



Update from LIGO Laboratory

LIGO-Virgo Collaboration Meeting

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Cascina, Italy

May 22-25, 2007





LIGO Laboratory Update Outline

- Progress with the science run
- Enhanced LIGO
- Advanced LIGO
- Outreach



S5 Run Status

- In November 2005 began S5 ...
 - Effective range for $1.4 M_{\odot} + 1.4 M_{\odot}$ neutron pair coalescence
 - for 4 km IFOs: ~ 10 Mpc
 - for 2 km IFO: ~ 5 Mpc
- Today ...
 - Ranges are now $> 50\%$ greater than at start of run ...
 - H1 - up to 16 Mpc peak
 - L1 - up to 15 Mpc peak
 - H2 - almost 8 Mpc peak
 - Duty factors (weekly averaged) regularly exceed 80%
- Virgo is now observing jointly!
 - SR1 -- Effective 18 May



S5: Progress to date

Range trend over run

QuickTime™ and a
TIFF (LZW) decompressor
are needed to see this picture.

Histogrammed minute trends in range

Minute trends in range

QuickTime™ and a
TIFF (LZW) decompressor
are needed to see this picture.



S5: Progress to date Through 15 May 2007

Cumulative Up-time



QuickTime™ and a
TIFF (LZW) decompressor
are needed to see this picture.

QuickTime™ and a
TIFF (LZW) decompressor
are needed to see this picture.

Projected End of S5

1 Year Coincident Observation

QuickTime™ and a
TIFF (LZW) decompressor
are needed to see this picture.

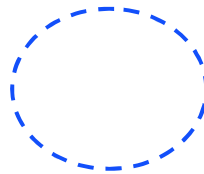
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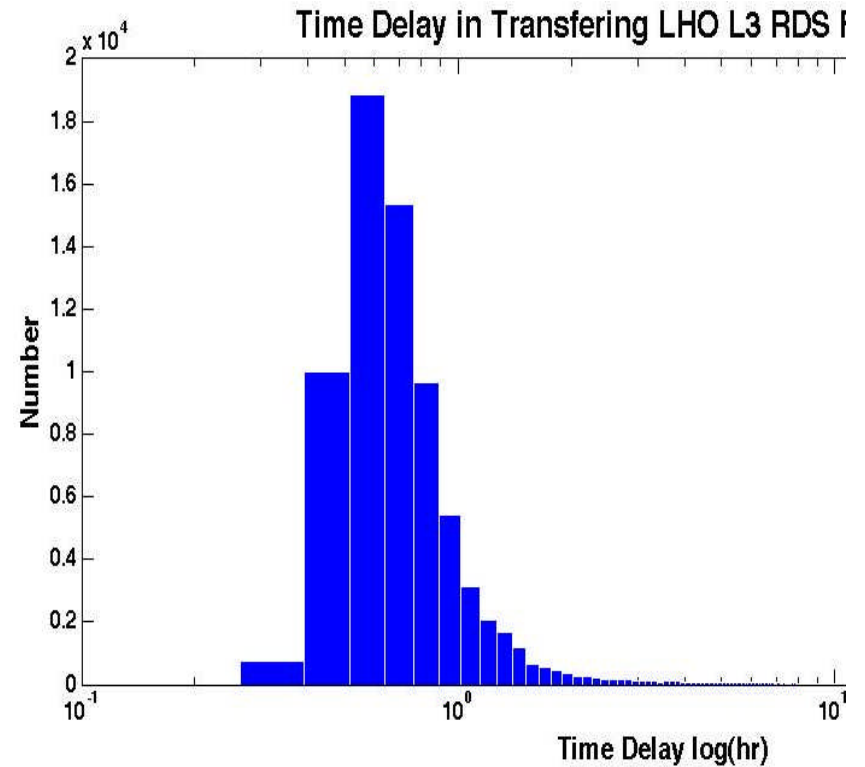
S5: Progress to date Data and Computing

QuickTime™ and a
TIFF (LZW) decompressor
are needed to see this picture.



