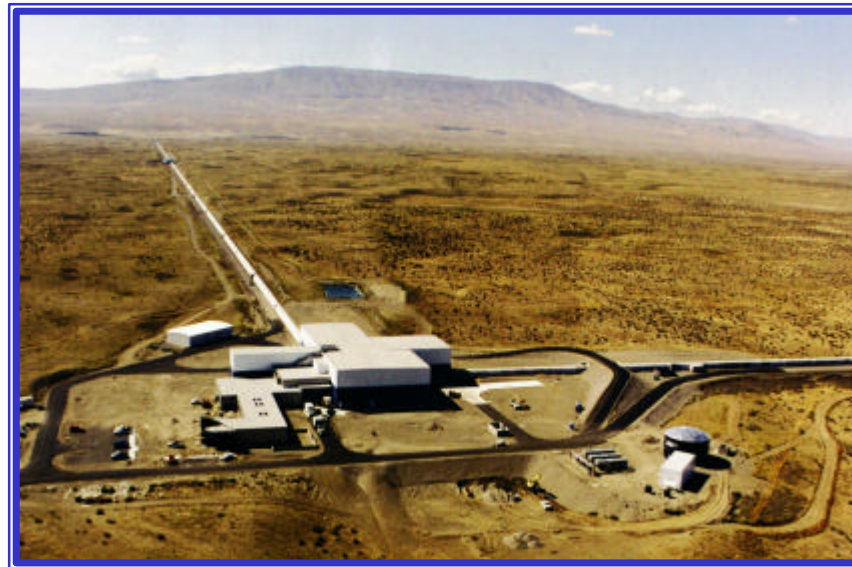


Commissioning at Hanford

Stan Whitcomb



Program Advisory Committee

12 December 2000

LIGO Livingston Observatory



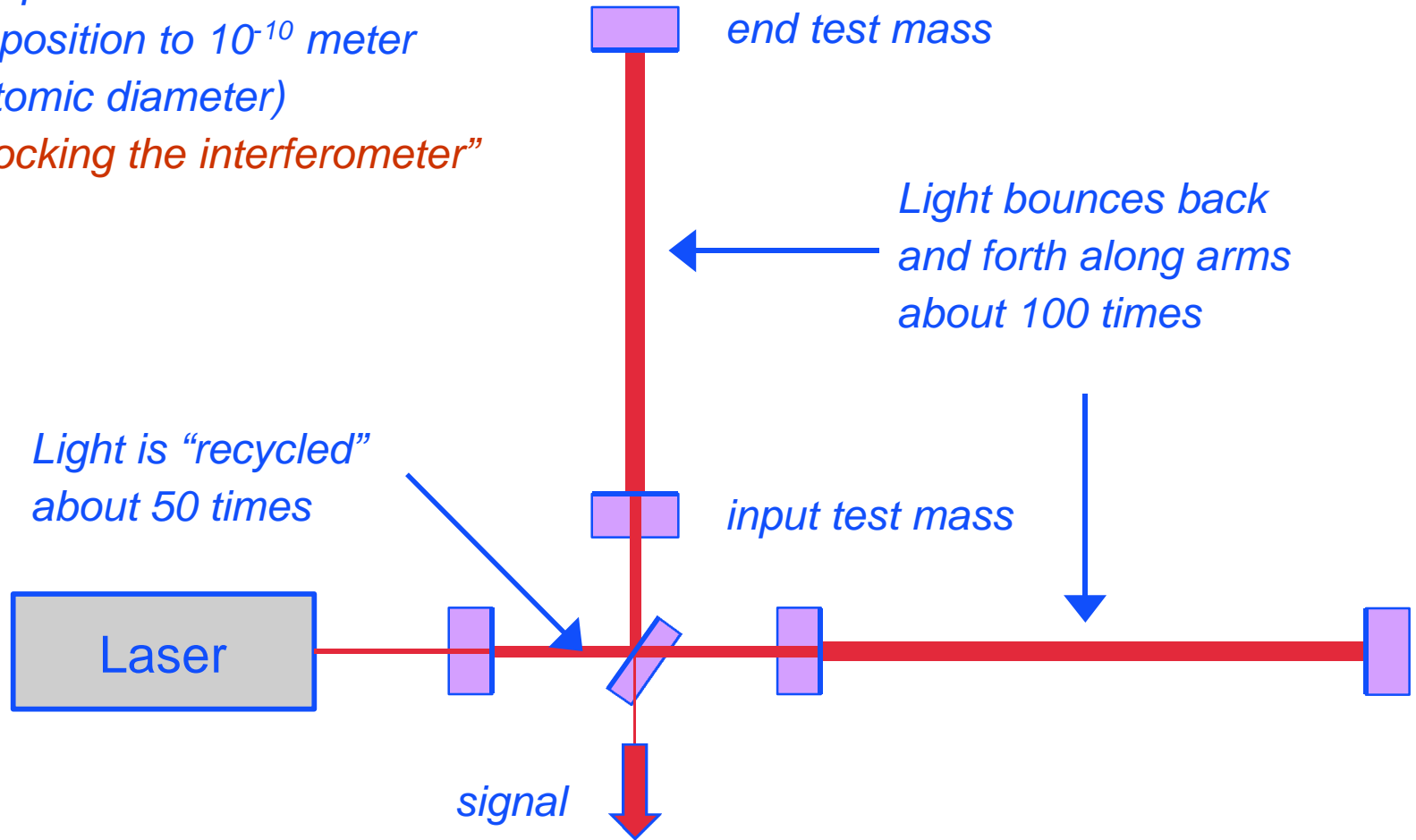
Installation Summary

- ✦ All in-vacuum installation complete for LHO 2km
- ✦ ASC and LSC electronics installed for 2 km interferometer, and under test
- ✦ Delayed installation of 4km LHO interferometer pending solution to scattered light problem
 - » Seismic isolation complete
 - » PSL installation underway
 - » In-vacuum optics scheduled for early 2001
- ✦ DAQ/Control Network infrastructure complete
 - » Generally reliable, but still a few minor problem areas
 - » Still verifying correct signal hook-ups

