

High power optical components for Enhanced and Advanced LIGO

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- eLIGO phase modulator
- AdvLIGO Mach-Zehnder
- eLIGO/AdvLIGO Faraday isolator

- After S5 LIGO will be upgraded to eLIGO
- Laser power will be increased to 30 W
- Electro-optic modulators (EOMs) must be replaced.
 - LiNbO₃ modulators would suffer from severe thermal lensing or might even break
- Faraday isolators (FIs) must also be replaced
 - Absorption in the FI leads to thermal lensing, thermal birefringence, and beam steering
- eLIGO devices (techniques) will be used in AdvLIGO

- eLIGO EOMs

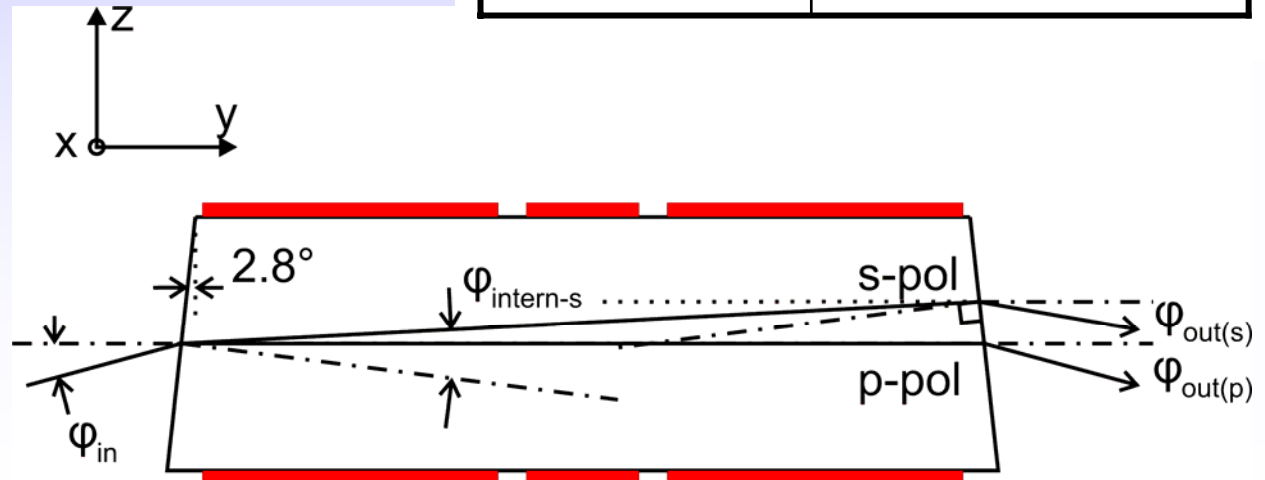
- Lithium niobate (LiNbO_3), used in initial LIGO, not satisfactory
 - Thermal lensing / Damage / Residual absorption
- Choose RTP (rubidium titanyl phosphate - RbTiOPO_4) as EO material
 - RTP has significantly lower absorption and therefore thermal lensing.
- Use custom made housing to separate the crystal housing from the housing for the resonant circuit.
Advantage: Resonant frequencies can be changed without disturbing the optical alignment.
- Use wedged crystals to reduce spurious amplitude modulation
Additional advantage: EOM acts as polarizer



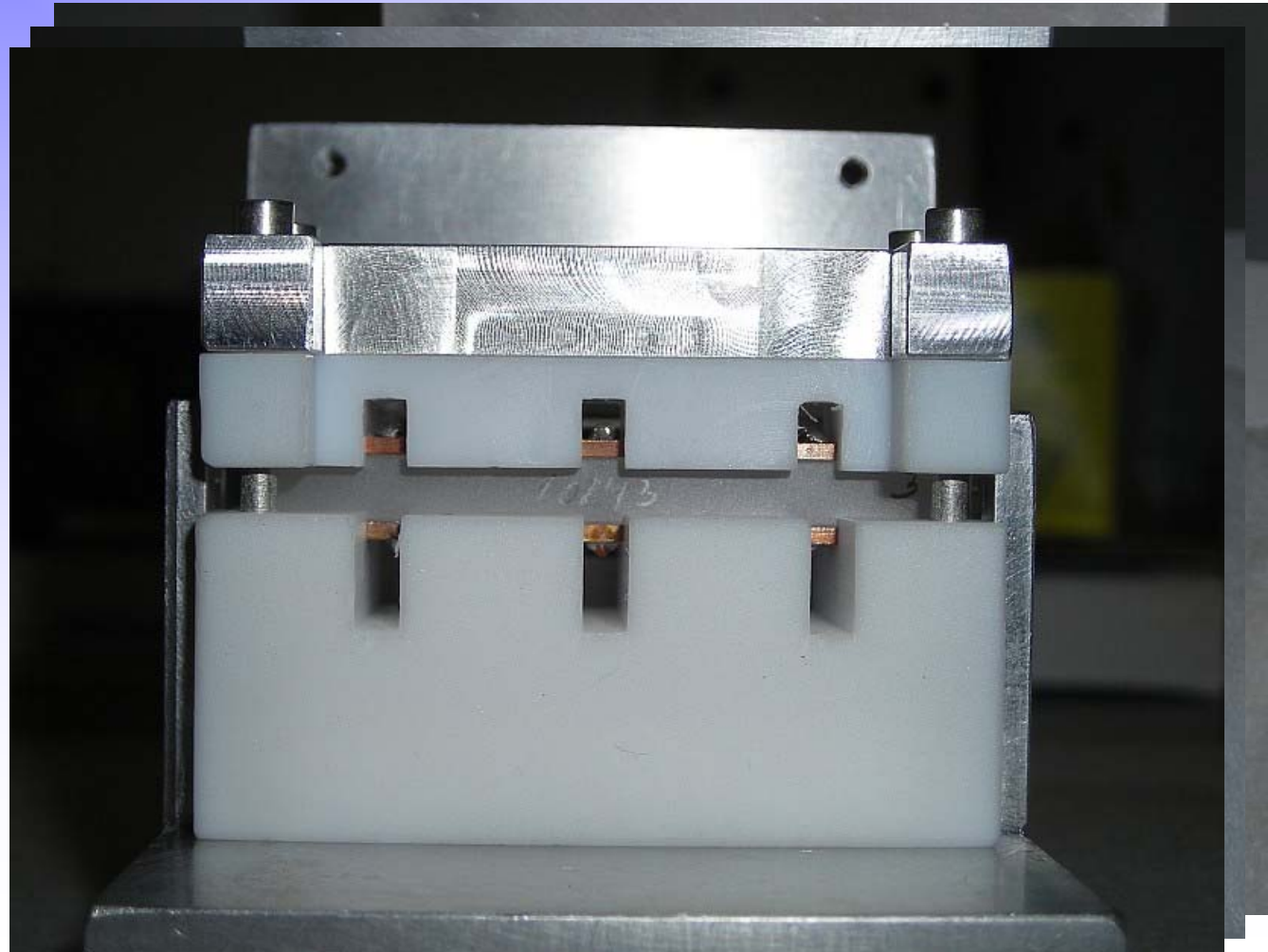
- AR coatings (< 0.1%) on crystal faces.

