

Improvement of the MGAS Filter Damping Performance

Alberto Stochino



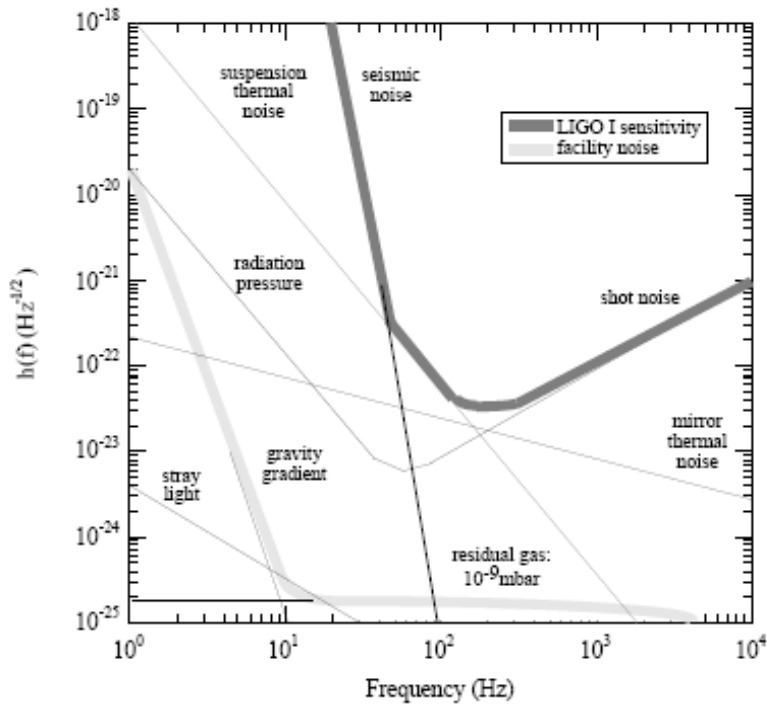
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SURF Student

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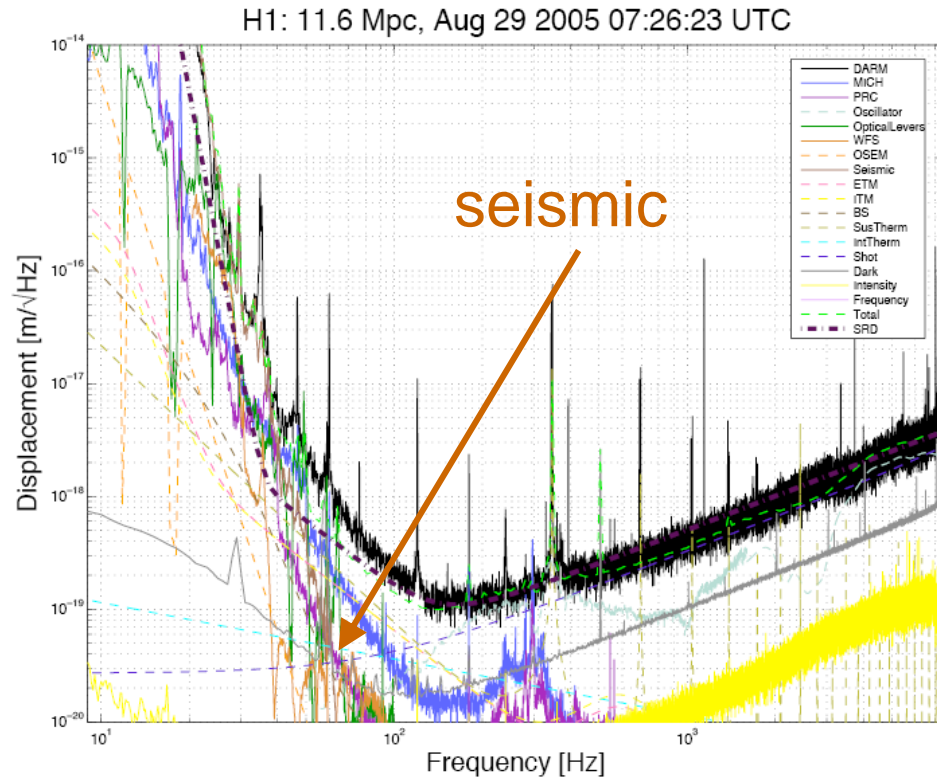
Seismic Noise

It is the disturbance resulting from the displacement of the mirrors due to the ground motion



Noise Budget

LIGO-G050484-00-R

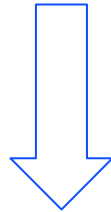


LIGO R&D

Vertical Motion of the Mirrors

The vertical and horizontal degrees of freedom are not independent

- mechanical complexity
- non uniformity of the terrestrial gravitational field



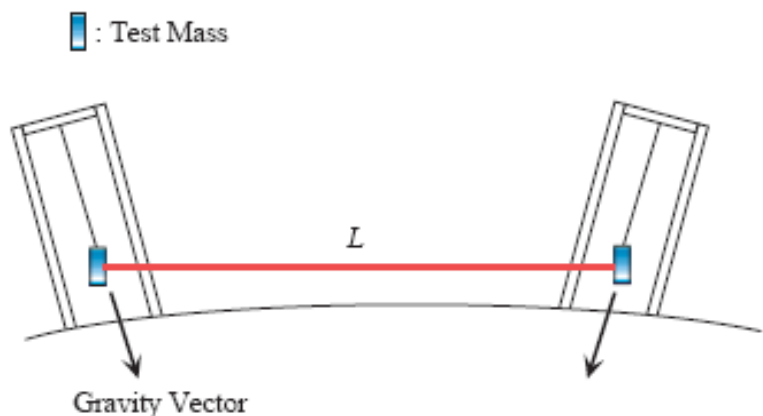
the motion of the mirrors in the vertical direction can change the arm length of the interferometer

Vertical Attenuation

- ground motion $\sim 10^{-6}$ m
- expected GW signal $\sim 10^{-18}$ m
- required attenuation factor $\sim 10^{-12}$

→ vertical to horizontal couplings must not be limiting factors to the horizontal sensibility

Gravitational Coupling



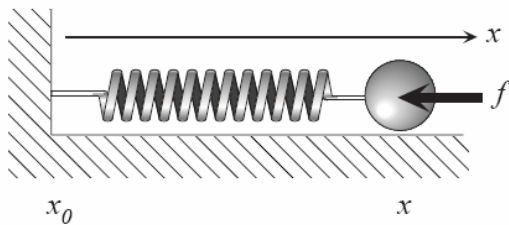
$$\Delta x = \alpha \Delta z$$

$$\alpha = L/r_0 \quad \sim < 10e-3$$

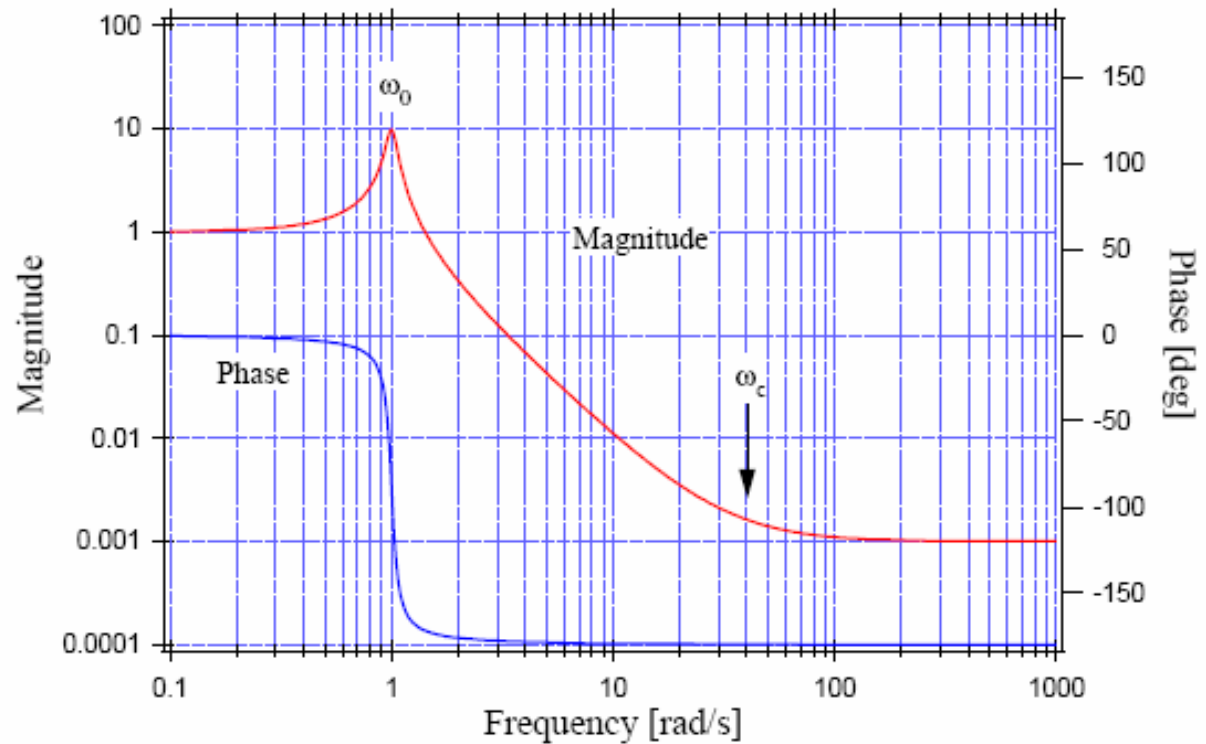
negligible respect to other couplings $\sim 1\%$

Passive Mechanical Filters

Massive spring

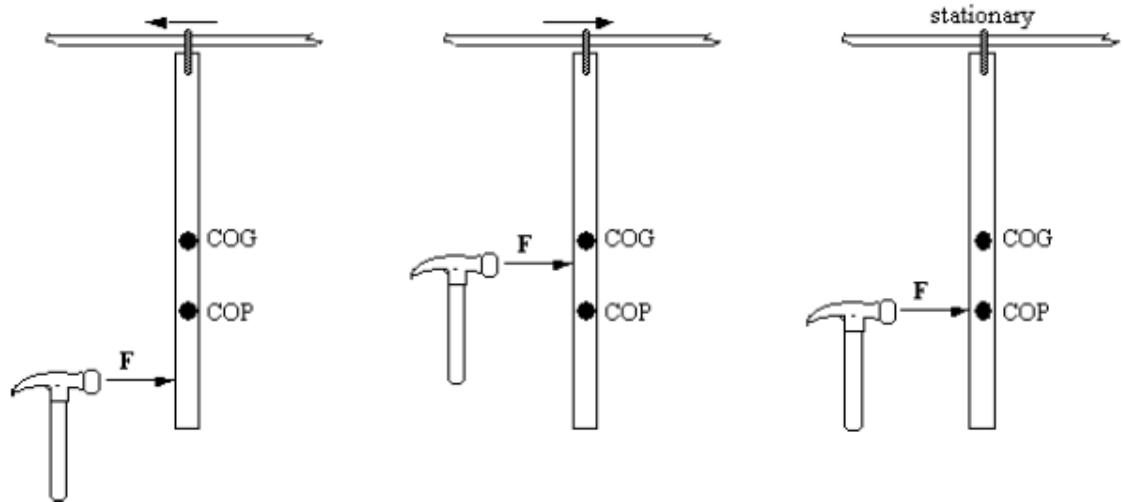
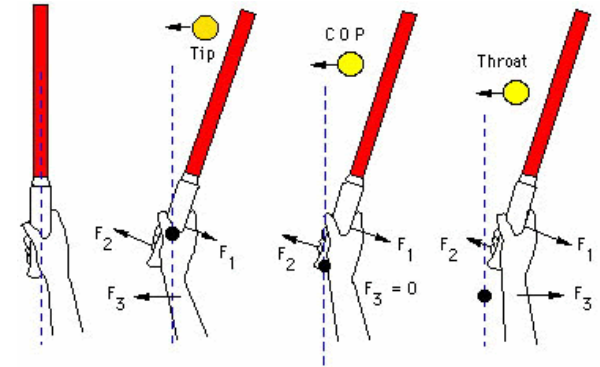
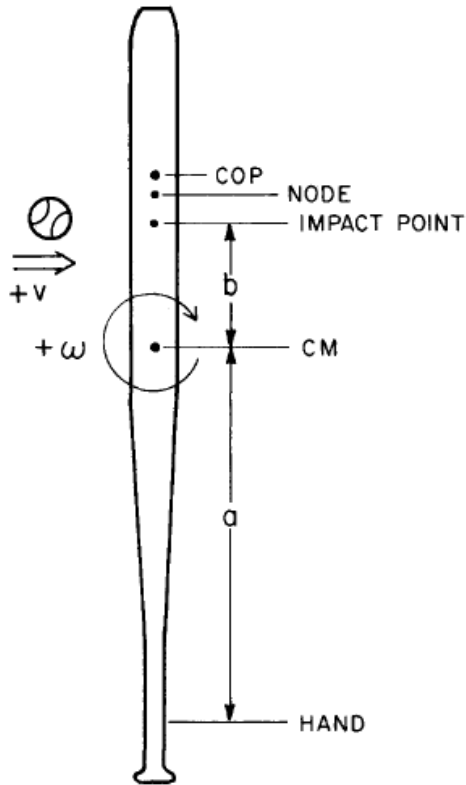