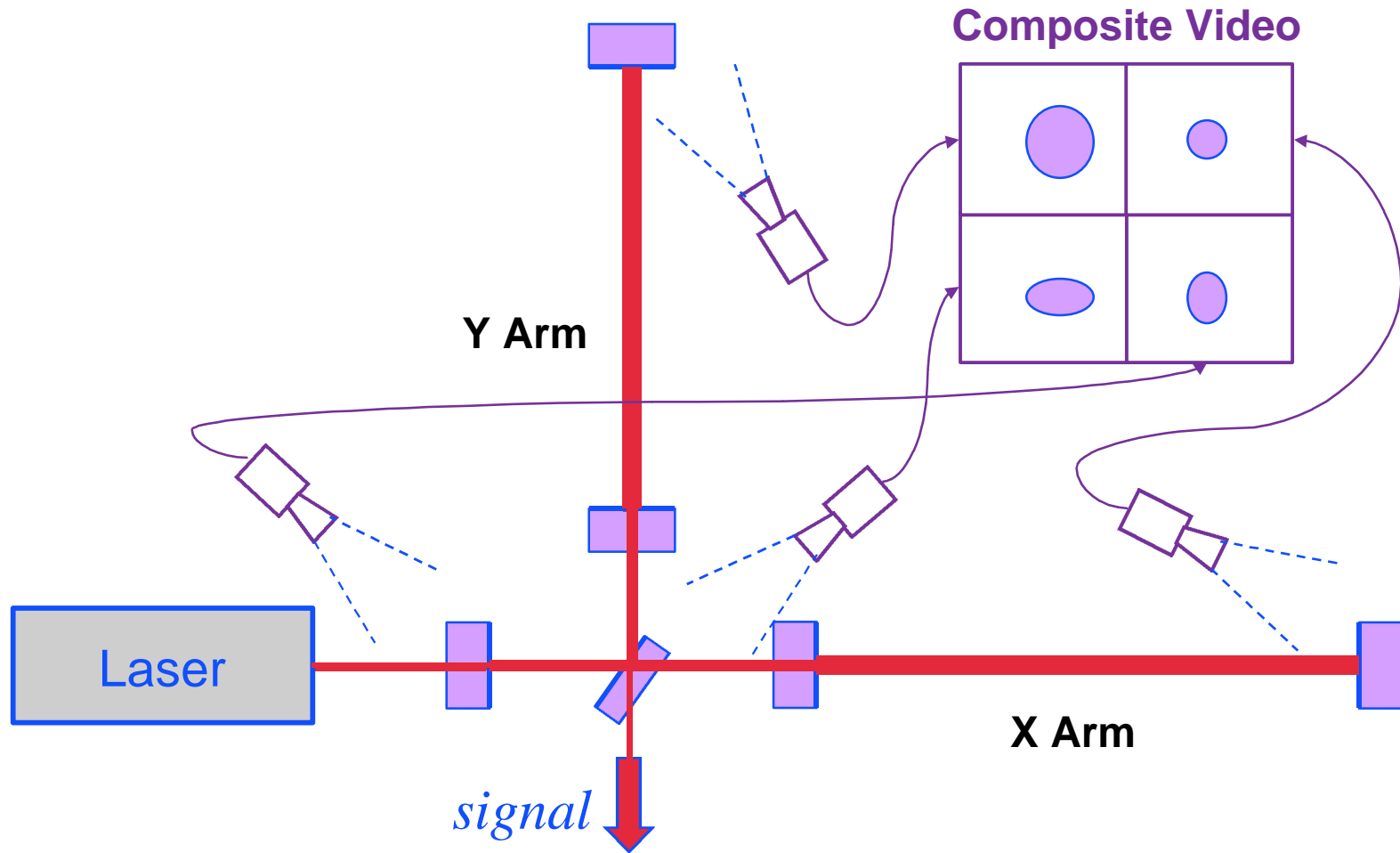


First Lock for a LIGO Interferometer

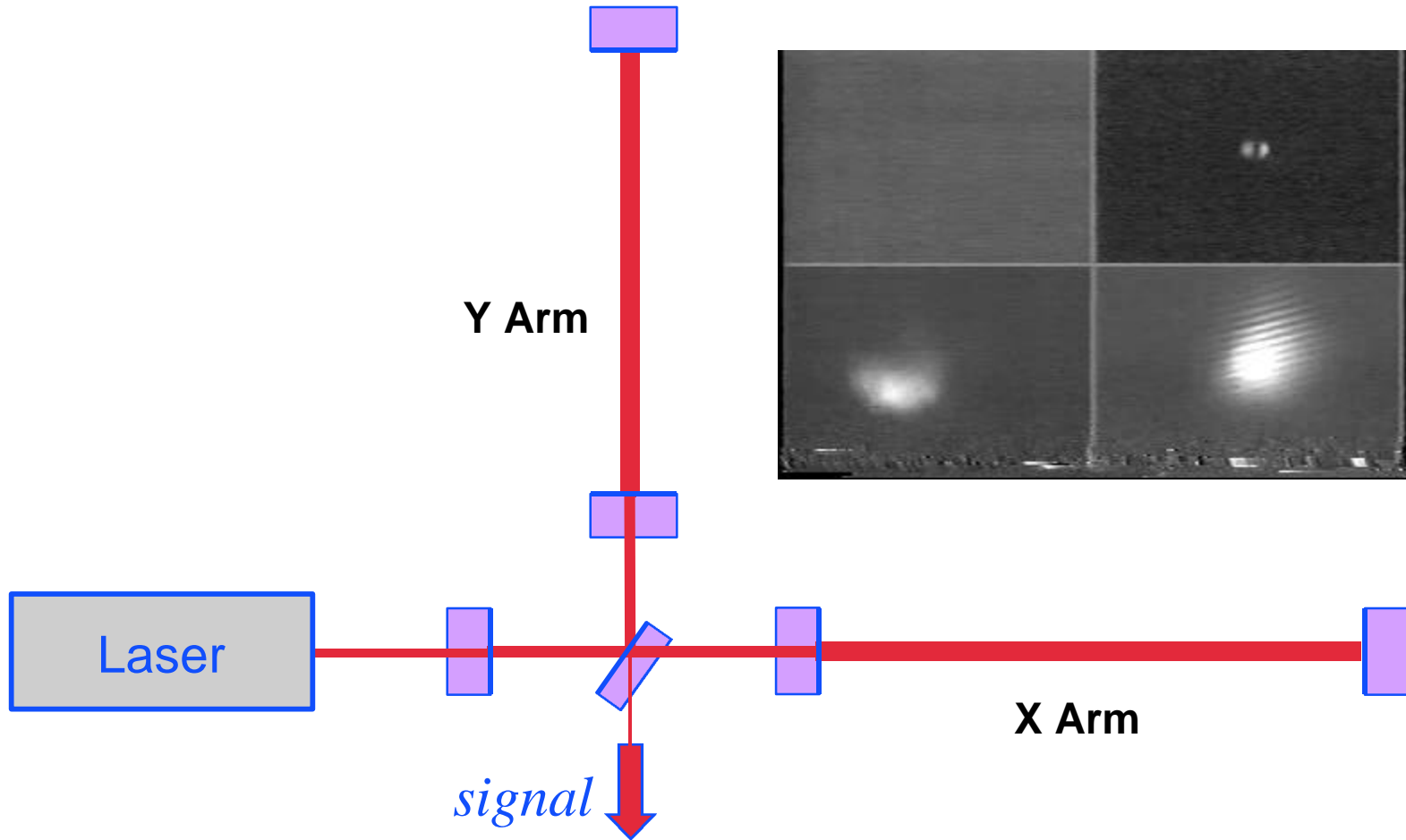
Requires test masses to be held in position to 10^{-10} meter (atomic diameter)
"Locking the interferometer"