- Wedged crystal separates the polarizations and acts as a polarizer.
 - This avoids cavity effects and reduces amplitude modulation.

Polarization	Angle [degrees]
p	4.81
s	4.31



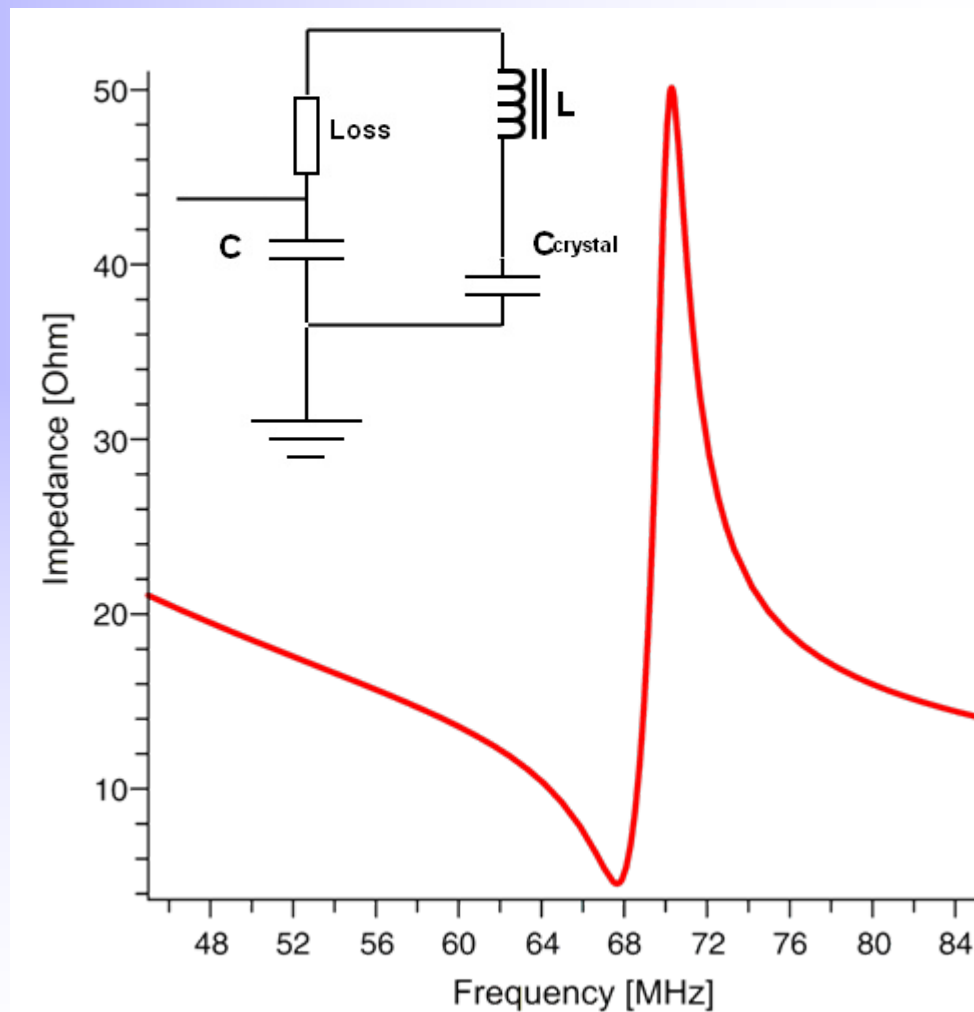
- Use one crystal but three separate pairs of electrodes to apply three different modulation frequencies at once.