$$H_x \equiv \frac{\omega_0^2(1+i\phi) + \frac{M}{m}\omega^2}{\omega_0^2(1+i\phi) - \omega^2 + i\frac{\gamma}{m}\omega}$$



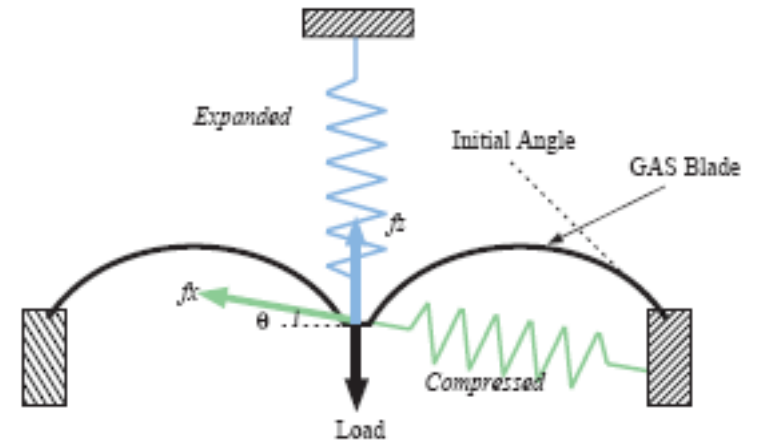
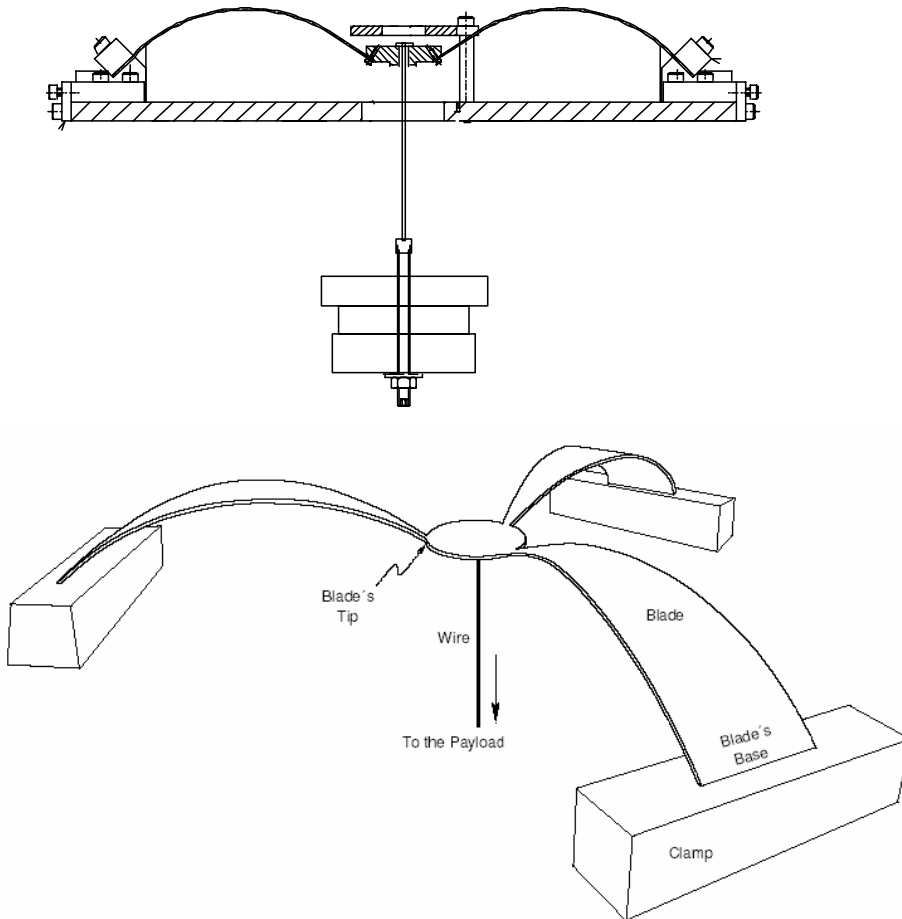
The C.O.P. Effect limitation on attenuation performance

Center Of Percussion



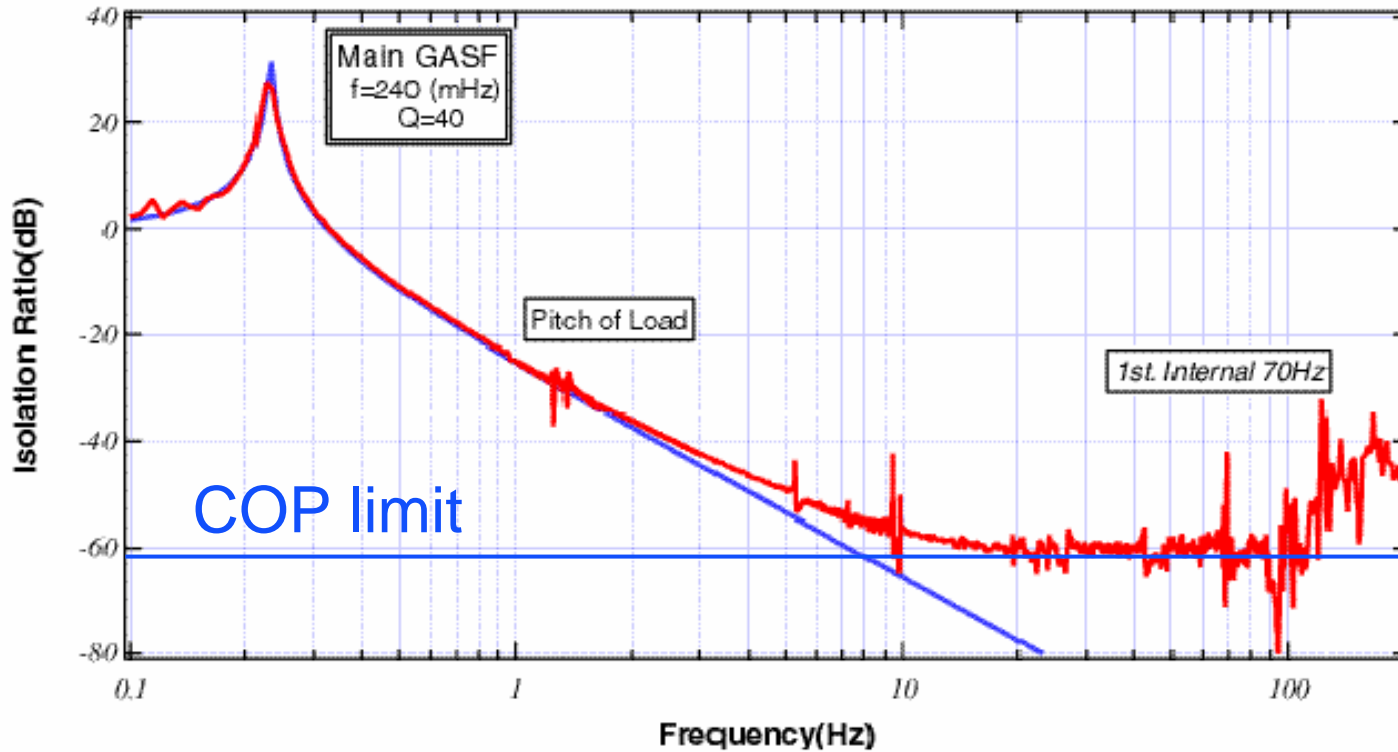
MGAS Filter

Linear Model for GAS Springs



GAS Filter Limit

Transfer Function



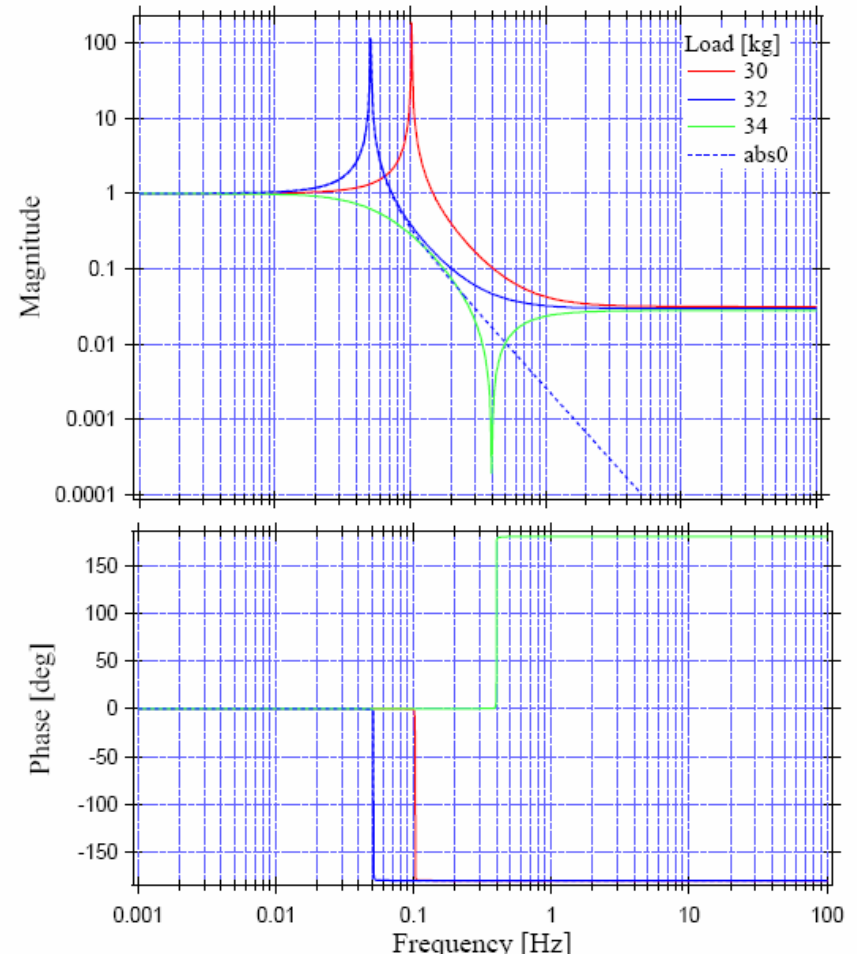
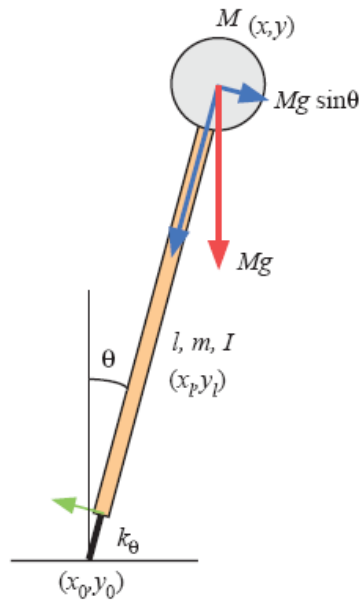
(Kenji Numata)



