

### Waveform Consistency Test in Burst Detection

Laura Cadonati LSC meeting, Hannover August 20, 2003

LIGO-G030438-00-Z



### Test goal

The LIGO Burst Search pipeline uses *Event Trigger Generators* (ETGs) to flag times when "something anomalous" occurs in the strain time series

 $\Rightarrow$  burst candidate events ( $\Delta t$ ,  $\Delta f$ , SNR)

Events from the three LIGO interferometers are brought together in coincidence (time, frequency, power).

In order to use the full power of a coincident analysis:

- » Are the waveforms consistent? To what confidence?
- » Can we suppress the false rate in order to lower thresholds and dig deeper into the noise?

Cross correlation of coincident events



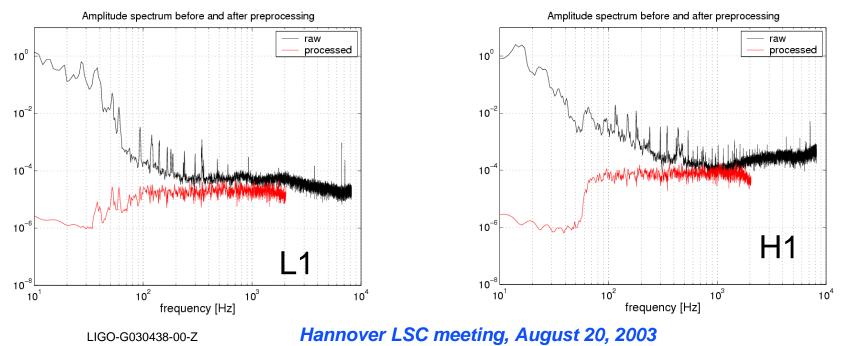
# Data Conditioning

Decimate and high-pass few seconds of data around event  $\Rightarrow$ 100-2048 Hz

Remove predictable content (effective whitening/line removal): train a linear predictor error filter over 1 s of data (1 s before event start),

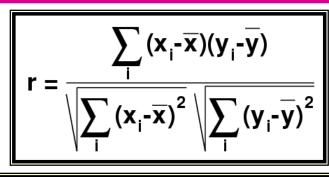
apply as second order sections model using zero-phase filtering - described in S. Chatterji's talk on data conditioning (LIGO-G030439)

 $\Rightarrow$ emphasis on transients, avoid non-stationary, correlated lines.





#### r-statistic



Linear correlation coefficient or normalized cross correlation for the two series  $\{x_i\}$  and  $\{y_i\}$ 

<u>NULL HYPOTHESIS</u>: the two (finite) series  $\{x_i\}$  and  $\{y_i\}$  are <u>uncorrelated</u>

⇒ Their linear correlation coefficient (**Pearson's r**) is normally distributed around zero, with  $\sigma = 1/\text{sqrt}(N)$  where N is the number of points in the series (N >> 1)

#### S = erfc (|r| sqrt(N/2))

double-sided significance of the null hypothesis i.e.: probability that |r| is larger than what measured, if {x<sub>i</sub>} and {y<sub>i</sub>} are uncorrelated

#### $C = -\log_{10}(S)$

confidence that the null hypothesis is FALSE  $\Rightarrow$  that the two series are correlated

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### **Delay and Integration Time**

#### What delay?

Shift {y<sub>i</sub>} vs {x<sub>i</sub>} and calculate:  $r_k$ ;  $S_k$ ;  $C_k$ ...then look for the maximum confidence  $C_M$ Time shift for  $C_M$  = delay between IFOs Shift limits: ±10 ms (LLO-LHO light travel time)

