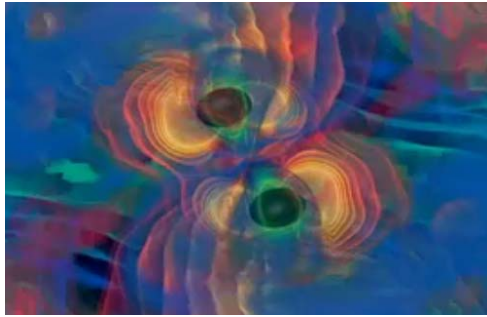


LIGO

LSC



Gravitational Waves: a new window to the Universe

Gabriela González

Physics and Astronomy, Louisiana State University

For the LIGO Scientific Collaboration

February 22, 2007



Annual
Conference

NATIONAL SOCIETY OF BLACK PHYSICISTS

NSBP

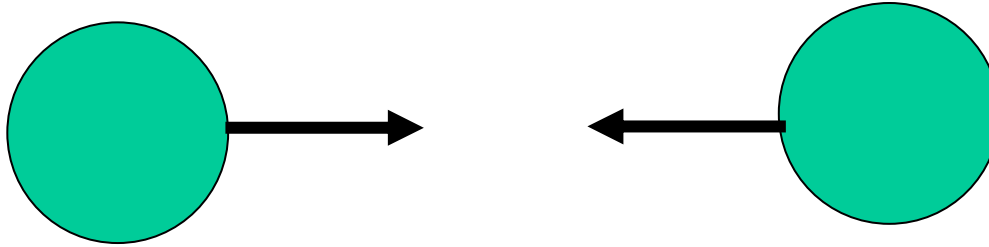


National Society of Hispanic Physicists

NSHP

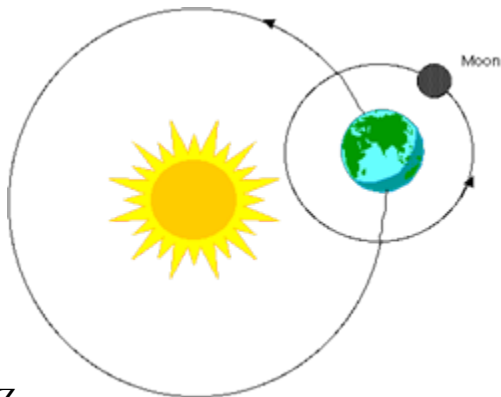


Newton's gravitation



Newtons' law: $F = Gm_1m_2/r^2$

Explains why things fall down,
and planetary motion.

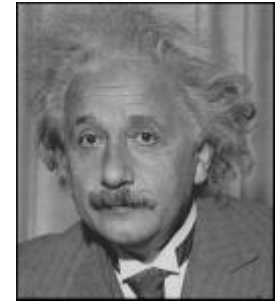


ckTime™ and a



Einstein's gravitation

When masses move, they wrinkle the space time fabric, making other masses move...

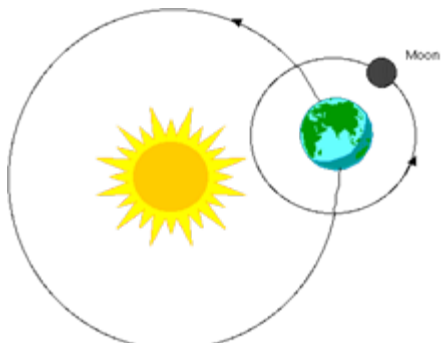


ckTime™ and a
↓

QuickTime™ and a
YUV420 codec decompressor
are needed to see this picture.

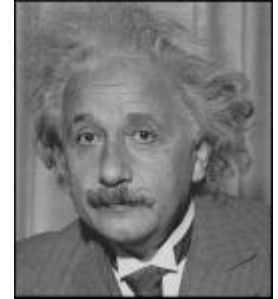
Explains just as well why things fall and planetary motion...

