



A Coherent Network Burst Analysis

Antony Searle (ANU)

in collaboration with

Shourov Chatterji, Albert Lazzarini,
Leo Stein, Patrick Sutton (Caltech),
Massimo Tinto (Caltech/JPL)



Overview



- **Eyes open search**
 - Capable of detecting unknown and unanticipated waveforms
- **Innately distinguishes between gravitational waves and glitches**
 - Networks of three or more detectors over-determine the two strain polarizations for an assumed gravitational wave direction
 - Can construct $N - 2$ *null streams* **exactly** orthogonal to the strain without **any** knowledge of the waveform
 - Anything affecting a null stream is **not** the postulated gravitational wave
 - A very powerful veto
- **Needs 3+ instruments**
 - Each with different locations and orientations
 - Such as H1, L1 and one of Virgo, GEO or TAMA

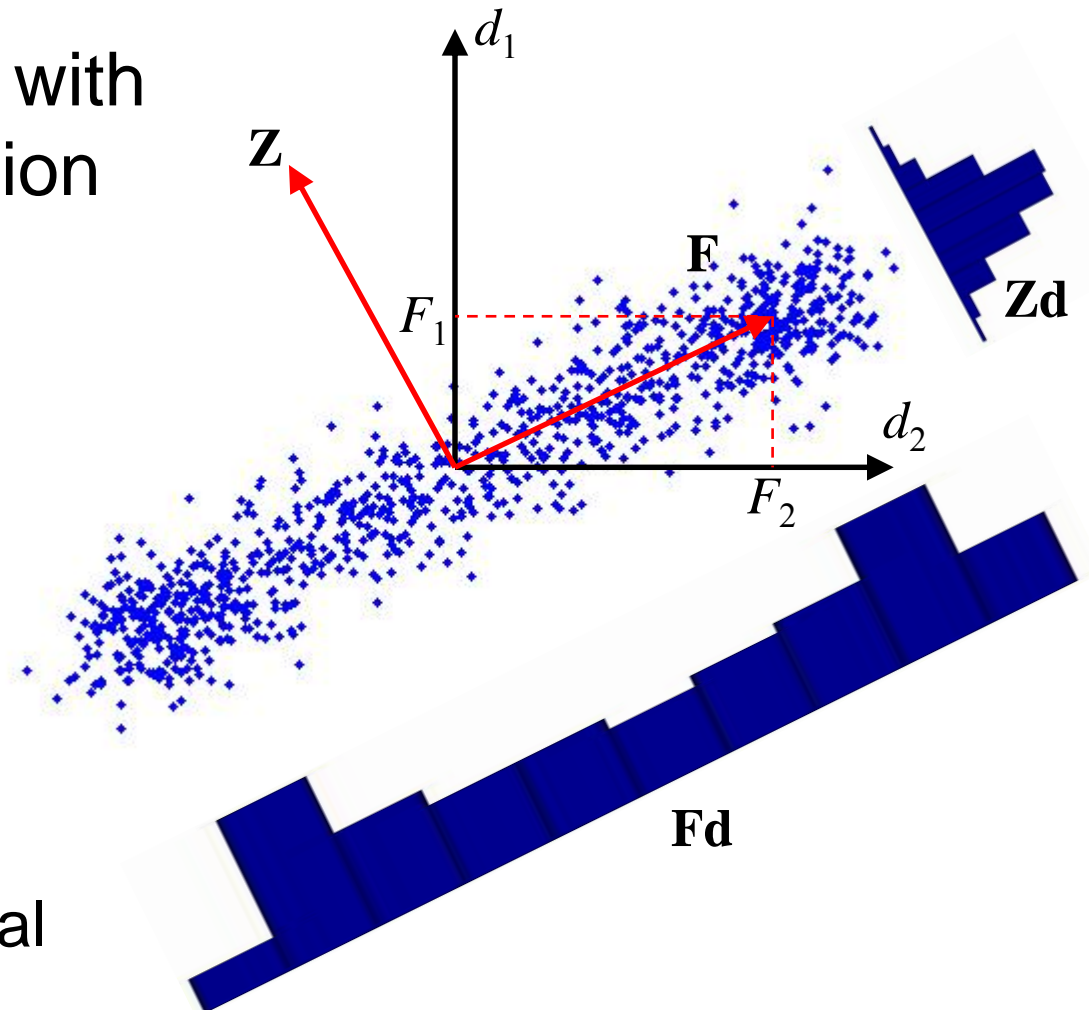
- The whitened output d_i of N detectors can be modelled by
 - Antenna patterns F_i
 - Strain h
 - Amplitude spectrum σ_i
 - White noise n_i
- The $N - 2$ linear combinations $(\mathbf{Zd})_j$ are orthogonal to strain and each other