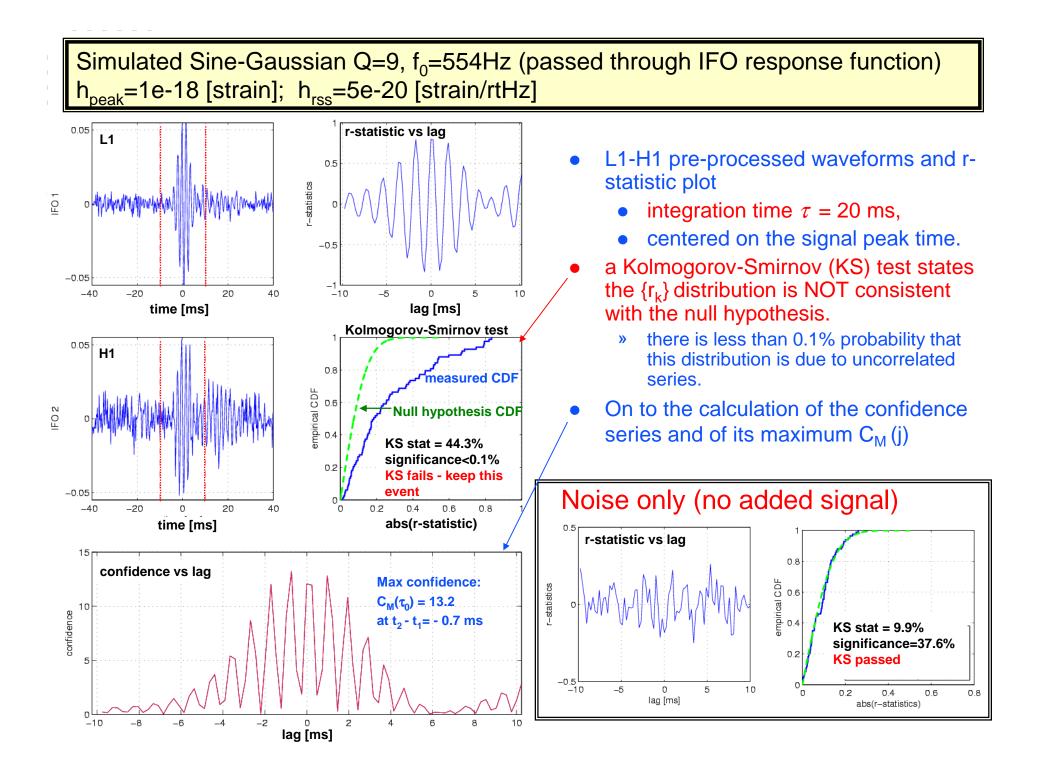
$$\mathbf{r}_{k} = \frac{\sum_{i} (\mathbf{x}_{i} - \overline{\mathbf{x}}) (\mathbf{y}_{i+k} - \overline{\mathbf{y}})}{\sqrt{\sum_{i} (\mathbf{x}_{i} - \overline{\mathbf{x}})^{2}} \sqrt{\sum_{i} (\mathbf{y}_{i+k} - \overline{\mathbf{y}})^{2}}}$$

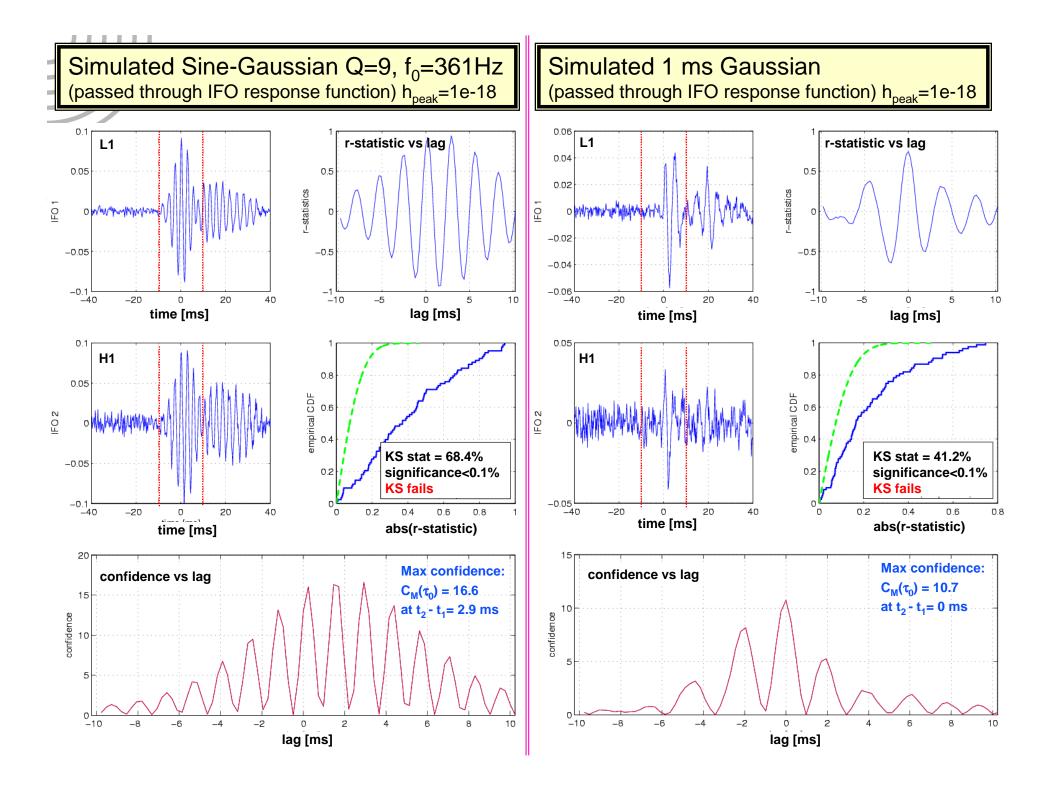
#### Integration time τ:

