

The Mesa Beam Initial Results



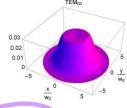
UNIVERSITY of GLASGOW

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LIGO-G050462-00-R

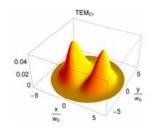
Caltech/ LIGO
LMA Lyon/ EGO
University of Glasgow



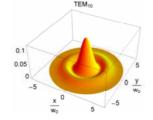


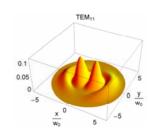
Introduction

- Thermal noise shall be a limiting factor for AdvLIGO
- Flat-topped beams reduce thermal noise
- A cavity has been constructed to generate and characterise the mesa beam

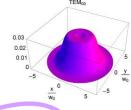


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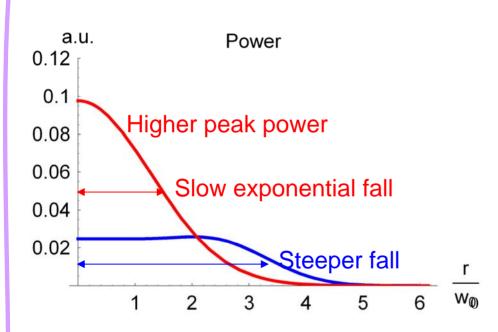








Profiles



• Profiles normalised for same integrated power

• Same diffraction losses

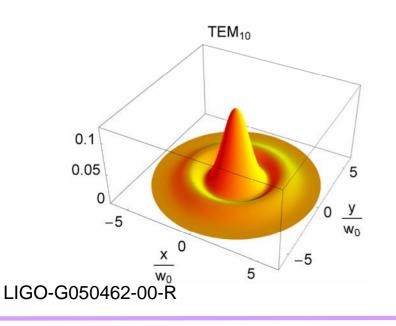
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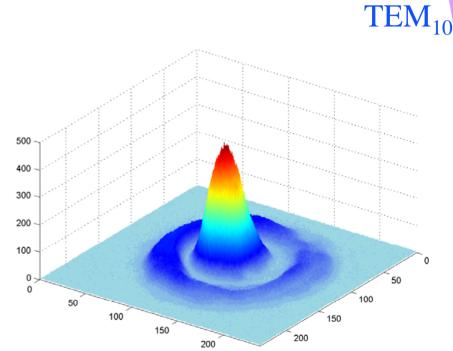


Results

LIGO R&D

- We have been able to lock to higher order modes
- These modes exhibit good agreement with theory