- As of 21 May -- LIGO-Virgo data *exchange started for $h(t)$* .
 - File replication to all LIGO Data Grid sites
 - Segment Database - for quick data quality lookups and data quality modifications using V1:Hrec_veto_dataQuality

History of LIGO data
transmission from sites ->CIT



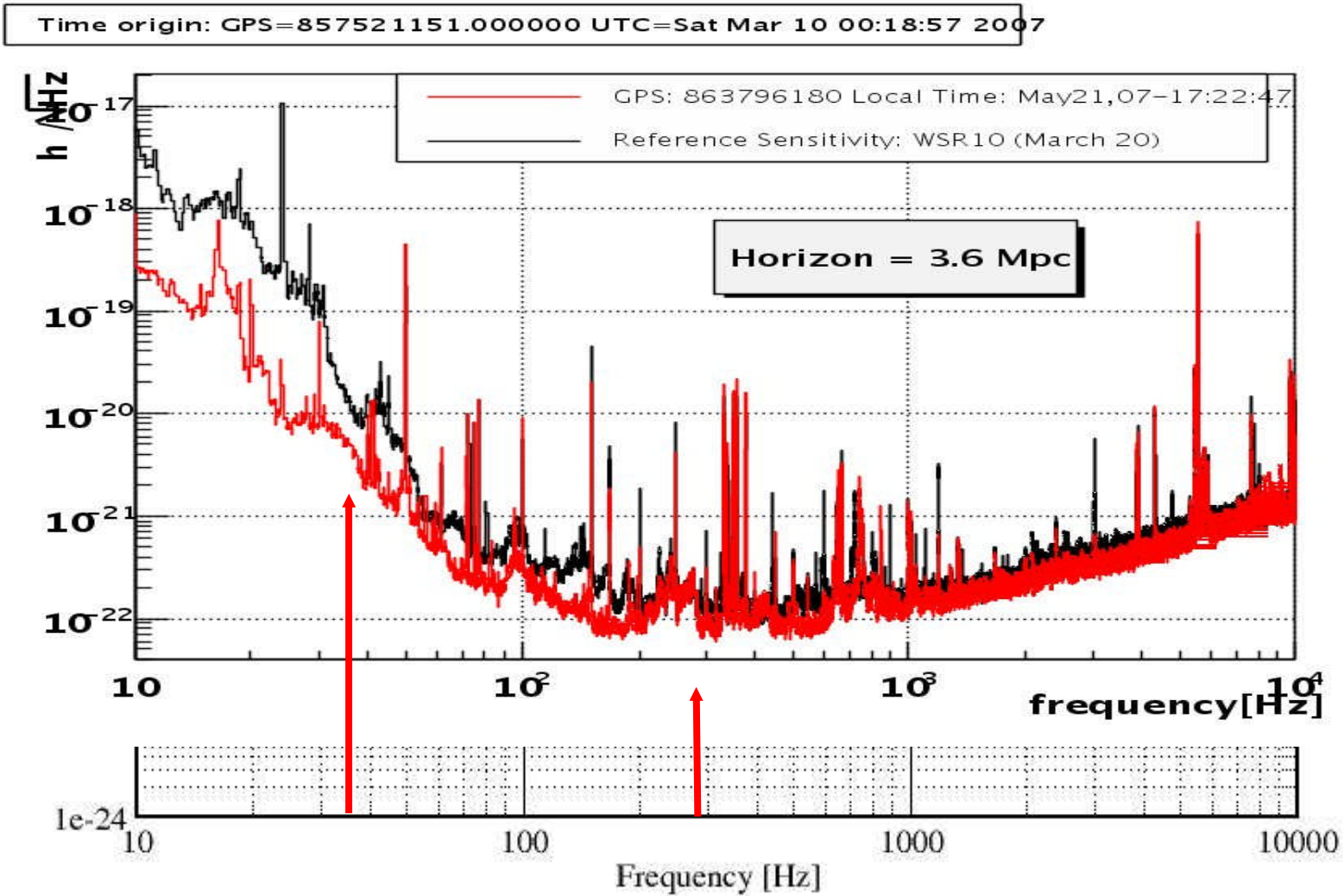


S5 beyond today

Beginning of a new era for the GW community

- 18 May marked the beginning of joint observation with Virgo.
- Culmination of *more than a decade* of communication, planning ...
 - First face-to-face with Virgo-LIGO took place at CIT 1996!
 - First discussion of vision for a common data format that would eventually enable common analysis.
- Level of cooperation and coordination is extremely high
 - Judged by frequency & duration of teleconferences ...
 - Technical challenges ahead ...
 - Addressed by joint committees to coordinate many key activities
 - Run planning, upgrade coordination
 - Data analysis
 - Resource management & sharing

