



Semi-analythic optimization of future interferometers

F.Ya.Khalili

October 26, 2007

LIGO-G070803-00-Z

- ① Motivation
- ② The method
- ③ Example 1: AdvLIGO
- ④ Example 2: Squeezed AdvLIGO
- ⑤ Conclusion

- ① Motivation
- ② The method
- ③ Example 1: AdvLIGO
- ④ Example 2: Squeezed AdvLIGO
- ⑤ Conclusion

Methods of enhancing AdvLIGO sensitivity

- ① Squeezed light
- ② Phase filter cavity(ies)
- ③ Amplitude filter cavity(ies)
- ④ Two-pass schemes
- ⑤ ... + any combinations

Methods of enhancing AdvLIGO sensitivity

- 1 Squeezed light
- 2 Phase filter cavity(ies)
- 3 Amplitude filter cavity(ies)
- 4 Two-pass schemes
- 5 ... + any combinations



- 1 Low-frequency (inspirals)
- 2 Broad-band (bursts)
- 3 Narrow-band (periodic)
- 4 ...

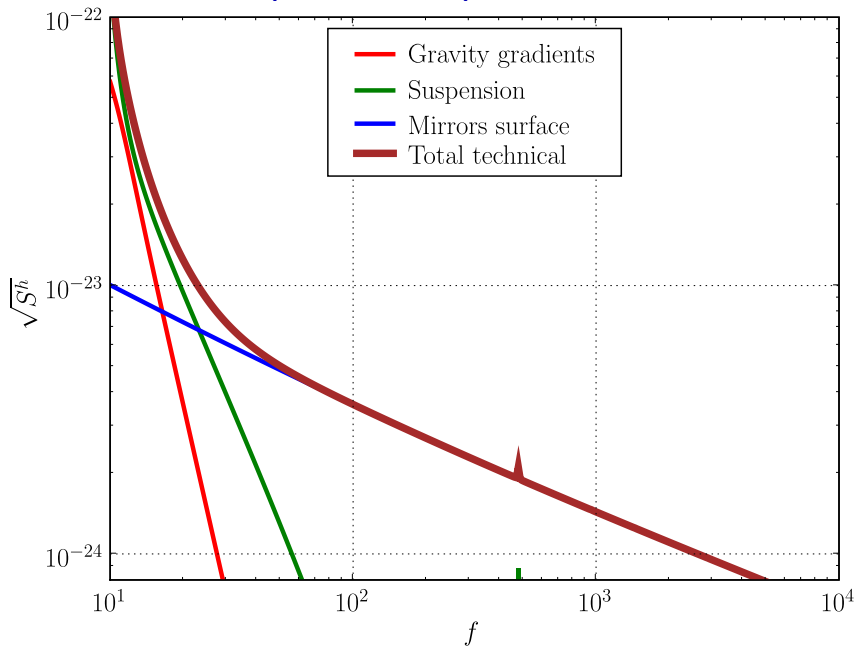
Parameters space

Universal constants

$$\hbar \quad \omega_p \quad L \quad M$$

Technical noises (tend to drift up with time)

Technical noises (bench 6.1)



Parameters space

Universal constants

$$\hbar \quad \omega_p \quad L \quad M$$

Technical noises (tend to drift up with time)

Variables

Arm cavities: $\gamma_{\text{arm}} \quad \delta_{\text{arm}}$

SR cavity: $\rho_{\text{SRC}} \quad \phi_{\text{SRC}}$

Pumping: $W_{\text{arm}} \quad \phi$

Squeezing: $r \quad \theta$

Filter cavity(ies) parameters

...

The motivation

We have lots of parameters to optimize.

The motivation

We have lots of parameters to optimize.

For each pair “configuration \otimes GW source type”,
numeric optimization can be performed.

The motivation

We have lots of parameters to optimize.

For each pair “configuration \otimes GW source type”, numeric optimization can be performed.

However, only (semi)analytical optimization is able to provide heuristics in new configurations search.