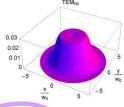
TEM₀₀

5 -5

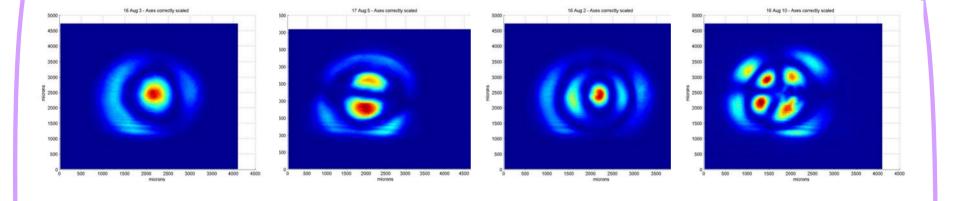
0.03

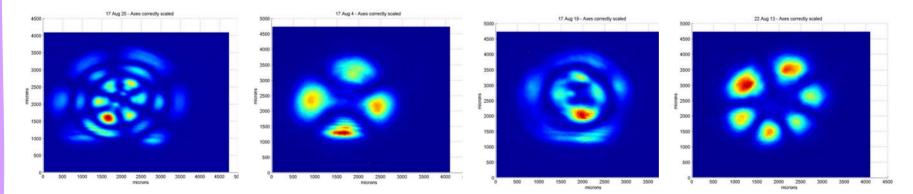
0.02

LIGO



Results - HOM

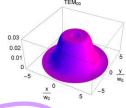




Diffraction around beam baffle eliminated

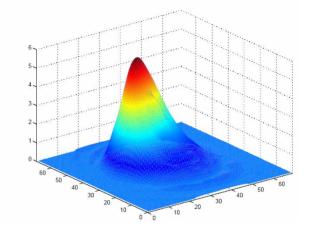
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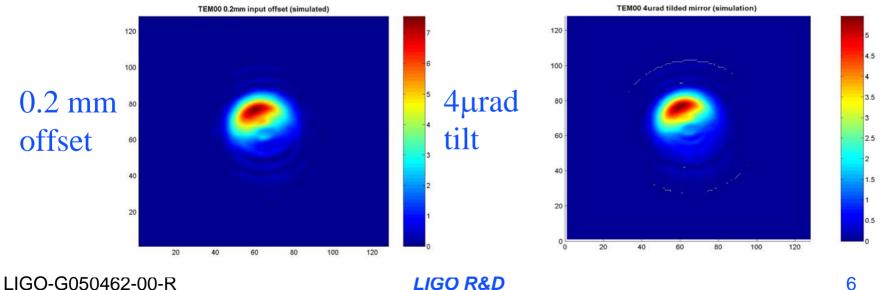




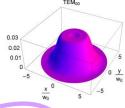
The Fundamental?

- Alignment is taxing
- We are unable to distinguish between translations and tilts of the input beam



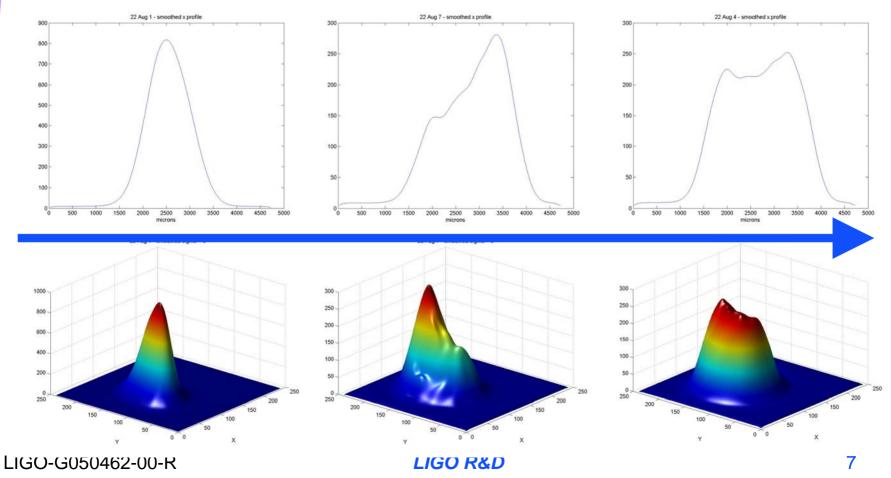




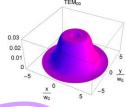


Improving Alignment

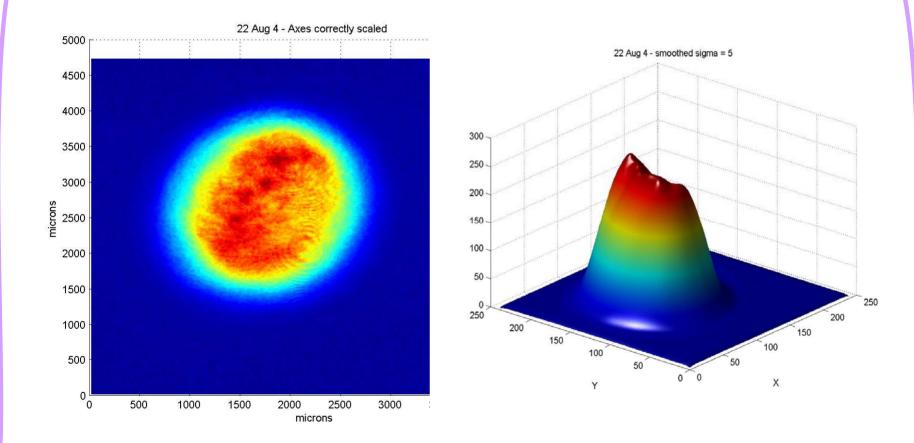
• The reference during alignment was changed from the intensity profile to the transverse mode spectrum





