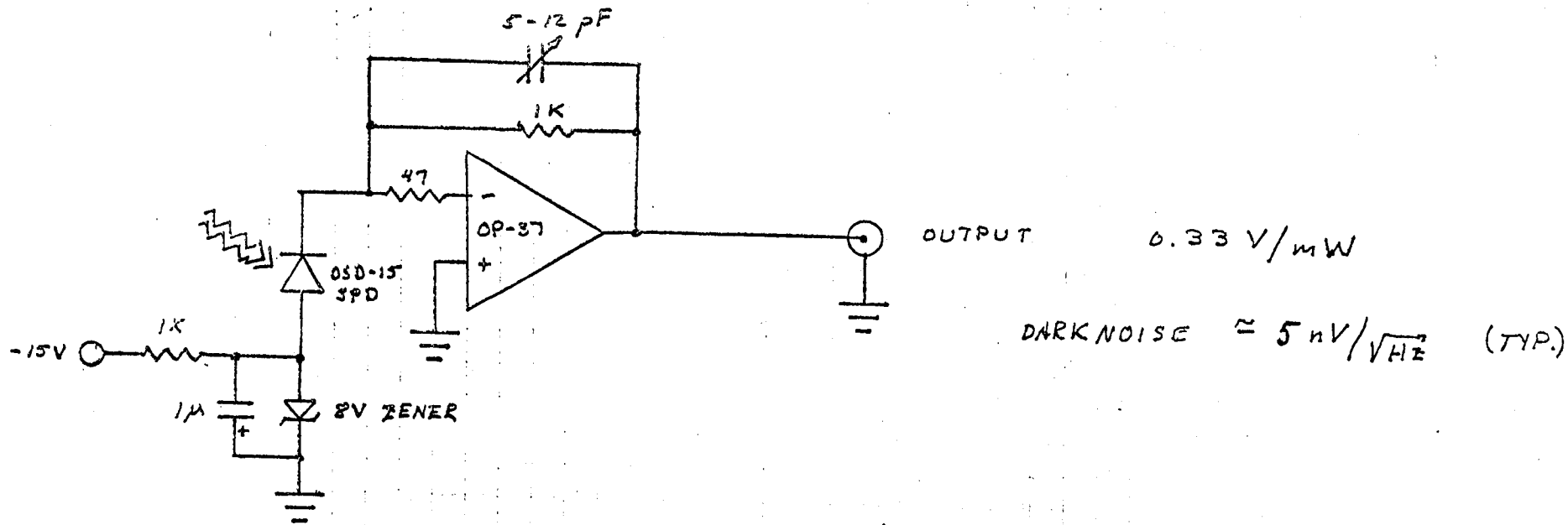
July 24, 1984
JE

Intensity Stabilizing Servo (The Noise Eater)