$$\begin{bmatrix} d_1 \\ d_2 \\ \vdots \\ d_N \end{bmatrix} = \begin{bmatrix} F_1^+ / \sigma_1 & F_1^\times / \sigma_1 \\ F_2^+ / \sigma_2 & F_2^\times / \sigma_2 \\ \vdots & \vdots \\ F_N^+ / \sigma_N & F_N^\times / \sigma_N \end{bmatrix} \begin{bmatrix} h^+ \\ h^\times \end{bmatrix} + \begin{bmatrix} n_1 \\ n_2 \\ \vdots \\ n_N \end{bmatrix}$$

$$\mathbf{Z}^T \equiv \text{null } \mathbf{F}^T$$

$$\begin{aligned} \mathbf{Zd} &= (\mathbf{ZF}) \cdot \mathbf{h} + \mathbf{Zn} \\ &= \mathbf{0} \cdot \mathbf{h} + \mathbf{Zn}, \end{aligned}$$

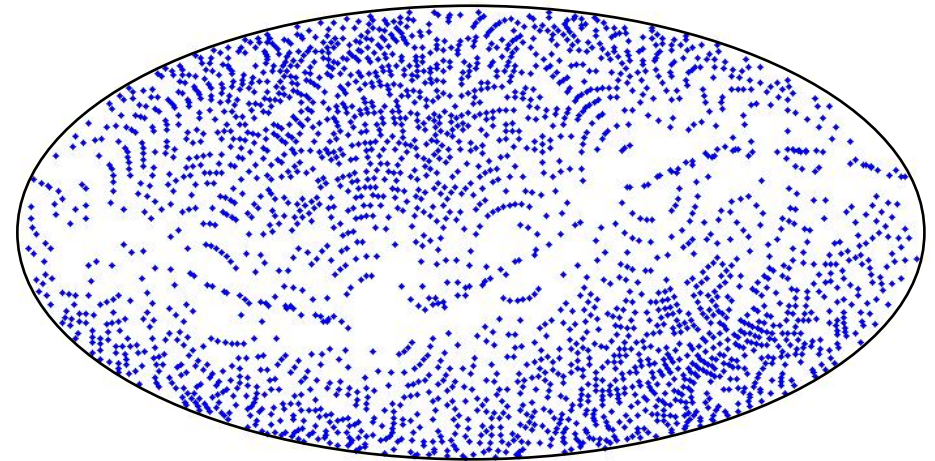
- Consider analogy with one fewer dimension
 - Detectors d_1, d_2
 - One polarization
 - Sensitivity F_1, F_2
 - Large strain h
- Null stream \mathbf{Z} is orthogonal to \mathbf{F}
 - \mathbf{Zd} is white
 - \mathbf{Fd} estimates signal



- Every direction Ω on the sky has different
 - Null stream coefficients \mathbf{Z}
 - Delays Δt_i for detector at \mathbf{x}_i

$$c\Delta t_i = -\mathbf{x}_i \cdot \Omega$$

- Sample the sky with some limited mismatch
 - Template placement problem
 - Affected by network geometry