Steps to Locking the Interferometer



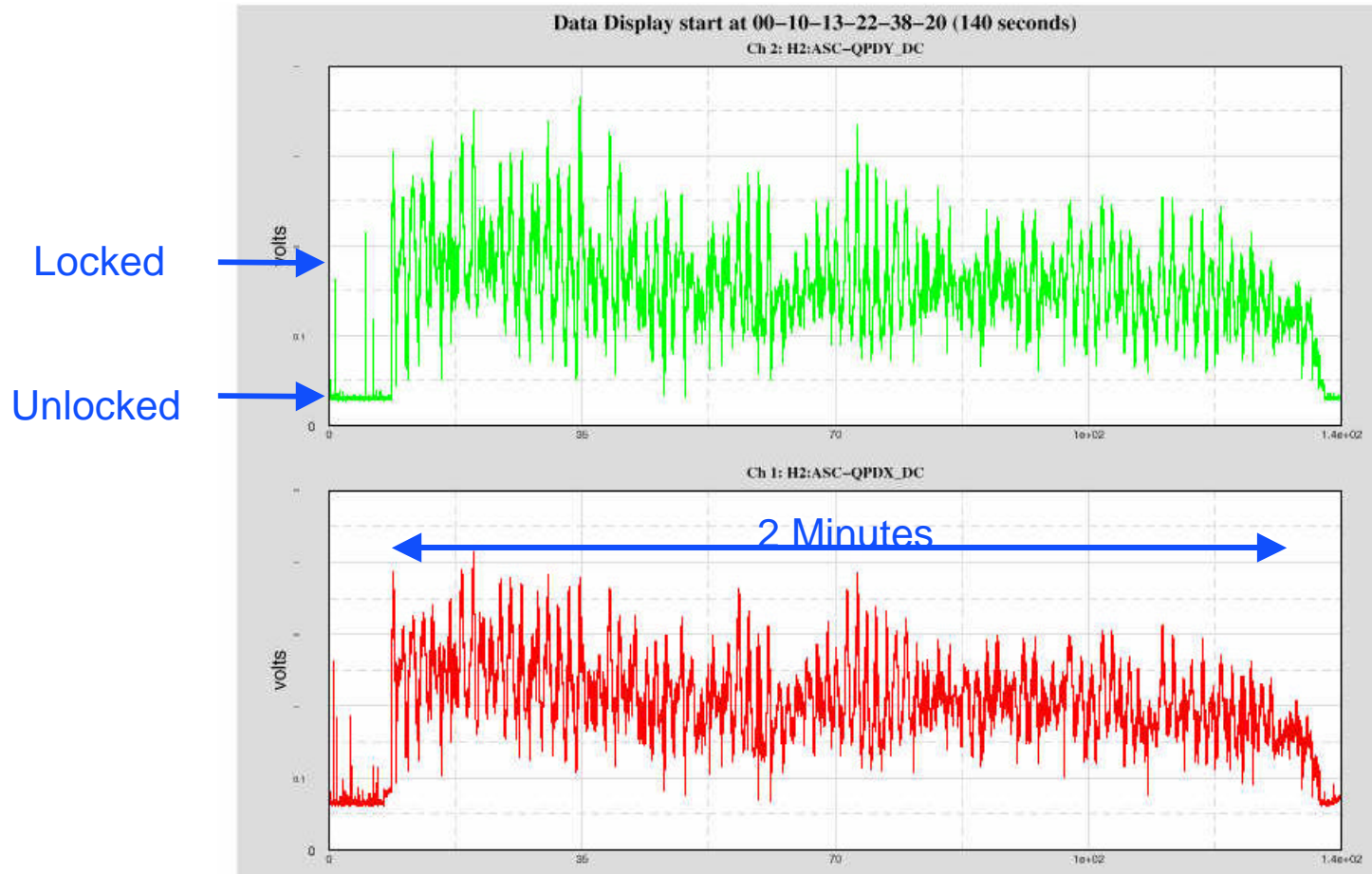
Watching the Interferometer Lock