LIGO S5 Sensitivity -- 2006 June





S5 -- Running with Virgo Joint Run Planning Committee (JRPC) *See talk later in this session*

- First F2F meeting of JRPC at LSC-Virgo meeting in Baton Rouge, March 2007
- Second earlier today
- Charged with specific short-term items:
 - Coordination for S5 joint running
 - Scenarios for post-S5 to S6 era
- Bi-weekly JRPC teleconferences
 - Most run coordination issues settled
 - Progress has been made on scenarios

Beyond S5



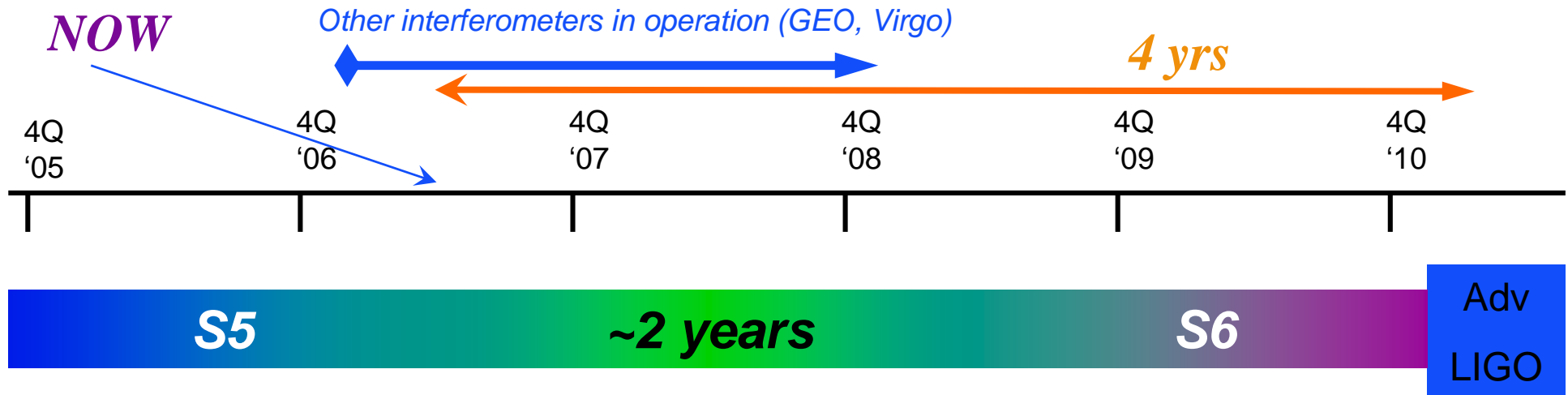
Enhanced LIGO and S6

See talk later in this session

- In 2005/2006 LIGO developed a concept to upgrade of the 4 km interferometers
- Laboratory commitment to upgrade made after August 2006 internal planning review
 - *SYNERGY WITH ADVANCED LIGO: UPGRADE WILL RETIRE RISK FOR A NUMBER OF SUBSYSTEMS*
 - *PSL, input optics, readout, seismic isolation, output mode cleaner, ...*
- e-LIGO team in place & working on the upgrade
 - Designs, prototyping, demonstration of proof-of-principle
 - Major hardware procurements
 - Seismic isolation for detection system
 - Lasers
 - Thermal compensation upgrade
 - Main laser (PSL) upgrade to 35W using Advanced LIGO front-end
 - Contributed by GEO
- *8-month status review successfully completed 16 May*