NOISE EATER SENSING DIODE

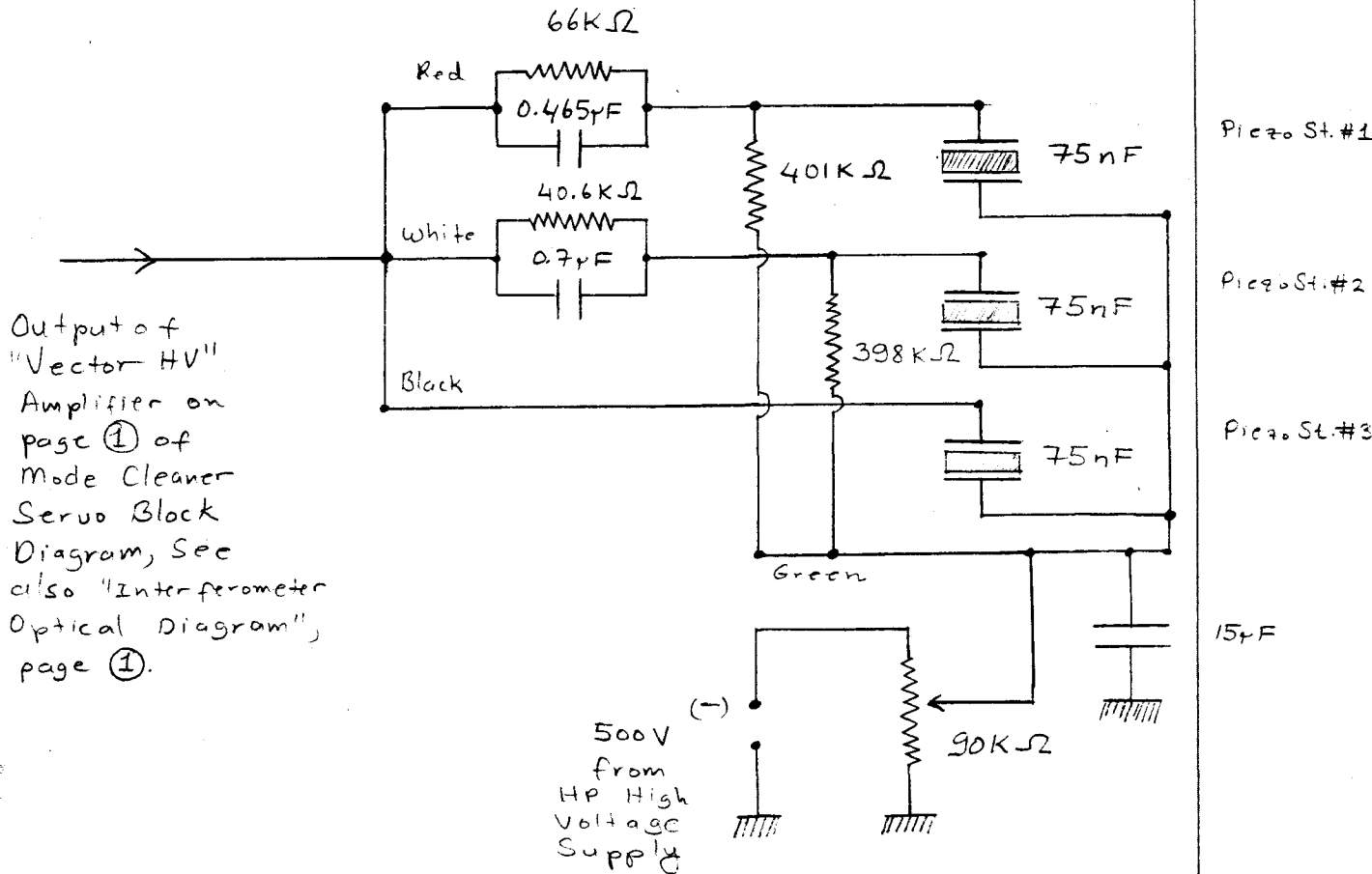
2/27/85 MEZ



23 Photodiode Amplifier

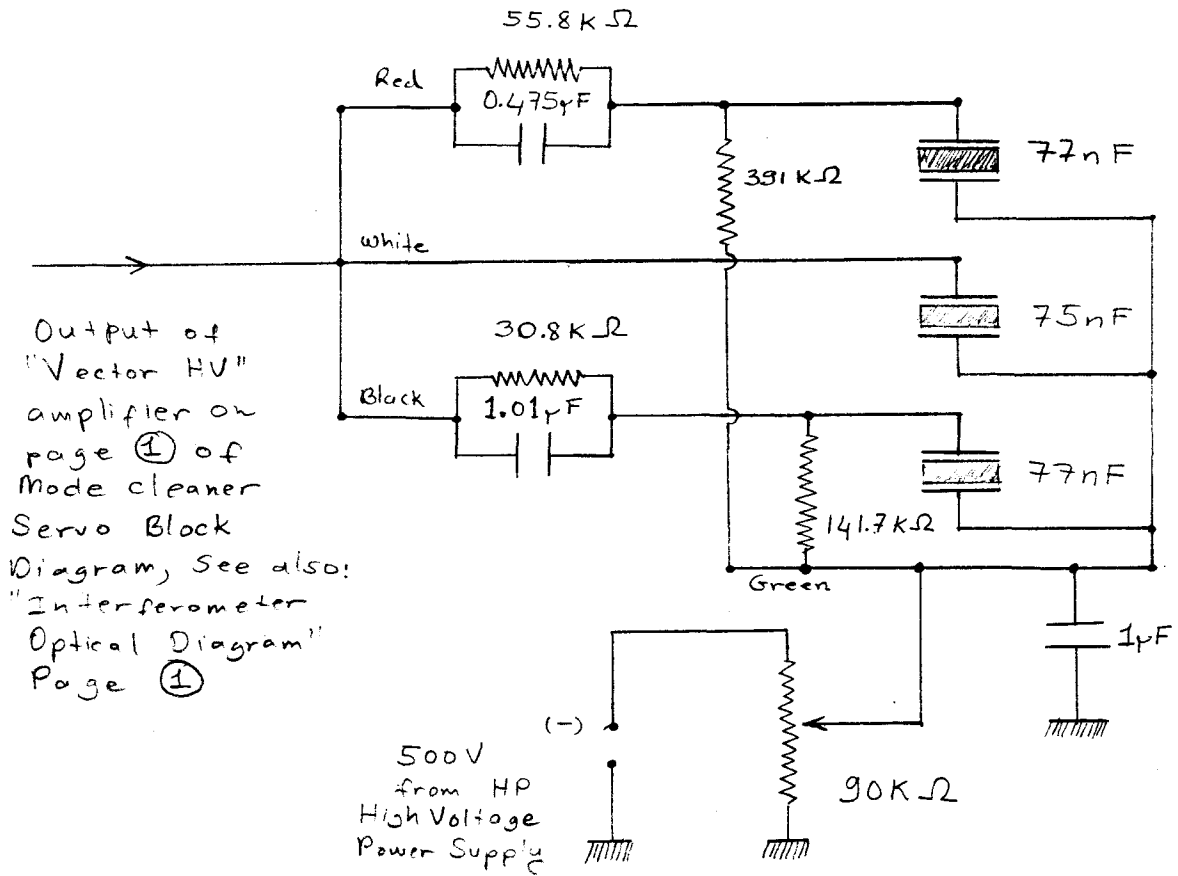
(See "The Intensity Stabilizing Servo" and the "Interferometer Optical Diagrams")

3 Piezo Stack Balancing Network (For the Laser "BARNEY")



Also on page ① of "mode cleaner servo Block Diagram".

3 Piezo Stack Balancing Network (For the laser "FRED")



Piezo St. #1

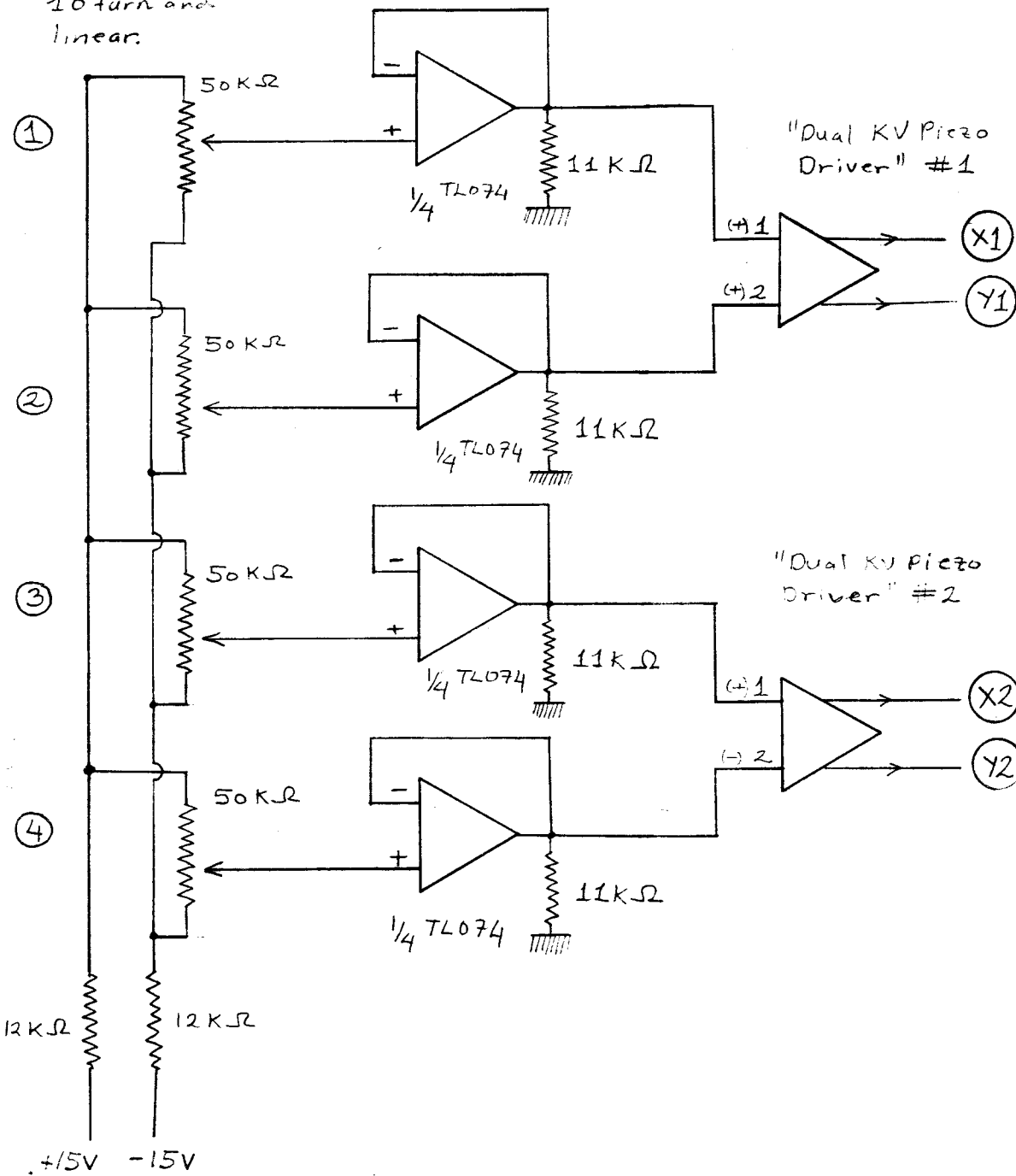
Piezo St. #2

Piezo St. #3

Also on page ① of "Mode Cleaner Servo Block Diagram"