Full Interferometer Locking

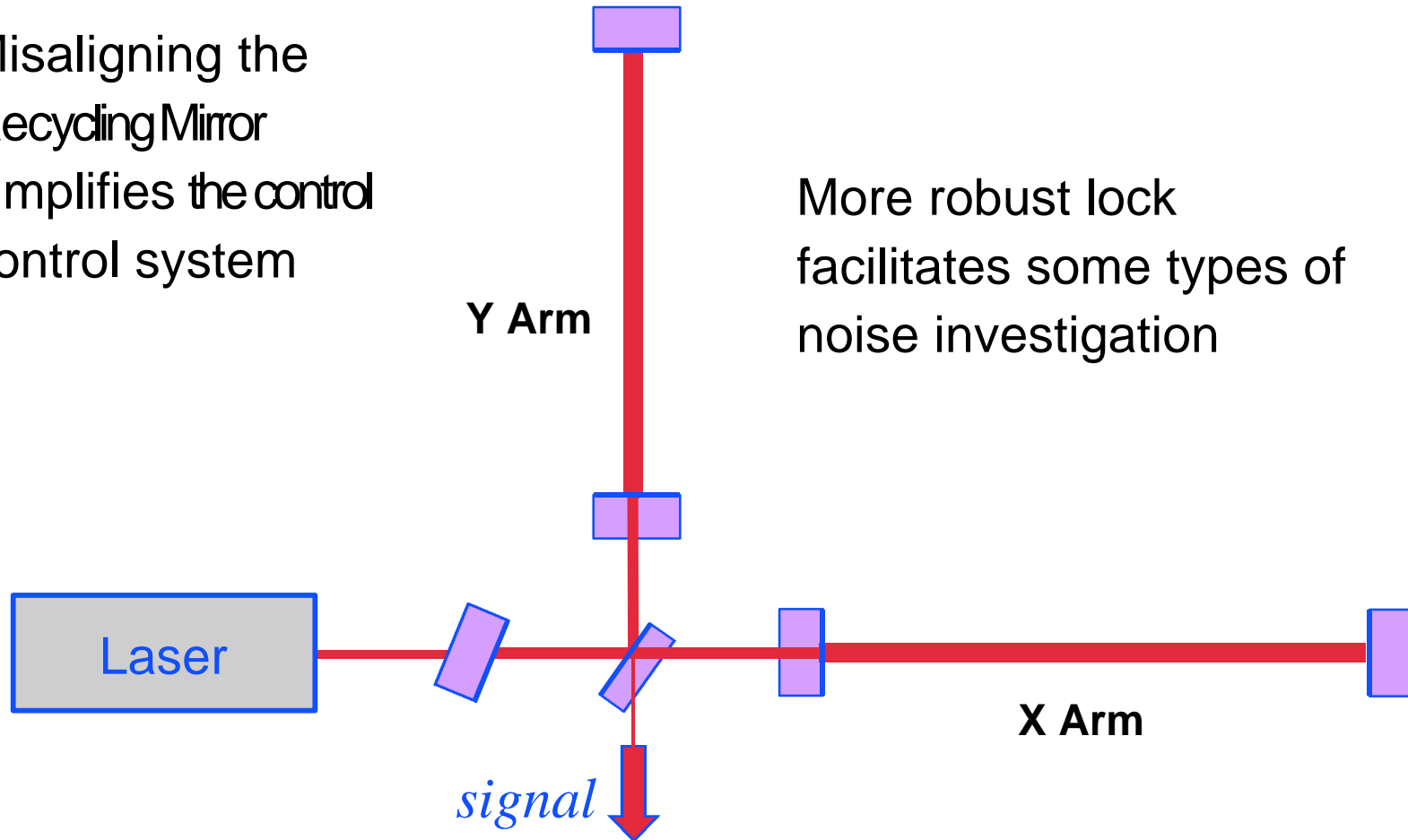
✦ Still a bit tenuous.....