Inverted Pendulum

Transfer Function

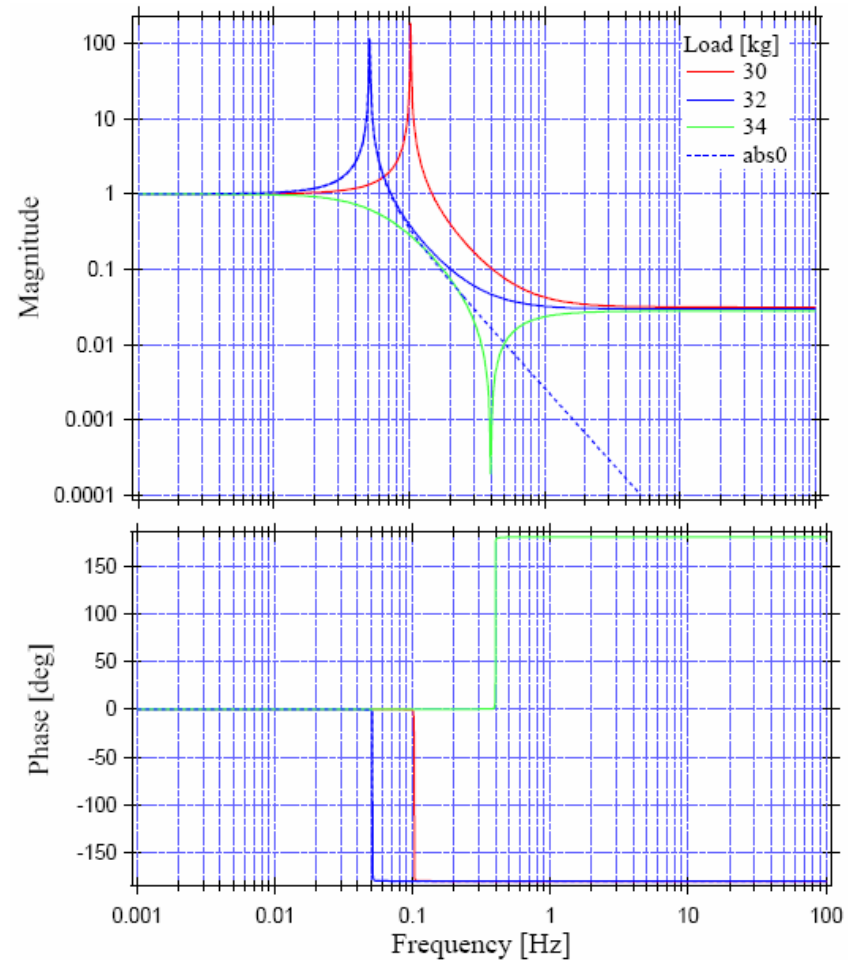
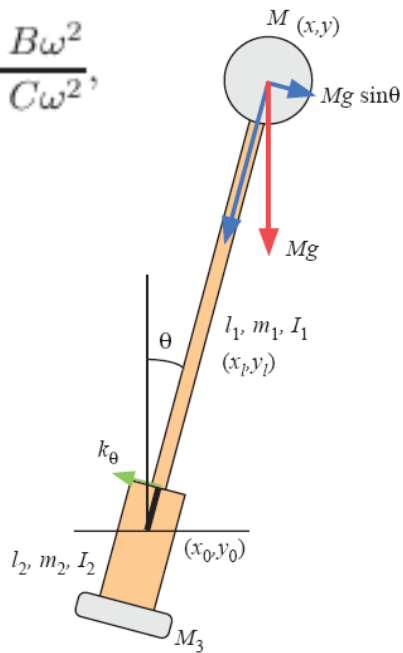
$$H_{IP}(\omega) = \frac{x}{x_0} = \frac{k_{\text{eff}} + (\frac{m}{4} - \frac{I}{l^2})\omega^2}{k_{\text{eff}} - (M + \frac{m}{4} + \frac{I}{l^2})\omega^2}$$



C.O.P. Displacement

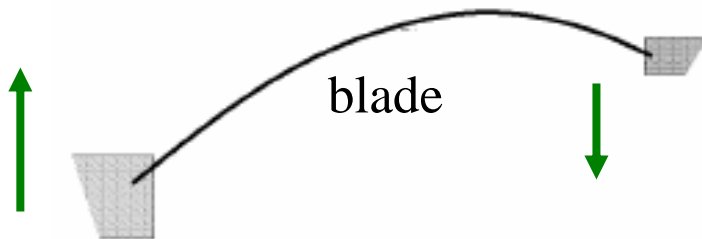
Transfer Function for the counterweighted IP

$$H_{IPCW}(\omega) = \frac{A + B\omega^2}{A - C\omega^2},$$

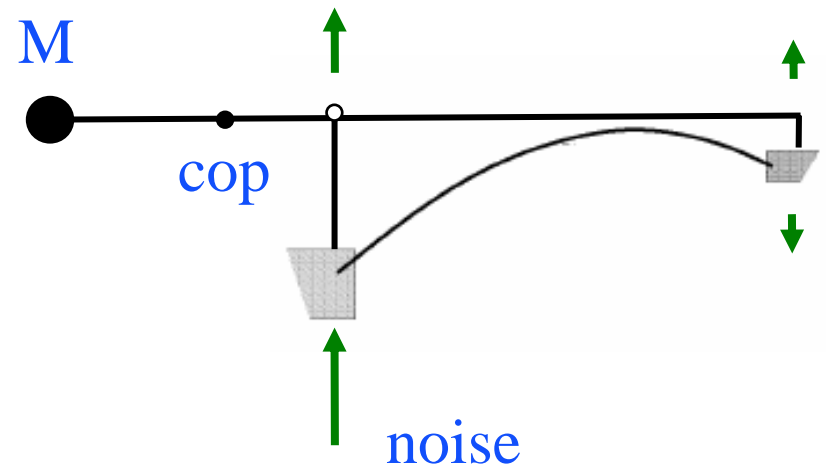


Extending the Idea to the GAS Springs

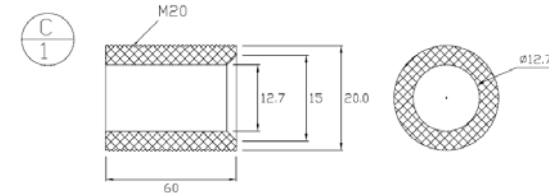
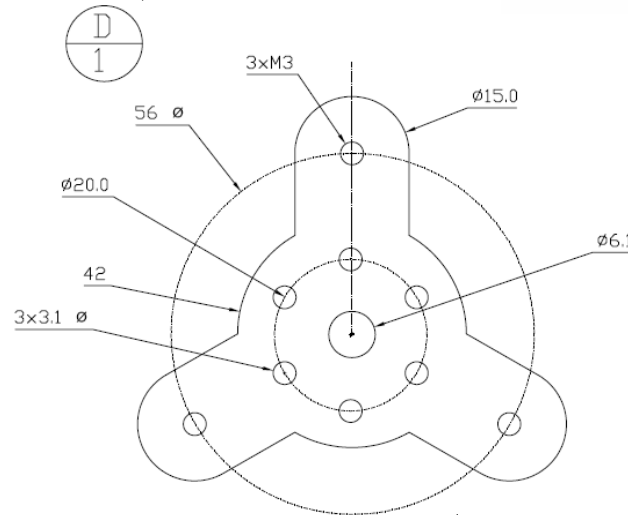
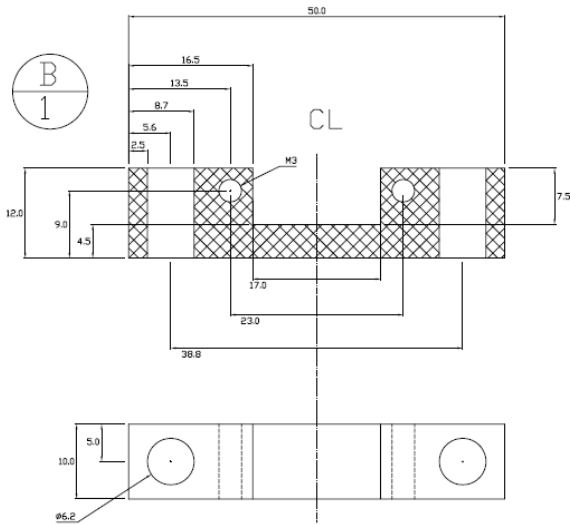
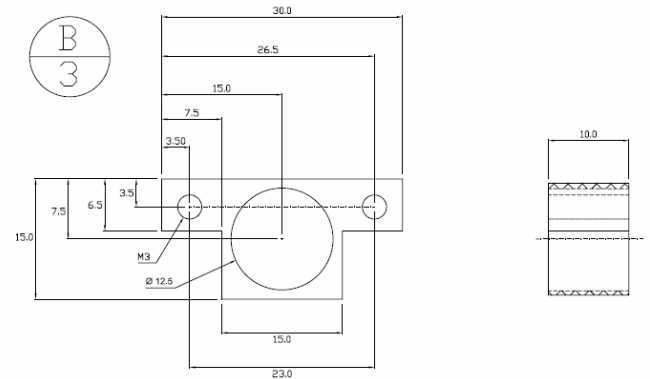
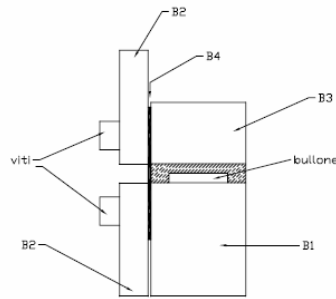
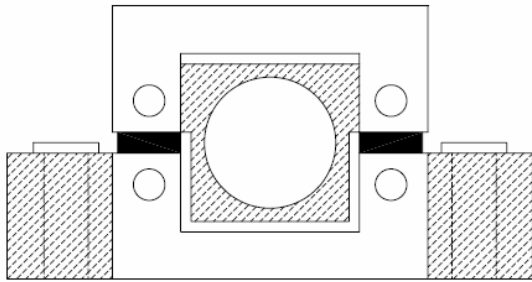
What we had before