sciencebulletins.amnh.org

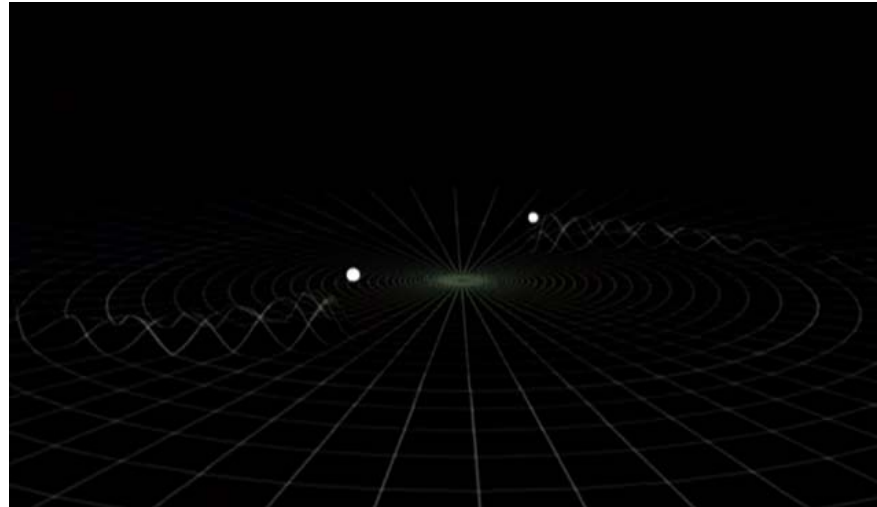


“Einstein’s” gravitation

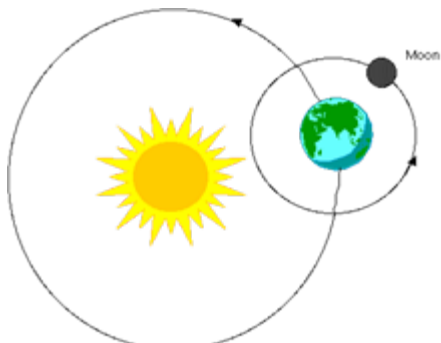
When masses move, they wrinkle the space time fabric, making other masses move...



ckTime^M and a
↓



Explains just as well why things fall and planetary motion...



Einstein’s messengers,
National Science Foundation video

.. but it also predicts **gravitational waves** traveling away from moving masses!



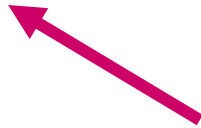
**Where do gravitational
waves come from?**

From stars living in galaxies...



LIGO

QuickTime™ and a
YUV420 codec decompressor
are needed to see this picture.



Supernova explosions

**Where do gravitational
waves come from?**

From stars living in galaxies...

LIGO

QuickTime™ and a
Animation decompressor
are needed to see this picture.



QuickTime™ and a
VU423 codec decompressor
are needed to see this picture.

Supernova explosions



Rotating stars (pulsars)

**Where do gravitational
waves come from?**

From stars living in galaxies...



LIGO

QuickTime™ and a
Animation decompressor
are needed to see this picture.

QuickTime™ and a
Animation decompressor
are needed to see this picture.



Supernova explosions

Rotating stars (pulsars)

**Where do gravitational
waves come from?**

QuickTime™ and a
Animation decompressor
are needed to see this picture.

QuickTime™ and a
YUV420 codec decompressor
are needed to see this picture.

Binary systems
(black holes, neutron stars)

Credit: John Rowe Animation/Australia Telescope National Facility, CSIRO

G070028-00-Z

QuickTime™ and a
Animation decompressor
are needed to see this picture.

QuickTime™ and a
Animation decompressor
are needed to see this picture.



Supernova explosions

Rotating stars (pulsars)

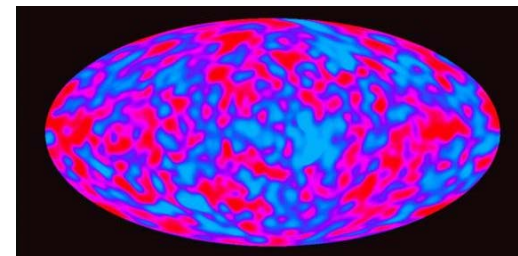
Where do gravitational waves come from?

QuickTime™ and a
Animation decompressor
are needed to see this picture.

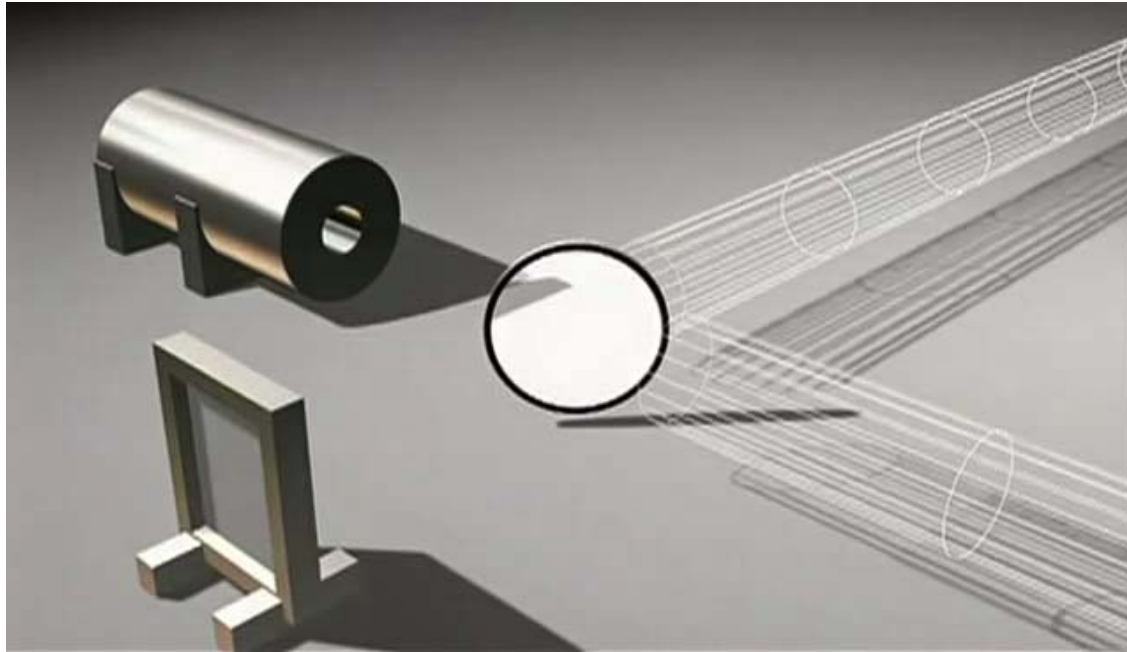
QuickTime™ and a
Animation decompressor
are needed to see this picture.