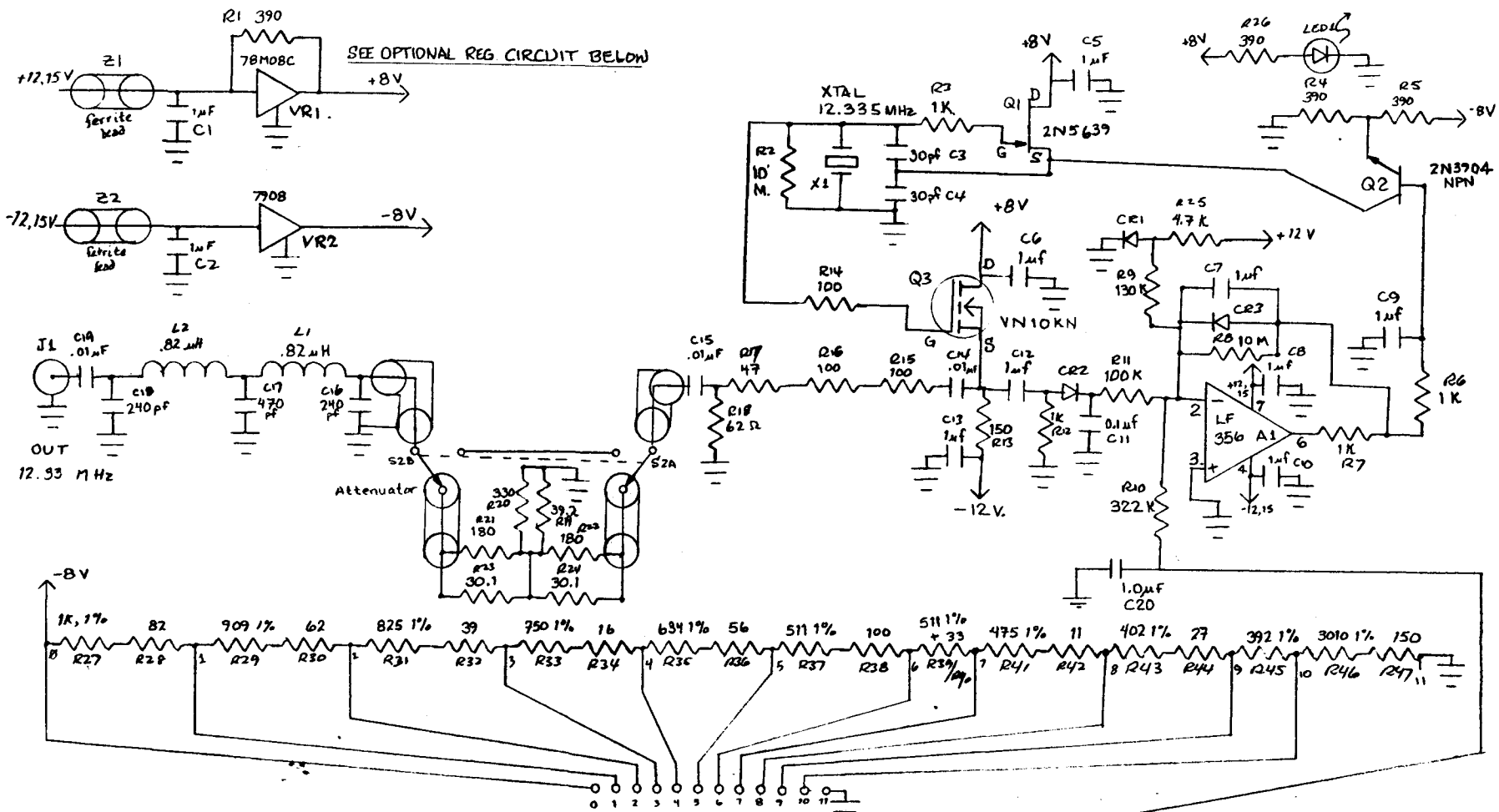
Mode Cleaner Alignment Fine Tuning

All pots are
10 turn and
linear.



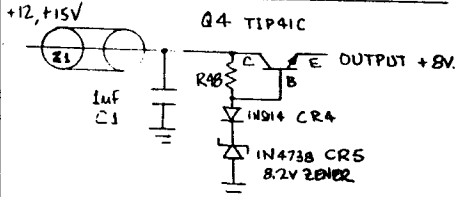
On page
① of
Interferomet
Optical
Diagram

On page
① of
Interferomet
Optical
Diagram



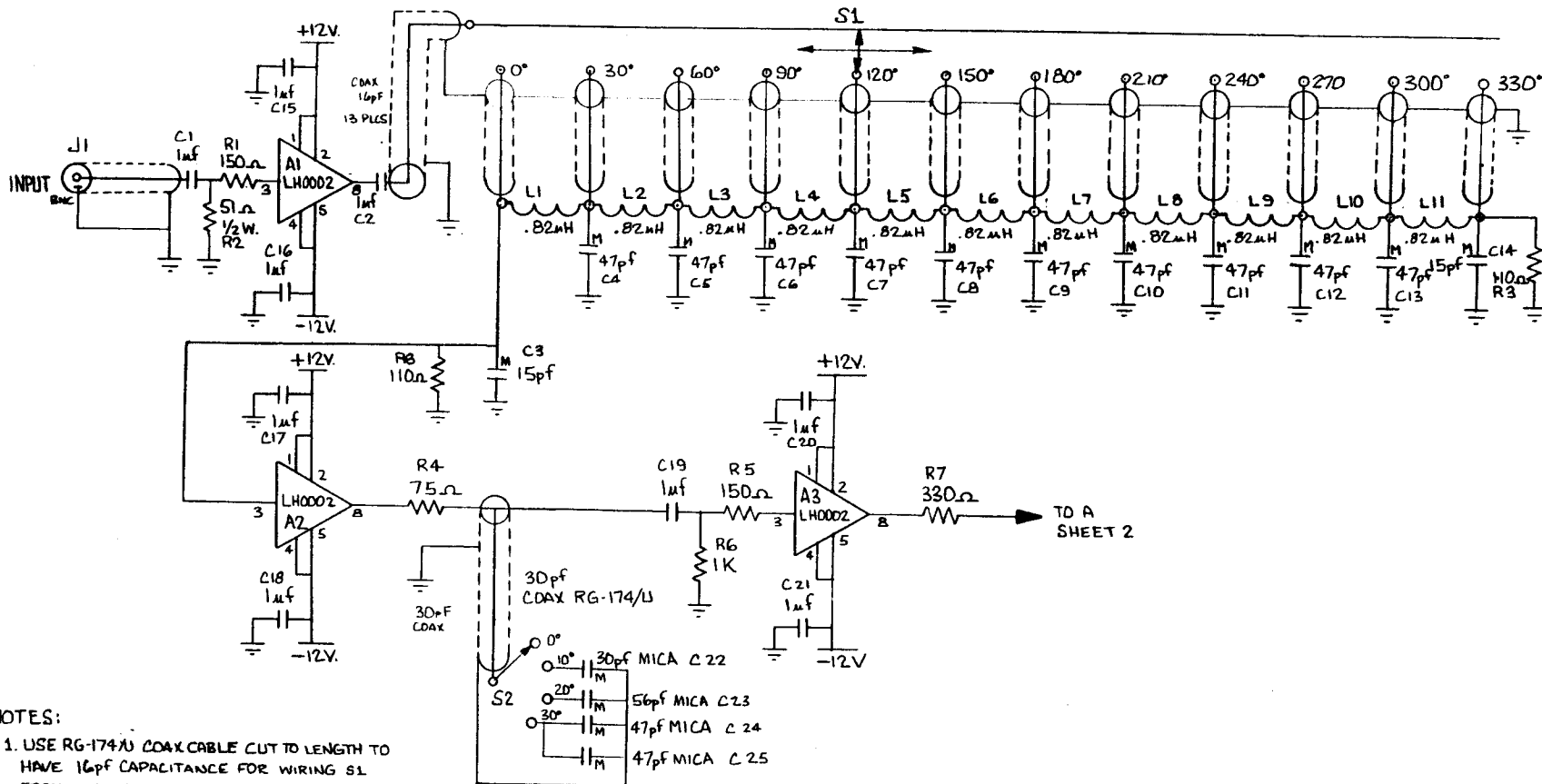
- L2
- A1
- LED1
- S2
- R4B
- Q4
- C20
- CR5

OPTIONAL +8V REGULATOR CIRCUIT



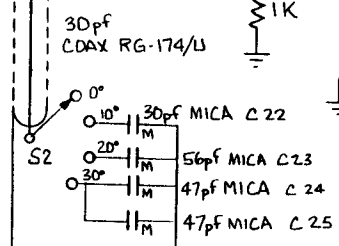
CALIFORNIA INSTITUTE OF TECHNOLOGY GRAVITATIONAL PHYSICS		
12.335 MHz OSCILLATOR		
DRAWN BY <i>E. Lindley</i>	DATE 8/5/87	DRAWING NO.
CHECKED BY	SCALE	
APPROVED BY	W.D.	