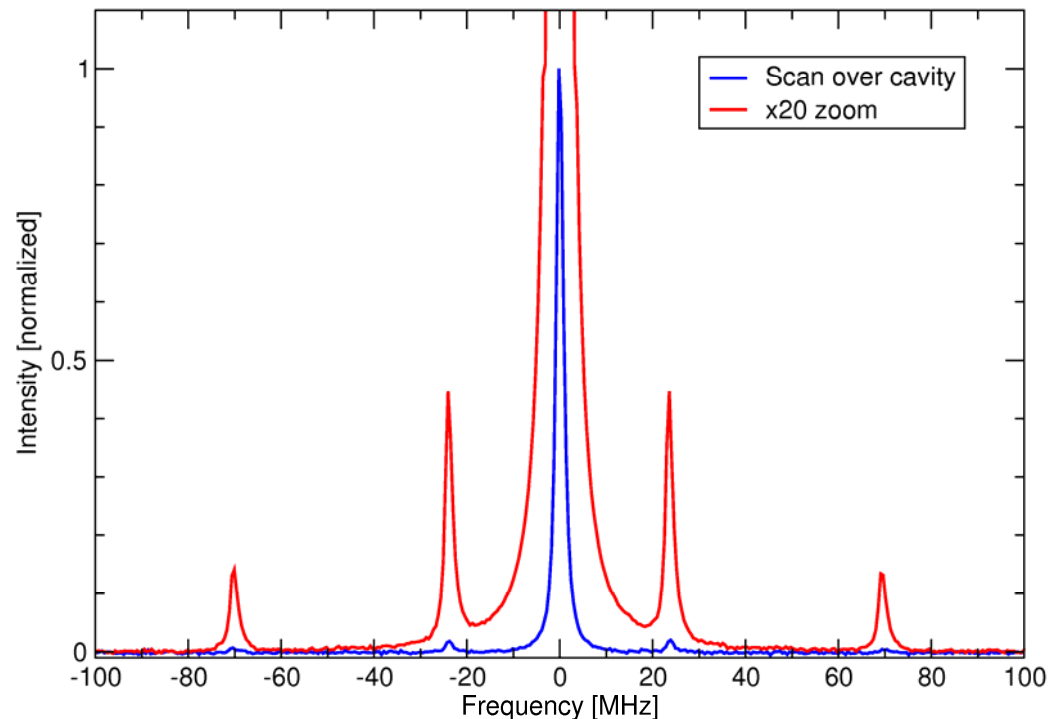
- Separate the crystal housing from the housing of the electronic circuits to maintain maximum flexibility.



- Impedance matching circuit in separate housing.
- Resonant circuit with $50\ \Omega$ input impedance.
- Current prototype has two resonant circuits:
 - 23.5 MHz and 70 MHz



- Sideband measurement with 10 V_{pp} drive into 23.5 MHz and 70 MHz input.
 - $m_{23.5} = 0.29$
 - $m_{70} = 0.17$

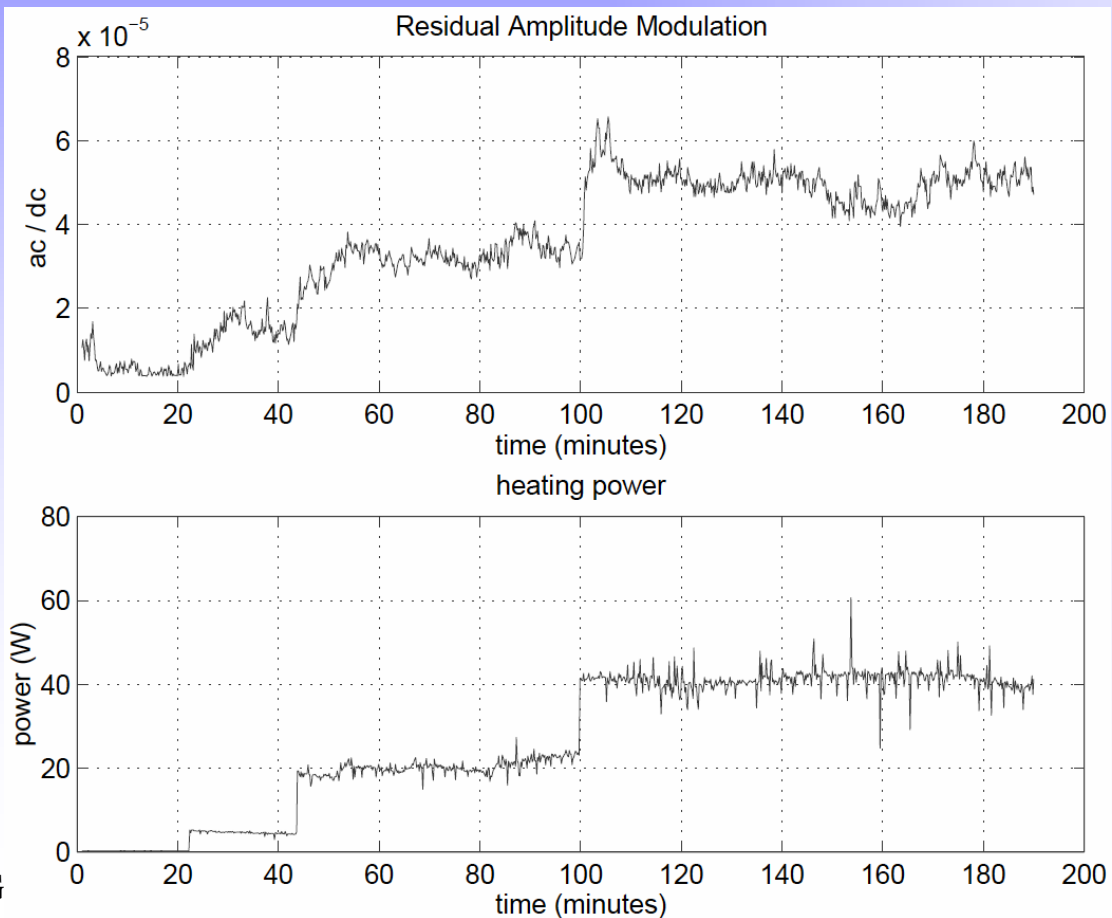


- Use a YLF laser was used to measure the thermal lensing.
 - Full Power = 42 W
 - Beam Waist = 0.5 mm (at RTP)
 - 4x4x40 mm RTP crystal