..and from the beginning of
the Universe!

Binary systems
(black holes, neutron stars)

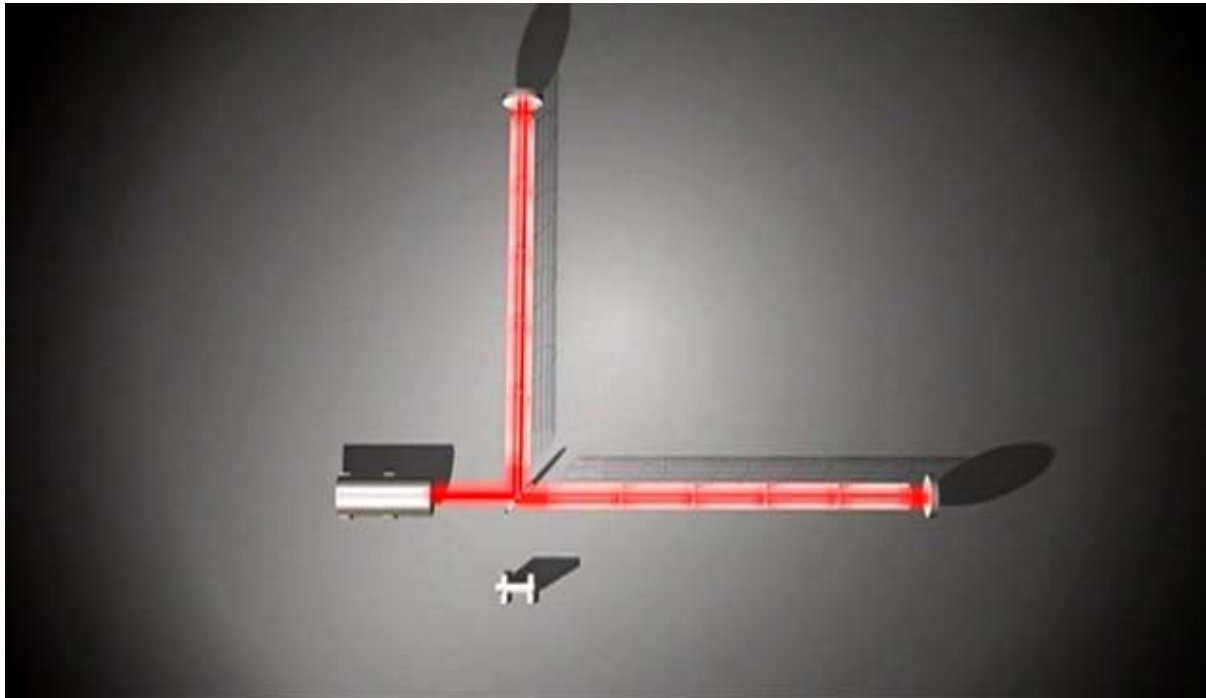


How to detect gravitational waves



Einstein's messengers,
National Science Foundation video

How to detect gravitational waves



Einstein's messengers,
National Science Foundation video

The LIGO logo features the word "LIGO" in a bold, black, sans-serif font. To the left of the text are several concentric, light gray circles of varying radii, suggesting a ripple or wave pattern.

LIGO



How to detect gravitational waves in Louisiana!

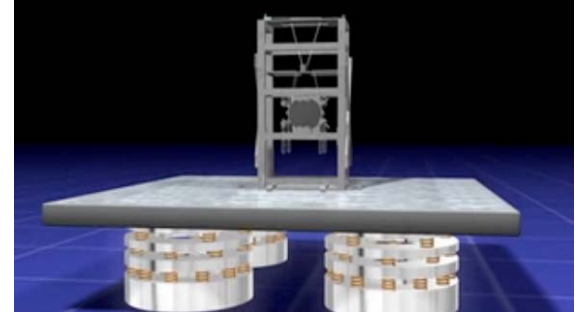
QuickTime™ and a
YUV420 codec decompressor
are needed to see this picture.

sciencebulletins.amnh.org

a difficult and fun experiment



QuickTime™ and a
Animation decompressor
are needed to see this picture.



QuickTime™ and a
Animation decompressor
are needed to see this picture.