UPDATED 10-24-88

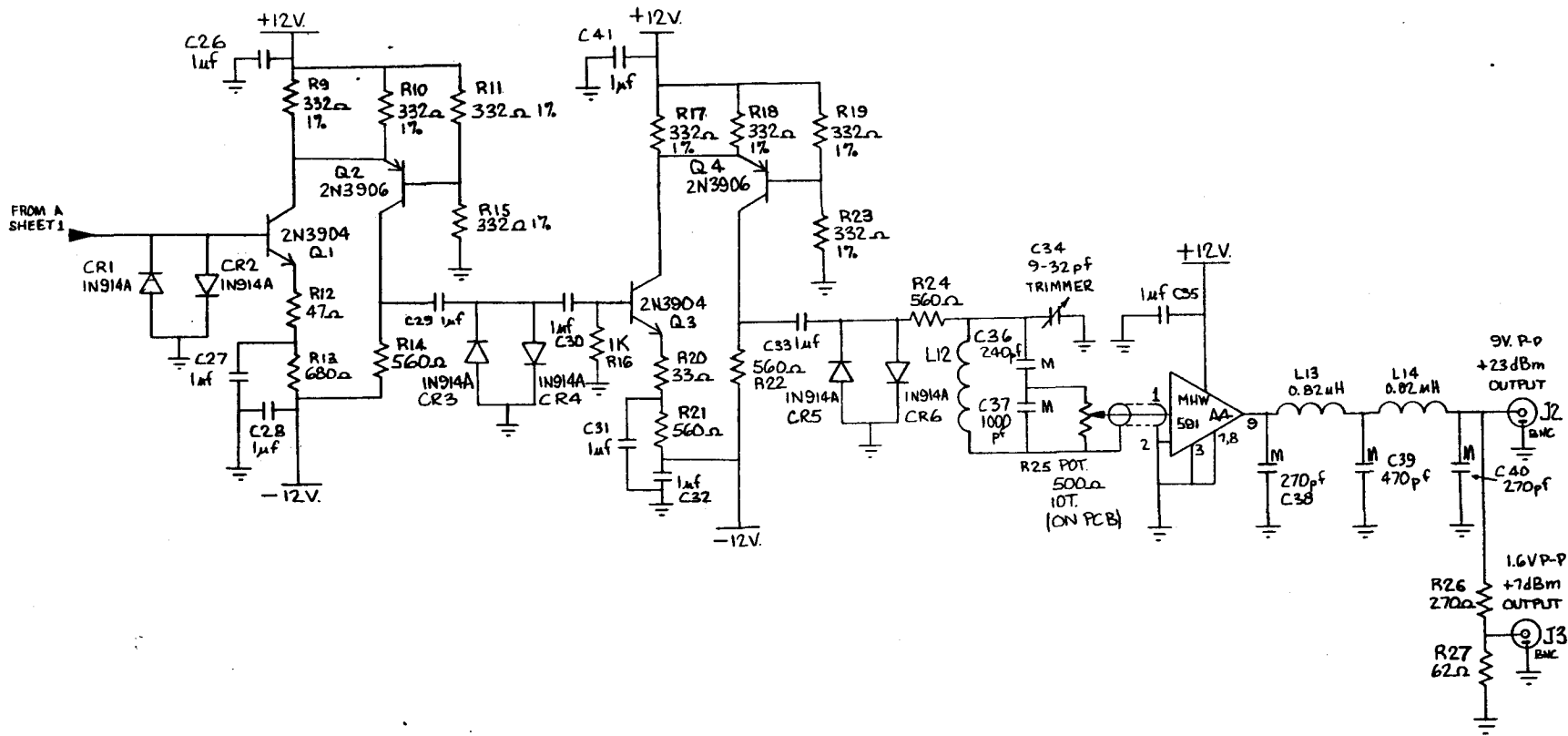


NOTES:

1. USE RG-174/U COAX CABLE CUT TO LENGTH TO HAVE 16pf CAPACITANCE FOR WIRING S1 FROM TERMINALS TO IND. $.82\mu\text{H}$
2. COAX FROM $75\Omega/1\mu\text{f}$ NODE (OUTPUT OF A2) TO S2 WILL BE CUT TO A LENGTH TO HAVE A 30pf CAPACITANCE. DRESS COAX AWAY FROM COMPONENTS ON P.C.B.
3. GND ALL COAX FROM ISO-BNC ON FRONT PANEL TO P.C. CARD GND. PLANE.



CALIFORNIA INSTITUTE OF TECHNOLOGY		
GRAVITATIONAL PHYSICS		
12.33 MHZ. LIMITER PHASE SHIFTER		
SHEET 1		
DRAWN BY B.T.	DATE 11-6-87	DRAWING NO.
CHECKED BY	SCALE NONE	
APPROVED BY	W.O.	-1



