

## *LIGO Overview*



*for the LIGO Scientific Collaboration*

35th COSPAR meeting

Paris

21 July 2004

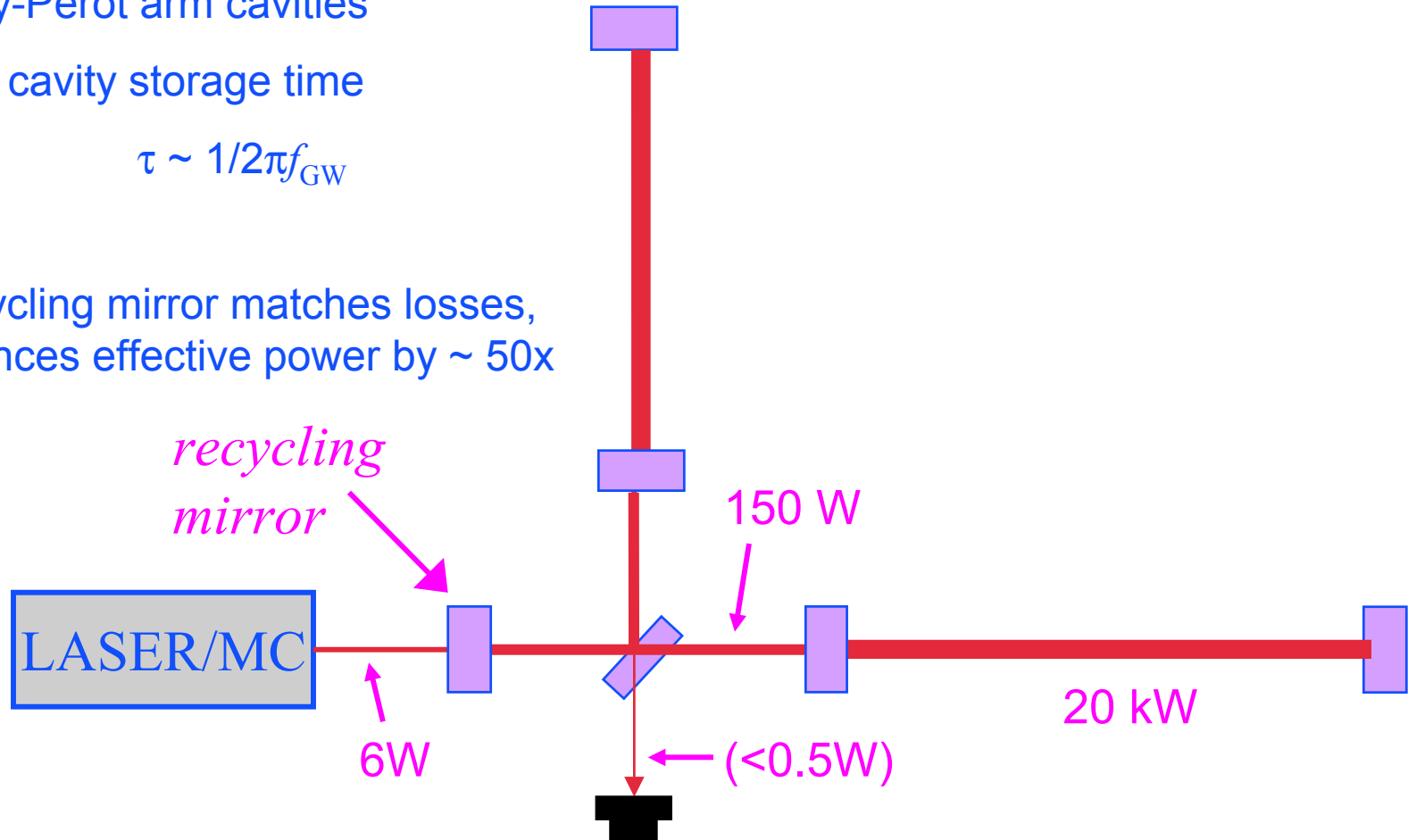


- Michelson interferometer with Fabry-Perot arm cavities

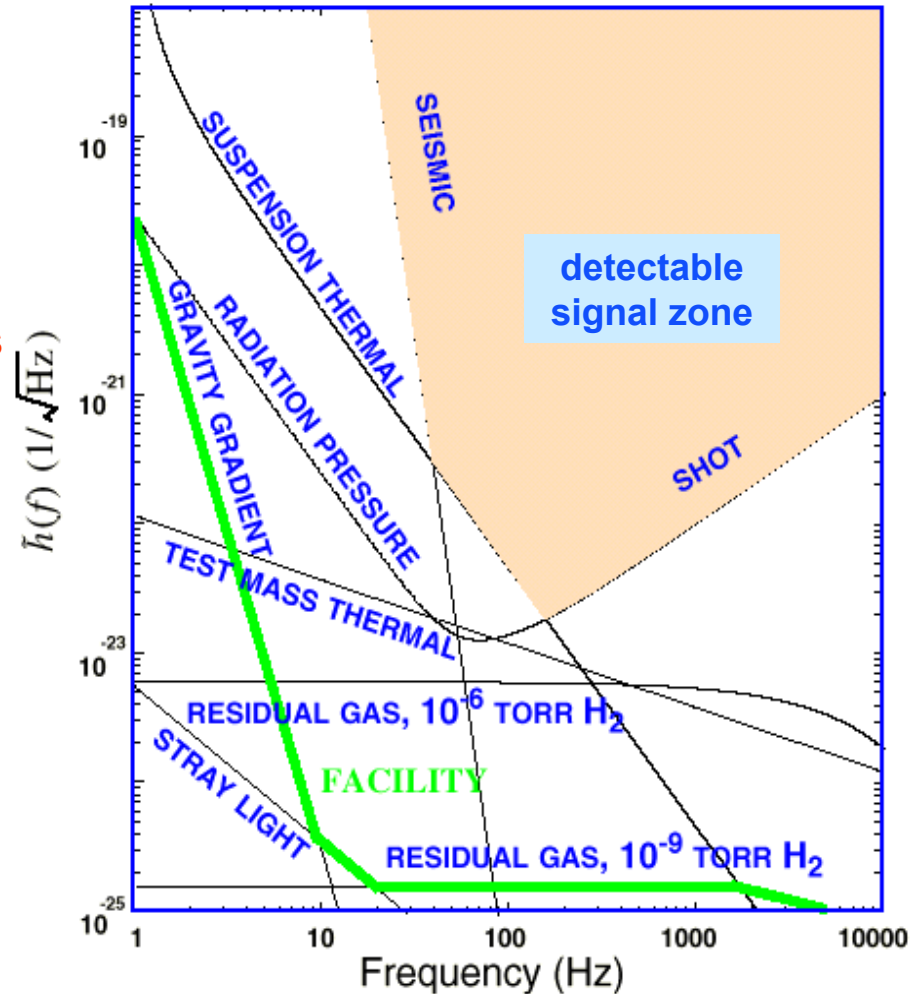