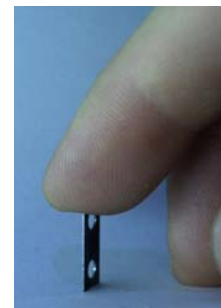
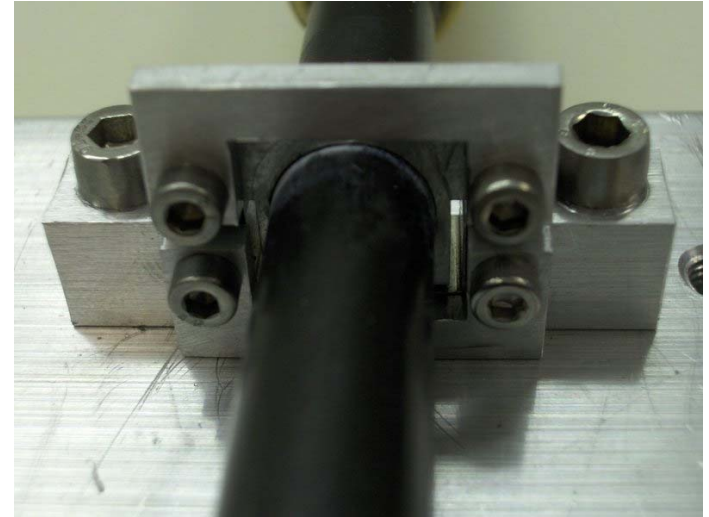
What we wanted to do



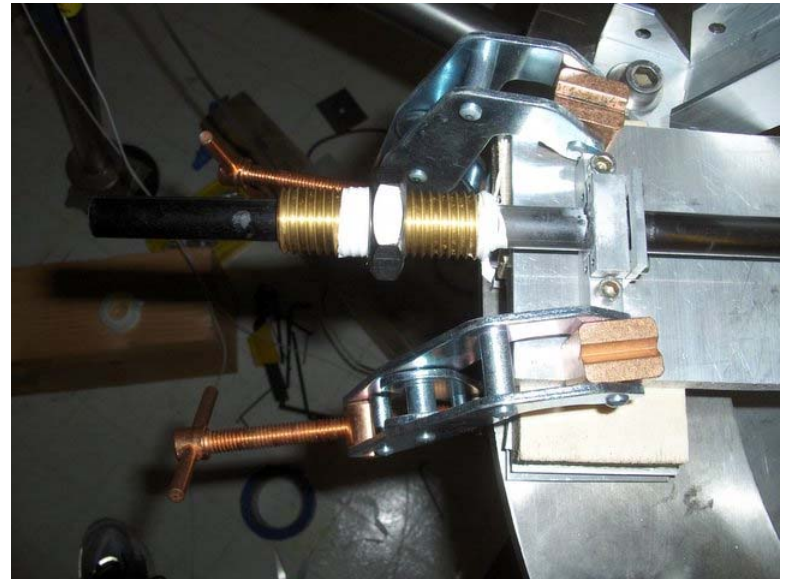
Prototype Design



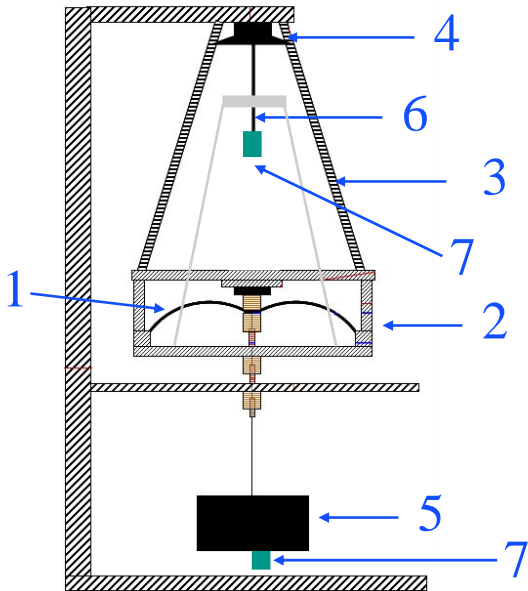
Making it Real



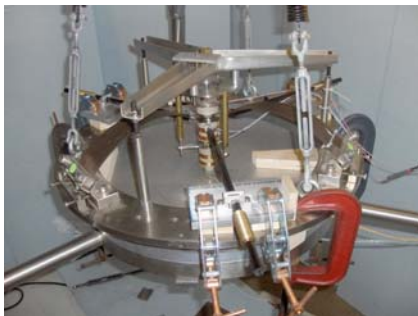
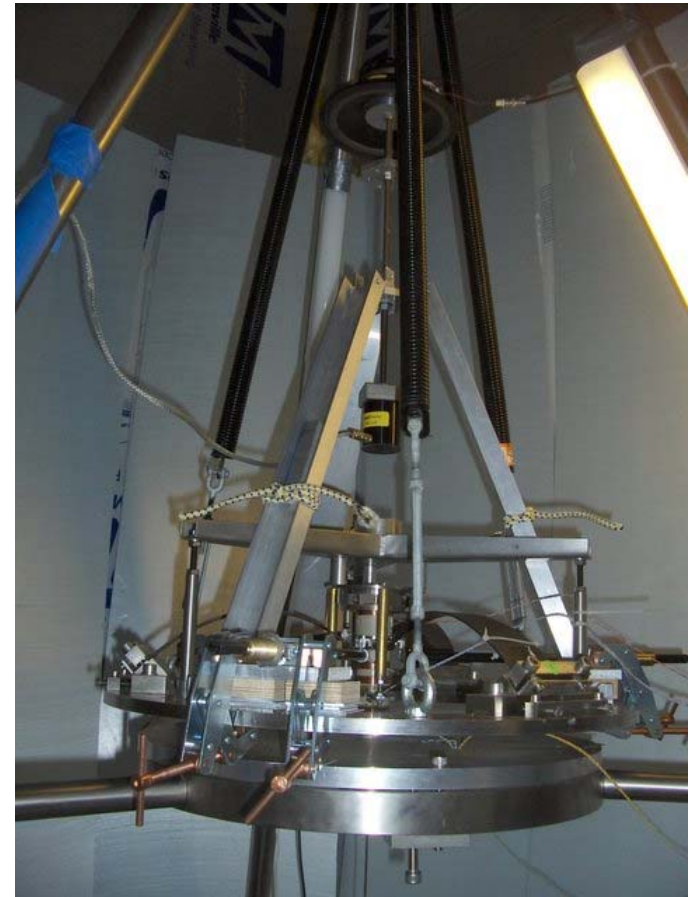
Tuning of the Counterweight



Experiment Setup



- 1) GAS blade
- 2) Filter body
- 3) Support spring
- 4) Voice-coil
- 5) Payload
- 6) Rigid connector
- 7) Accelerometer