- ① Motivation
- ② The method
- ③ Example 1: AdvLIGO
- ④ Example 2: Squeezed AdvLIGO
- ⑤ Conclusion

Reduction of the parameters space, stage 1

$$J = \frac{8\omega_p W_{\text{arm}}}{McL}; \quad \text{AdvLIGO: } J = (2\pi \times 100 \text{ s}^{-1})^3.$$

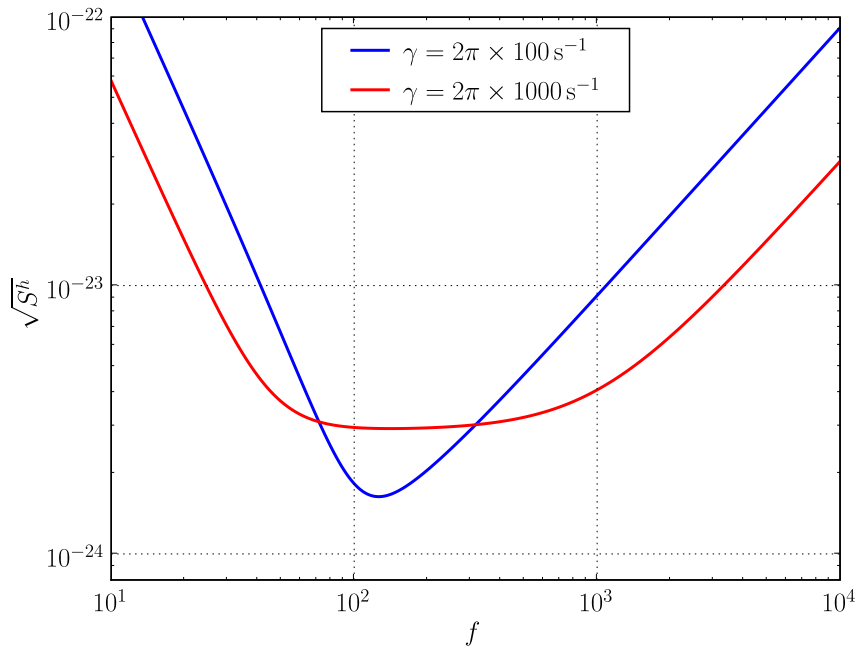
A.Buonanno, Y.Chen, PRD **67**, 062002 (2003):

$$\gamma = \gamma_{\text{arm}} \operatorname{Re} \frac{1 + \rho_{\text{SRC}} e^{2i\phi_{\text{SRC}}}}{1 - \rho_{\text{SRC}} e^{2i\phi_{\text{SRC}}}},$$

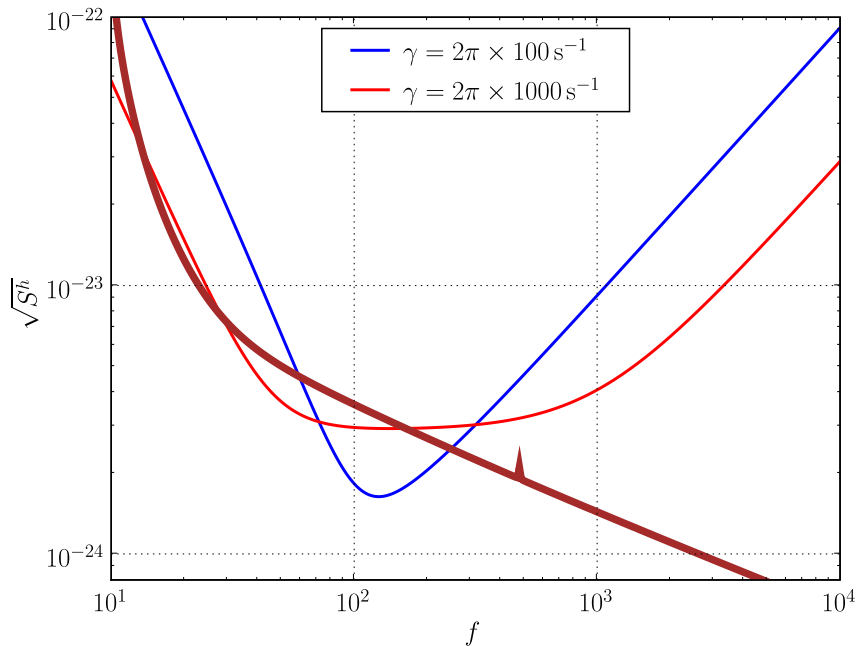
$$\delta = \delta_{\text{arm}} + \gamma_{\text{arm}} \operatorname{Im} \frac{1 + \rho_{\text{SRC}} e^{2i\phi_{\text{SRC}}}}{1 - \rho_{\text{SRC}} e^{2i\phi_{\text{SRC}}}}$$

$$\Leftrightarrow \Gamma = \sqrt{\gamma^2 + \delta^2}, \quad \beta = \arctan \frac{\delta}{\gamma}.$$