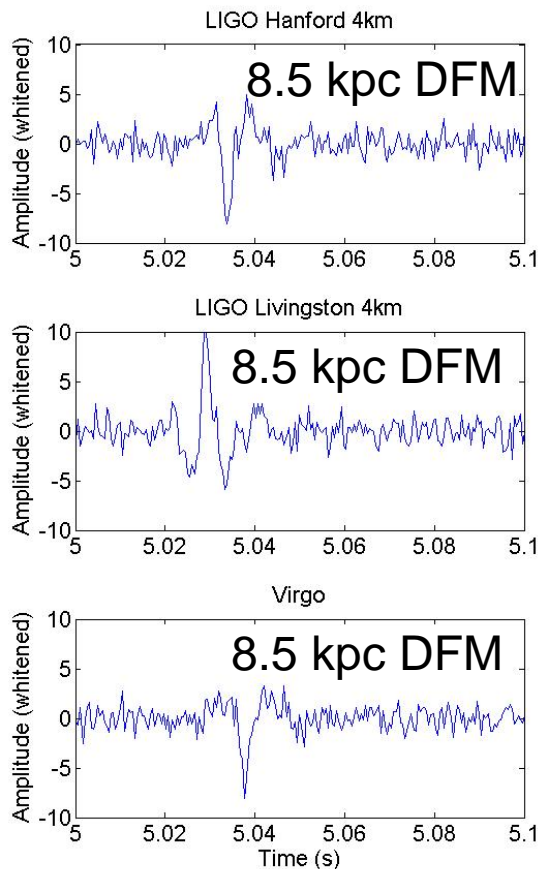
- Mollweide plot of 0.6 ms resolution map for HLV
 - Near-optimal
 - Low density on plane of HLV baselines

“Is the data consistent with noise plus a gravitational wave from some direction?”

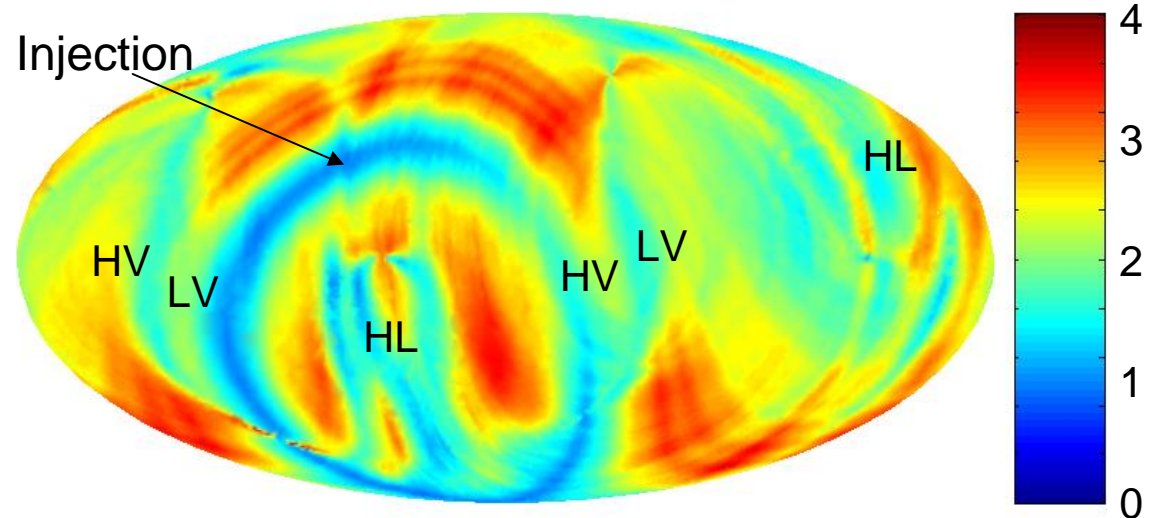
is equivalent to

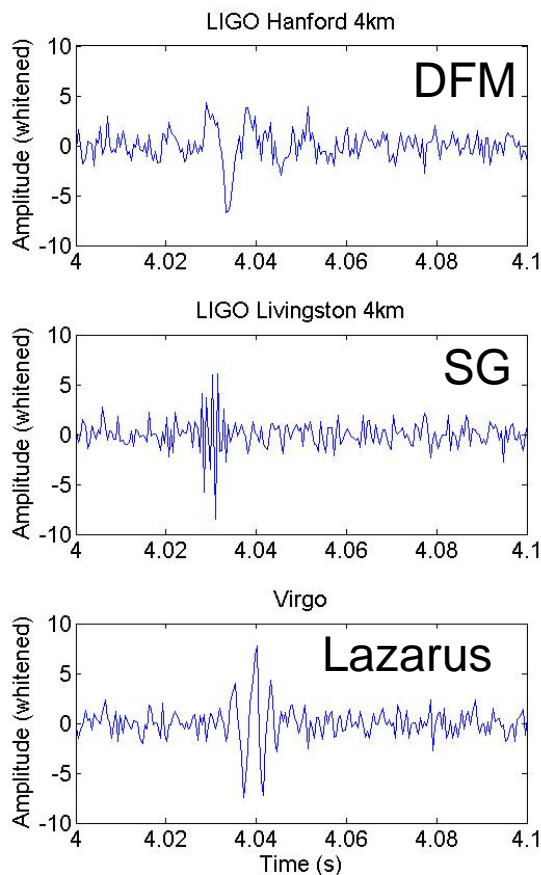
“Are the null streams for that direction consistent with noise?”