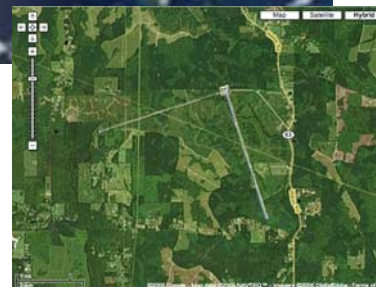
The LIGO project



Hanford, WA



Livingston, LA



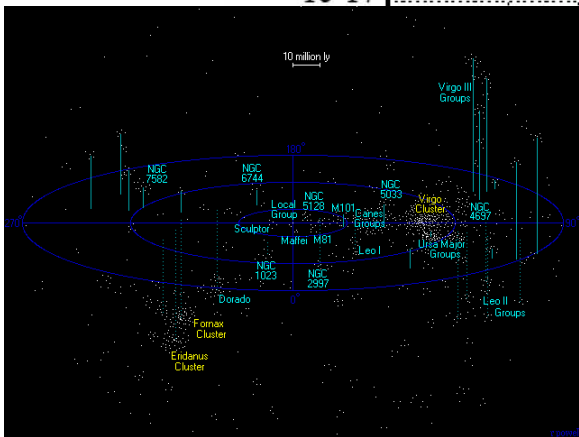
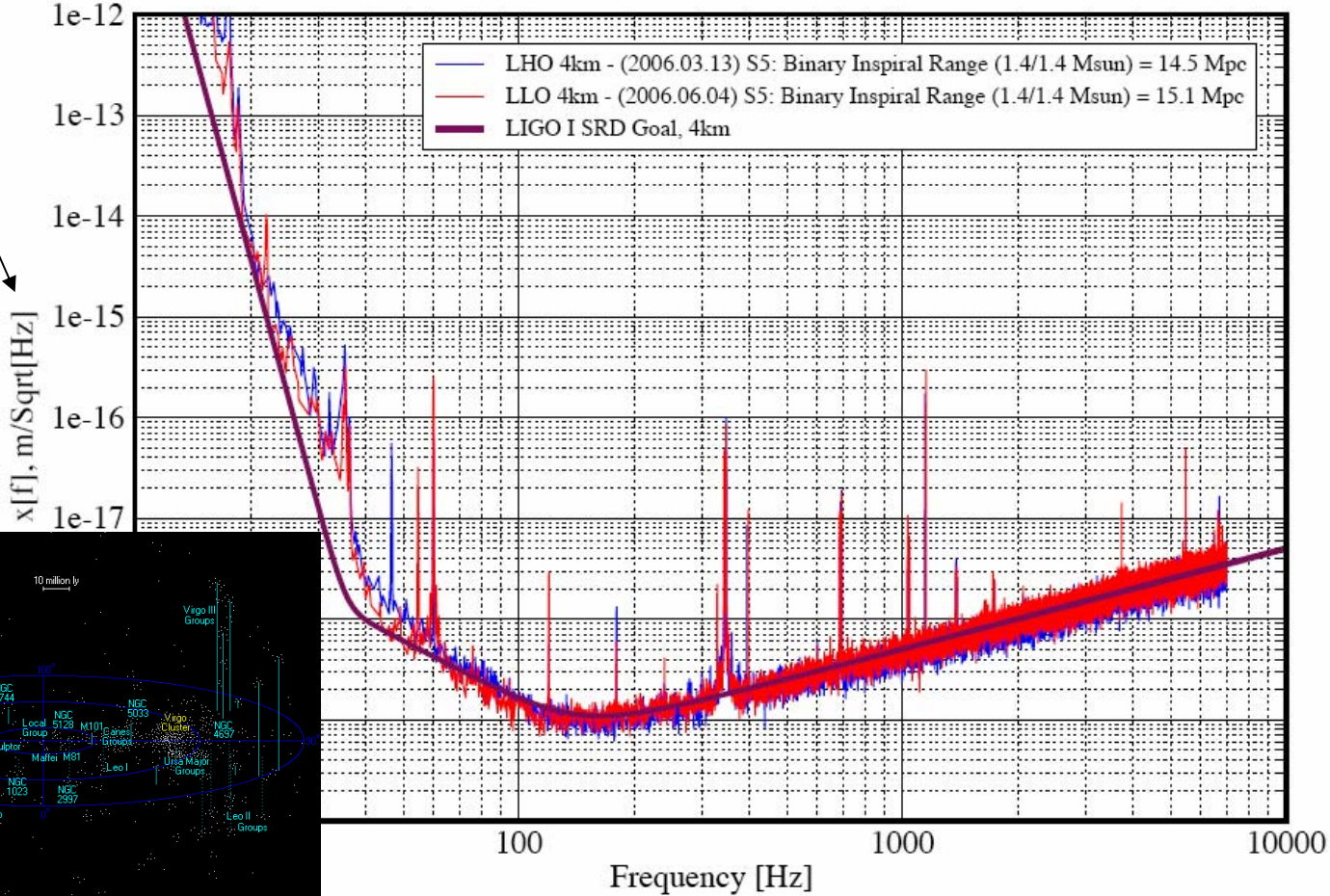
Hundreds of people working on the experiment and looking at the data:
LIGO Scientific Collaboration