Conventional vs. broadened



Conventional vs. broadened



Conventional vs. broadened

Due to the technical noises (mainly, mirrors surface fluctuations), the optimal γ should be large:

$$\gamma \gg J^{1/3} = 2\pi \times 100 \text{ s}^{-1}.$$

Reduction of the parameters space, stage 2

$$S^h(\Omega) = \frac{8}{m^2 L^2 \Omega^4} \left\{ [K_{\text{eff}}(\Omega) - m\Omega^2]^2 S_x(\Omega) + \frac{\hbar^2}{4S_x(\Omega)} \right\} + S_{\text{tech}}^h(\Omega).$$

Here

$S_x(\Omega)$: shot noise.

$S_{xF}(\Omega)$: cross-correlation of shot noise and radiation-pressure noise; depends on ϕ .

$K_{\text{eff}}(\Omega) = K(\Omega) + \frac{S_{xF}(\Omega)}{S_x(\Omega)}$: effective rigidity.

By some cryptic reason, without optical losses

K_{eff} is always real.

Reduction of the parameters space, stage 2

$$S^h(\Omega) = \frac{8}{m^2 L^2 \Omega^4} \left\{ [K_{\text{eff}}(\Omega) - m\Omega^2]^2 S_x(\Omega) + \frac{\hbar^2}{4S_x(\Omega)} \right\} + S_{\text{tech}}^h(\Omega).$$

If $\Omega \ll \Gamma$, then

$$S^h(\Omega) \approx \frac{8}{m^2 L^2 \Omega^4} \left\{ [K_{\text{eff}}(0) - m\Omega^2]^2 S_x(0) + \frac{\hbar^2}{4S_x(0)} \right\} + S_{\text{tech}}^h(\Omega),$$

i.e., all optical parameters are reduced to

$$S_x(0) \quad \text{and} \quad K_{\text{eff}}(0).$$

Reduction of the parameters space, stage 2

$$S^h(\Omega) = \frac{8}{m^2 L^2 \Omega^4} \left\{ [K_{\text{eff}}(\Omega) - m\Omega^2]^2 S_x(\Omega) + \frac{\hbar^2}{4S_x(\Omega)} \right\} + S_{\text{tech}}^h(\Omega).$$

If $\Omega \gtrsim \Gamma \gg J^{1/3}$, then

$$\begin{aligned} S^h(\Omega) &\approx \frac{8}{L^2} S_x(\Omega \rightarrow \infty) + S_{\text{tech}}^h(\Omega) \\ &= \frac{dS_x(\Omega \rightarrow \infty)}{d\Omega^2} \times \Omega^2 + S_{\text{tech}}^h(\Omega). \end{aligned}$$

i.e., all optical parameters are reduced to

$$\frac{dS_x(\Omega \rightarrow \infty)}{d\Omega^2}.$$

Reduction of the parameters space, stage 2

$$S^h(\Omega) = \frac{8}{m^2 L^2 \Omega^4} \left\{ [K_{\text{eff}}(\Omega) - m\Omega^2]^2 S_x(\Omega) + \frac{\hbar^2}{4S_x(\Omega)} \right\} + S_{\text{tech}}^h(\Omega).$$

Therefore, if $\Gamma \gg J^{1/3}$, then optimization for low-frequency sources gives

$$S_x(0) \quad \text{and} \quad K_{\text{eff}}(0).$$

and optimization for high-frequency ones gives

$$\frac{dS_x(\Omega \rightarrow \infty)}{d\Omega^2}.$$

- ① Motivation
- ② The method
- ③ Example 1: AdvLIGO**
- ④ Example 2: Squeezed AdvLIGO
- ⑤ Conclusion