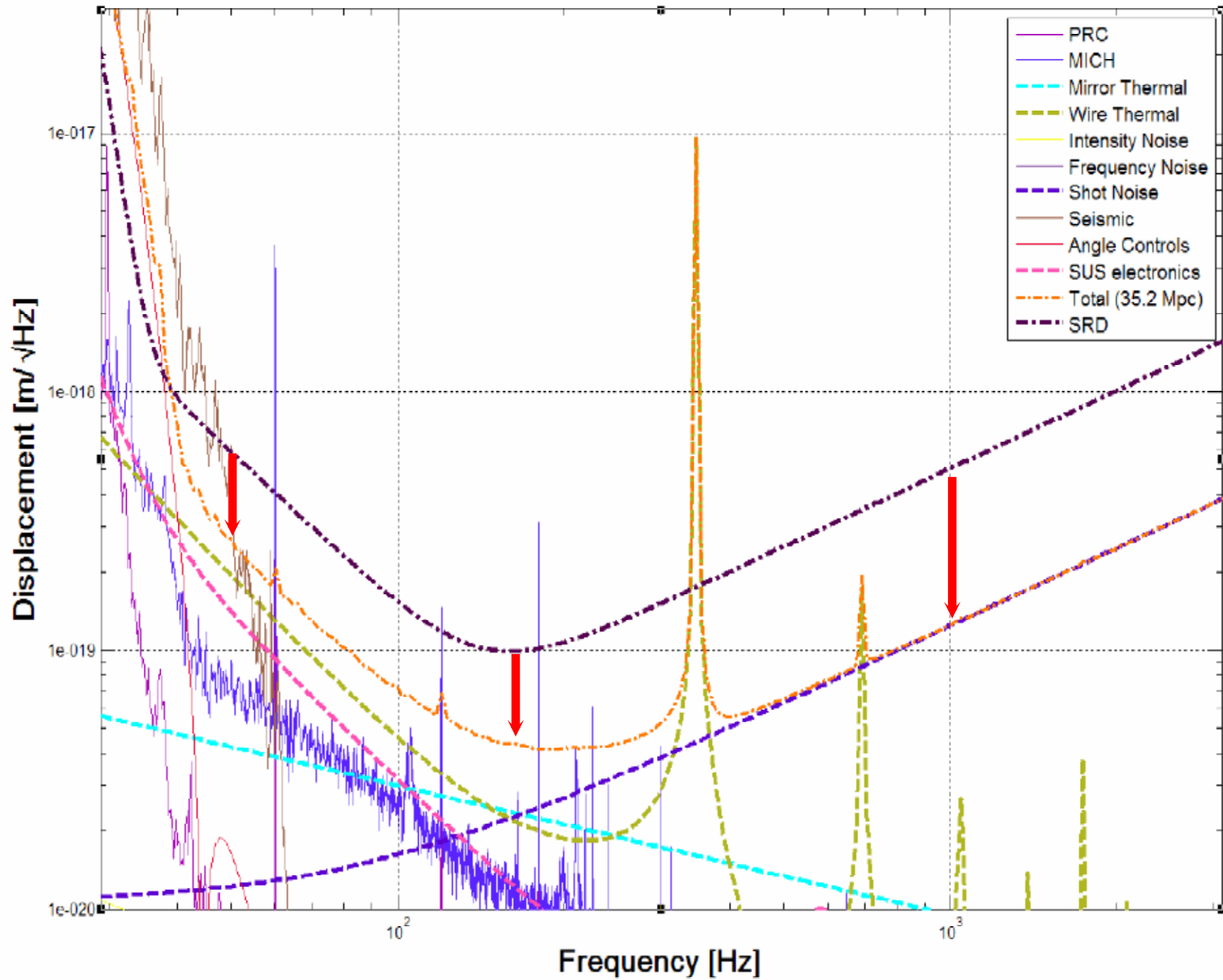


e-LIGO - The next three years



- Between now and AdvLIGO, there is some time to learn and improve and detect gravitational waves...
 - ~Few years of hardware improvements + ~1 ½ year of observations.
 - Factor ~2X in noise, factor ~5X-10X in event rate.
 - Better to spend debugging time before AdvLIGO to understand new systems planned for AdvLIGO...
 - AdvLIGO is a HUGE step in terms of interferometry!**