Axis	Focal length
X-axis	3.8 m
Y-axis	4.8 m

- compare with LiNbO3 (20 mm long):
 $f_{\text{thermal}} \sim 3.3 \text{ m @ } 10 \text{ W}$

- Measurement of RFAM, for RTP crystal with parallel faces (**previous** prototype @19.7 MHz, comparable with current LiNbO3 EOMs but better thermal properties)



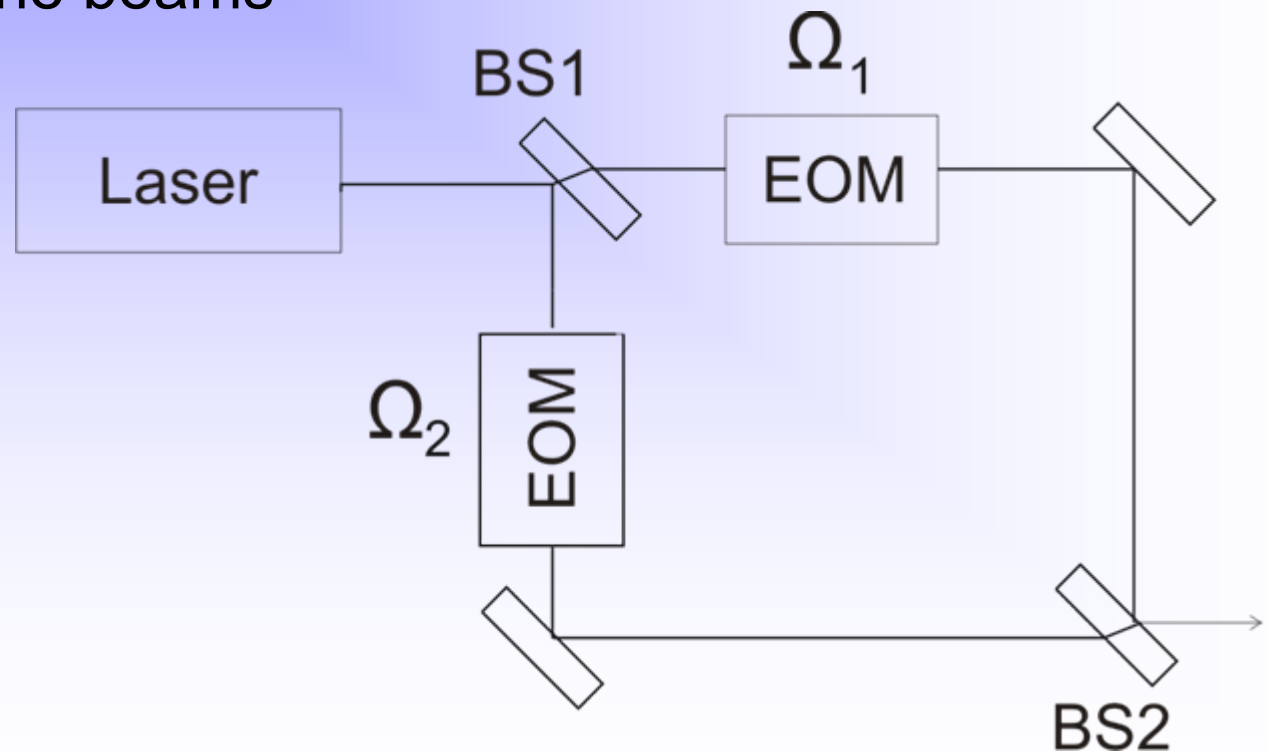
- Preliminary result for the **new** prototype:
 $\Delta I/I < 10^{-5}$ at Ω_{mod}
 $\Omega_{mod} = 25$ MHz
 $m = 0.17$