Current sensitivity

Displacement Sensitivity for the LIGO 4km Interferometers

Performance for S5 - June 2006 LIGO-G060292-00-E

$x = L_x - L_y$
 $L_{x,y} \sim 4 \text{ km!}$



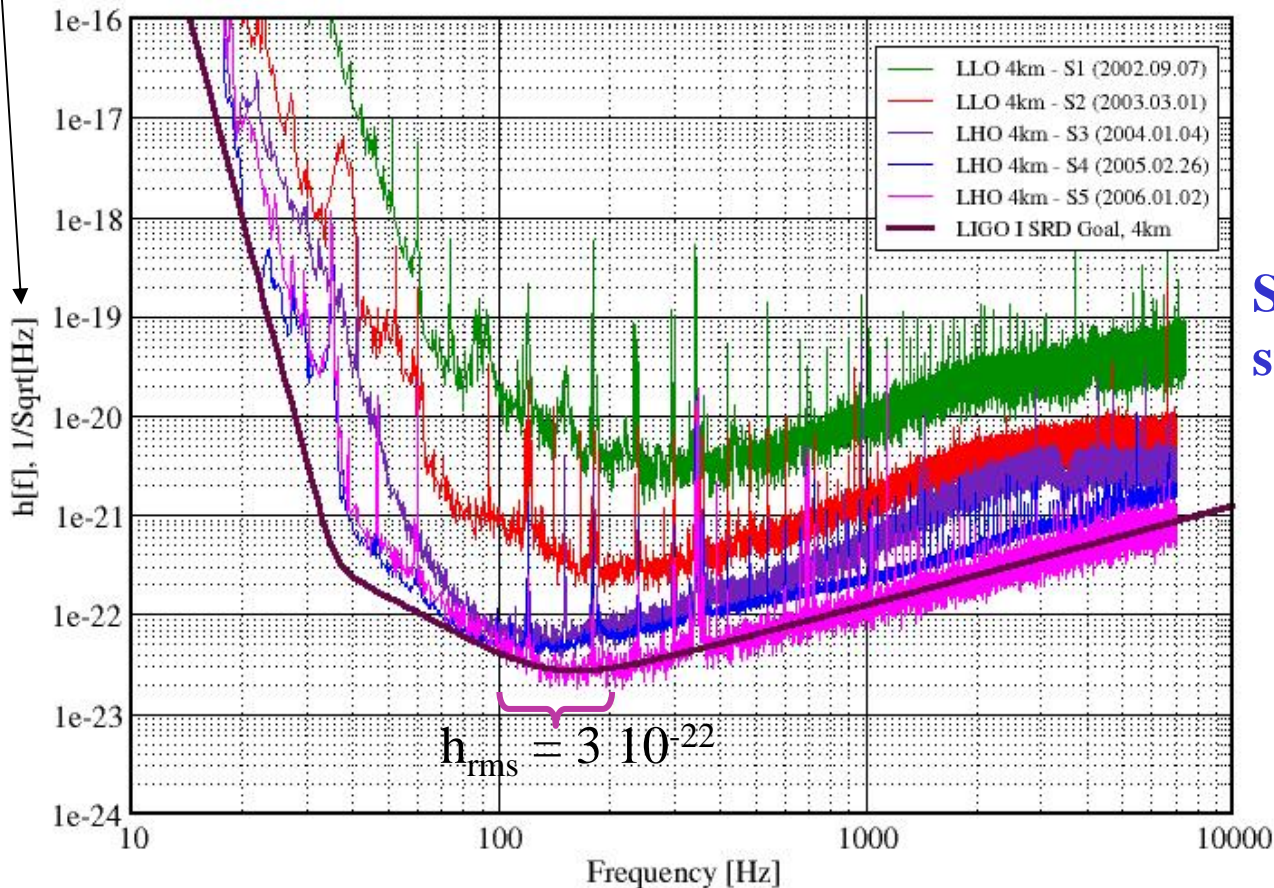
LIGO: Steady progress

$$x = \Delta L / L$$

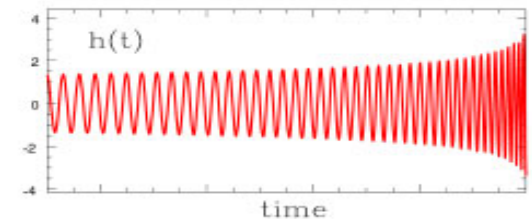
L = 4 km!

~~Best~~ Strain Sensitivities for the LIGO Interferometers

Comparisons among S1 - S5 Runs LIGO-G060009-01-Z

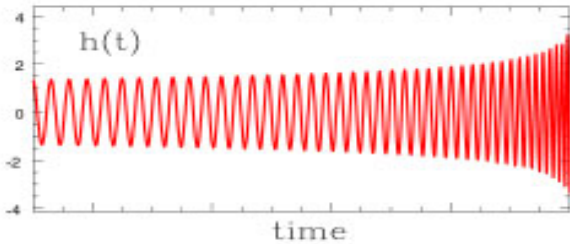
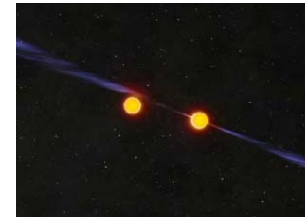


S5 started Nov 4, 2005, still ongoing!