Low-frequency

Numerical optimization of

$$\text{SNR}_{\text{NSNS}} = \int_0^{f_{\text{ISCO}}} \frac{f^{-7/3} df}{S^h(2\pi f)}$$

gives

$$\frac{MS_x(0)}{\hbar} = \frac{\Gamma}{4J \cos \beta \cos^2(\phi + \beta)} = (2\pi \times 48 \text{ s}^{-1})^{-2},$$
$$\frac{K_{\text{eff}}(0)}{M} = \frac{J[\cos \beta \sin 2(\phi + \beta) + \sin \beta]}{\Gamma} = (2\pi \times 51 \text{ s}^{-1})^2.$$

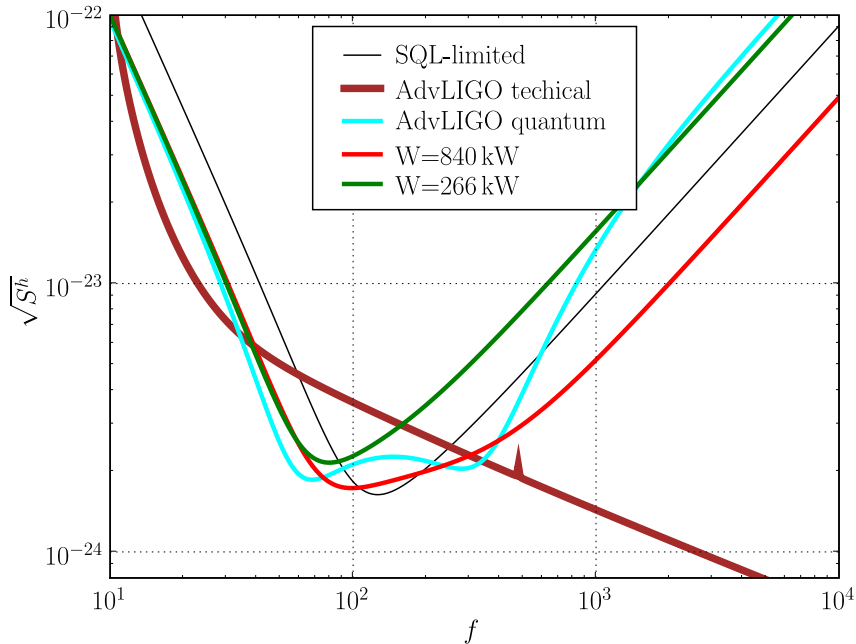
High-frequency

$$\frac{M}{\hbar} \frac{dS_x(\Omega \rightarrow \infty)}{d\Omega^2} = \frac{1}{4J\Gamma \cos \beta \cos^2 \phi} = \min .$$

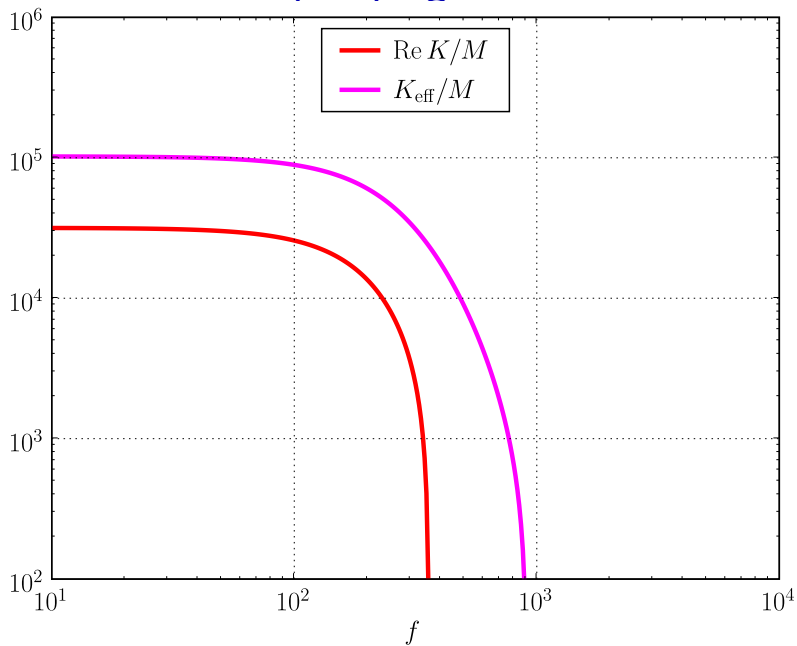
A minimization problem with 2 additional conditions. The solution is the following:

$$\frac{\Gamma}{J} \approx \frac{2300}{(2\pi \times 100)^3} \text{ s}^{-2}, \quad \beta \approx 0.29, \quad \phi \approx 0.08 .$$