- Arm cavity storage time

$$\tau \sim 1/2\pi f_{\text{GW}}$$

- Recycling mirror matches losses, enhances effective power by  $\sim 50x$



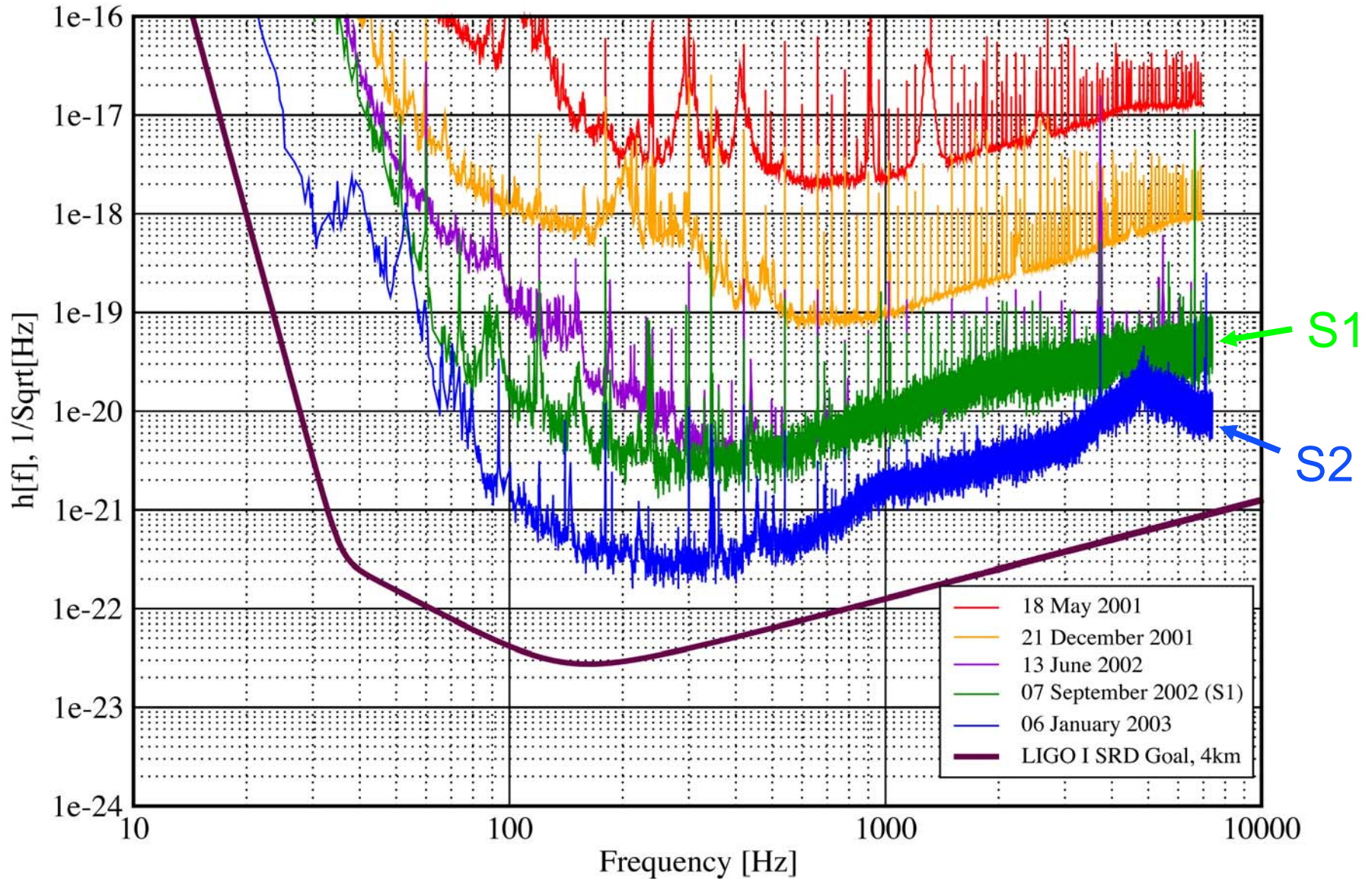
- "Fundamental" limits (with then-current technology) determined design goals
  - **seismic** at low frequencies
  - **thermal** at mid frequencies
  - **shot noise** at high frequencies
  
- **Facility limits** much lower to allow improvement as technology matures
  
- Other "technical" noise not allowed above 1/10 of these (by design, anyway...)