eLIGO - sensitivity goal

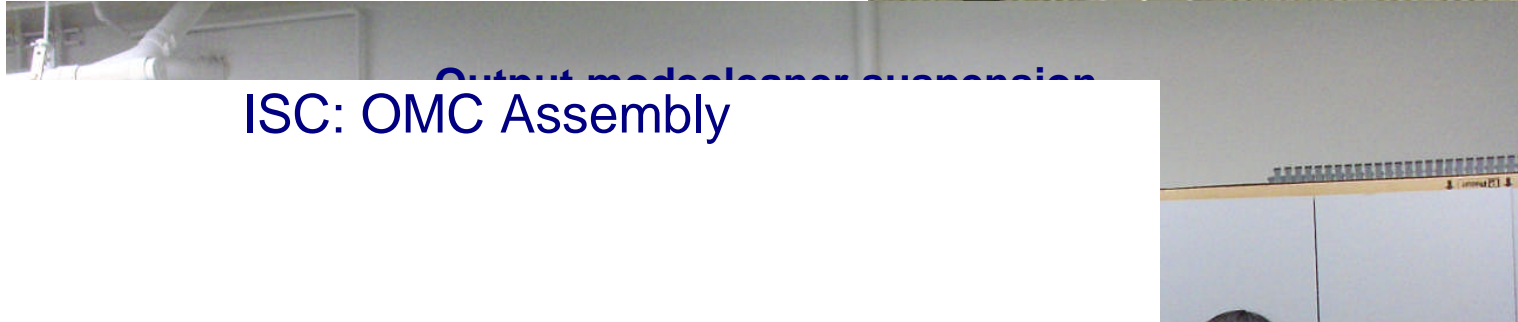




eLIGO Hardware Design & Fabrication

Adv.LIGO stiff active seismic isolation system

High-power Faraday Isolator



ISC: OMC Assembly

QuickTime™ and a TIFF (LZW) decompressor are needed to see this picture.



LZH Adv.LIGO laser 35W front end





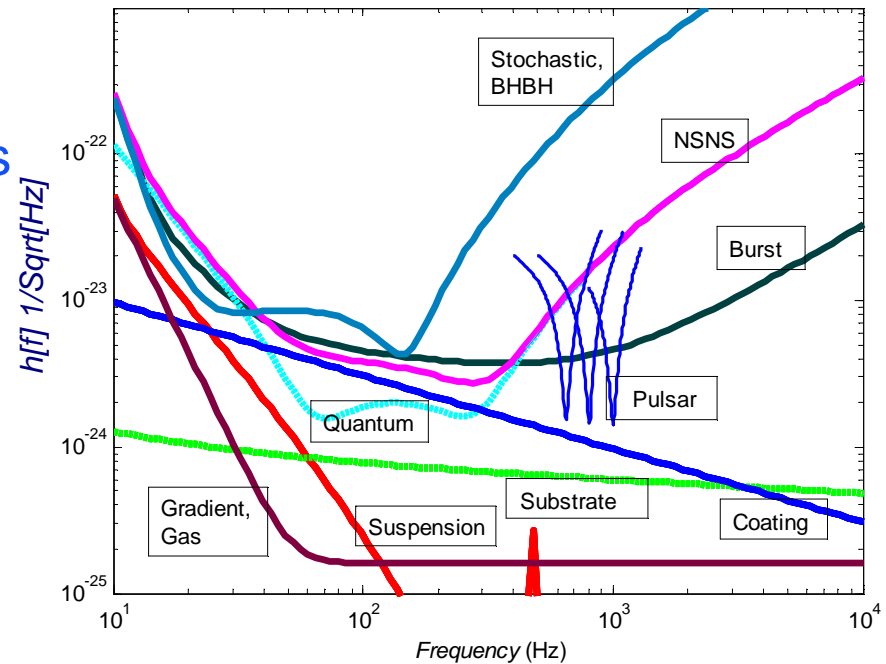
Advanced LIGO



- Reminder ...
 - Second generation of detectors in LIGO
 - Factor ~10X in amplitude sensitivity
 - Factor ~4X lower frequency 'wall'

- Quantum Limited at most frequencies
 - Recombined Fabry-Perot Michelson
 - ~20X higher input power
 - Signal recycling → tunable

- Gravitational gradient, thermal noise limits
 - 40 kg fused silica masses
 - Fused silica suspension
 - Aggressive seismic isolation





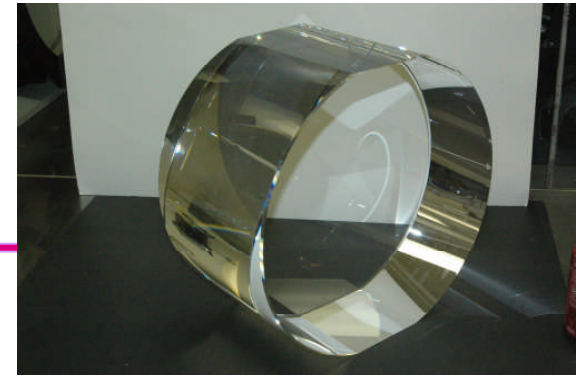
LIGO Advanced LIGO Status, Trajectory

- Team includes many LSC members plus, important capital contributions from UK and Germany
- On track to *start* Adv. LIGO Construction Project in FY08 (1 Oct 2007)
- Final Baseline Review at NSF prior to authorization 5, 6 June 2007
 - Preparatory internal reviews (re)confirmed cost, schedule planning are stable
- The only NSF Major Facility *start* in FY08 in the Office of Management and Budget request
 - Cost and schedule provided by LIGO, accepted by NSF & codified by OMB
- Breach vacuum in 2010 (end of e-LIGO)
- Start commissioning Advanced LIGO in 2013