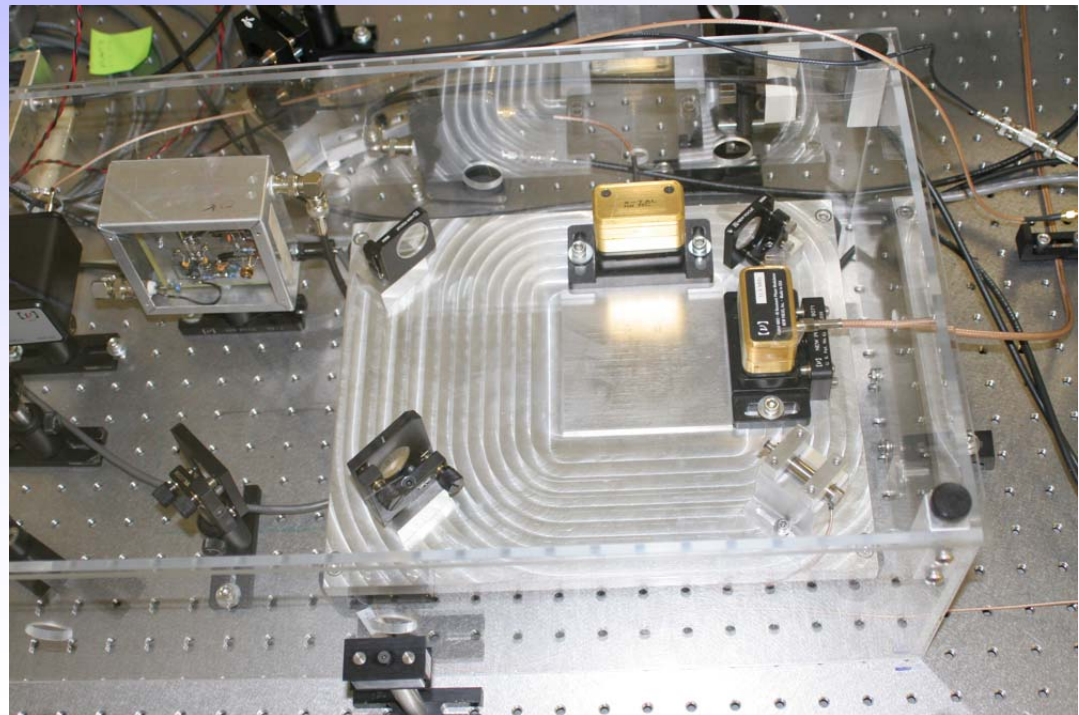
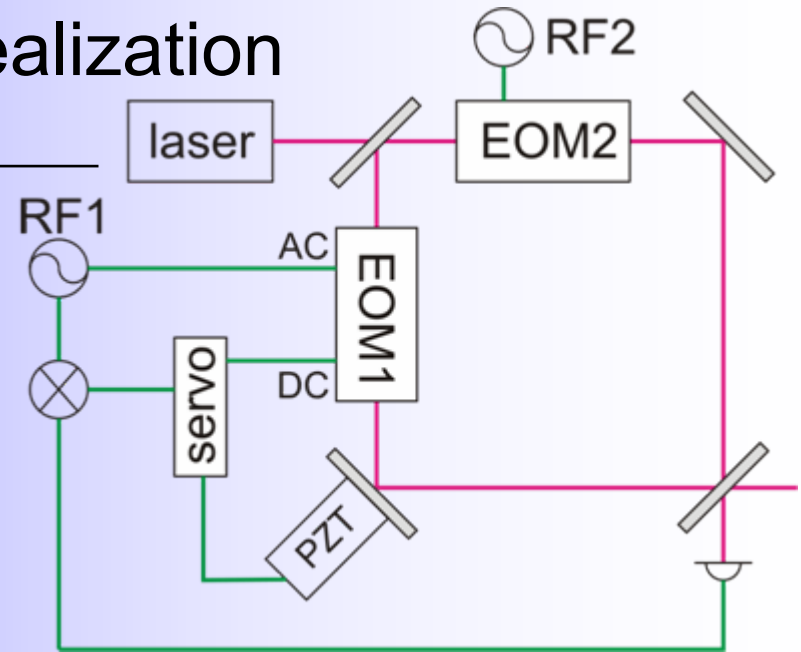
LIGO AdvLIGO Mach-Zehnder (parallel) modulation

- Not really a high power issue, but needs to be addressed also.
- Objective:
 - Solve the sidebands on sidebands problem by using parallel modulation.
 - Currently used in the 40m prototype
- Problems:
 - Sideband power reduced by a factor of 4
 - Additional intensity noise at modulation and mixing (sum/difference) frequencies
 - Excess intensity, frequency and sideband noise is possible depending on the stability of the MZ and the corner frequencies of the MZ stabilization loop.
- Only address the last point for now ..

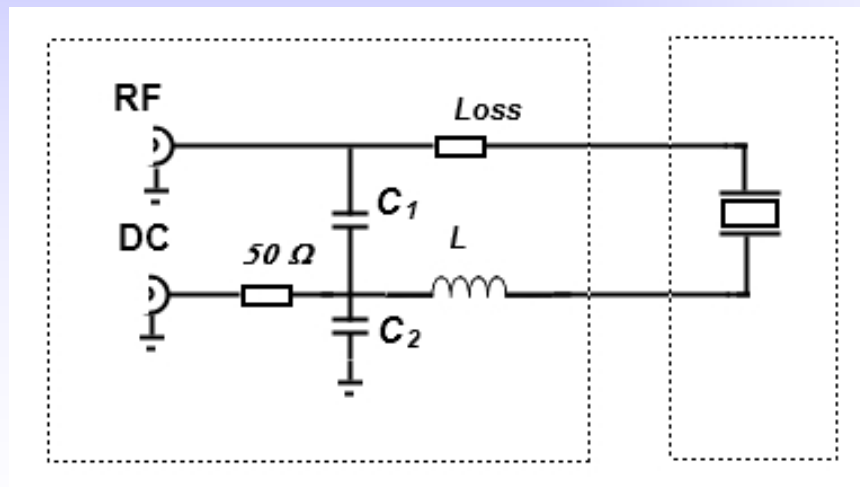
- Parallel modulation with two modulation frequencies
- Avoid the sideband-on-sideband problem by separating the beams