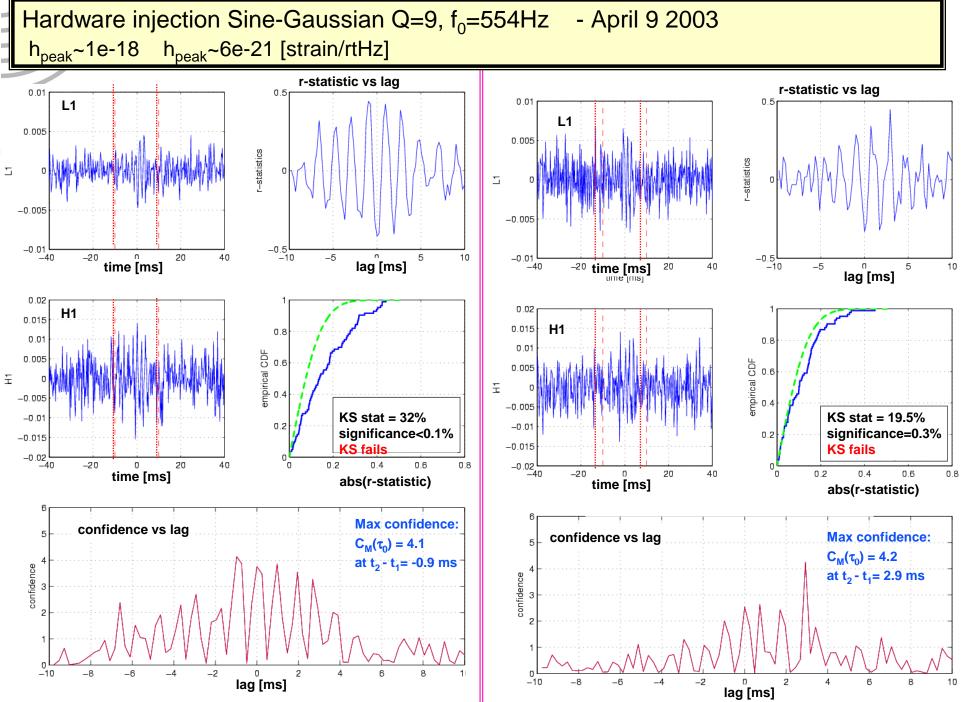
How long?