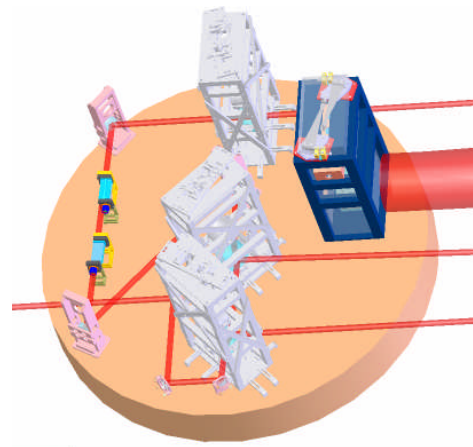
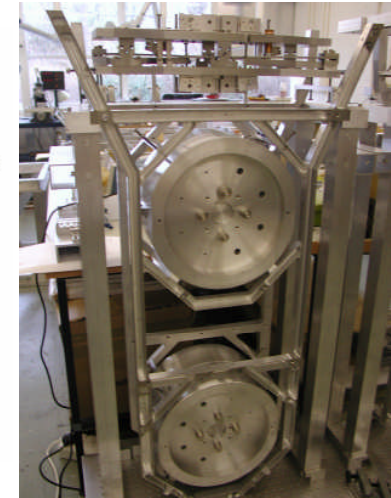
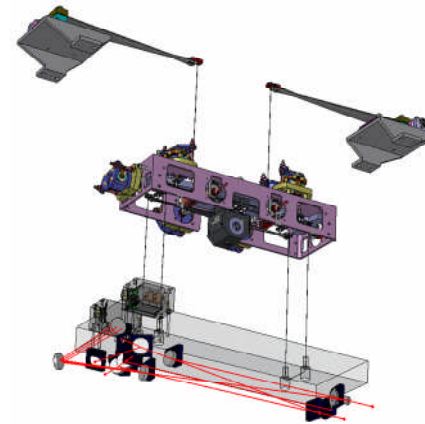