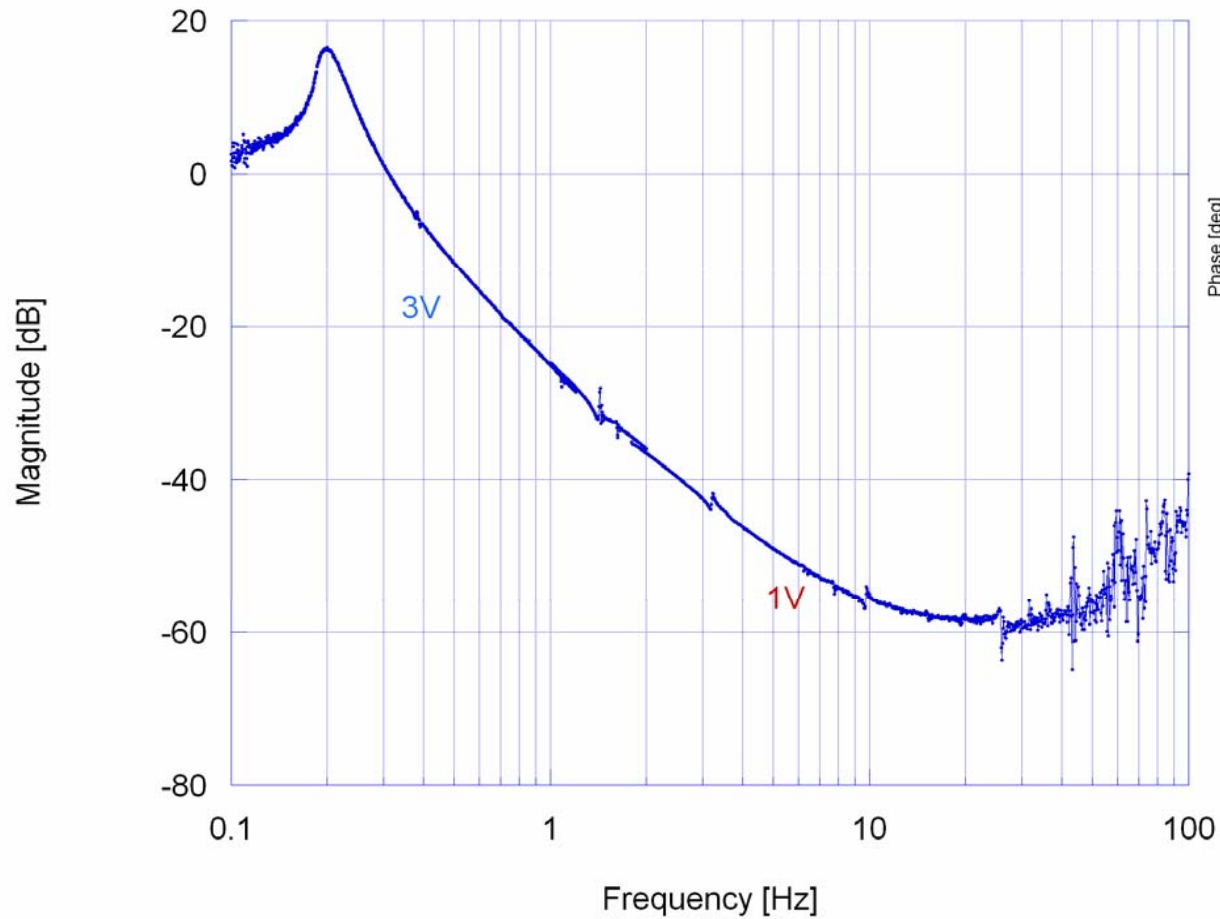


LIGO-G050484-00-R

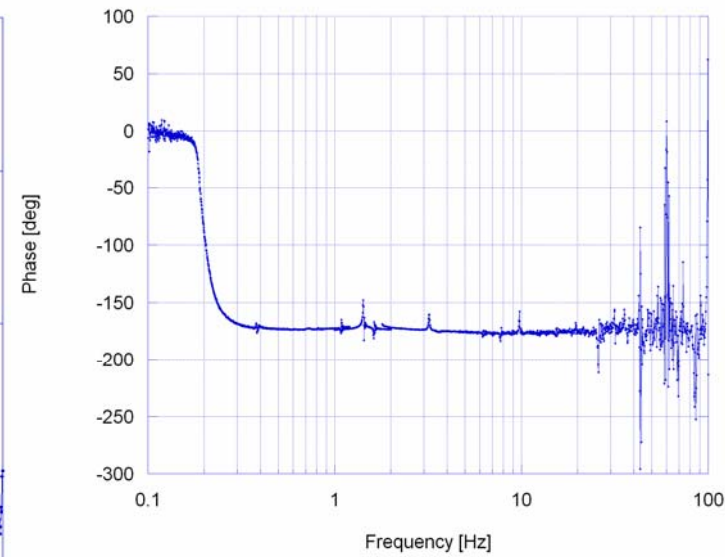
LIGO R&D

The Initial TF

Transfer Function

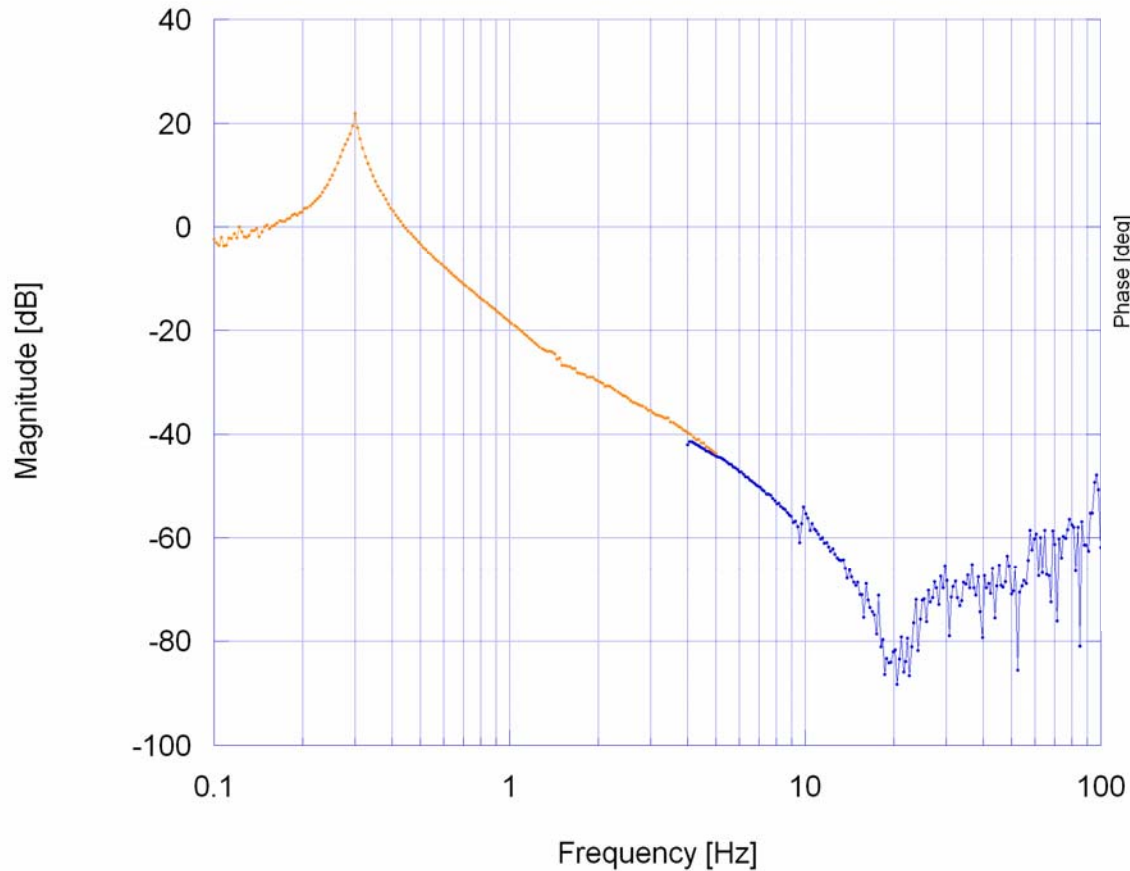


Phase

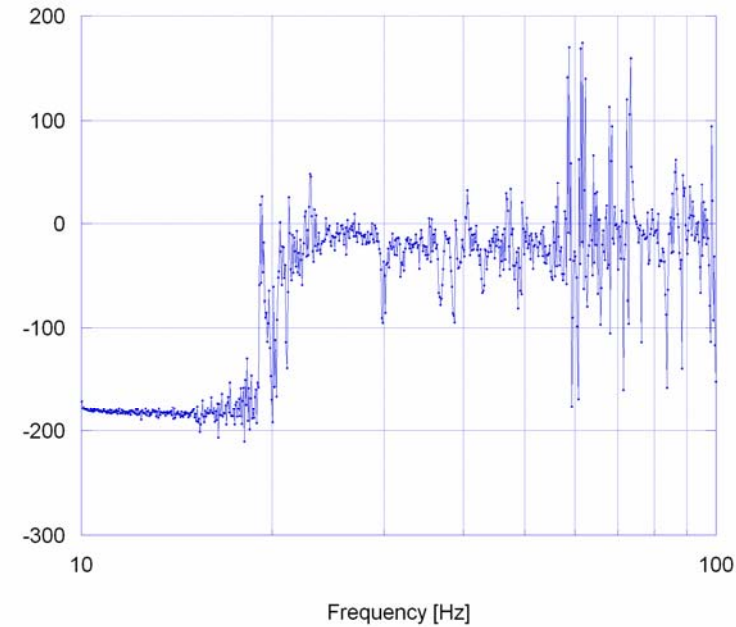


Overcompensation Three Booms No CWs

Transfer Function

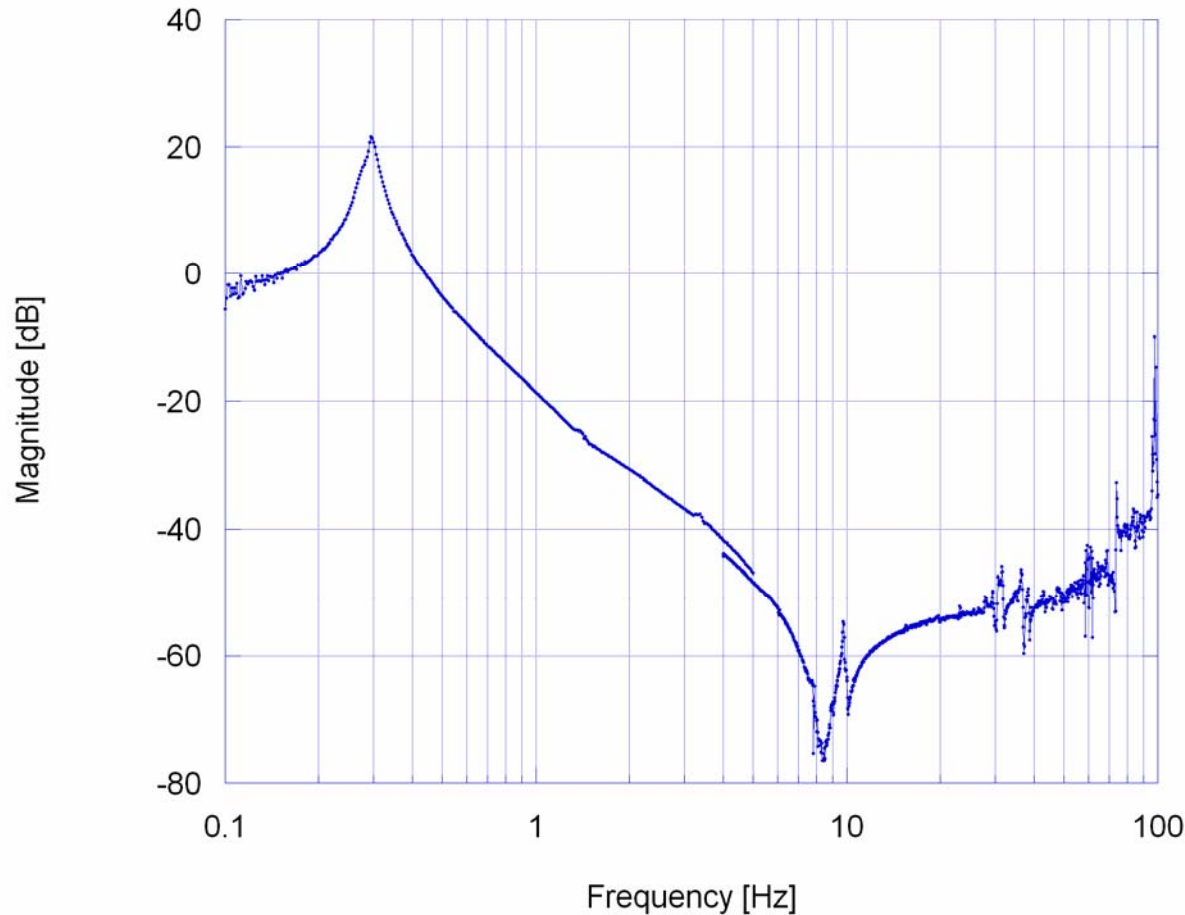


Phase

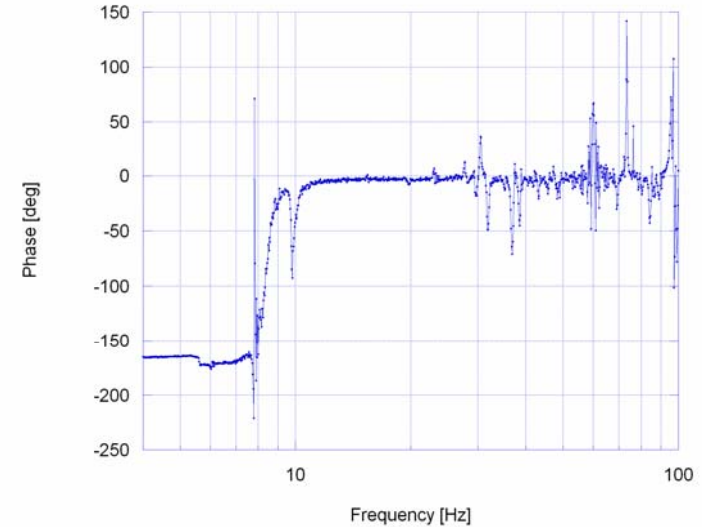