- » If too small, we lose waveform information and the test becomes less reliable
- » If too large, we wash out the waveform in the cross-correlation

Test different  $\tau$ 's and do an OR of the results (20ms, 50ms, 100ms)



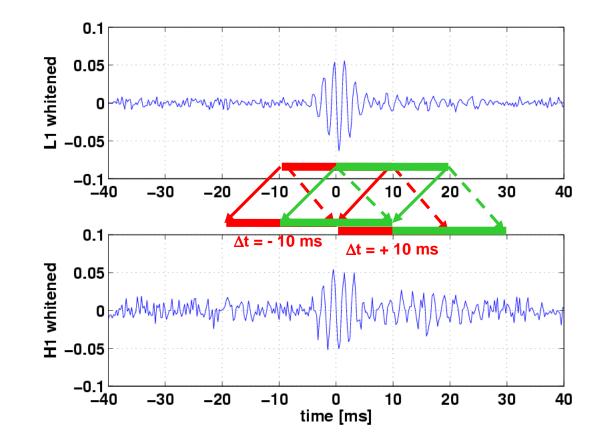


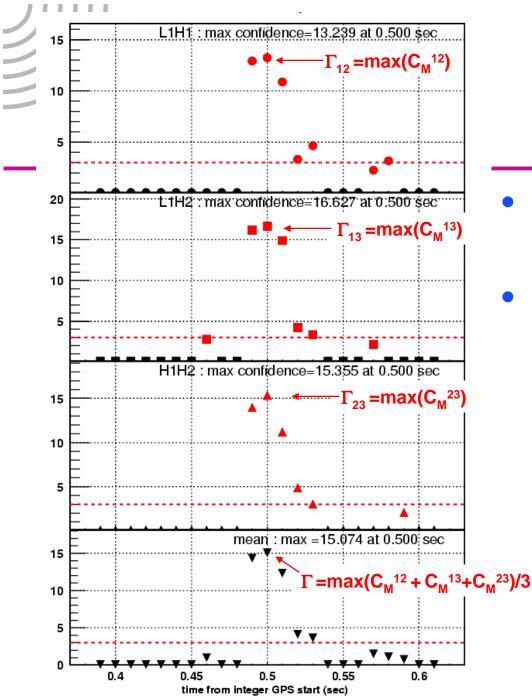




### Scanning the Trigger Duration $\Delta T$

- » Partition trigger in  $N_{sub}=(2\Delta T/\tau)+1$  subsets and calculate  $C_M(j)$  $(j=1.. N_{sub})$
- » Use  $\Gamma_{ab} = \max_{j}(C_{M}(j))$  as the correlation confidence for a pair of detectors over the whole event duration





# C<sub>M</sub>(j) plots

- Each point: max confidence C<sub>M</sub>(j) for an interval τ wide (here: τ = 20ms)
- Define a cut (pattern recognition?): 2 IFOs: Γ=max<sub>j</sub>(C<sub>M</sub>(j)) > β<sub>2</sub> 3 IFOs: Γ=max<sub>i</sub>(C<sub>M</sub><sup>12</sup>+ C<sub>M</sub><sup>13</sup>+ C<sub>M</sub><sup>31</sup>)/3 > β<sub>3</sub>

In general, we can have  $\beta_2 \neq \beta_3$  $\beta=3$ : 99.9% correlation probability



#### **Open Issues**

#### Calibration

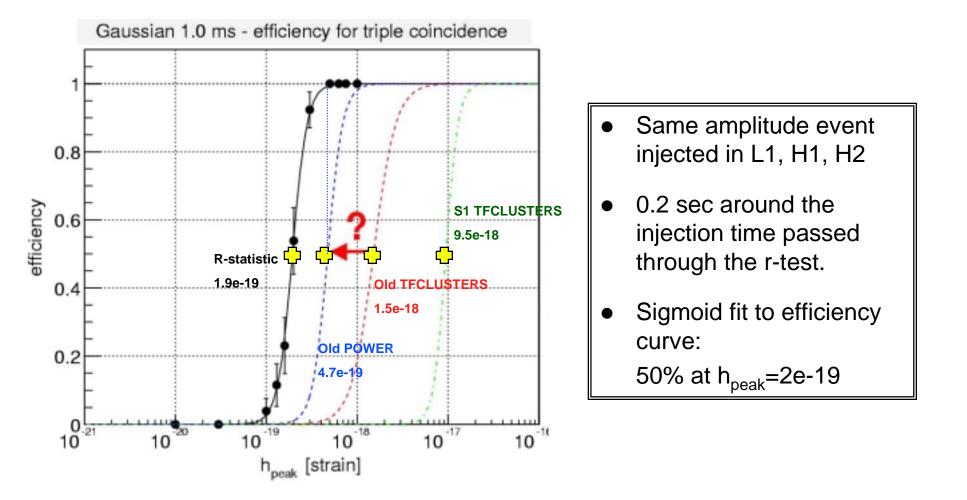
» Needed to account for waveform distortions due to frequency-dep calibration