- Use a χ^2 test
 - Test that the total energy E_{null} of the null streams is consistent with white noise

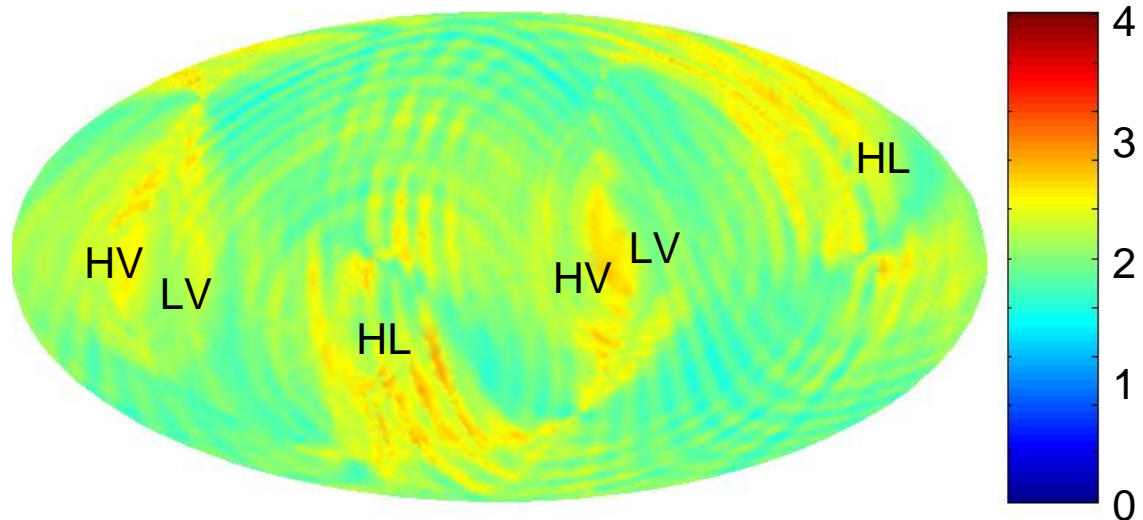


- Inject a gravitational wave
- Null stream energy **consistent** with noise at correct direction
 - Signal cancels out





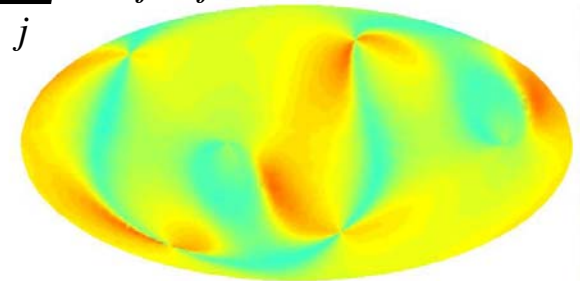
- Inject three different waveforms (a 'glitch')
 - Consistent times, energies
- Nowhere **consistent** with noise



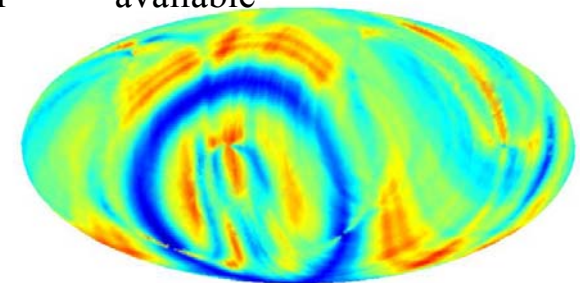
- Null stream energy E_{null} may be split into two parts
- Available energy $E_{\text{available}}$
 - “Diagonal” terms
 - Weighted sum of detector energies
 - Broad features on sky map
- Correlation energy $E_{\text{correlation}}$
 - “Off-diagonal” terms
 - Weighted sum of pair-wise detector correlations
 - Fringes on sky map