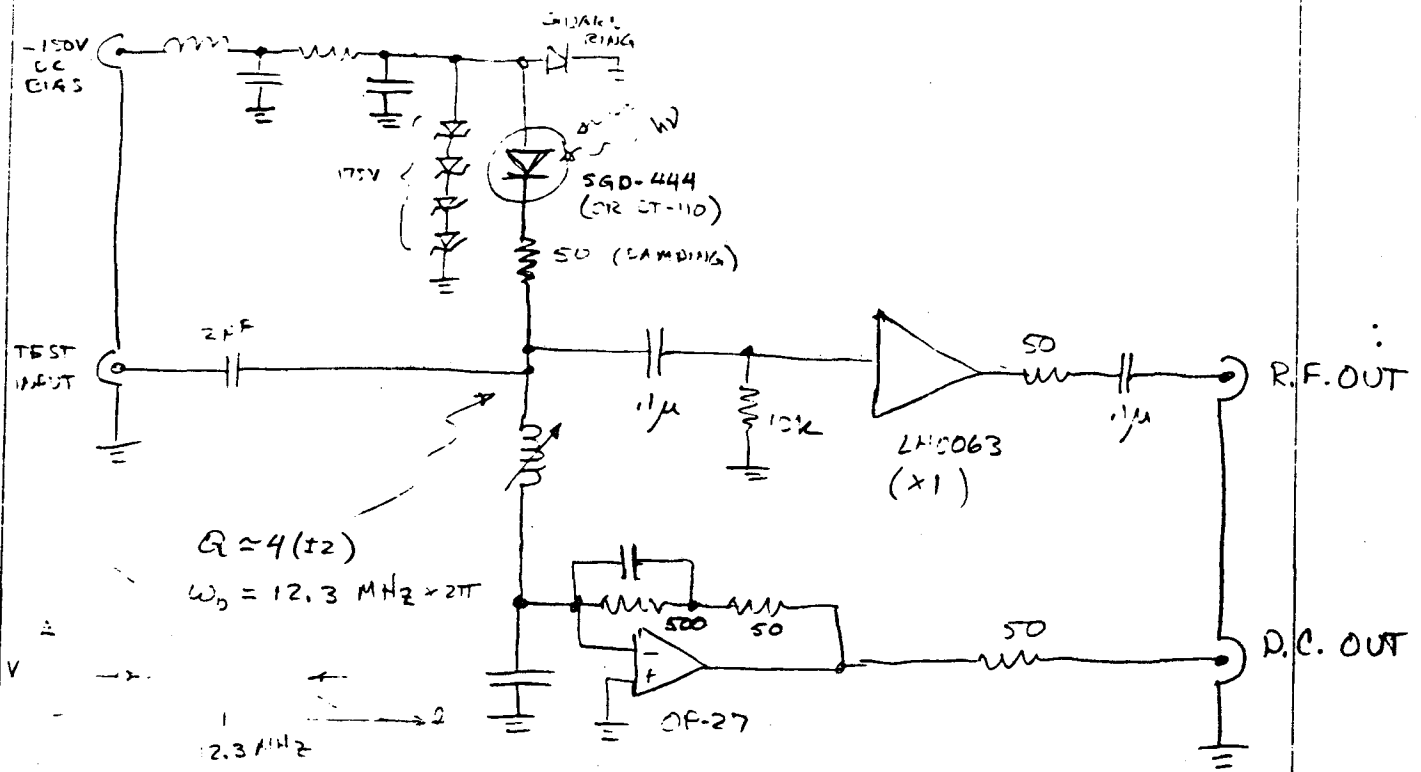
A4
Q4
J3
L14
CR6
C40
R27

LAST NO. USED R27

CALIFORNIA INSTITUTE OF TECHNOLOGY GRAVITATIONAL PHYSICS		
12.33 MHz. LIMITER PHASE SHIFTER SHEET 2		
DRAWN BY B.T.	DATE 11-6-87	DRAWING NO.
CHECKED BY	SCALE	-2
APPROVED BY	W.G.	

11/14/88 WSE

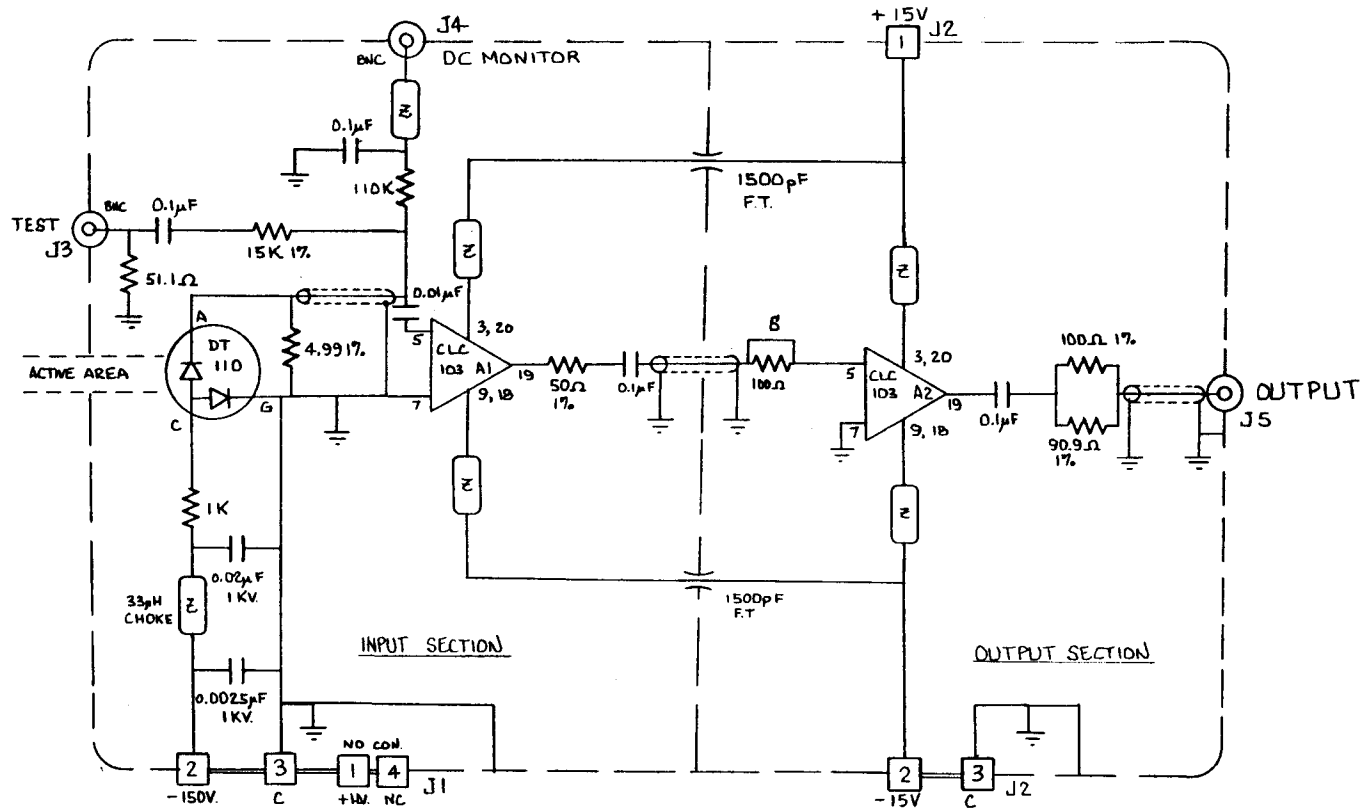
MODE CLEANER SERVO PHOTODIODE / BUFFER (STANDARD CALTECH FRONT END)



$Q \approx 4(t_2)$
 $\omega_0 = 12.3 \text{ MHz} \times 2\pi$

R.F. DARK NOISE \approx SHOT NOISE AT 1mW, 5145 \AA

The First and The Second Arm
Photodiodes are the same as this
one. JG. July 25, 1989