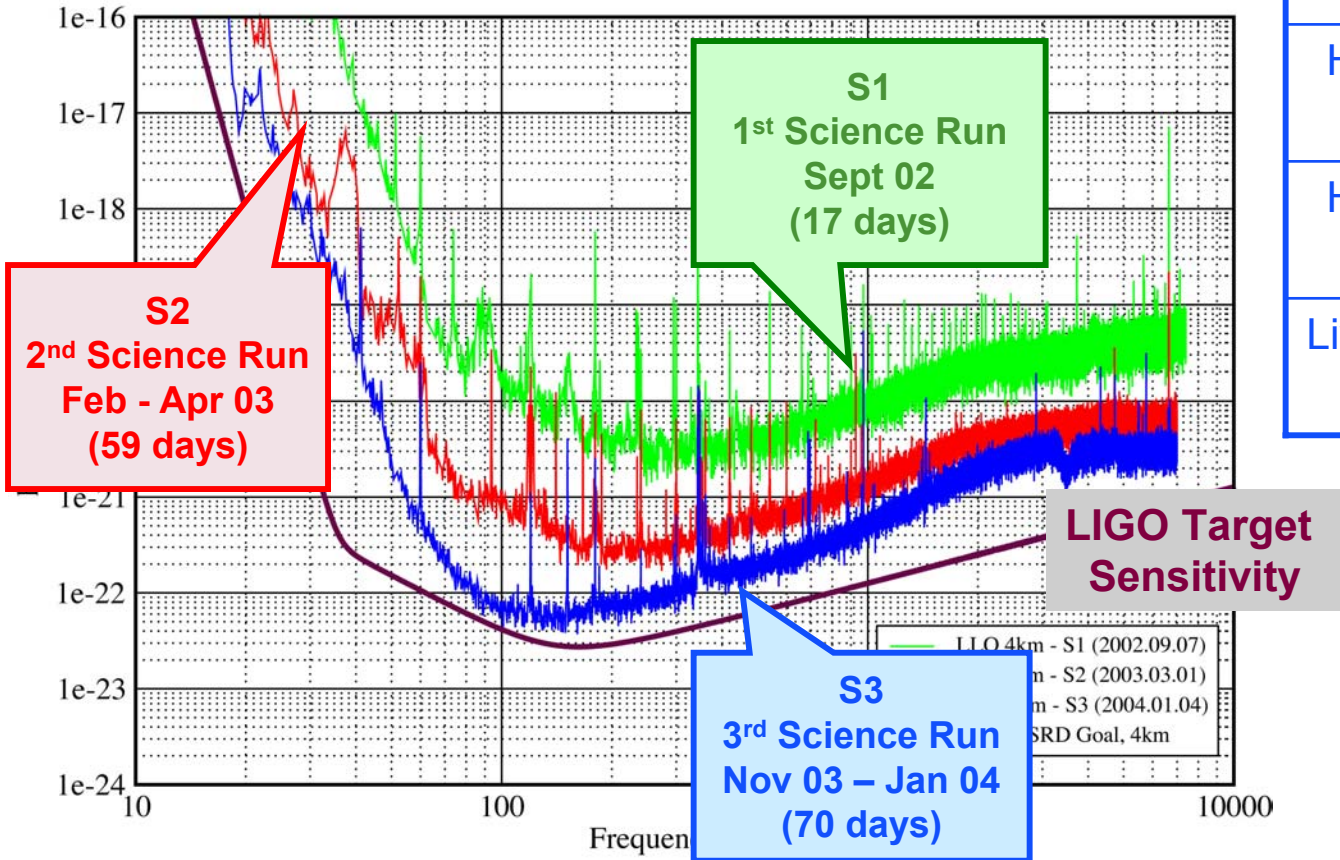
# Strain Sensitivity for the LLO 4km Interferometer

31 January 2003

LIGO-G030014-00-E



Best Strain Sensivities for the LIGO Interferometers  
 Comparisons among S1, S2, S3 LIGO-G030548-02-E