$$E_{\text{null}} = \sum_i^{N-2} \sum_j^N \sum_k^N Z_{ij} Z_{ik} d_j d_k$$

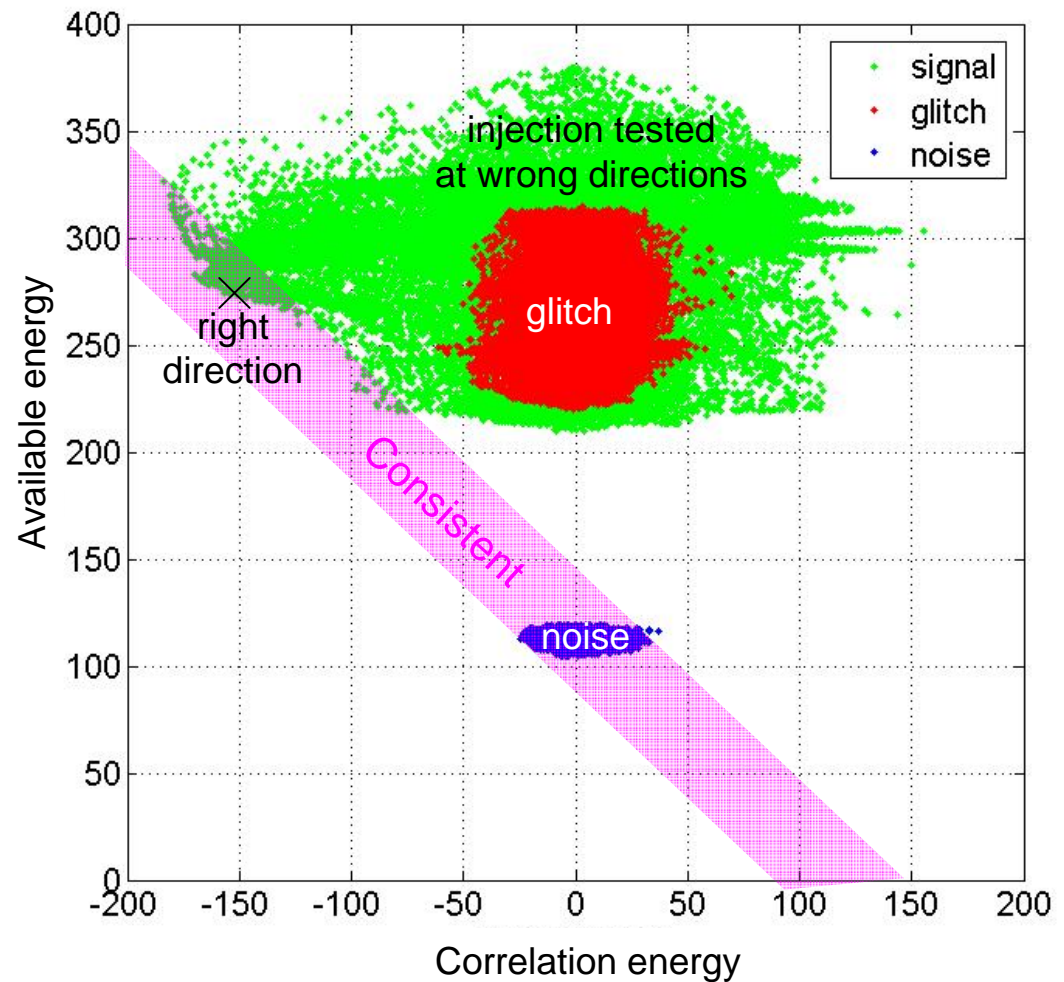
$$E_{\text{available}} = \sum_i^{N-2} \sum_j^N (Z_{ij} d_j)^2$$



$$E_{\text{correlation}} = E_{\text{null}} - E_{\text{available}}$$



- **Energy** in the detectors boosts up the plot
- **Correlation** in the detectors broadens across the plot
- **Cancellation** when
 - Consistent with gravitational wave
 - Right direction on sky