Progress

Technical advances

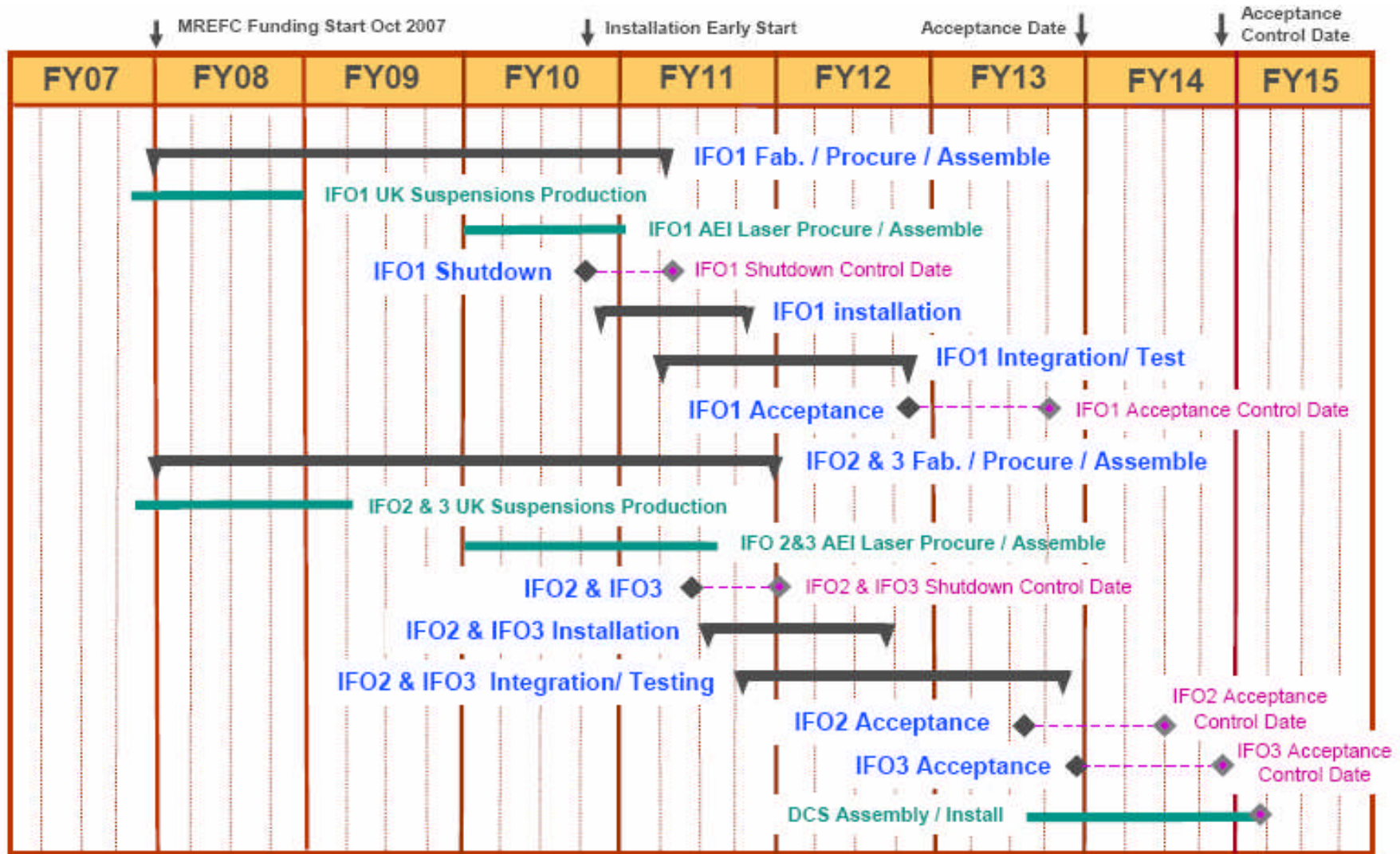


- Full scale prototyping of mechanical systems
- Tests of 'DC readout'
- Laser production
- Understanding of coating scatter
- Systems design





Schedule



2/2/07

LIGO - G070303-01-M

End Project Construction



Education & Outreach

LHO:

- >2800 total visitors; >1100 students (34% under-represented minorities)
- Summer immersion program for teachers for inquiry-based classroom instruction skills
- e-mentors to 6th local students
- >75% LHO staff participation
- Statewide WA Leadership Assistance Award for Science Education Reform
- Developing an *E-Lab* interactive tool for students to build research projects with LIGO PEM data

LLO – Since LIGO Science Education Center opened November 13th, 2006:

- > 2400 student visitors
(1 out of state group, 1 group >90 persons)
- > 225 teachers received Prof. Devel. Training at LIGO SEC
- 19 Special Events (star party, SPS Zone 10 mtng, family science night at schools, etc)
- Partners & participants: Southern University, Exploratorium, LA GEAR UP, Tulane University, Southeastern Louisiana University, LA Tech





Outreach

To scientific community

- **GWIC Thesis Prize -- replaces LIGO Thesis Prize**
 - 8 nominated theses (4 countries, 4 different projects, 5 experimental, 3 theoretical/data analysis)
 - Selection committee of 8 representing different GW projects and expertise
 - Winner announced 22 May 2007
 - Yoichi Aso (University of Tokyo)
 - "Active Vibration Isolation for a Laser Interferometric Gravitational Wave Detector using a Suspension Point Interferometer"
 - Prize to be awarded at Amaldi meeting in Sydney
- **American Astronomical Society (AAS) session on GWs**
 - **Encouraged to Proposed** a Special Session for the January 2008 meeting in Austin TX
 - 5 talks, 90 minutes,
 - Emphasis: Astronomy results and the evolution of a global network
 - Should hear this summer if proposal is accepted.



Summary

- LIGO is operating in a science mode at design sensitivity
 - 1st long science run is ~87.3% complete
 - Virgo has started SR1 and joined S5 and data are flowing!
- Near term vision: Enhanced LIGO upgrade 2008 - 2010
 - Improve by factor ~ 2X (*w.r.t.* S5) in $h[f]$ in 2009
 - S6 run - last of initial LIGO era
- Longer term -- poised for beginning of Advanced LIGO construction
 - Improve by factor 10X (*w.r.t.* S5) in $h[f]$ ~2014
- LIGO is having a positive, significant impact in both the local (WA, LA) public communities