Three Booms + 2x3 CWs

Transfer Function



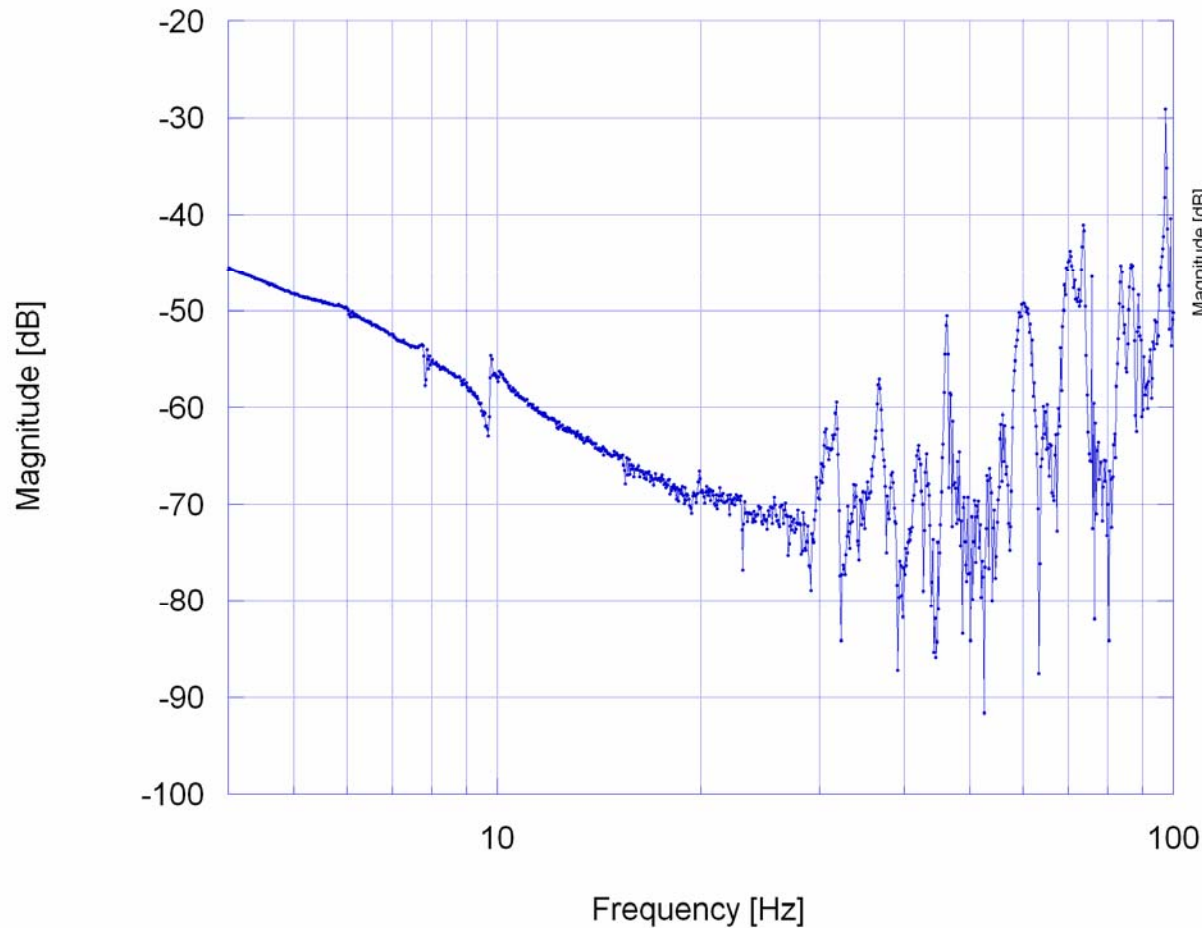
Phase



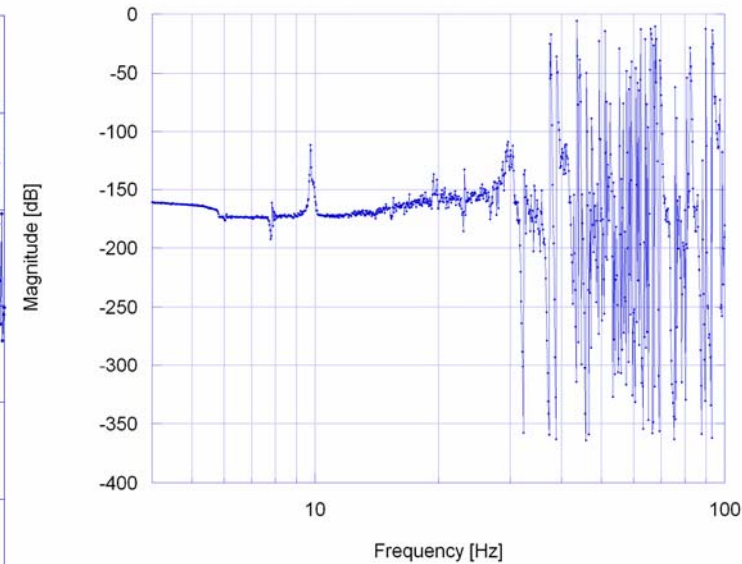
Note the
phase advance

Undercompensation Two Booms No CWs

Two Booms

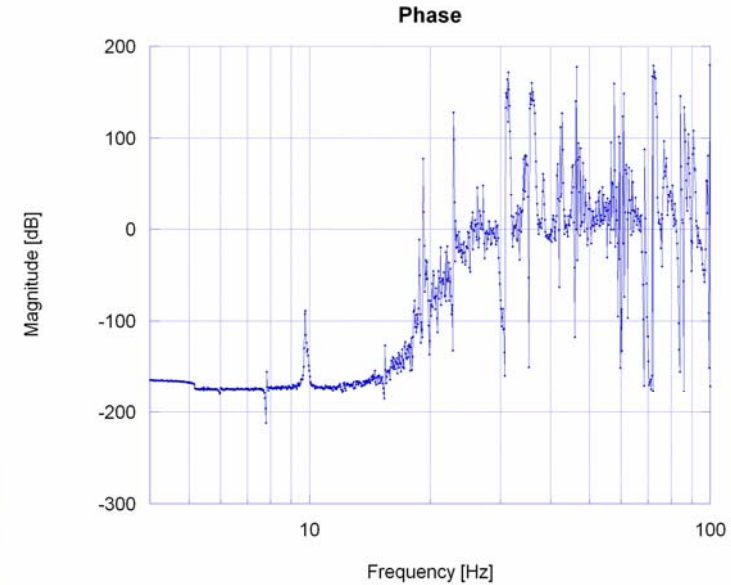
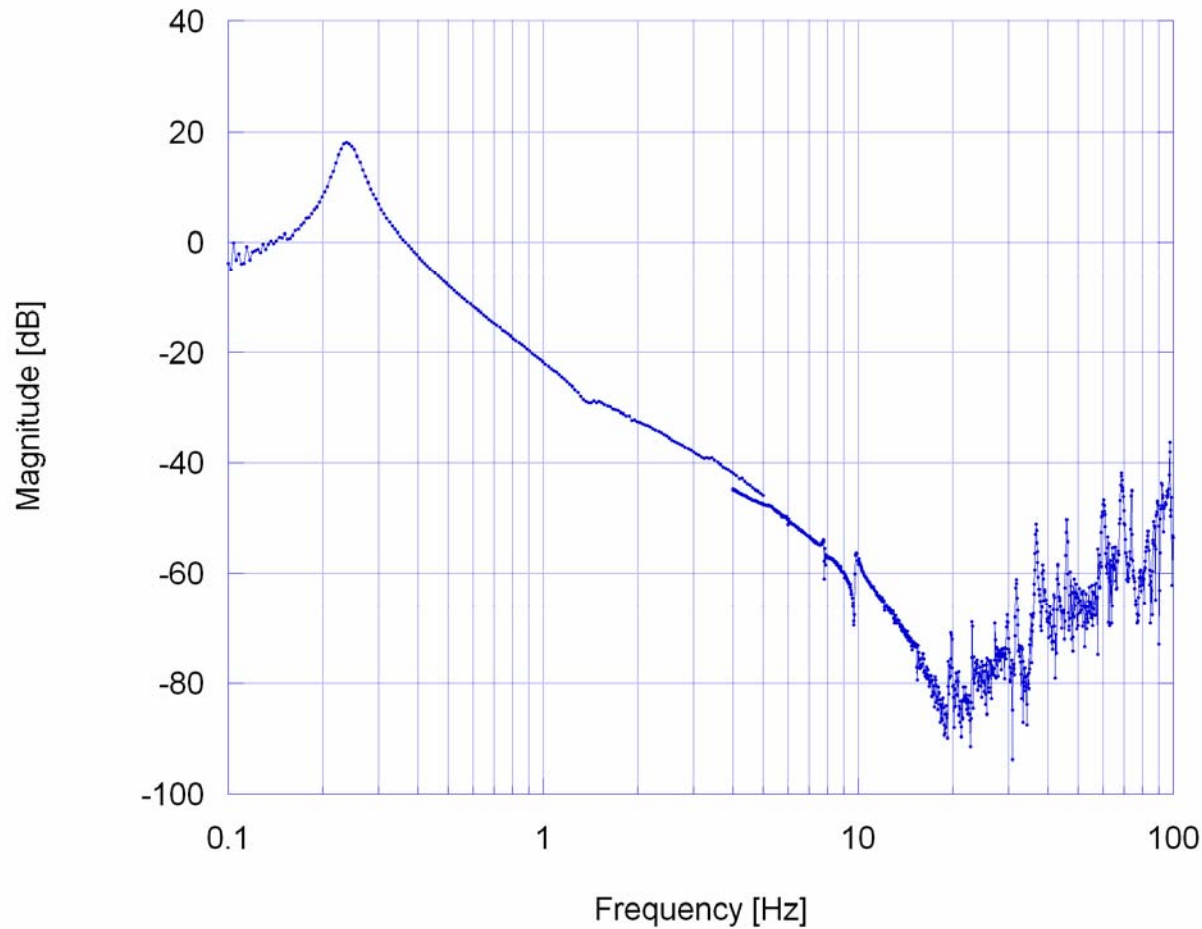


Phase



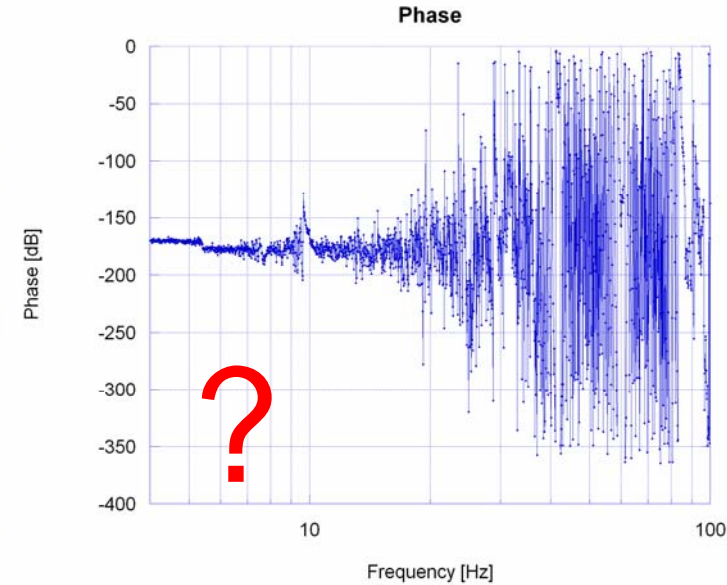
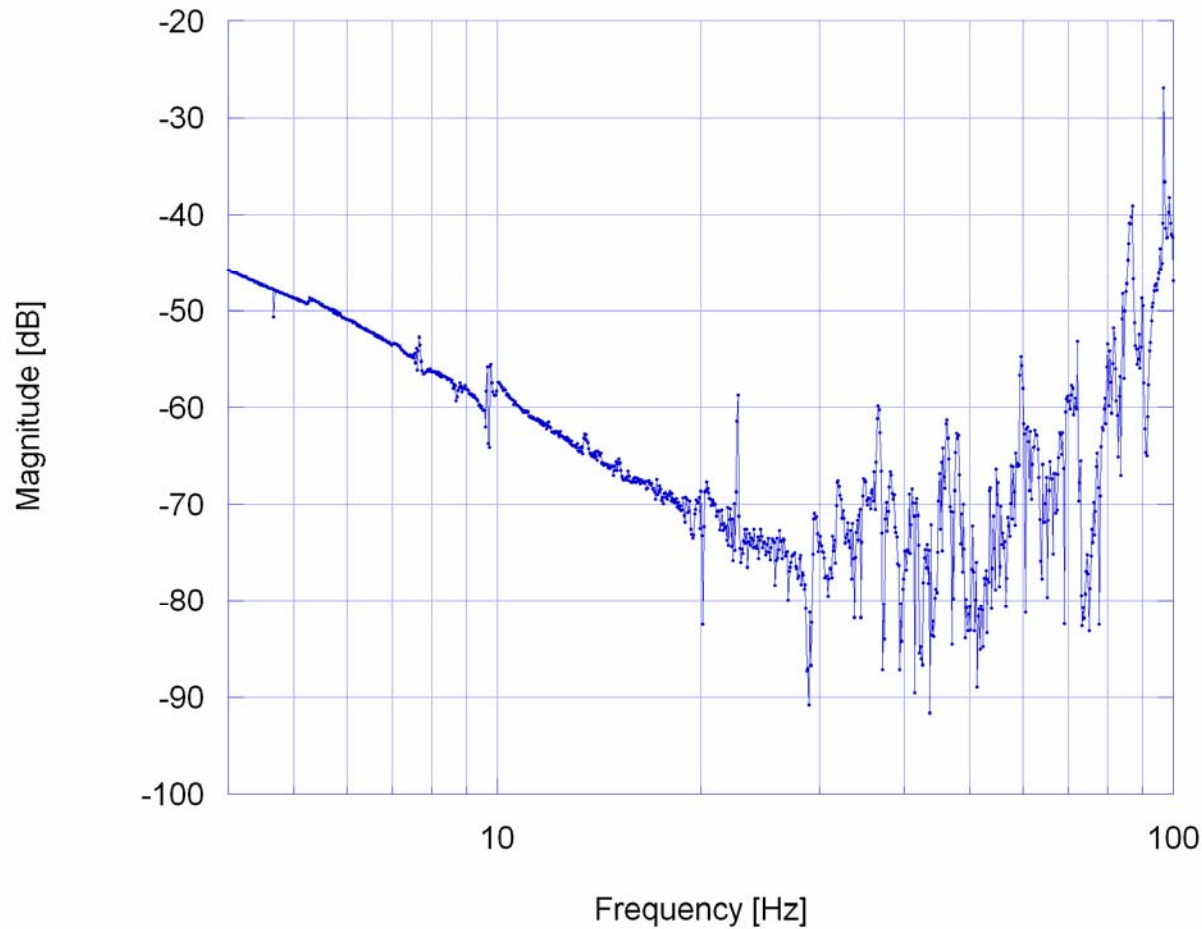
Two Booms + 1 CW