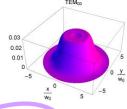


The Mesa Beam

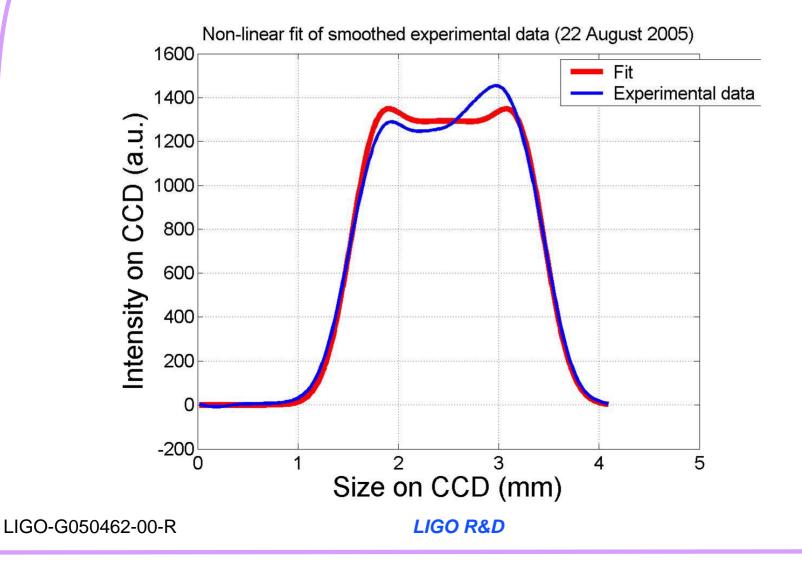


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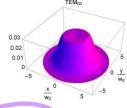
LIGO



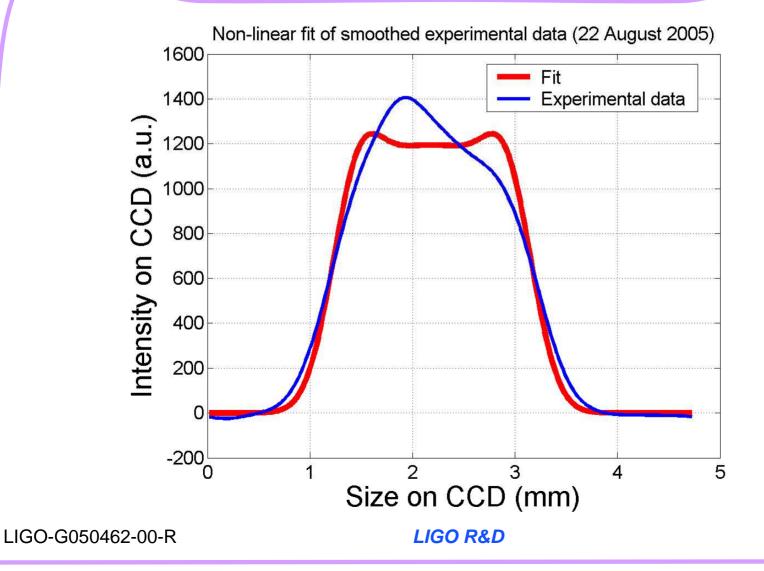
Non-Linear Fit X



LIGO



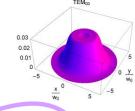
Non-Linear Fit Y

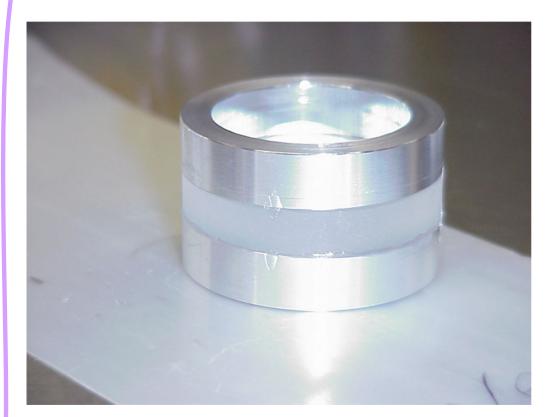


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How Was This Achieved?

