

LISA: Status and Recent Work

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JILA - University of Colorado

Aspen Winter Conference on
Gravitational Waves and their Detection

Aspen, CO

8 February 2001

LIGO-G010083-00-M

Outline

- Overview
- Programmatic advances
- R&D
 - European
 - U.S.
- Demonstration flights
- Conclusions

Introduction

- The LISA concept
 - 3 spacecraft forming two interferometers with 5×10^6 km long arms, benign environment
 - Measure between free-falling test masses with drag-free spacecraft control to limit unwanted disturbances
 - Measurement band of 10^{-4} to 1 Hz, $10^{-23} / \sqrt{\text{Hz}}$ best strain sensitivity in a 1 yr integration
- Sources
 - Galactic binaries: WDs, NSs, stellar mass BHs ($1-10^2$)
 - Coalescence of intermediate mass black holes (10^2-10^4)
 - Coalescence of supermassive black holes (10^5-10^7)
 - Stellar mass black holes into SMBHs
 - Cosmological background(s)

LISA Researchers

The LISA Mission Definition Team (US)

E. S. Phinney	California Institute of Technology (Chair)
B. Allen	University of Wisconsin
J. W. Armstrong	Jet Propulsion Laboratory
P. L. Bender	University of Colorado
E. A. Boldt	NASA Goddard Space Flight Center
O. Blaes	University of California Santa Barbara
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R. L. Byer	Stanford University
T. E. Chupp	University of Michigan
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Supported by

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R. E. Spero	Jet Propulsion Laboratory
G. Walushka	NASA Goddard Space Flight Center

The LISA Science Study Team (Europe)

T. Edwards	Rutherford Appleton Laboratory (Study Manager)
R. Reinhard	ESTEC (Study Scientist, Acting)
A. Brillet	University de Paris Sud
I. Ciufolini	University of Rome
A. M. Cruise	University of Birmingham
C. Cutler	Albert Einstein Institute
K. Danzmann	University of Hannover
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J. Hough	Glasgow University
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T. Sumner	Imperial College
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S. Vitale	University of Trento
H. Ward	Glasgow University
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Supported by

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J. Gonzalez	ESTEC
B. Kent	Rutherford Appleton Laboratory
S. Marcuccio	Centrosazio
S. Peskett	Rutherford Appleton Laboratory
P. C. E. Roberts	Cranfield University
R. Turner	Rutherford Appleton Laboratory

Programmatic Advances in Past Year

- NASA OSS Strategic Plan
 - Recommended new start in SEU, start around 2006
 - Launch 2010, data 2012-2013
- NRC Decadal Survey
 - Highest priority new start in intermediate class
- ESA selection
 - Joint cornerstone
 - Smart 2 demo flight
- Project startup
 - Draft U.S. management plan involving GSFC and JPL
 - Key positions being filled
 - ESA/NASA discussions
 - Science Team formation

Looming Programmatic Events

- ESA/NASA agreements
- Joint management plan
- Science team startup
- Cosmic Journeys

Technology Development

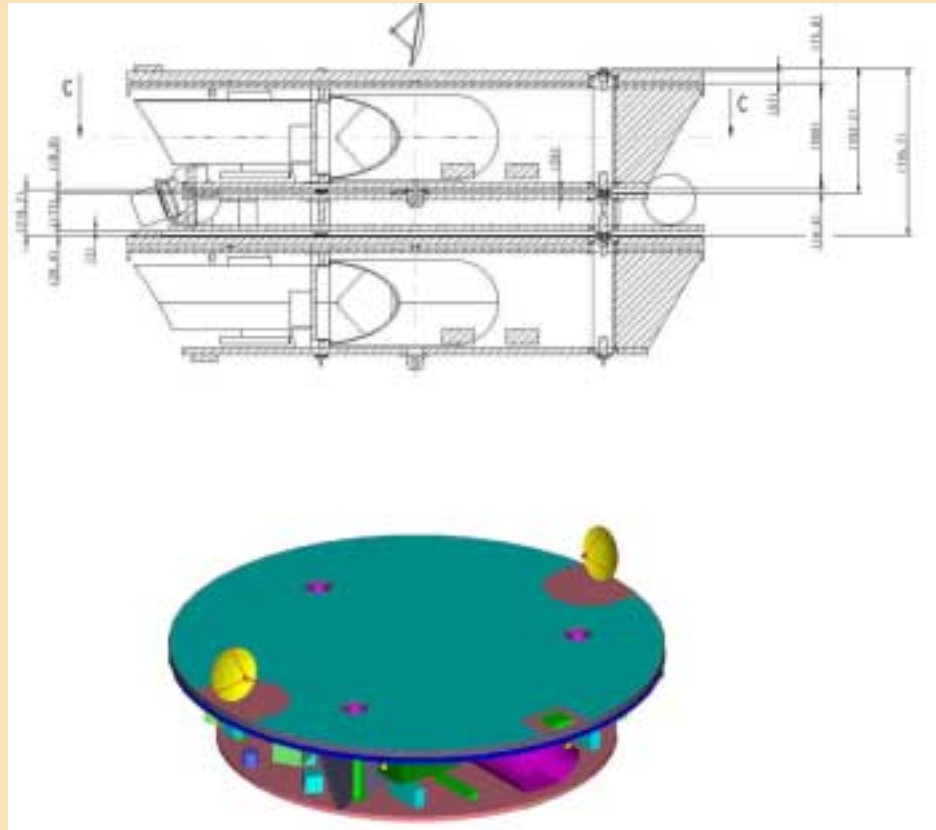
- Disturbance reduction system
 - Gravitational sensor: $3 \times 10^{-15} \text{ m/s}^2 / \sqrt{\text{Hz}}$, $10^{-4} - 10^{-2} \text{ Hz}$
 - μN thrusters: 50 μN max, 20 μN constant, 0.1 $\mu\text{N} / \sqrt{\text{Hz}}$, 10 yr lifetime
 - Drag-free control: 10 nm/ $\sqrt{\text{Hz}}$, 3 nrad/ $\sqrt{\text{Hz}}$
- Interferometry
 - Space-qualified components
 - Fringe timing: $2 \times 10^{-5} \text{ cycles} / \sqrt{\text{Hz}}$, $10^{-3} - 1 \text{ Hz}$
 - Stabilized laser: frequency stability of 30 Hz/ $\sqrt{\text{Hz}}$ at 1 mHz, power stability of $2 \times 10^{-4} / \sqrt{\text{Hz}}$