AdvLIGO, coherent pumping



AdvLIGO, coherent pumping



AdvLIGO, coherent pumping

Pseudo-resonance (the effective rigidity K_{eff}) is created mostly by the cross-correlation term S_{xF} ; the real rigidity K is much smaller.

Estimates of h_{SQL}/h

Low-frequency (NSNS):

$$\frac{h_{\text{SQL}}}{h} = \sqrt{\frac{\int_{f_c}^{f_{\text{ISCO}}} \frac{f^{-7/3}}{S^h(2\pi f) + S_{\text{tech}}^h(2\pi f)} df}{\int_{f_c}^{f_{\text{ISCO}}} \frac{f^{-7/3}}{S_{\text{conv}}^h(2\pi f) + S_{\text{tech}}^h(2\pi f)} df}}$$

High-frequency:

$$\frac{h_{\text{SQL}}}{h} = \sqrt{\frac{S_{\text{conv}}^h(2\pi f) + S_{\text{tech}}^h(2\pi f)}{S^h(2\pi f) + S_{\text{tech}}^h(2\pi f)}} \Bigg|_{2\pi f \gg \gamma}$$

Estimates of h_{SQL}/h

	W	NSNS	$\gtrsim 1$ kHz
SQL-limited	840 kW	1	1
AdvLIGO	840 kW	1.18	0.50
Broadened	266 kW	1.07	0.58
	840 kW	1.17	1.85

- ① Motivation
- ② The method
- ③ Example 1: AdvLIGO
- ④ Example 2: Squeezed AdvLIGO**
- ⑤ Conclusion

Low-frequency

The same numerical optimization of low-frequency sensitivity can be used:

$$\frac{MS_x(0)}{\hbar} = \frac{\Gamma[\cosh 2r + \sinh 2r \cos(\phi + \theta + 2\beta)]}{4J \cos \beta \cos^2(\phi + \beta)} \\ = (2\pi \times 48 \text{ s}^{-1})^{-2},$$

$$\frac{K_{\text{eff}}(0)}{M} = \text{a long formula which does not fit here} \\ = (2\pi \times 51 \text{ s}^{-1})^2.$$

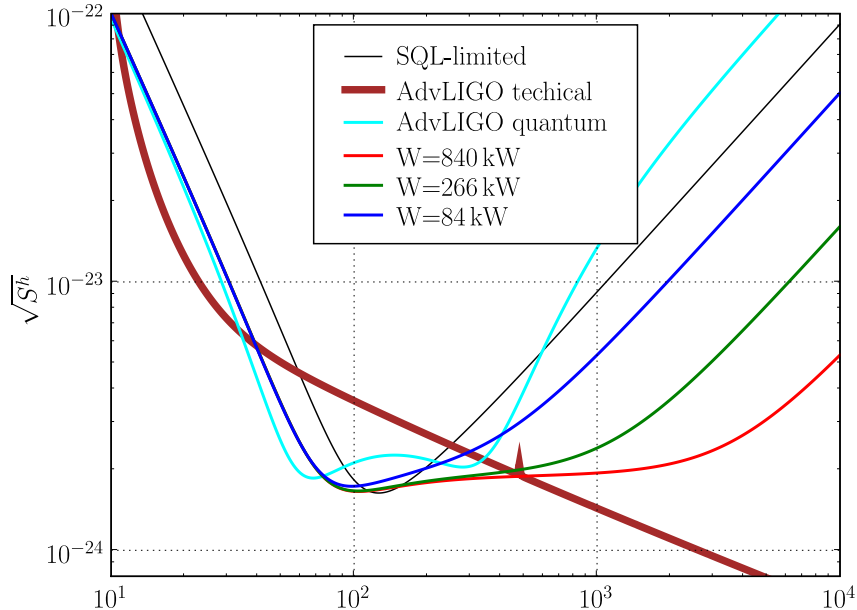
High-frequency

$$\frac{M}{\hbar} \frac{dS_x(\Omega \rightarrow \infty)}{d\Omega^2} = \frac{\cosh 2r + \sinh 2r \cos(\phi + \theta)}{4J\Gamma \cos \beta \cos^2 \phi} = \min .$$

