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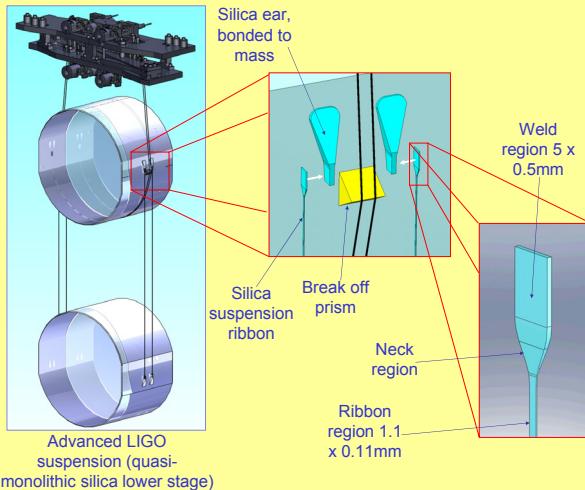
Production and characterisation of fused silica suspension ribbons for Advanced LIGO



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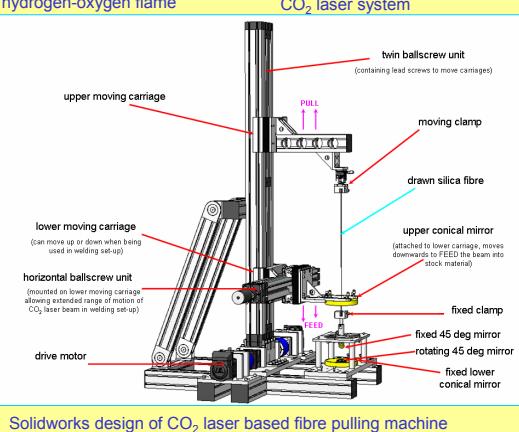
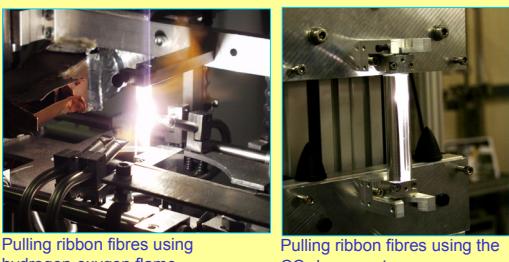
Ultra low thermal noise suspensions

Quasi-monolithic mirror suspension technology was developed for the GEO600 interferometric detector to minimise thermal noise effects. Circular cross section fused silica fibres are welded to silica ears, which are silicate bonded onto the sides of the test masses. Upgrade to Advanced LIGO will extend this technology, further reducing thermal noise, by use of rectangular cross section ribbon fibres.



CO₂ laser production of silica fibres

Fused silica fibres for GEO600 were manufactured by heating silica rods using a hydrogen-oxygen flame and pulling them down into fibres. For the upgrades to LIGO, however, increased tolerances on ribbon fibre dimensions are required. A new fibre pulling system, based on heating with a CO₂ laser, was designed to give improved control.

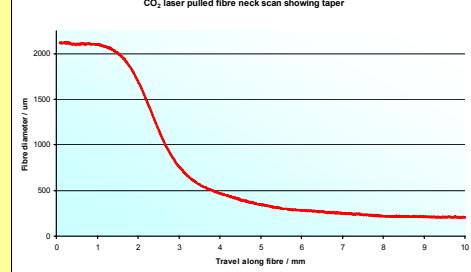
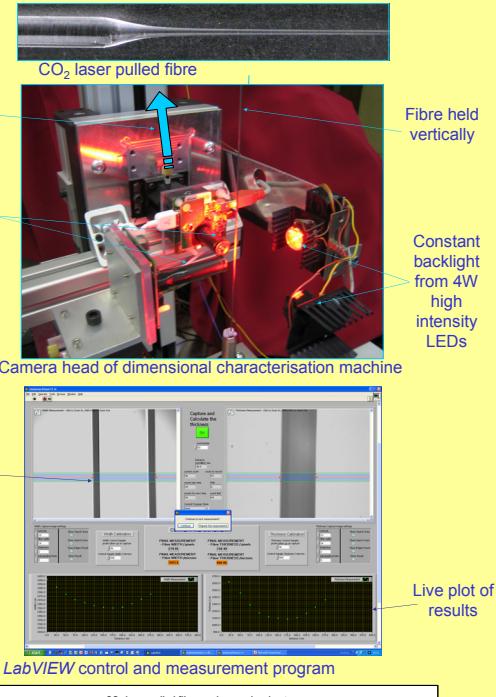


Solidworks design of CO₂ laser based fibre pulling machine

Dimensional characterisation

In conjunction with the pulling machine, a non-contact ribbon dimension measurement system has been developed. This system uses two Firewire web-cameras with high magnification lenses to focus on the thick and thin ribbon dimensions. Ribbons require accurate dimensional measurements to ensure they are within the required tolerances, and measurement of neck shapes is important in understanding the effective pendulum length, and quantifying suspension thermal noise.

The cameras are mounted on a motorised carriage and traverse the length of the fibre. A LabVIEW control program performs an edge detection measurement to obtain fibre cross sectional dimensions. Particular emphasis is placed on the neck region with measurements performed at up to every 1.5 μm, providing ample resolution to allow detailed dimension profiles to be obtained.



Typical plot of diameter of CO₂ laser pulled fibre

Both the CO₂ pulling and dimensional characterisation machines are at an advanced development stage, and duplicate machines will be installed at the LIGO LASTI test facility at M.I.T. over coming months in preparation for fabrication of the first Advanced LIGO thermal noise prototype.

We would like to thank STFC, the University of Glasgow, SUPA, EGO, Leverhulme Trust and collaborators at LIGO for their support in these projects

