Status of TAMA300



National Astronomical Observatory, The University of Tokyo, The Institute for Cosmic Ray Research, Institute for Laser Science, KEK-High Energy Accelerator Research Organization, Osaka City University, Osaka University, Yukawa Institute of Theoretical Physics



Specification





Facilities





Past data taking (DT)

	Period	Obs. Time	Main Target
DT1	1999 8/6 ~ 8/7	11h	Establishment of calibration
DT2	1999 9/17 ~ 9/20	31h	First event search
DT3	2000 4/20 ~ 4/23	13h	Improved sensitivity
DT4	2000 8/21 ~ 9/4	167h	100-h data
DT5	2001 3/2 ~ 3/10	111h	24-h full-time observation
DT6	2001 8/1 ~ 9/20	1038h	1000-h data
DT7	2002 8/31 ~ 9/2	25h	Recycling
DT8	2003 2/14 ~ 4/14	1158h	International coincidence run



Lock system for recycled FPMI





Frequency stabilization

Extended BW of MC servo300kHz600kHz (~2002/10)Extended BW of δL + servo20kHz40kHz



Alignment control

Test mass mirrors: Wave front sensing at 15MHz (BW 60Hz) Recycling mirror: Mechanical modulation at 60Hz & 70Hz (BW 1Hz)





Alignment control for recycling mirror



Sensitivity in DT8

Reduced non-stationary fluctuation owing to wide band filters for the alignment control

Improved strain sensitivity: $h = 2.7 \times 10^{-21} / \text{Hz}^{1/2} @2 \text{kHz}$



Orientation control

Local control on the injection bench (QPD1, 2) to keep the incident beam to MC Grobal control using the transmitted lights (QPD3, 4) to keep the beam centering of arms QPD4 M 2 QPD3 **B**S -Q PD Laser

Automation

- Self-switching sub-system (Laser & MC)
- MC frequency stabilization
- MC alignment control
- Laser intensity stabilization
- Orientation control
- Injection-lock servo of laser
- Digital switching using PC & LabVIEW
- Lock acquisition
- Manual mirror alignment
- IFO status monitoring



MC Lock



MC Alignment



Data Taking 8

Target Full-scale coincidence run with LIGO More than 1000-h data (>70%) Period Feb. 14 14:00 ~ Apr. 14 22:00 (1424h) Shift 94 people (2 people×3 shift /day) •Observation time: 1158h

Duty cycle: 81.3%

•Best sensitivity: 8×10⁻¹⁹m/rHz@1.5kHz

Longest lock time : 20.5h



Observable distance for GW from inspiral NS binaries (SNR=10)



Time table in DT8





Disturbance of the observation

- •Construction: Difficult to keep the lock
- Wave due to depression: Difficult to acquire the lock
- •Laser: Sometimes to unstable mode
- •Misc.: DAQ trouble, Electricity shutdown... Duty rate with Seismic motion





Construction around the site







Data analysis

Talk on Wednesday!

Inspiraling compact binaries H. Takahashi Black hole ring-downs Y. Tsunesada Burst waves M. Ando



Future plan

Improvement of sensitivity Investigation of the noise from Michelson part at <1kHz Power recycling Alignment control of RM using WFS High gain (G=10) recycling Seismic attenuation system (SAS) For low frequency (0.1 \sim 10Hz) R&D with Caltech and Univ. of Pisa Installation in early 2005 Observation Shared run with experiments **Crewless operation** Online real-time analysis



R&D in Univ. of Tokyo





Summary

- Recycled FPMI with alignment control for RM
- Improved frequency stabilization: BW=600kHz (MC loop), 40kHz (δL+ loop)
- Improved sensitivity: $h = 2.7 \times 10^{-21} / \text{Hz}^{1/2} @2 \text{kHz}$
- Items for observation: auto lock, orientation control
- •DT8: observation time=1158h, Duty cycle=81.3%
- •SAS: Installation in early 2005