#### • Time resolution

- » Depends on waveform
- » Affected by detector response function
  - Implement phase calibration to match IFOs?
- » Affected by pre-processing filters
  - Do a second test pass with less aggressive filters?
- » At the moment, no use is made of the delay time in assessing the confidence
  - Implement a T-statistic test?

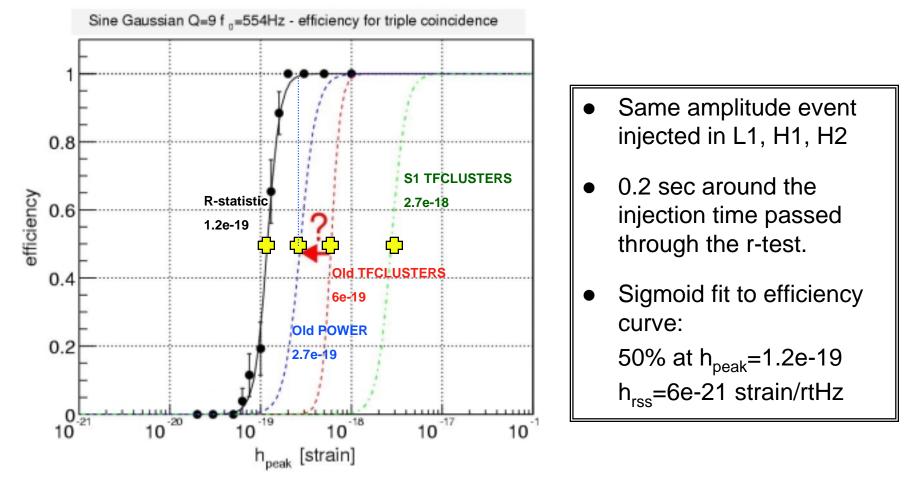


## Triple-coincidence efficiency for 1 ms gaussians





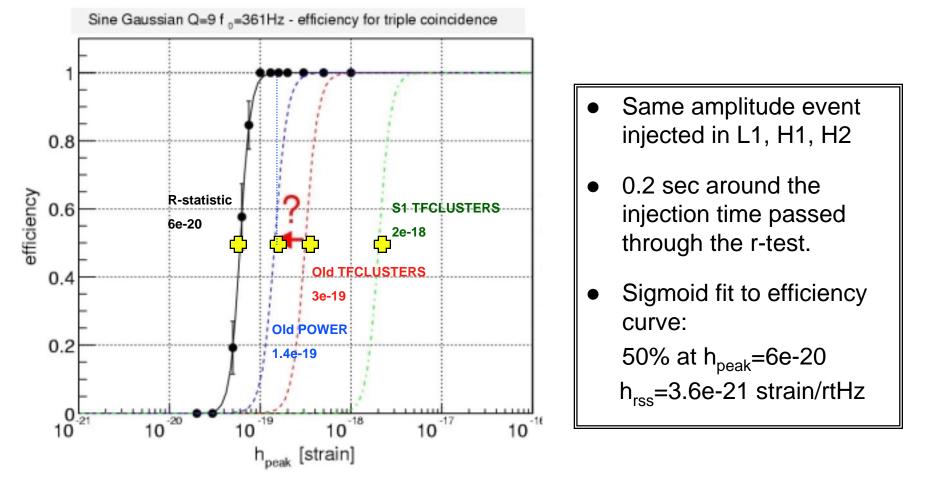
## Triple-coincidence efficiency for 554 Hz Q=9 sine gaussians



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## Triple-coincidence efficiency for 361 Hz Q=9 sine gaussians



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