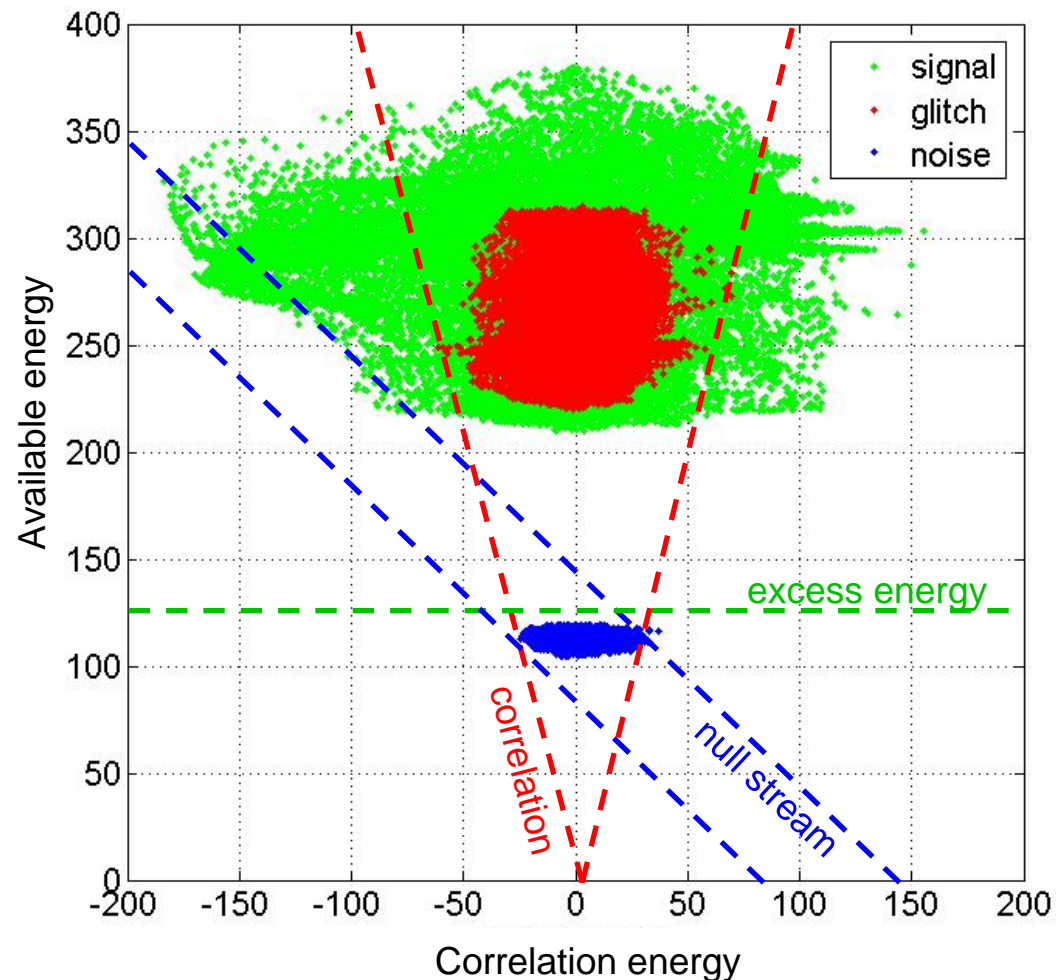


Characterization



- How does it compare to existing methods?
 - Will reject energetic and even correlated glitches
 - Won't reject a gravitational wave
 - Won't reject background noise and small glitches
 - **Complementary to existing tests**
- To form a search, must combine it with some other test(s)
 - **What is the population of small glitches that pass the null test, and how can we eliminate them?**

- Use **excess energy** to trigger
- Require **correlation**
- Use the **null stream** to identify gravitational wave candidate events





Real world problems



- **Nonstationarity,**
- **Calibration errors**
and
- **Direction mismatch**
 - Null stream will not exactly cancel signal, so there will be residual excess energy
- **Computational cost**
 - May be practical as triggered search only
- **Duty cycle**
 - Requires at least three different sites taking data
- **Glitch population**
 - How correlated?



Future directions



- Performing large scale Monte-Carlo simulations
 - MATLAB pipeline
 - `lsc-soft/matapps/src/searches/burst/coherent-network`
 - Test against real glitches
 - Compare with other tests
- Preferentially detect “physical” waveforms?
 - Maximum entropy methods?
- More work on statistics
 - Bayesian interpretation?
 - Pattern recognition on sky maps?