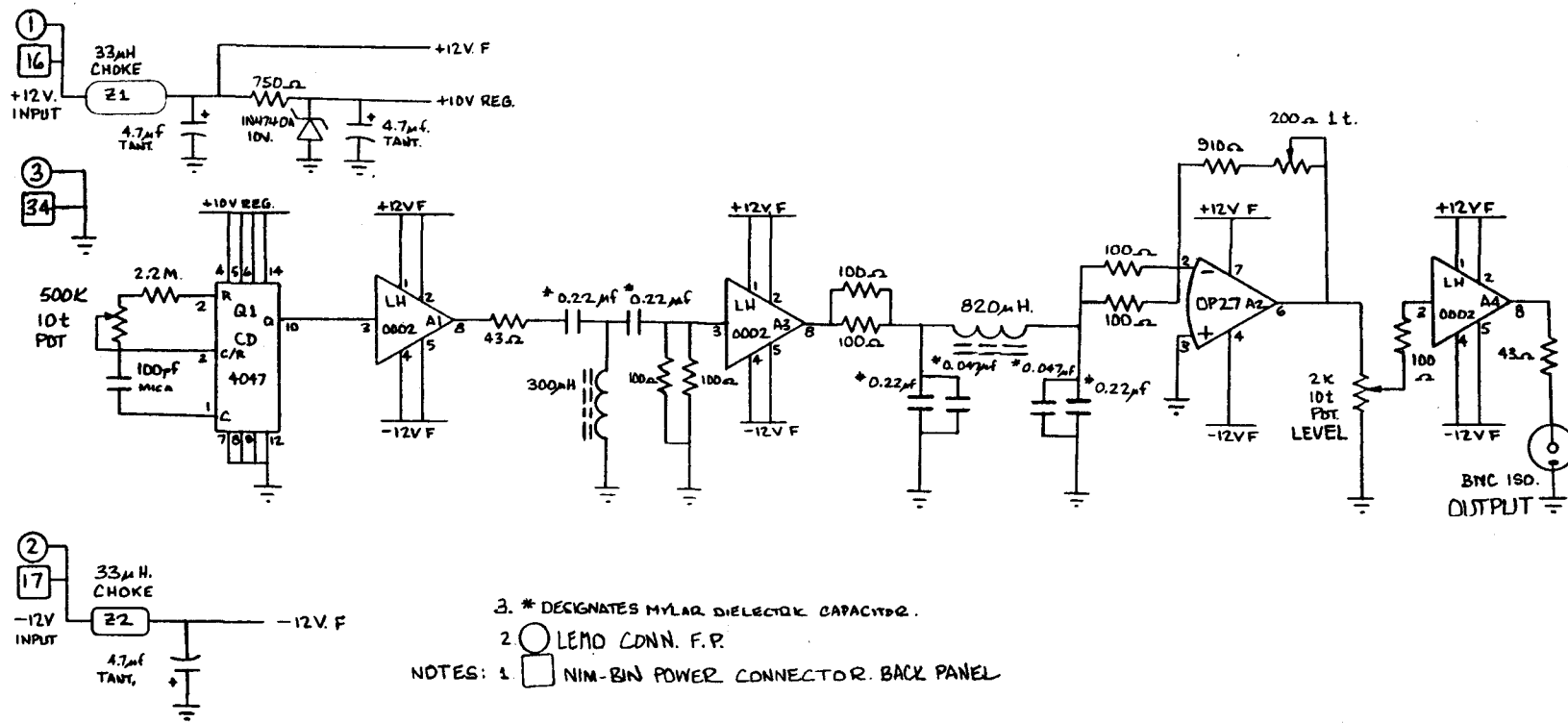
- NOTE 1. COMPONENTS MARKED "Z" ARE 33μH CHOKE.
 2. RESISTORS 1/4W 5% UNLESS MARKED.
 3. J1 IS 4-PIN LEMO.
 4. J2 IS 3-PIN LEMO

FAST PHOTODIODE AMP. 3/29/88

CALIFORNIA INSTITUTE OF TECHNOLOGY
 GRAVITATIONAL PHYSICS

FAST PHOTODIODE AMPLIFIER SHT. 1

DRAWN BY B.T.	DATE 3/29/88	DRAWING NO.
CHECKED BY	SCALE	- 1
APPROVED BY	W.O.	

