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**New Folder Name** SHOT NOISE Comparison

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# Comparison of Non-recombined and Recombined Interferometer Shot Noise Sensitivity

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## Non-recombined Fabry-Perot

The displacement sensitivity for the system in its present, non-recombined configuration is<sup>1</sup>

$$\tilde{x}_{\text{NR}}(f) = \frac{L}{\pi\tau_e} \sqrt{\frac{3}{8}} \sqrt{\mathcal{M}} \left[ \frac{\lambda h}{cP} (1 + [f/f_k]^2) \right]^{1/2} \quad (1)$$

with the modulation function defined by

$$\mathcal{M} = \frac{1}{3} \left[ \frac{M^{-1} + A^2 J_0^2 - 2A J_0^2 + 2A J_0 J_2}{M A^2 J_0^2 J_1^2} \right]$$

Here  $\lambda$  is the optical wavelength,  $P$  is the total bright-fringe power,  $f$  = gravity wave signal frequency,  $M$  is the mode matching fraction (in the absence of modulation and mirror coupling mismatches),  $A$  represents the field inside the cavity and is given by  $A = 2/(1 + \mathcal{L}/T_1)$ ,  $L$  is the length of one arm,  $\tau_t = L/c$ ,  $\tau_e$  is the cavity energy storage time,  $f_k \equiv 1/(4\pi\tau_e)$ , with  $\tau_e = 2\tau_t [T_1 + \mathcal{L}]^{-1}$ , and the Bessel functions  $J_0, J_1, J_2$  are evaluated at the modulation index.

## Recombined Sensitivity

The sensitivity for recombined but not recycled interferometers in Fabry-Perot and delay-line configurations is<sup>2</sup>

$$S_{h_{\text{FR}}}(f) = \frac{2\hbar c\lambda}{P} \left( \frac{1}{2BL} \right)^2 \left[ 1 + (2\pi BLf/c)^2 \right] \quad (2)$$

$$S_{h_{\text{DL}}}(f) = \frac{2\hbar c\lambda}{P} \left( \frac{1}{2BL} \right)^2 \left[ \frac{2\pi BLf/c}{\sin(2\pi BLf/c)} \right]^2 \quad (3)$$

where  $\lambda = (\text{optical wavelength})/2\pi$ ,  $\tilde{x}(f) = L\sqrt{S_h(f)}$ ,  $2\pi BL/c \equiv 1/f_k$ . For the Fabry-Perot,  $1/2BL = 1/4c\tau_e$ , and for the delay-line,  $B$  is the number of round trips in each arm.

<sup>1</sup>Shot Noise in the Caltech Gravitational Wave Detector—The mid-1984 Configuration, Stanley E. Whitcomb. There is a copy on file with annotations by R.S., draft dated 23 January 1990. The dependence on modulation index and mode matching is in agreement with a paper to be published by Niebauer et. al., Phys. Rev. A

<sup>2</sup>From the Article by Kip S. Thorne in *300 Years of Gravitation*, eds. S.W. Hawking and W. Israel, Cambridge University Press, p. 330 (1987) page 424, Equation 115.