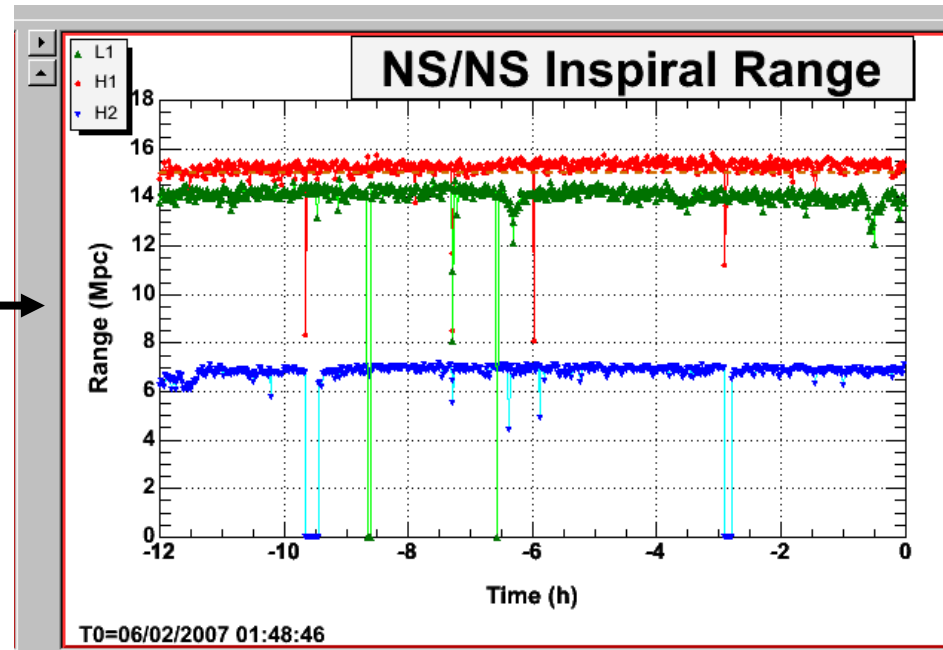
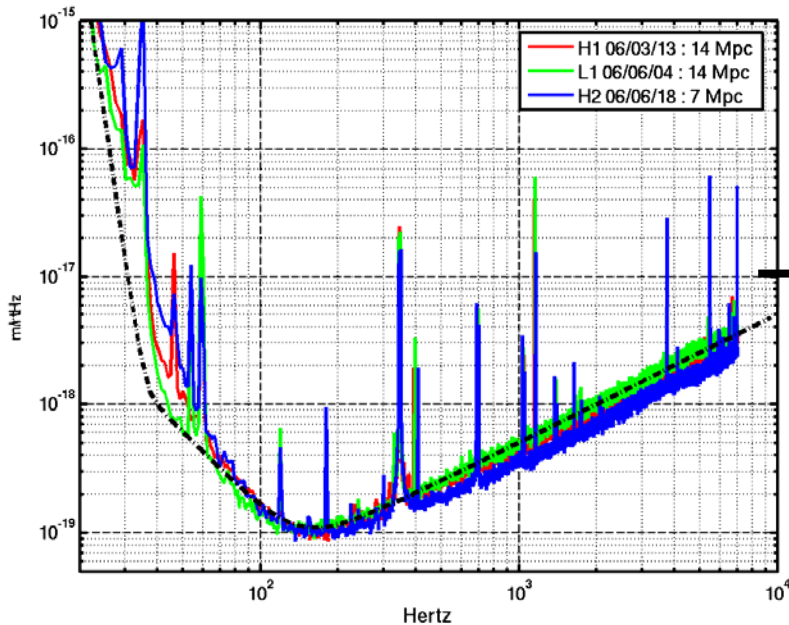


A binary neutron star system in the Virgo cluster (~15 Mpc away) produces GWs of amplitude $h \sim 10^{-21}$ on Earth, with freqs < 1400 Hz.

Looking for Binary systems



Can translate strain amplitude into (effective) distance

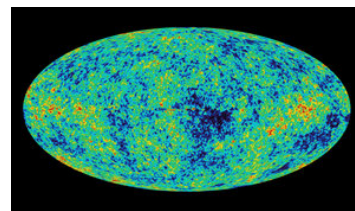
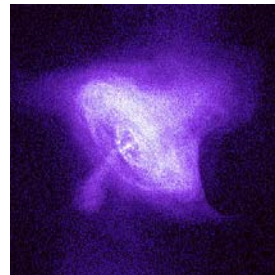
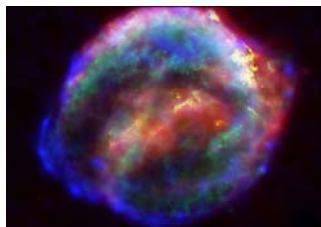


If system is optimally located and oriented, we can see even further: we are surveying hundreds of galaxies!

Electronic logs are public!
www.ligo.caltech.edu

Predictions are difficult... many unknowns!

- Rotating stars: how lumpy are they?
- Supernovae, gamma ray bursts: how strong are the waves (and what do they look like)?
- Cosmological background: how did the Universe evolve?
- Binary black holes: how many are there? What masses do they have?
- Binary neutron stars: from observed systems in our galaxy, predictions are up to 1/3yrs, but most likely one per 30 years, at LIGO's present sensitivity.
- From rate of short GRBs, much more optimistic predictions for BNS and BBH rates? Ready to be tested with S5!