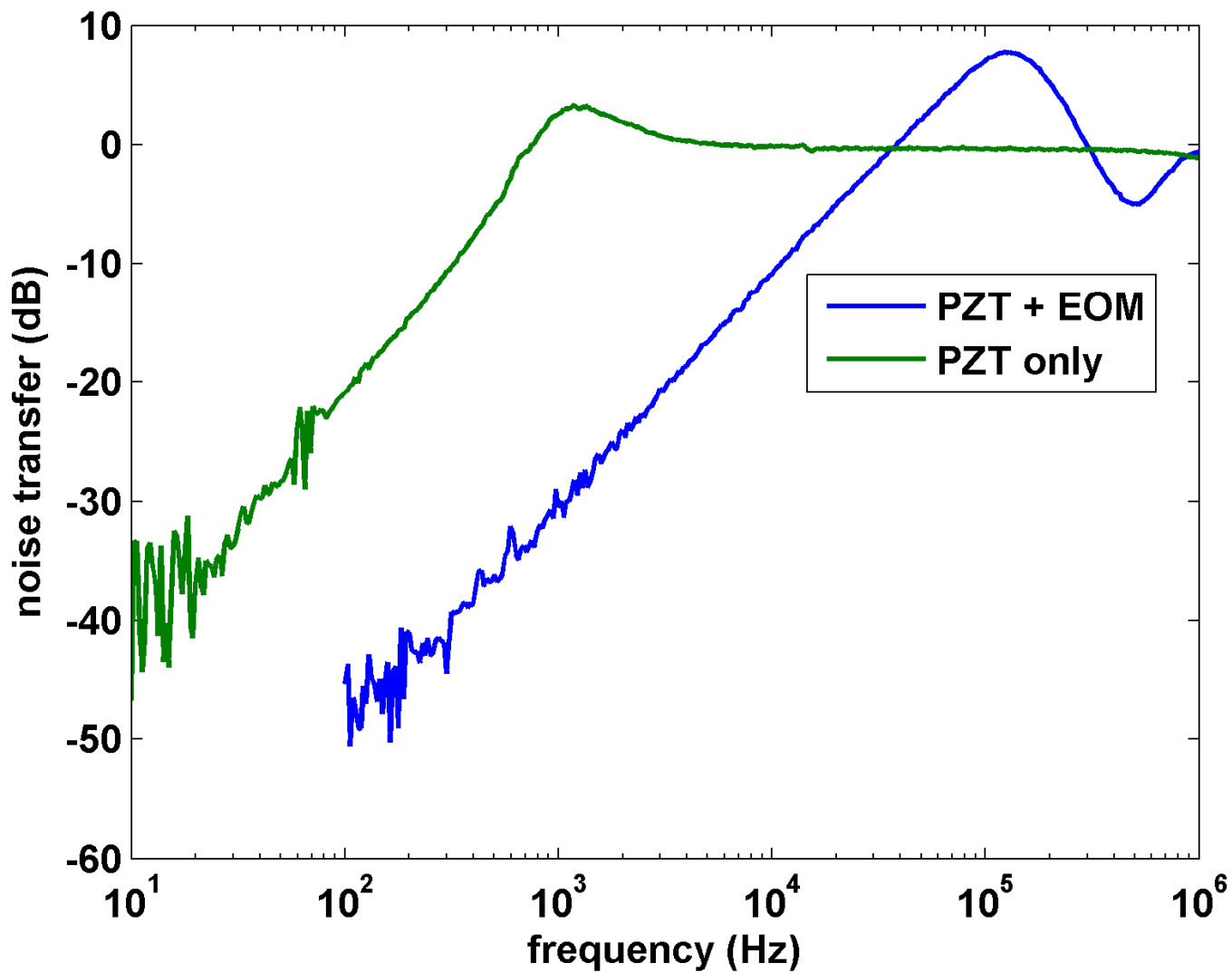


- Slow length control with big dynamic range with PZT
- Fast phase control with phase correcting EOM
- Stable mechanical “quasi-monolithic” design
- Reduce environmental effects with a Plexiglas enclosure.
- Modulation at 25 MHz and 31.5 MHz



- To realize the fast phase correcting without using an additional EOM a slightly modified resonant circuit was used.
 - Simultaneous modulation at resonant frequency
 - DC phase changes up to 1 MHz possible



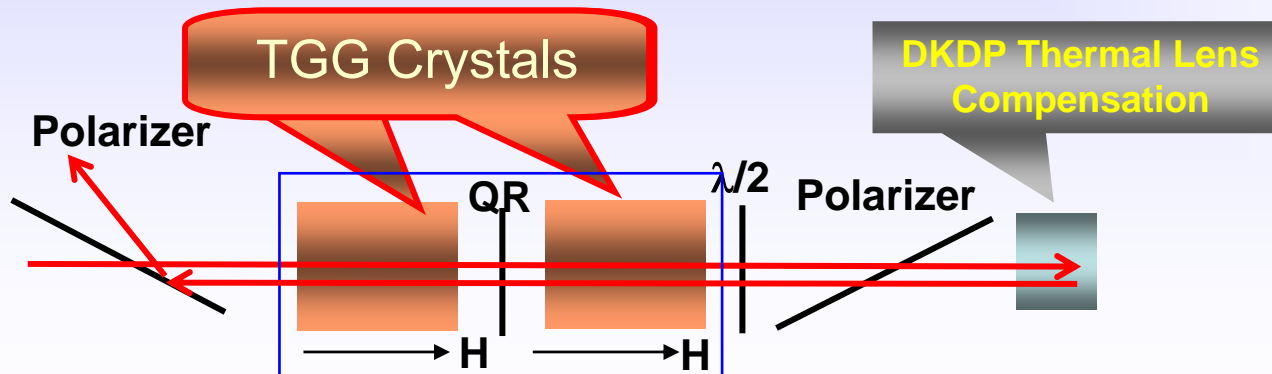


- Low noise performance of the PZT control (driven directly out of an OpAmp provides $\sim 4 \mu\text{m}$ dynamic range)
- Fast phase correcting EOM currently limits the unity-gain frequency to 50 kHz but is only limited by the current servo electronics