Hanford 4km	69%
Hanford 2km	63%
Livingston 4 km	22%*

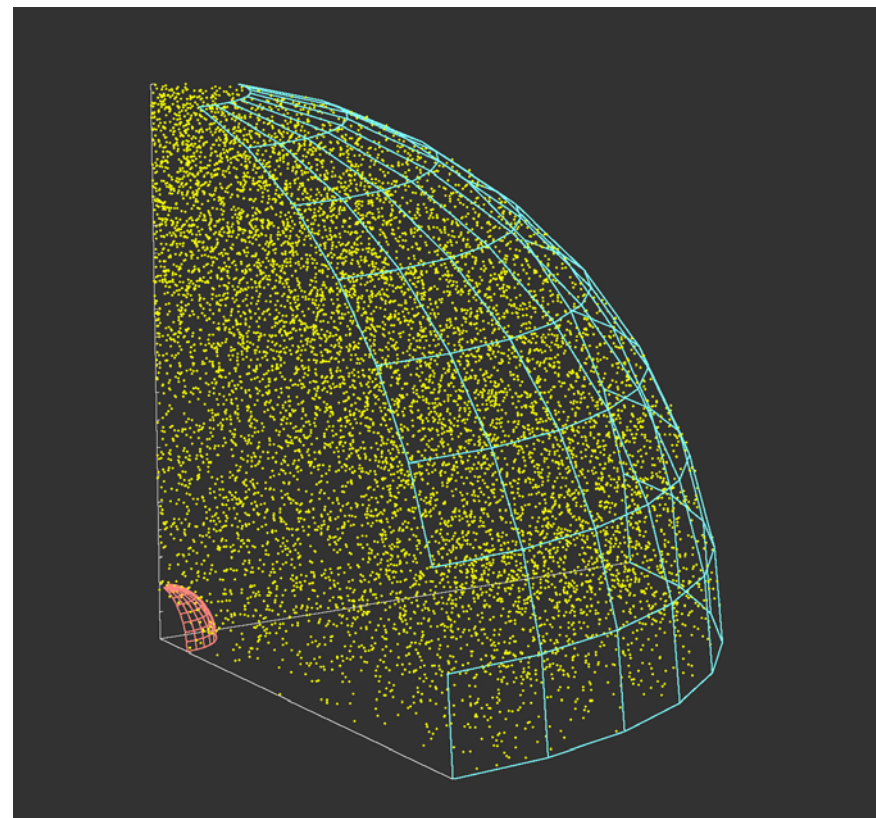
\*Limited by high ground noise—  
 active seismic upgrade currently underway

# LIGO Upper Limit Papers Using S1 Data

Papers by the LIGO Science Collaboration (~370 authors, 40 institutions):

- “*Detector Description and Performance for the First Coincident Observations between LIGO and GEO*”, Nucl. Inst. Meth A, **517**, 154-179 (2004)
- “*Setting upper limits on the strength of periodic gravitational waves using the first science data from the GEO600 and LIGO detectors*” gr-qc/0308050, Phys Rev D, April 15, 2004
- “*Analysis of LIGO data for gravitational waves from binary neutron stars*”, gr-qc/0308069, Phys Rev D, June 15, 2004
- “*First upper limits from LIGO on gravitational wave bursts*”, gr-qc/0312056, Phys Rev D, May 15, 2004
- “*Analysis of First LIGO Science Data for Stochastic Gravitational Waves*”, gr-qc/0312088, accepted by Phys Rev D

- Improved detector
  - » Must be of significance for astrophysics
  - » At the limits of reasonable extrapolations of detector physics and technologies
  - » Realizable, practical, reliable
  - » Neither too early nor too late
- Advanced LIGO:
  - ~2.5 hours = 1 year of Initial LIGO
  - » Volume of sources grows with cube of sensitivity
  - » >10x in sensitivity; ~ 3000 in rate



- » Begin installation: 2008?
- » Operational: 2011?