Recombined Michelson with F-P Arms

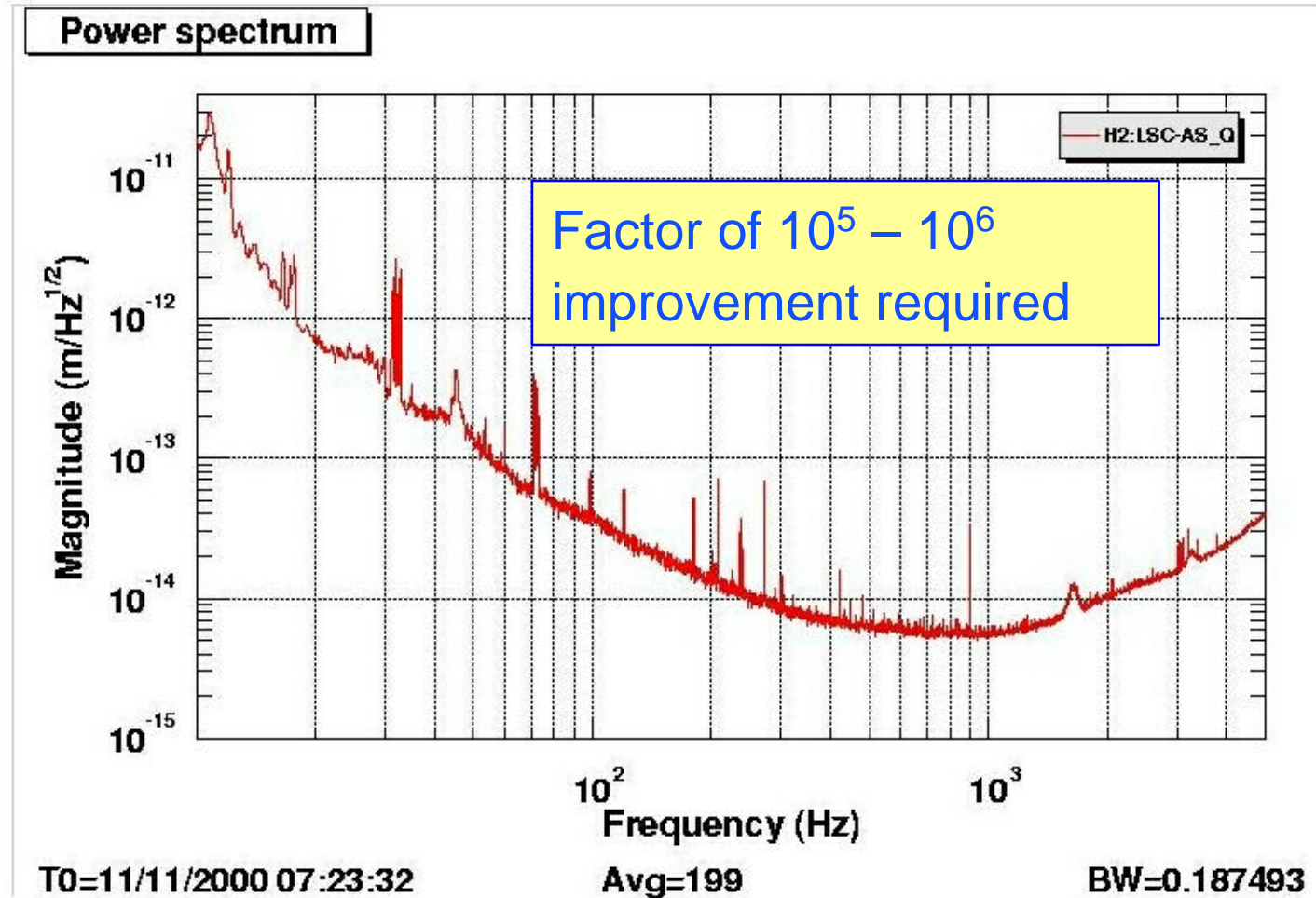
Misaligning the Recycling Mirror simplifies the control control system