European R&D Activities

- To be described in more detail by David Robertson, Thursday morning
- Phase A Industrial Study, and satellite studies
- Interferometry at Glasgow
- Laser stabilization at Hannover
- Gravitational sensor design, prototyping and testing at Trento
- Control system analysis at Trento
- In-FEEP noise measurements at Siebersdorf
- In-FEEP thrust measurements at ONERA
- Photodiode investigation at Birmingham
- FEEP testing at ESTEC

ESA Phase A Study

- Reviewed system design and trades
- Refined spacecraft and payload designs
 - Structural analysis
 - Thermal analysis
 - Gravitational analysis
 - Accommodation study
- Reviewed operational modes
- Modeled various drag-free control strategies
- Mass increase



U.S. R&D Activities

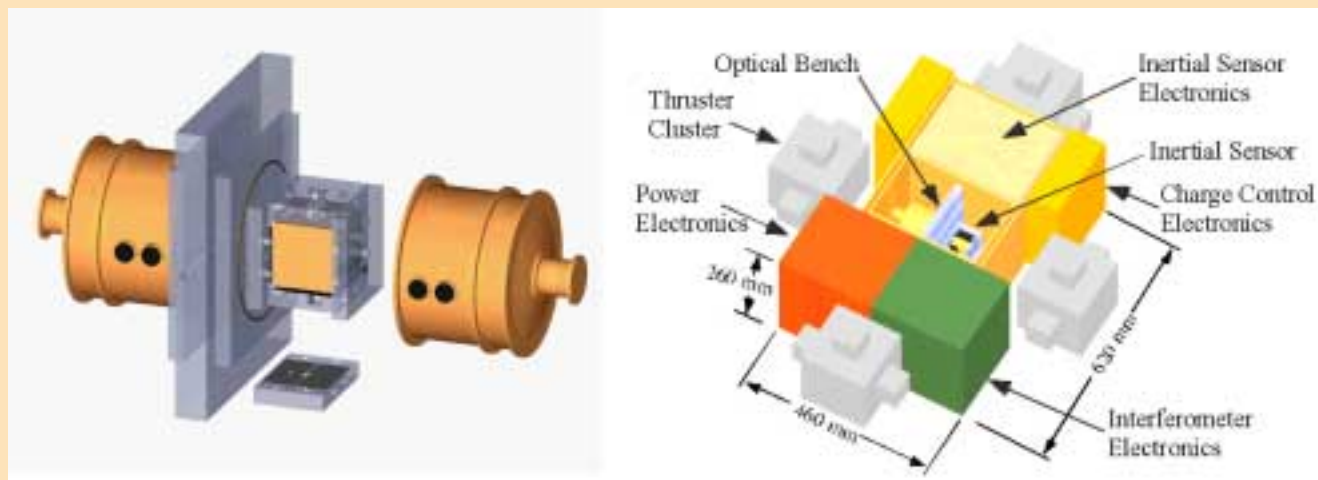
- Fringe timing at JILA (Scott Pollack, Friday afternoon)
- Control system and error analysis at JILA
- Noise correction at JPL (M. Tinto, Tuesday morning)
- Noise correction at Washington/JILA (P. Bender, Tuesday morning)
- Micronewton thruster testing at GSFC (S. Merkowitz, Friday afternoon)
- Colloidal thrusters at JPL
- Gravitational sensor design at Stanford

Demonstration Flights

- Gravitational sensor can't be tested on ground
- Two gravitational sensors, thrusters and servo control
- An interferometer to measure between sensors
- Drift-away orbit
- Test of residual acceleration within order of magnitude required by LISA and precision stationkeeping
- Lock spacecraft to one test mass, compare residual acceleration between two test masses
- Monitor stationkeeping
- Test noise models