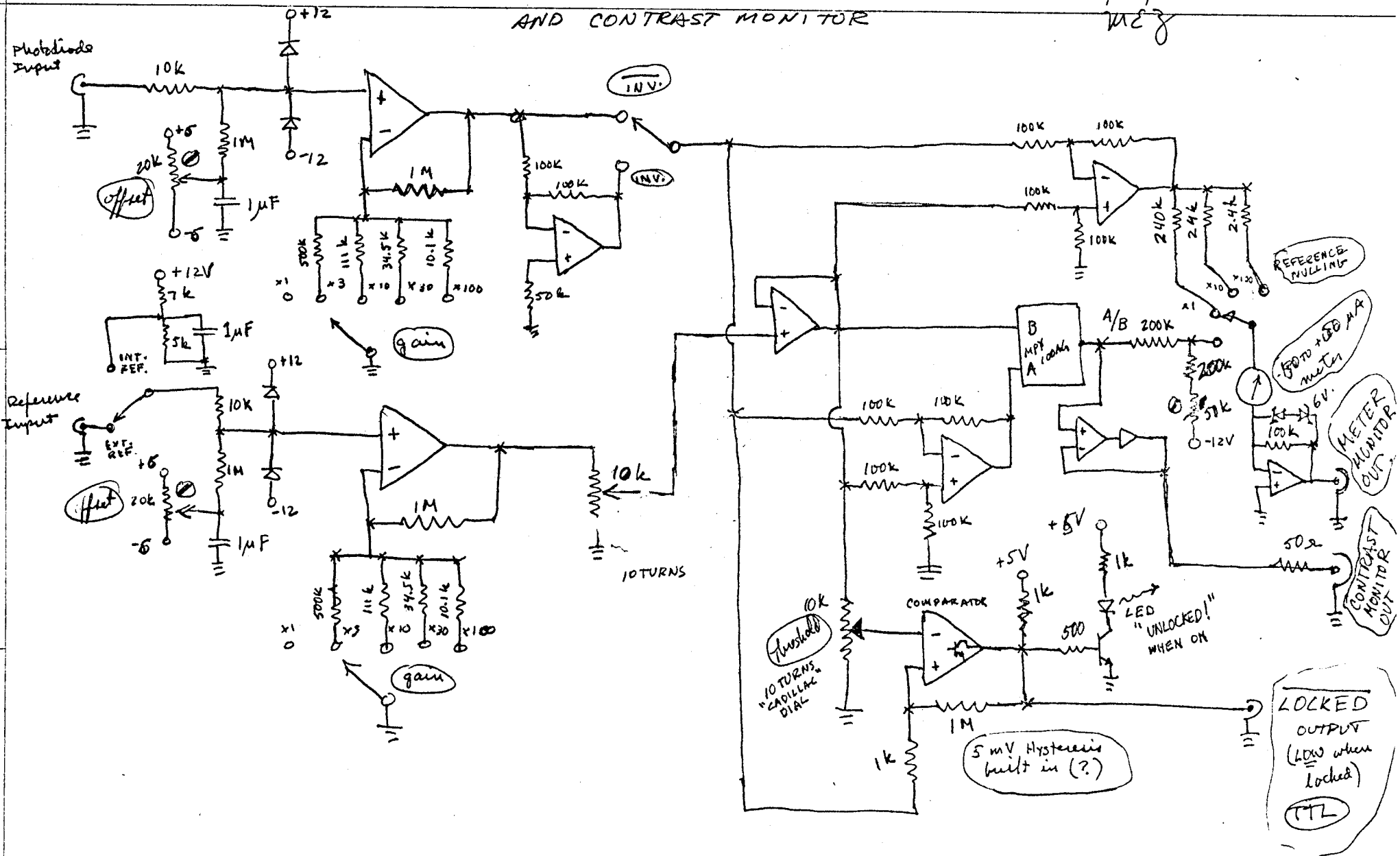


3. * DESIGNATES MYLAR DIELECTRIC CAPACITOR.
 2. ○ LEMO CONN. F.P.
 NOTES: 1. □ NIM-BIN POWER CONNECTOR. BACK PANEL

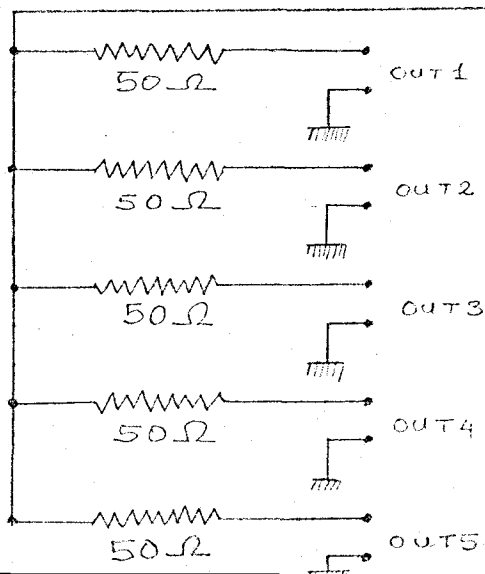
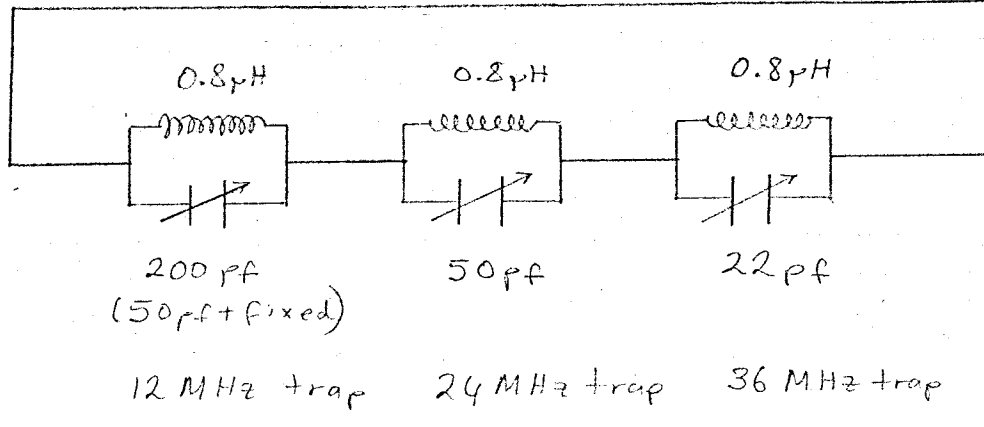
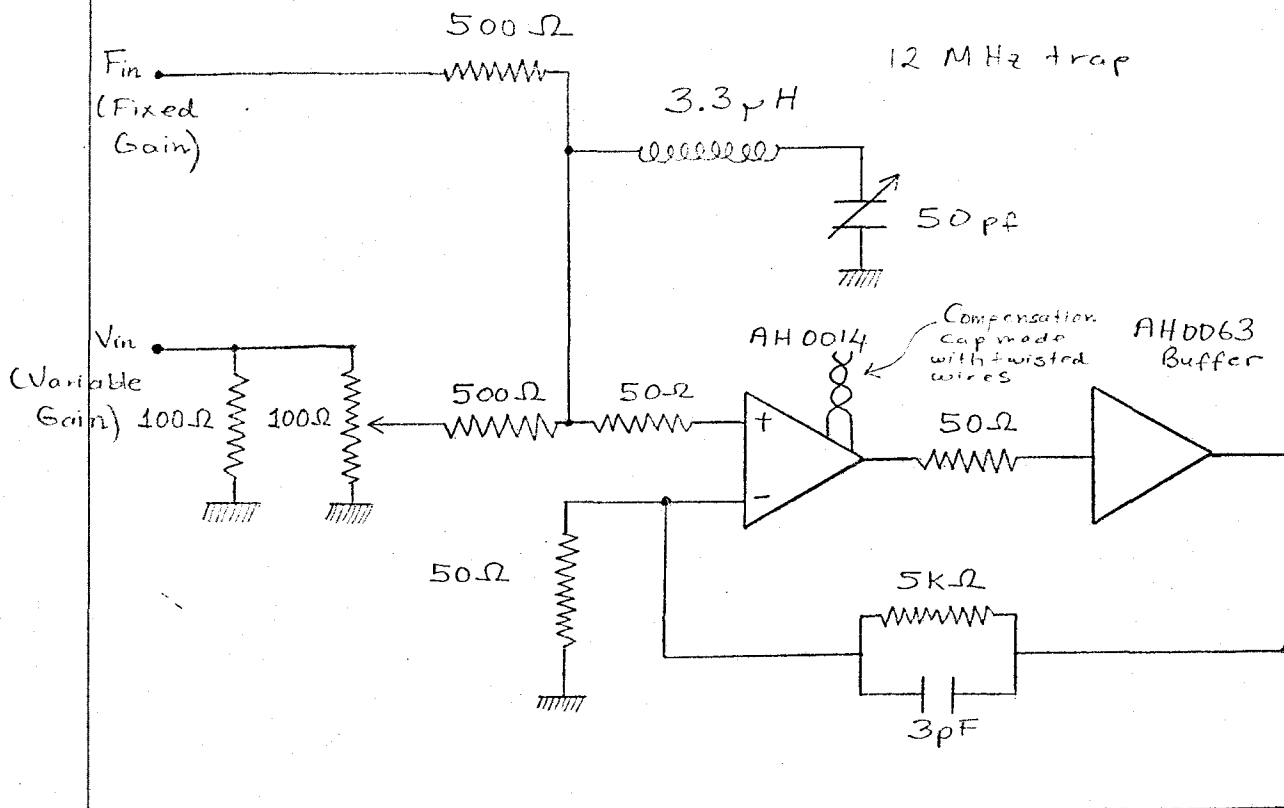
CALIFORNIA INSTITUTE OF TECHNOLOGY GRAVITATIONAL PHYSICS		
COMB CALIBRATOR		
DRAWN BY BT	DATE 9-24-87	DRAWING NO. -1
CHECKED BY	SCALE	
APPROVED BY	W.G.	