More robust lock facilitates some types of noise investigation





Recombined Interferometer Spectrum





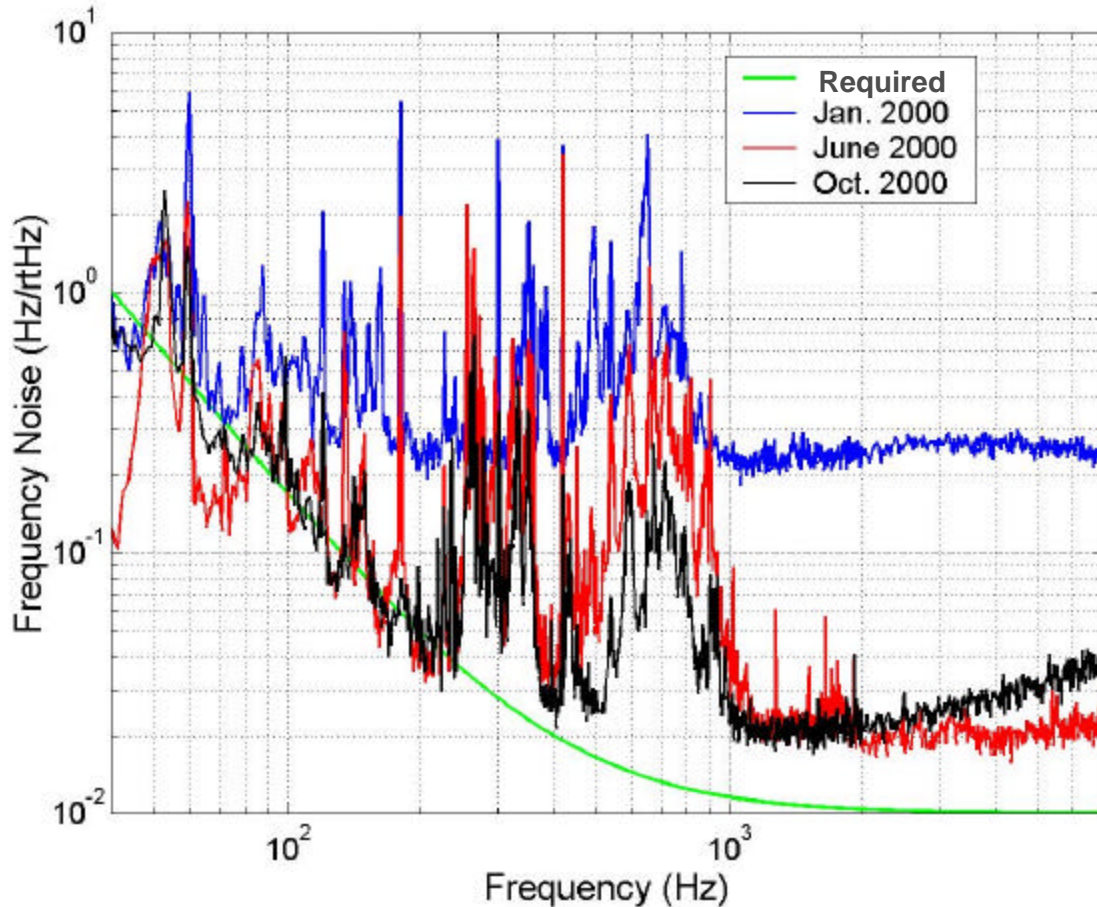
Beginning to Work on Noise

- ✦ Identified electronics noise (ADC) as dominant noise at high frequencies (due to low input laser power)
- ✦ Laser frequency noise dominates in mid frequency band (stabilization from arm common mode not yet implemented)
- ✦ Low frequencies seismic noise?
- ✦ Many resonant features to investigate and eliminate



Cont. to Work on Subsystem Performance

- ✦ Example: Prestabilized Laser
- ✦ Laser stability an important contributor to LIGO sensitivity
- ✦ Steady improvement in laser noise performance
 - » electronics
 - » acoustics
 - » vibrations
 - » elimination of excitation sources