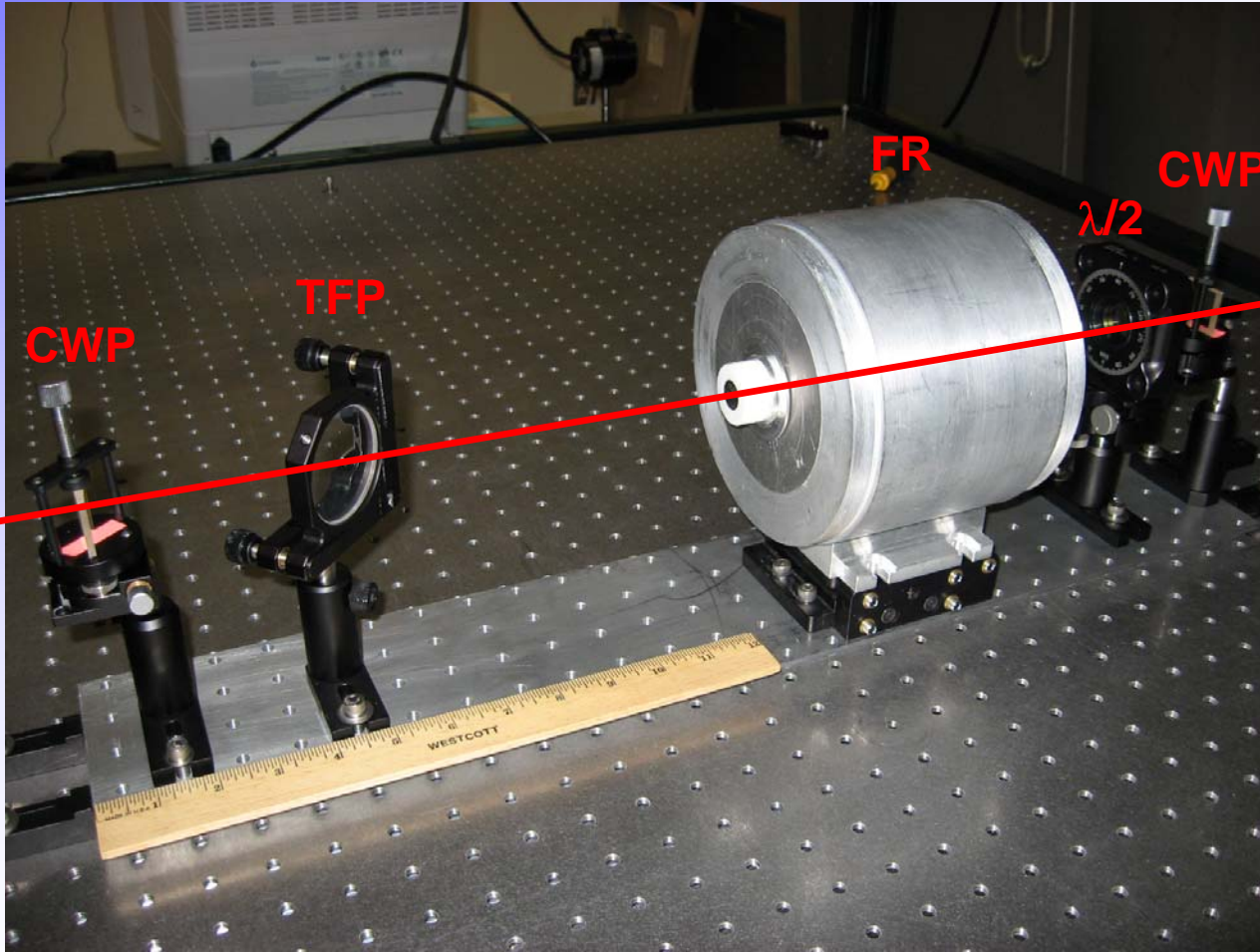
- Objective:
 - Strong suppression of back reflected light.
 - eLIGO ~ 30 W
 - AdvLIGO ~ 130 W
 - Minimal thermal lensing
 - Minimal thermal beam steering

Designed and parts supplied by IAP/UF

- Faraday rotator (FR)
 - Two 22.5° TGG-based rotators with a reciprocal 67.5° quartz rotator between
 - Polarization distortions from the first rotator compensated in the second.
 - $\frac{1}{2}$ waveplate to set output polarization.
 - Thermal lens compensation *via* negative dn/dT material: deuterated potassium dihydrogen phosphate, KD_2PO_4 , or ‘DKDP’).
- Calcite wedges or TFP polarizers are possible