- Neutron Star Binaries:**

Initial LIGO: $\sim 10\text{-}20$ Mpc \rightarrow

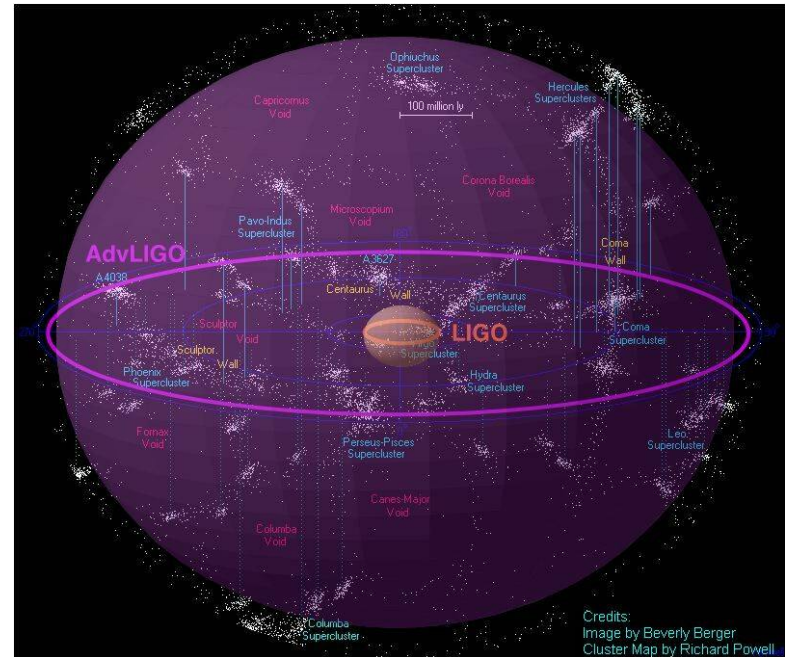
Advanced LIGO: $\sim 200\text{-}350$ Mpc

Most likely rate: 1 every 2 days !

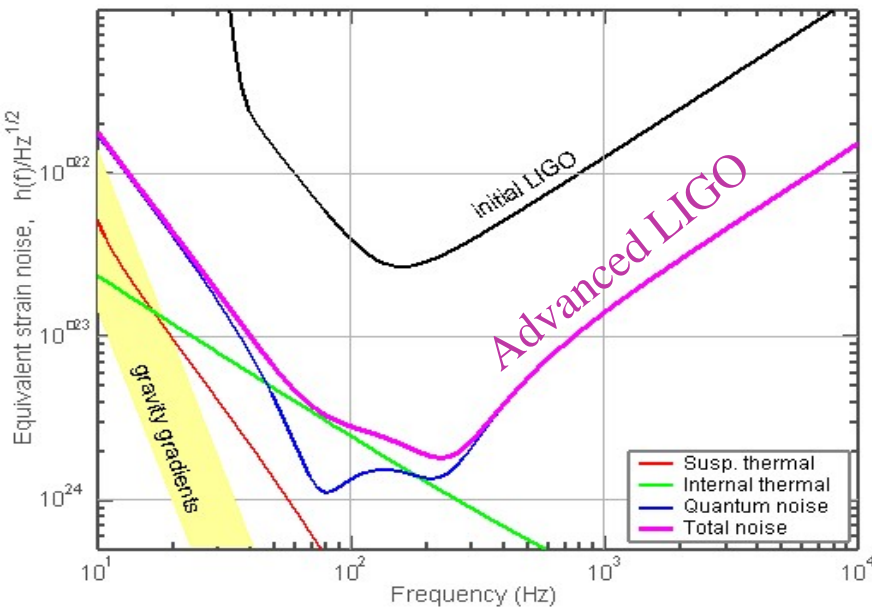
- Black hole Binaries:**

Up to $30 M_{\odot}$, at ~ 100 Mpc

\rightarrow up to $50 M_{\odot}$, in most of the observable Universe!