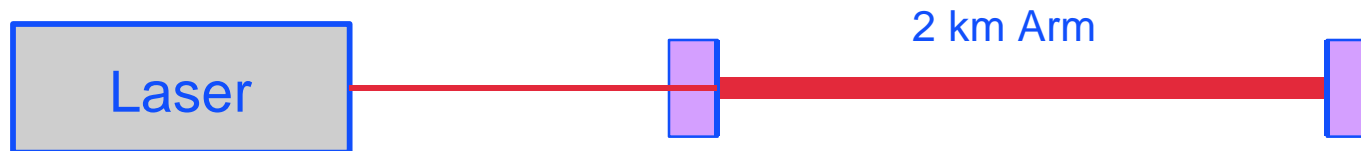




Progress Toward Robust Operation

- ✦ Different measure of interferometer performance (in contrast with sensitivity)
 - » Interferometer lock duration goal is 40 hours,
- ✦ Prestabilized laser
 - » Two years continuous operation with ~20% loss in power
 - » Locks to reference cavity and premodecleaner for months
- ✦ Mode Cleaner
 - » Locks for weeks at a time, reacquires lock in few seconds
- ✦ Data Acquisition and Control
 - » Data Acquisition and IOCs (Input Output Controllers) routinely operate for weeks-to-months without problems
 - » Tools in place for tracking machine state: AutoBURT, Conlog

Extending the Lock on a Single Arm



✦ Start with Y Arm

- » 12/1/99 Flashes of light
- » 12/9/99 0.2 seconds lock
- » 1/14/00 2 seconds lock
- » 1/19/00 60 seconds lock
- » 1/21/00 5 minutes lock

✦ Change to X Arm

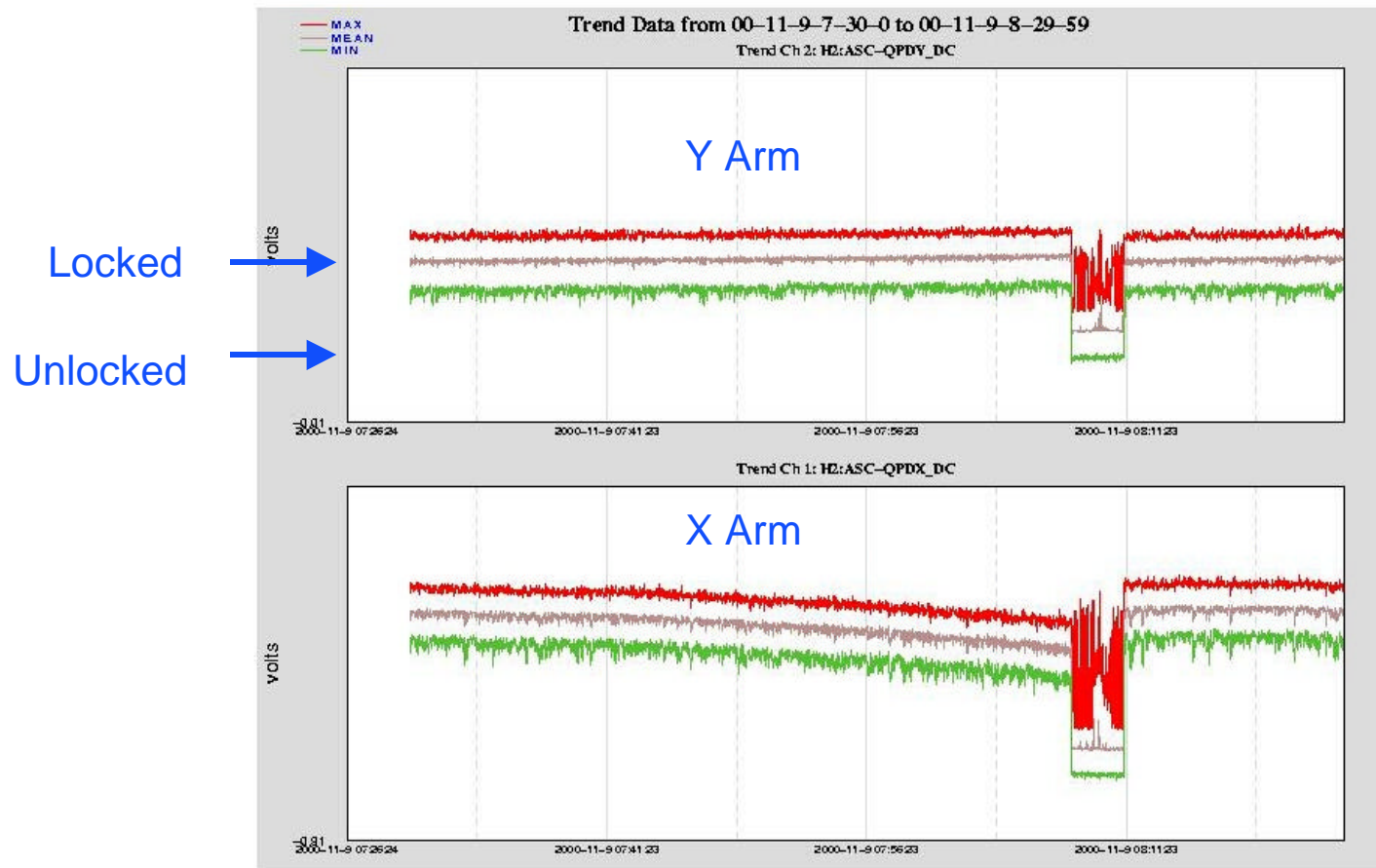
- » 2/12/00 18 minutes lock
- » 3/4/00 90 minutes lock
- » 3/26/00 10 hours lock

