

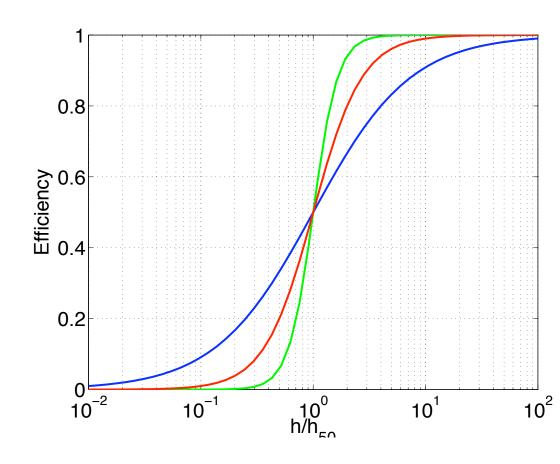
Tuning event trigger generators for burst searches

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What should guide event trigger tuning?

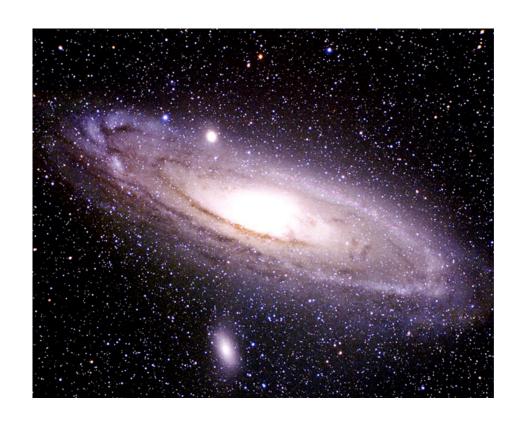
- Goal of event trigger generator: identify events
- Tuning generally involves several parameter choices
 - » Tuning affect false rate
 - » Tuning affects efficiency in multiple ways
 - h₅₀, "slope", etc.
- What should tuning optimize?





A Tuning Goal

- Maximize detected event rate at fixed false rate
- Fixed false rate?
 - » High h₅₀, shallow slope; low h₅₀, steep slope
- Detected event rate?
 - » "Physical" events: i.e., drawn from representative spatial distribution
 - » The greater the increase in sources with volume the more favored a shallow slope (reaching to low signal strength and greater distances)



LIGO

Quantitative Criterion

- θ = source population model characterization
 - » E.g., galaxy model, binary synthesis model
 - » $P(\theta)$ prior over pop models (could be δ function)
- ξ = source characterization in model θ
 - » E.g., supernova spin, mass, orientation, etc.
 - » $P(\xi|\theta)$ distribution or prior over source params
- ψ = tuning parameters
 - » E.g., TFCLUSTER black pixel prob, etc.
- $\epsilon(\xi\theta\psi)$ = fraction of events ξ in source model θ detected for tuning param ψ
- Monte Carlo compute figure of merit $\epsilon_0(\psi)$ proportional to number detected events

$$\epsilon_0(\psi) = \int_{\theta} d\theta \, P(\theta) \int_{\xi} d\xi \, P(\xi|\theta) \epsilon(\xi\theta\psi)$$



Summary

- Shape of detection efficiency curve depends on tuning parameters. What goal should guide choice of parameters?
- Proposal: Maximize the detected rate of "physical" events at fixed false alarm rate
 - » Physical: expected source distribution (galactic, isotropic, etc.), source parameters
- Will investigate this tuning on BlockNormal burst search