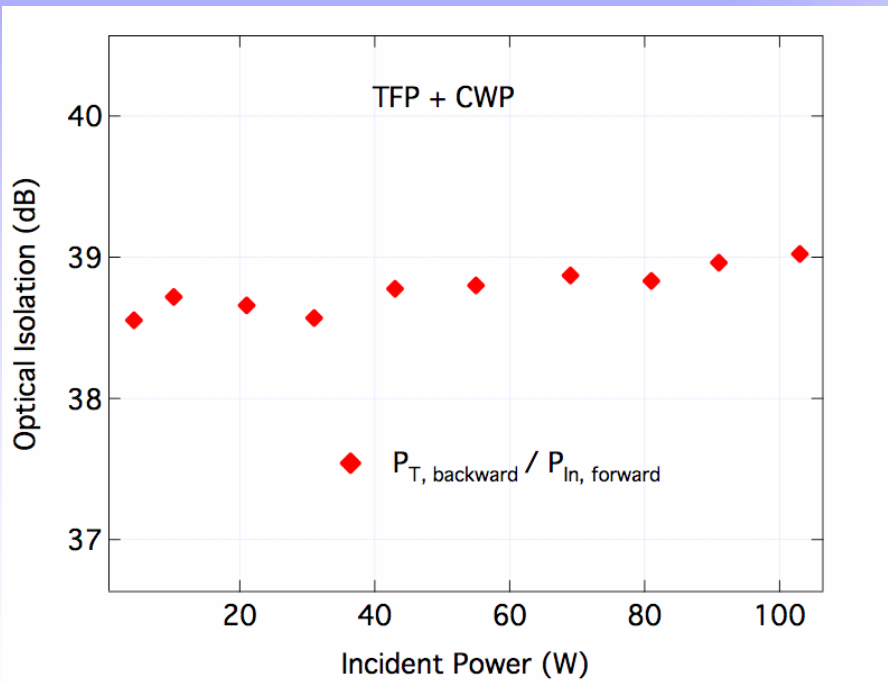


FI set up at LLO

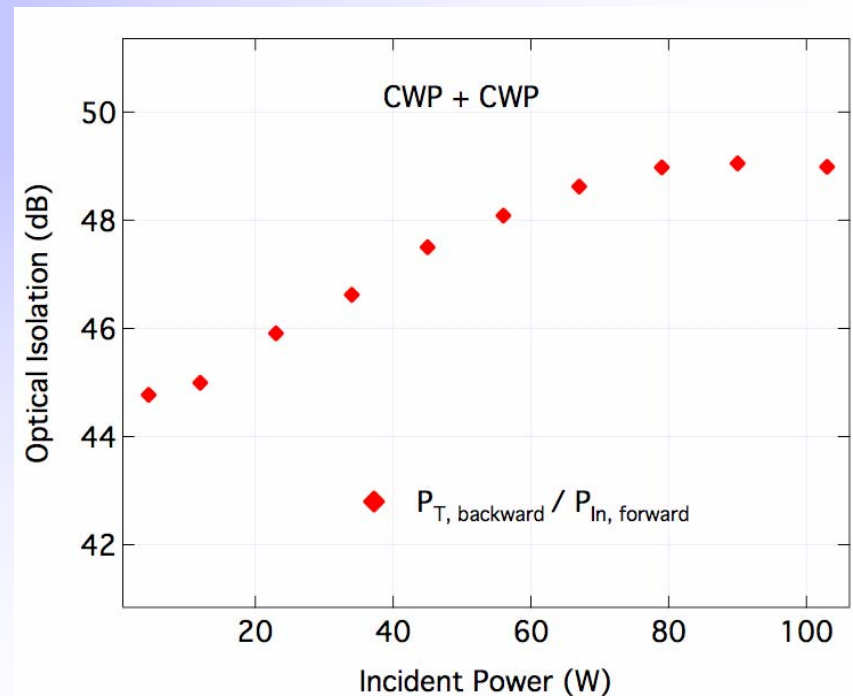


- Suppression is affected by the polarizers:

TFP and calcite polarizer



Two calcite polarizers



- Thermal lensing is compensated by DKDP
- Beam steering is measured to be smaller than (@ 100 W)
 - 80 μrad for two calcite wedges
 - 50 μrad for the TFP / calcite wedge setup

- Everything seems to be on track!



LIGO

Supplementary material

Properties	Units	RTP	RTA	KTP	LiNbO ₃
dn_x/dT	10 ⁻⁶ /K	-	-	11	5.4
dn_y/dT	10 ⁻⁶ /K	2.79	5.66	13	5.4
dn_z/dT	10 ⁻⁶ /K	9.24	11.0	16	37.9
κ_x	W/Km	3		2	5.6
κ_y	W/Km	3		3	5.6
κ_z	W/Km	3		3	5.6
α	cm ⁻¹	< 0.0005	< 0.005	< 0.005	< 0.05
Q_x	1/W	-	-	2.2	4.8
Q_y	1/W	0.047	0.94	2.2	4.8
Q_z	1/W	0.15	1.83	2.7	34

Properties	Units/conditions	RTP	RTA	LiNbO ₃
Damage Threshold	MW/cm ² ,	>600	400	280
n_x	1064nm	1.742	1.811	2.23
n_y	1064nm	1.751	1.815	2.23
n_z	1064nm	1.820	1.890	2.16
Absorption coeff. α	cm ⁻¹ (1064 nm)	< 0.0005	< 0.005	< 0.005
r_{33}	pm/V	39.6	40.5	30.8
r_{23}	pm/V	17.1	17.5	8.6
r_{13}	pm/V	12.5	13.5	8.6
r_{42}	pm/V	?	?	28
r_{51}	pm/V	?	?	28
r_{22}	pm/V			3.4
$n_z^3 r_{33}$	pm/V	239	273	306
Dielectric const., ϵ_z	500 kHz, 22 °C	30	19	
Conductivity, σ_z	$\Omega^{-1}\text{cm}^{-1}$, 10 MHz	$\sim 10^{-9}$	3×10^{-7}	
Loss Tangent, d_z	500 kHz, 22 °C	1.18	-	