Two Booms + 1 CW outer (just in one)



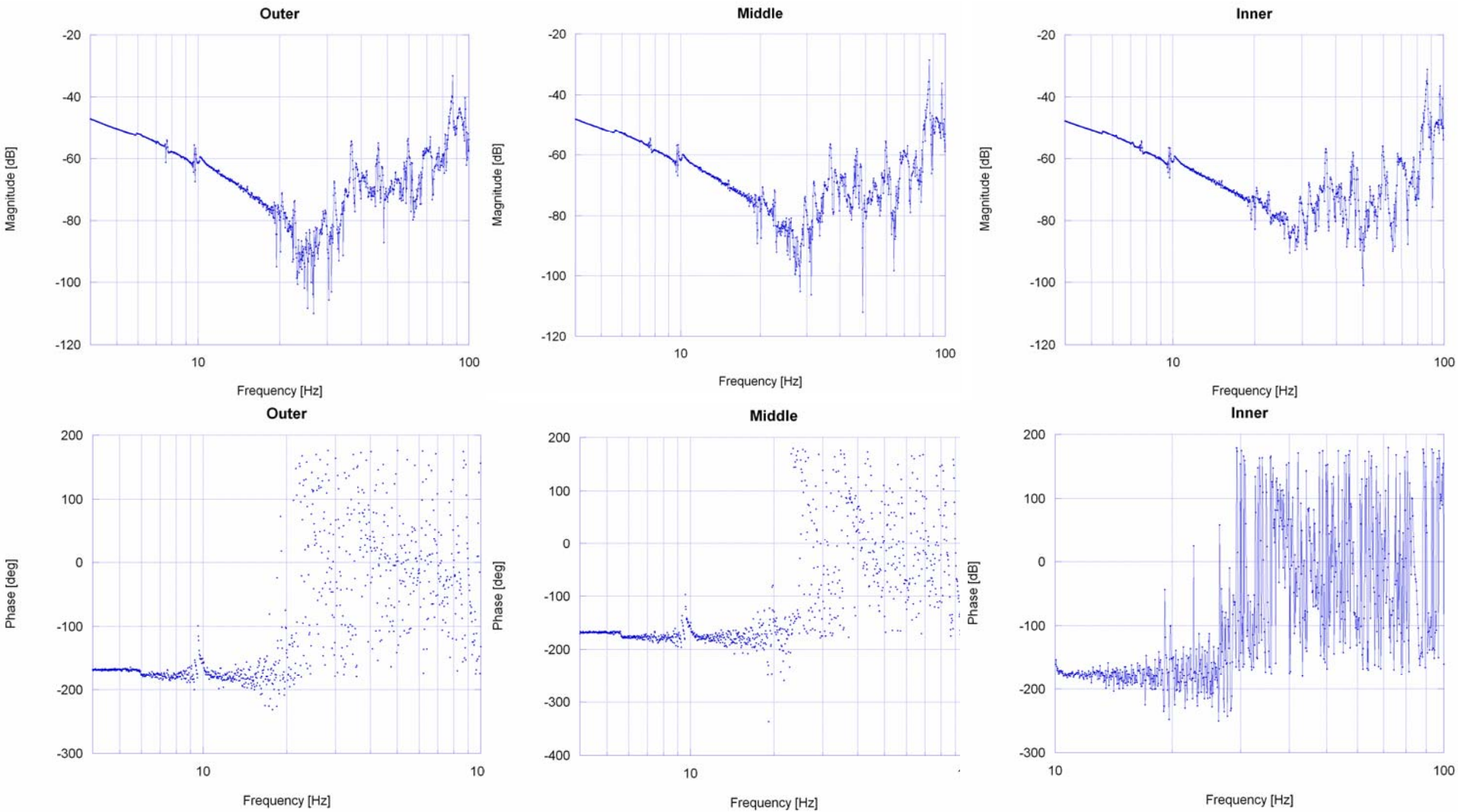
Two Booms + $\frac{1}{2}$ CW

Two Booms 1/2 CW (Bungee Cords)