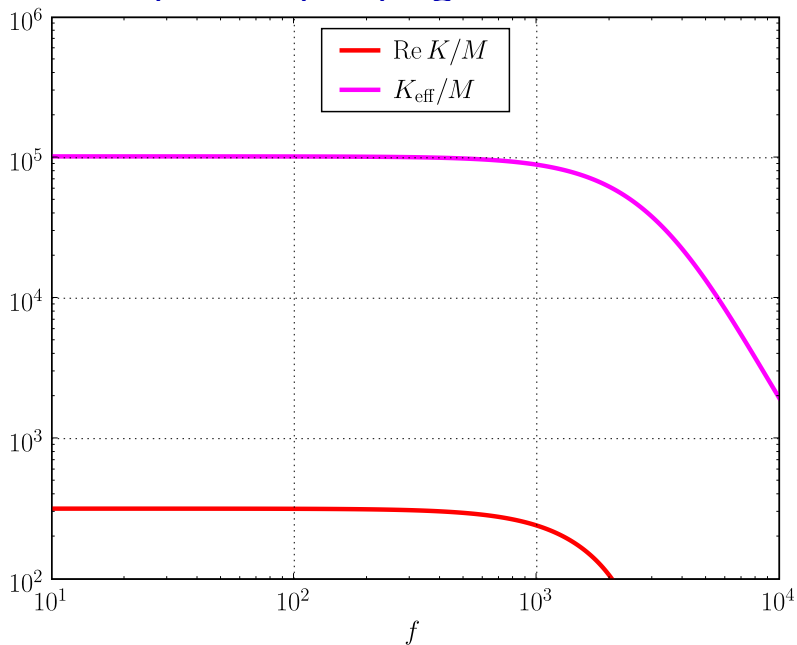
Again, a minimization problem with 2 additional conditions. The solution is the following:

$$\begin{aligned} \frac{\Gamma}{J} &\approx \frac{2.1 \times 10^4}{(2\pi \times 100)^3} \text{ s}^{-2}, & \beta &\approx 0.03, \\ \phi &\approx -0.05, & \theta &\approx -1.52. \end{aligned}$$

AdvLIGO, squeezed pumping



AdvLIGO, squeezed pumping



AdvLIGO, squeezed pumping

Pseudo-resonance (the effective rigidity K_{eff}) is created almost completely by the cross-correlation term S_{xF} ; the real rigidity K is negligibly small.

Estimates of h_{SQL}/h

	W	NSNS	$\gtrsim 1$ kHz
SQL-limited	840 kW	1	1
AdvLIGO	840 kW	1.18	0.50
Broadened	266 kW	1.07	0.60
	840 kW	1.17	1.89
Sqz.+Broad.	84 kW	1.17	1.80
	266 kW	1.21	5.67
	840 kW	1.22	16.9

- ① Motivation
- ② The method
- ③ Example 1: AdvLIGO
- ④ Example 2: Squeezed AdvLIGO
- ⑤ Conclusion

Conclusion

- Squeezed light in combination with the broadened interferometer configuration allows to obtain sensitivity, comparable with the AdvLIGO sensitivity at low frequencies, and significantly better — at high frequencies, using much smaller optical power.

Conclusion

- Squeezed light in combination with the broadened interferometer configuration allows to obtain sensitivity, comparable with the AdvLIGO sensitivity at low frequencies, and significantly better — at high frequencies, using much smaller optical power.
- In order to obtain the necessary interferometer bandwidth $\gamma \sim 10^4 \text{ s}^{-1}$ with narrow-band arm cavities:

$$T_{\text{ITM}}^2 = 5 \times 10^{-4} \Rightarrow \gamma_{\text{arm}} = 93.75 \text{ s}^{-1},$$

the SR mirror reflectivity have to be increased up to $R_{\text{SR}} \gtrsim 1 - 10^{-2}$.

To do:

Configurations with filter cavity(ies).