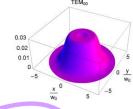




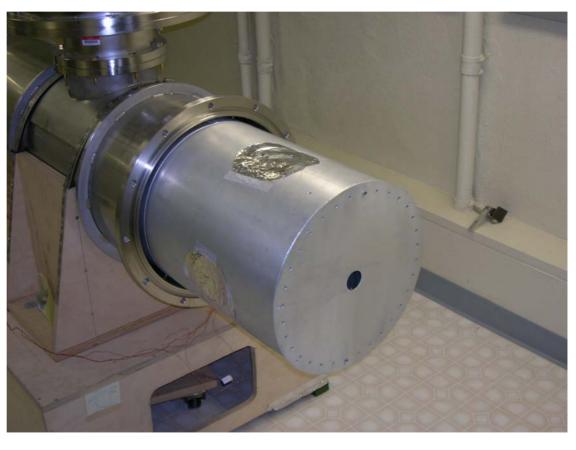
- We have reinforced flexible mirrors with aluminium rings
- Thicker substrates have been ordered



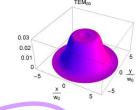
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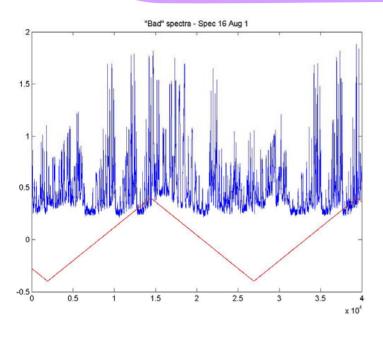


- Improved atmospheric isolation
- Better stability 'in lock'



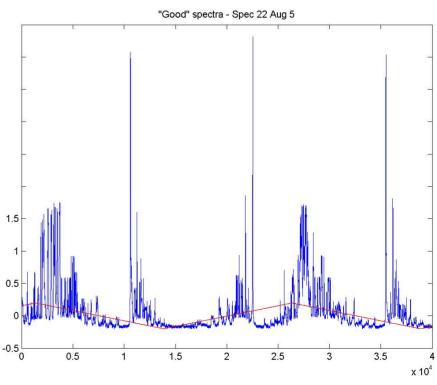
How Was This Achieved?





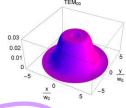
• More power in the fundamental mode





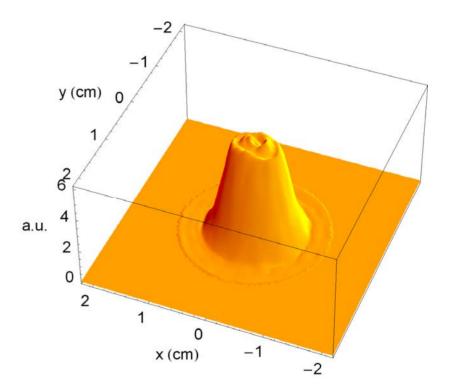
LIGO





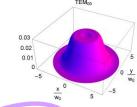
Optimal Profile

- Even with aluminium rings the flat mirrors are far from perfect
- Thicker flat mirrors have been ordered
- New mounts being designed





Further Work With This Set Up



- Improve profile using new flat mirrors
- Repeatability/ stability
- Quantify alignment/ mirror figure error tolerances cf. theory
- Test other two MH mirrors
- Characterise high order modes
- Long term design and build half of a nearly concentric MH Cavity

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