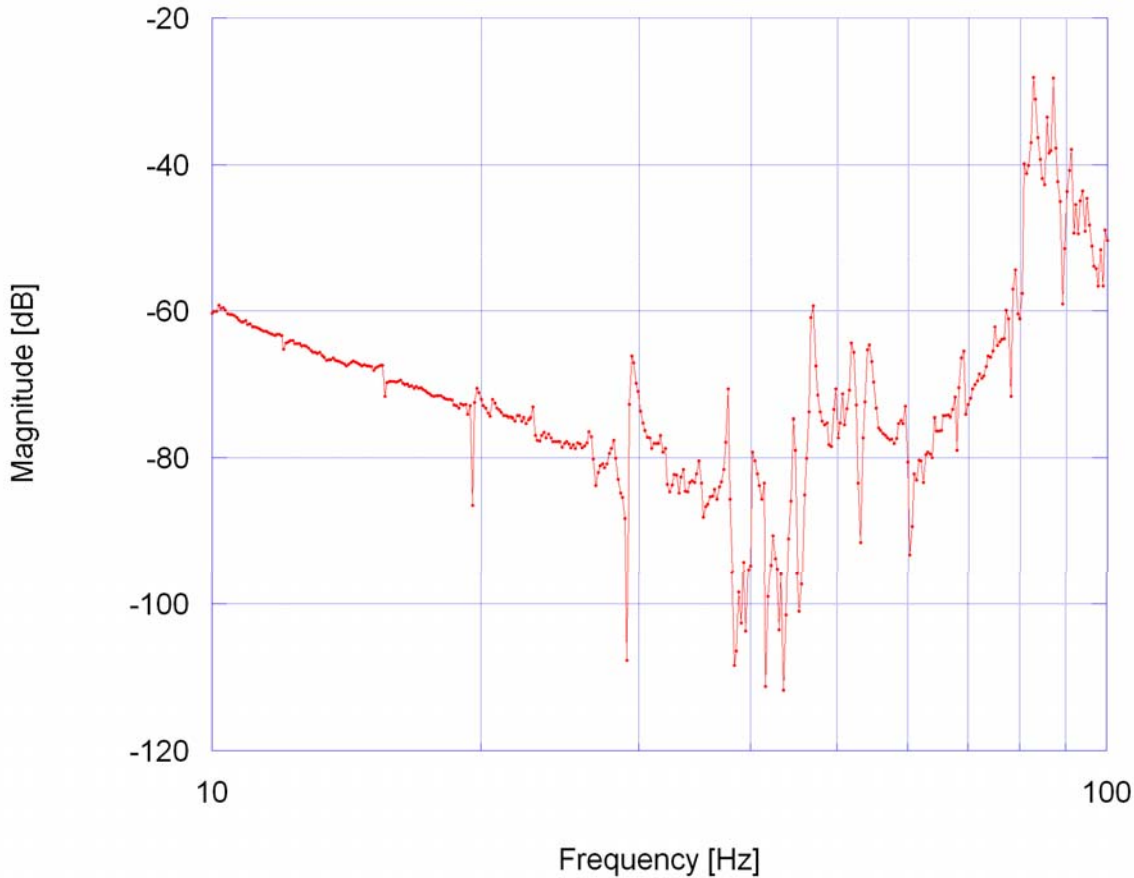


Two Booms 1 CW CW positions comparison

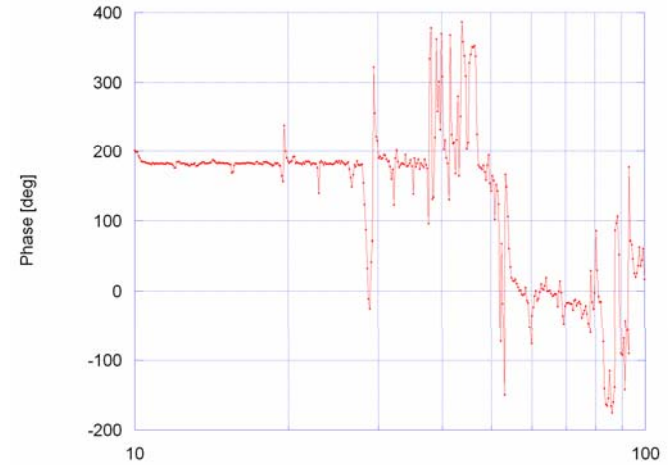


Two Booms 1 CW Inner

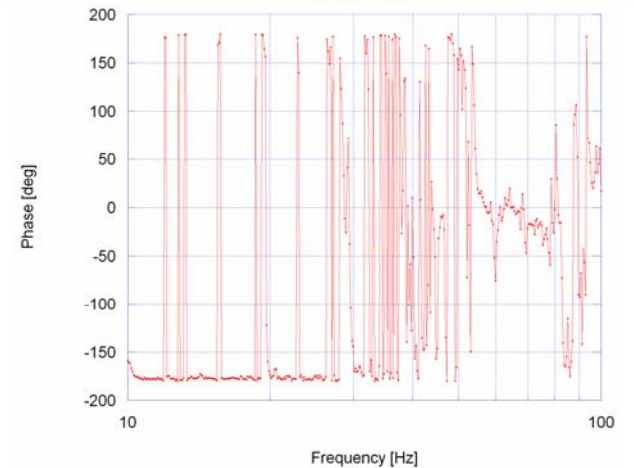
Transfer Function - Two Booms 1CW Inner



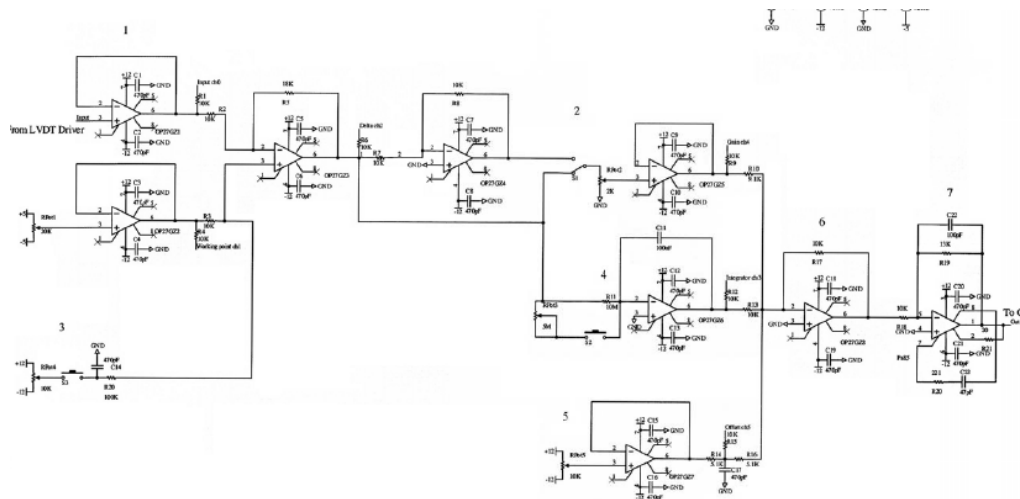
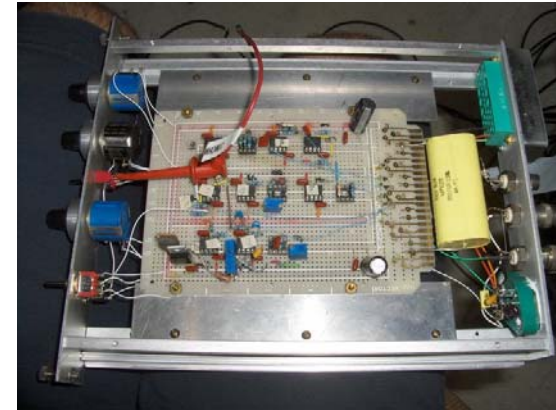
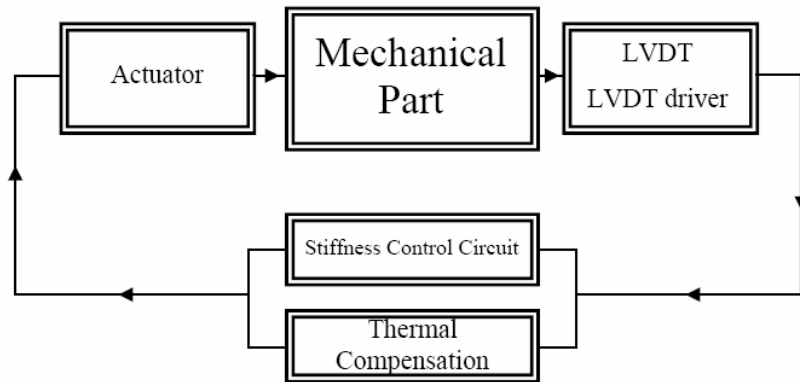
Phase Corrected



Raw Phase

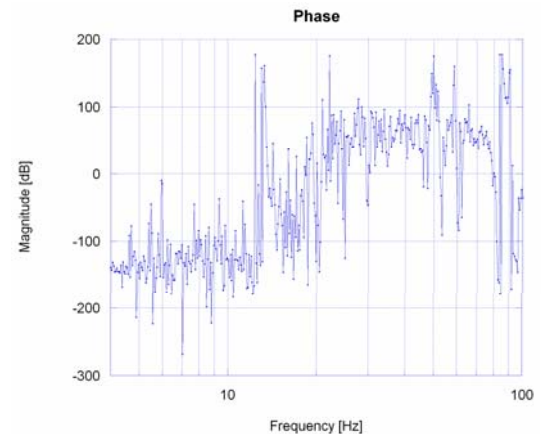
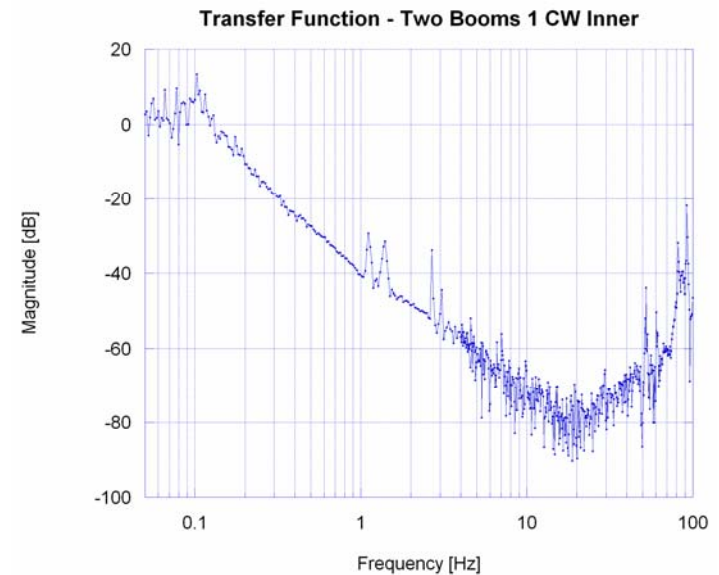
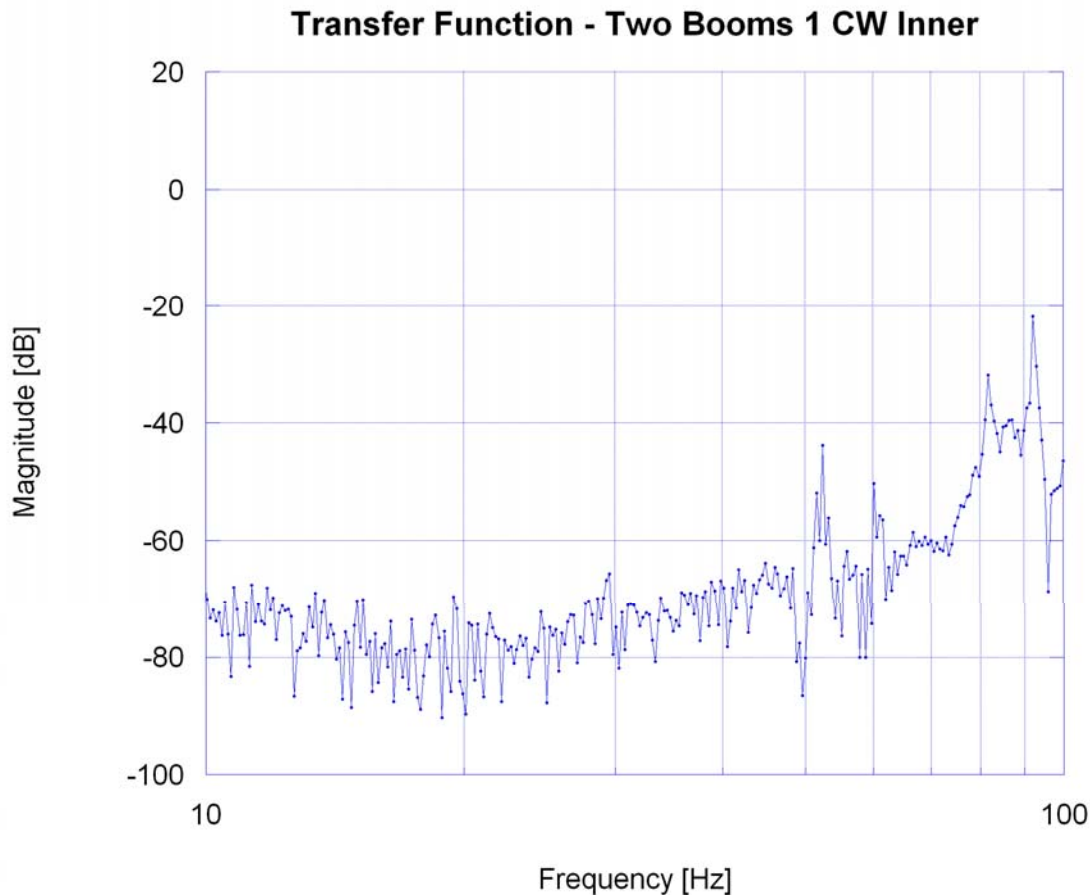


E.M. Antispring Circuitry



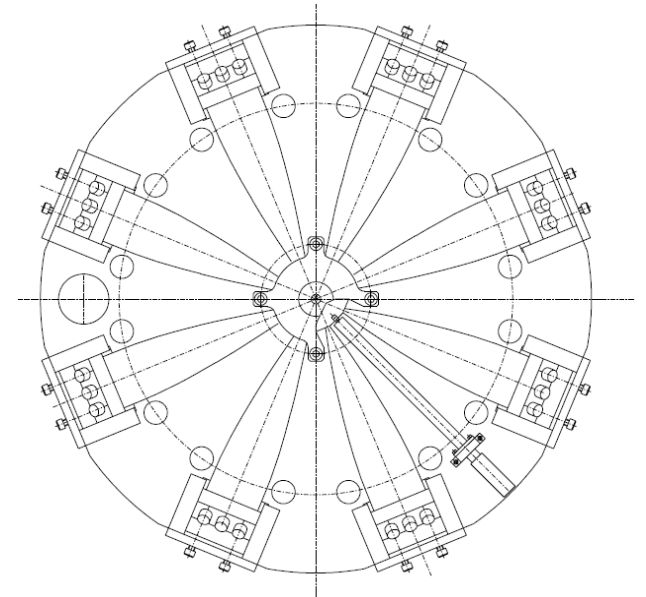
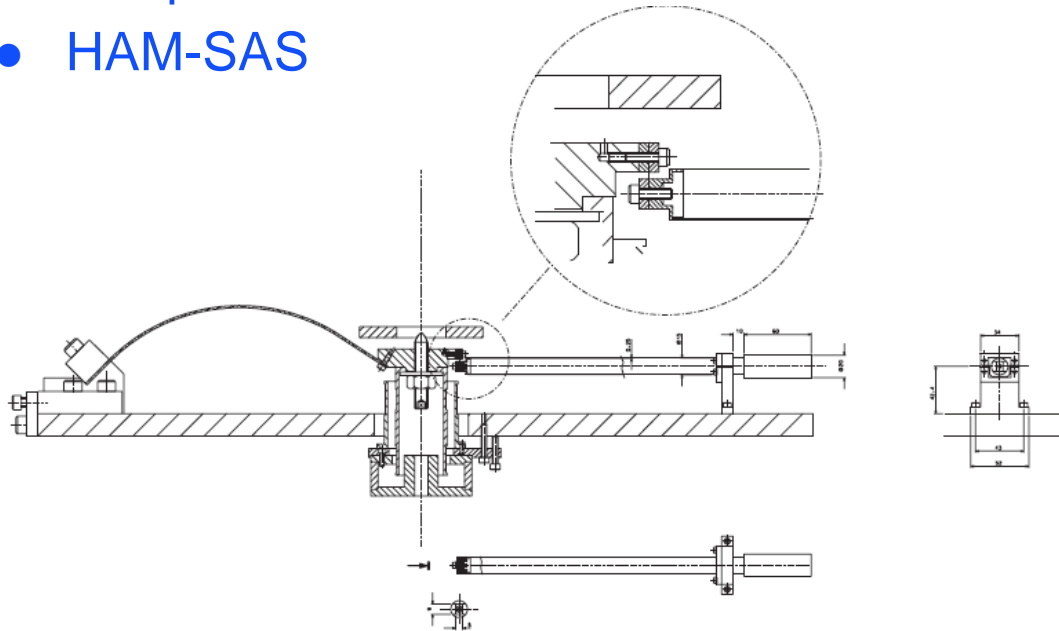
E.M. Antispring

Two Booms 1 CW Inner



Next

- Lower the noise at high frequencies
- Deepen the TF and Phase data analysis
- Find the optimal setup for the best compensation
- HAM-SAS



Acknowledgments

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