Credits: Image by Beverly Berger Cluster Map by Richard Powell

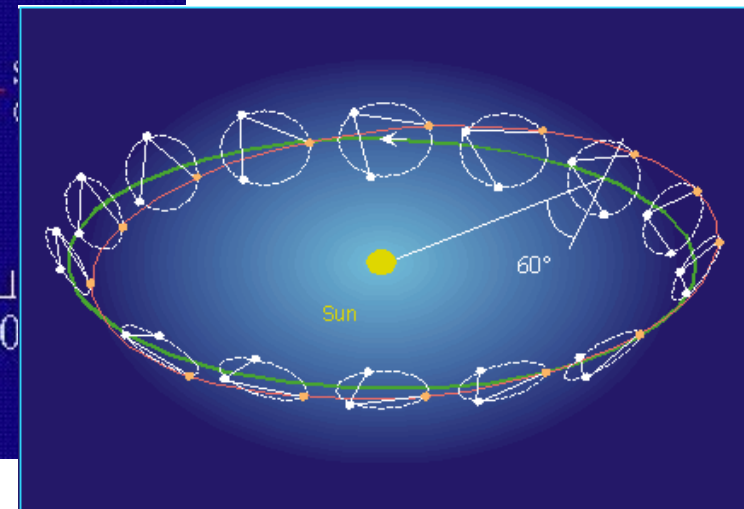
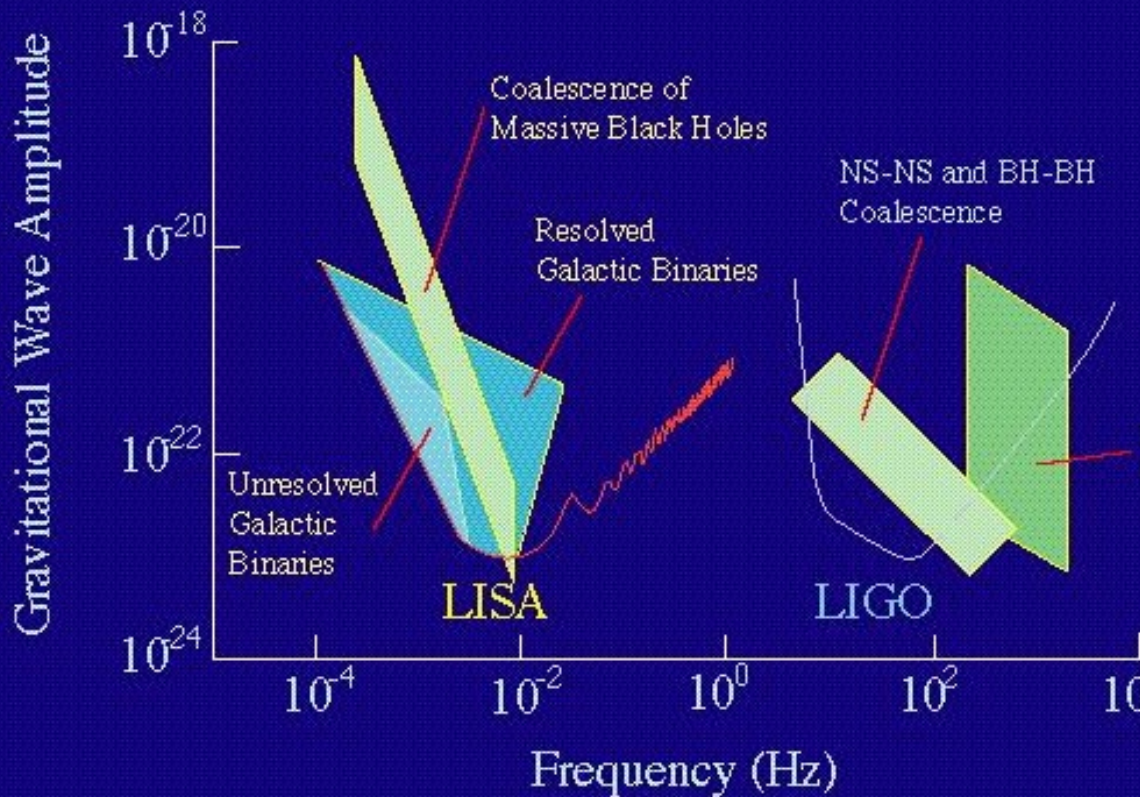


x10 better amplitude sensitivity
 \Rightarrow **x1000** rate= $(\text{reach})^3$
 \Rightarrow 1 year of Initial LIGO
 $<$ 1 day of Advanced LIGO !

Planned NSF Funding in FY'08 budget (being discussed right now!).

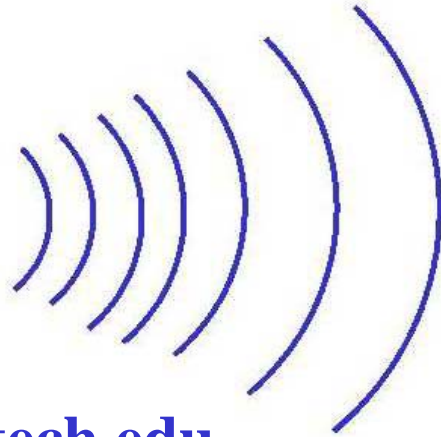
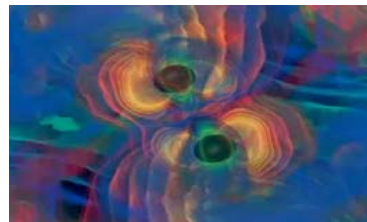
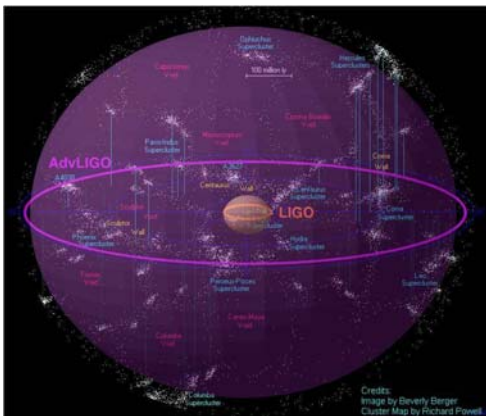
LIGO

Space interferometer: **LISA**



These are very exciting times!

- We are taking data at unprecedented sensitivity, and we are searching for gravitational waves.
- We are getting ready for Advanced LIGO.
- We are preparing ourselves for a direct observation of gravitational waves: *not if, but when!*
- LIGO detectors and their siblings will open a new window to the Universe: what's out there? **Join us to find out!**



www.ligo.caltech.edu
www.ligo.org