LOCKUP DISCRIMINATOR AND CONTRAST MONITOR

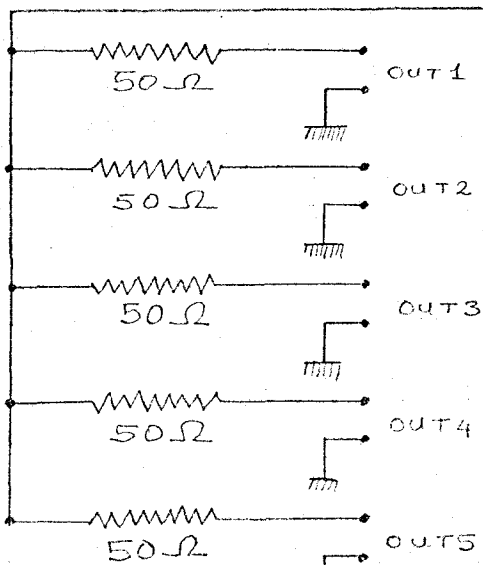
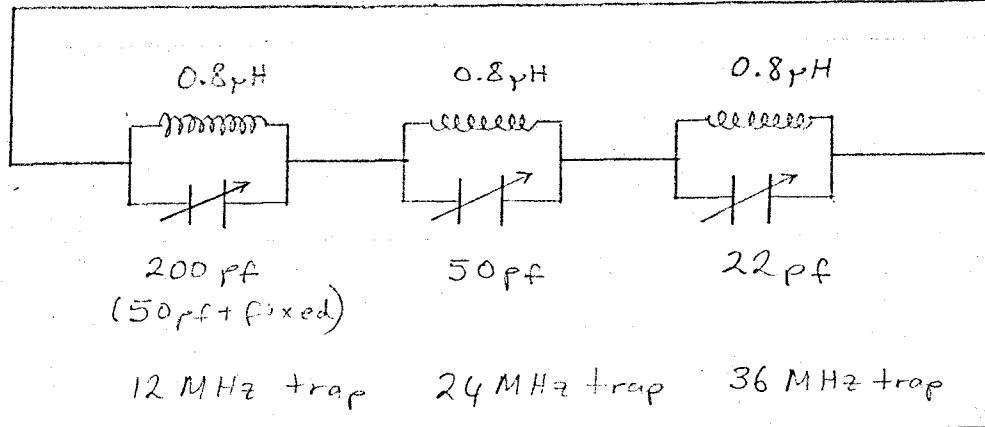
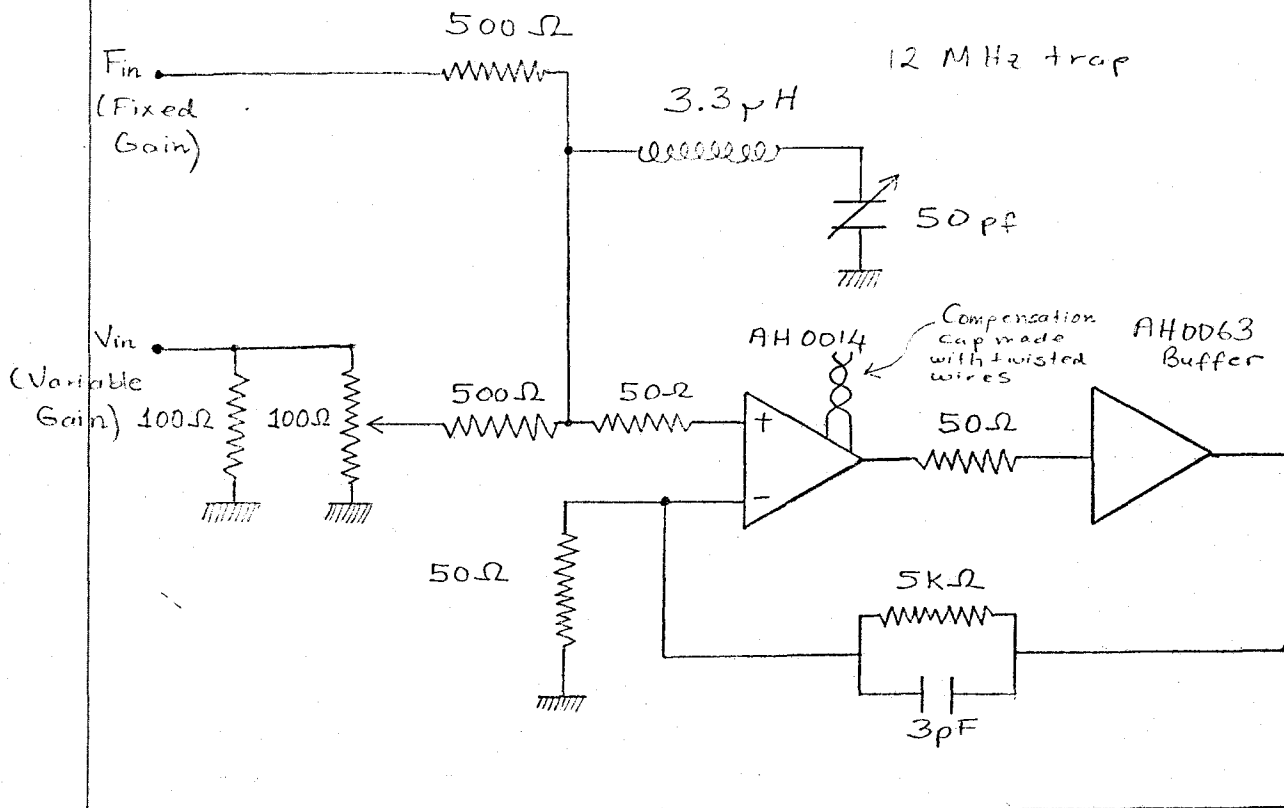
7/17/86
MEZ



"Fast 1" Amplifier



"Fast 1" Amplifier

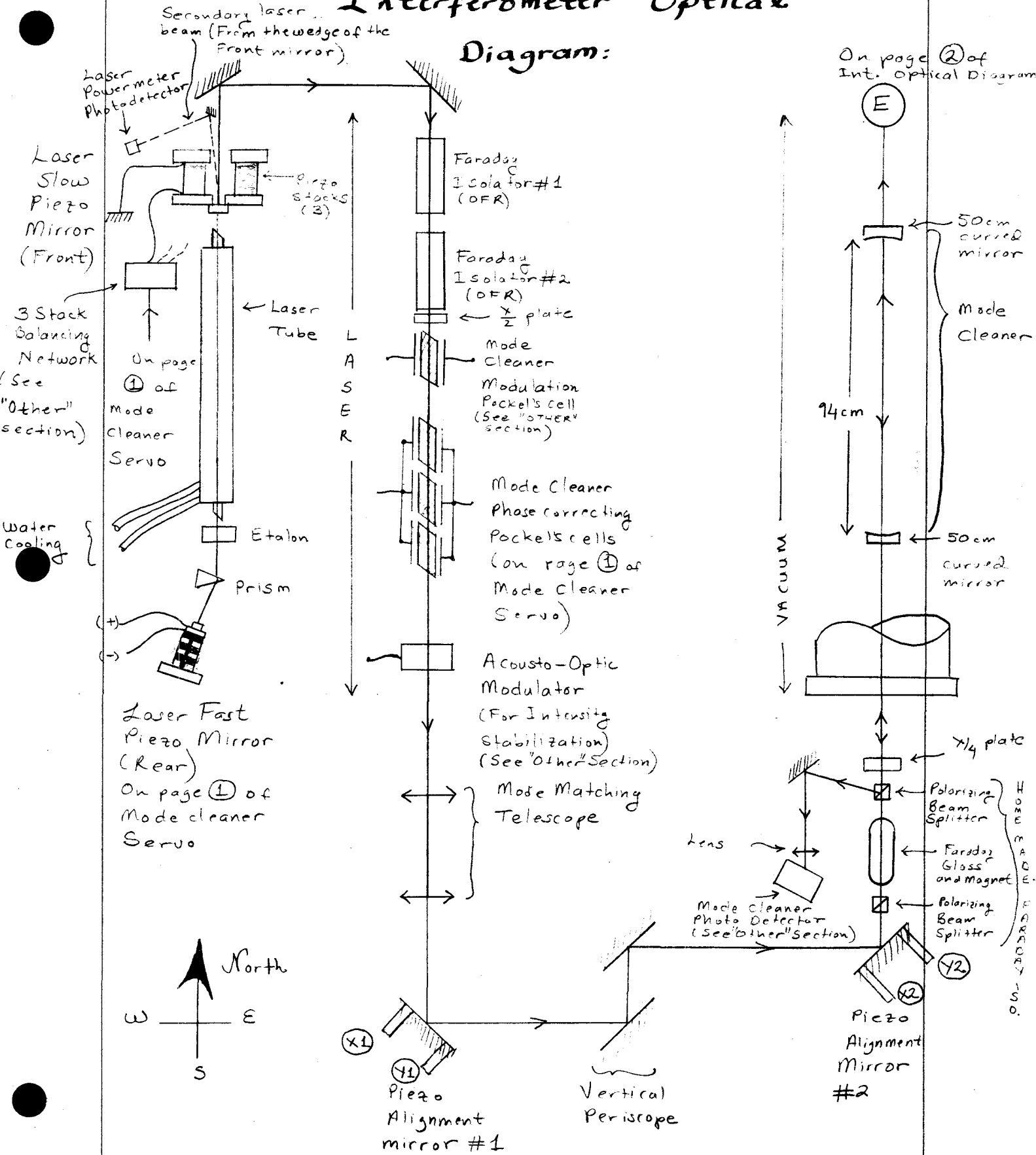


BATCH
START

STAPLE
OR
DIVIDER

Interferometer Optical

Diagram:



On page ② of Int. optical Diagram.