Result of :
 -automatic alignment system
 -tuning electronics
 -reduction of noise sources



Recombined Michelson Robustness

Randomly chosen hour from recent engineering run





Engineering Runs

- ✦ Engineering Runs are a key part of our commissioning plan
 - » Test interferometer stability, reliability
 - » Well defined dataset for off-site analysis
 - » Develop procedures for later operations
 - » Means to include the broader LSC in detector commissioning

- ✦ First Engineering Run (E1) in April 2000
 - » Single arm operation with wavefront sensing alignment
 - » 24 hour duration
 - » Lots of interest and good intentions, but rather limited follow-through on planned analysis

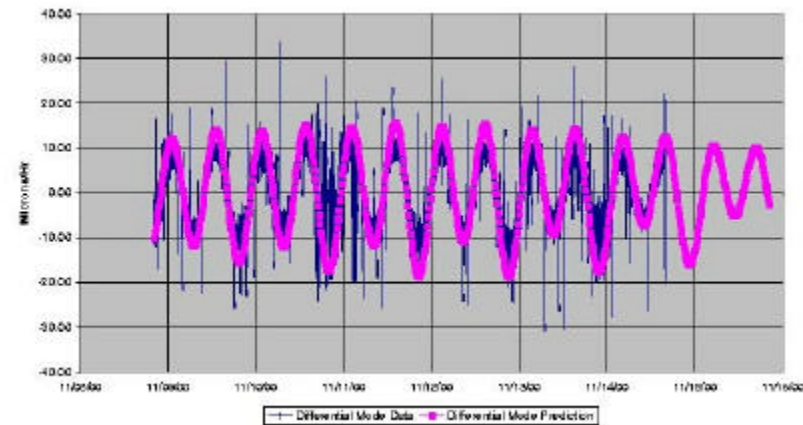
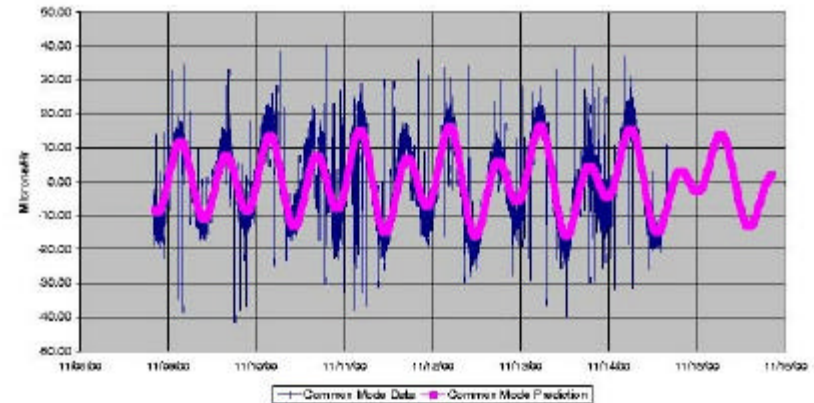


Second Engineering Run (E2)

- ✦ **November 2000**
 - » One week of around the clock operation
 - » Approximately 35 scientists participated on site
- ✦ **Recombined Michelson with Fabry-Perot arms**
 - » Misaligned recycling mirror to make for more robust locking
 - » Typical locked stretches 30 – 90 minutes (longest ~ 3 hours)
 - » >90% duty cycle
- ✦ **Organized around 14 detector investigations**
 - » Earthtides, frequency noise, calibration, noise stationarity, seismic noise, noise bursts, line tracking, ...
 - » More analysis during the run than for entire E1
- ✦ **Major test of DAQ system**
 - » Successfully transferred 2 terabytes of data to CACR archive

Earth tide Investigation

- ✦ Observed in earlier E1 Run, but predictions had unexplained time shift
- ✦ ~200 microns P-to-P
- ✦ Main cause of loss of lock for long arms in E2 run
- ✦ Input to design of tidal actuator needed for eventual long lock durations
- ✦ Common mode (both arms stretch together) and differential mode (arms stretch by different amounts)



- ✦ “Locking the Interferometer” marks a major transition for LIGO

- » Commissioning a full interferometer, not individual subsystems

“First Lock”
in the Hanford
Observatory
control room

- ✦ Steady Progress

- ✦ Locking duration
- ✦ Noise performance
- ✦ Operational experience

- ✦ Still a long way to go!