ESA Smart 2

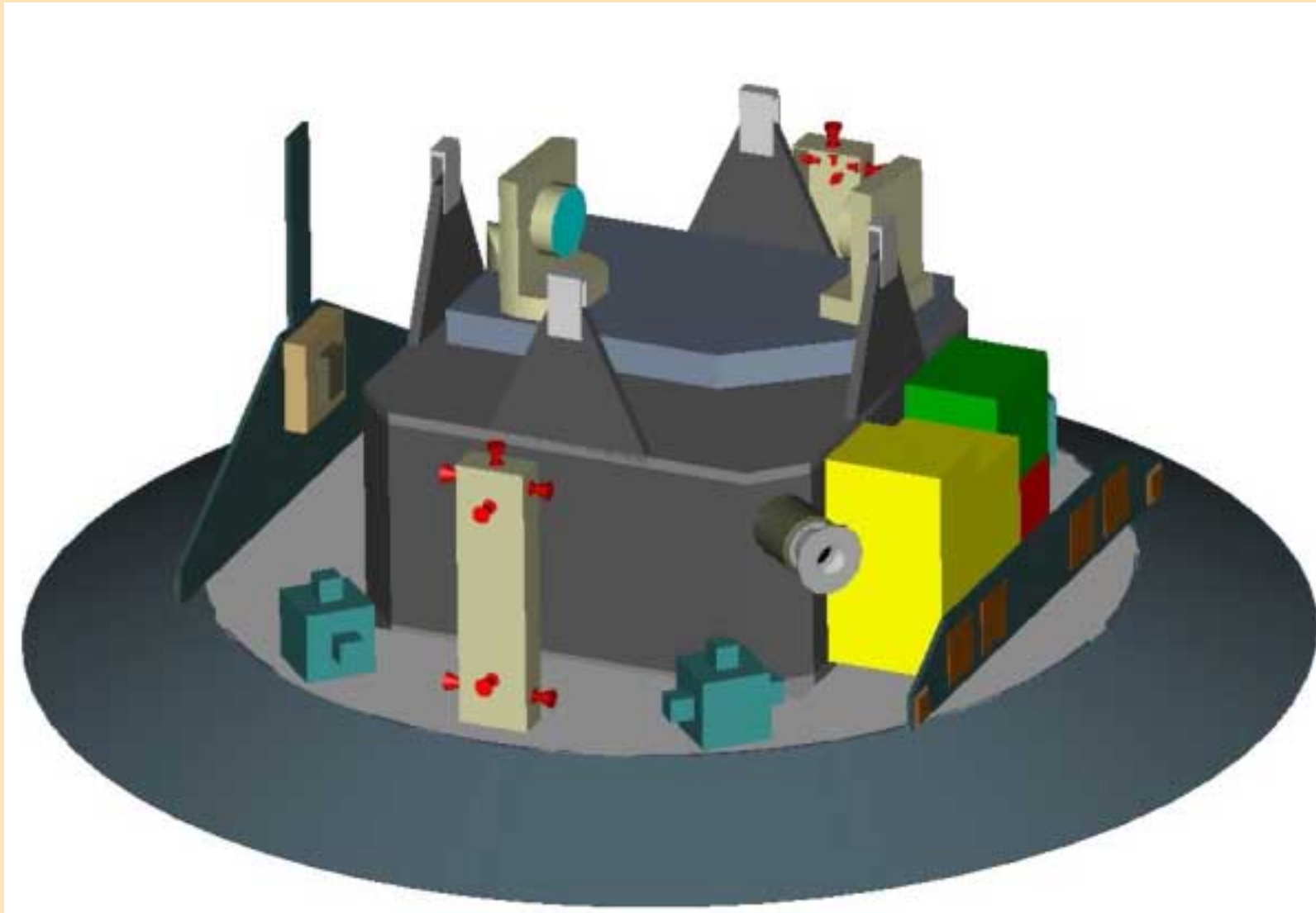
- LISA Test Package
- IRSI-Darwin test of formation-flying
- Two satellites
- Selected mission
- Phase A ITT in April '01
- Scheduled for launch in late '06
- ~120 Meuro allocated



U.S. Space Technology 7

- The New Millennium Program tests technology that benefits multiple missions. US technology only.
- ST7 is a competition between 4 concepts testing a technology suite
- LV not included, piggyback on either Smart 2 or Starlight
- Starlight (nee DS-3, ST-3) is a TPF precursor testing formation-flying and star light interferometry, now in Phase B, two satellites, mid '06 launch.
- Phase A Study ends Nov. '01, technology selection in Dec.
- \$52 M cost cap
- Launch in '05.

LISA Test Package on Starlight Spacecraft



Summary

- The LISA concept has made great strides toward becoming an established project. Still requires approval by Congress.
- Modest R&D efforts are progressing in both Europe and U.S. Awaiting significant funding for technology development.
- There are promising opportunities for a LISA technology demonstration flight.