

Inspiral, Merger, Ringdown

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LIGO-G080415-00-Z



MEETINGS

NONE OF US IS AS DUMB AS ALL OF US.

NR + DA



POTENTIAL

NOT EVERYONE GETS TO ~~BE AN ASTRONAUT~~ WHEN THEY GROW UP.
Quantize Gravity

Data Analysis

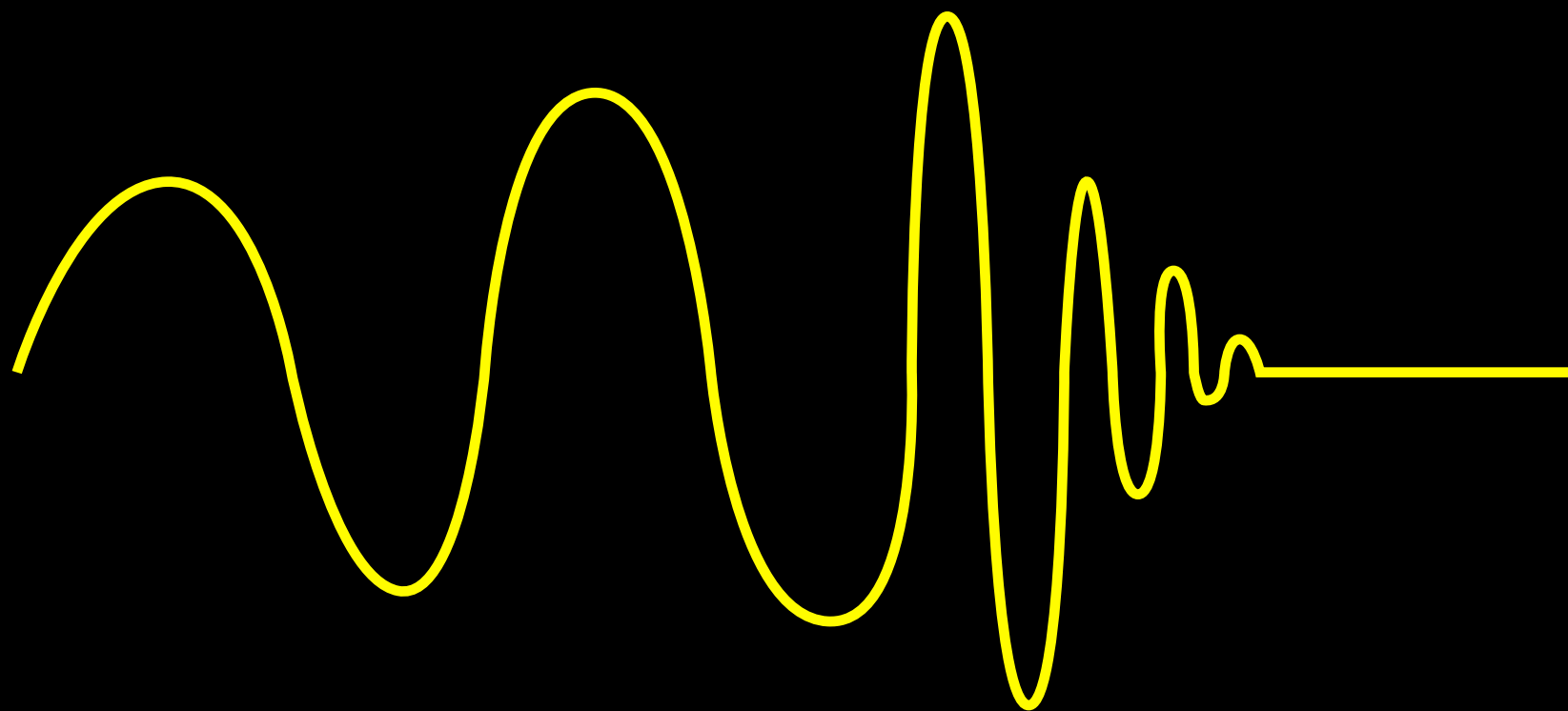
Outline

- Inspiral
- Merger
- Ringdown
- Outlook

Inspiral Merger Ringdown

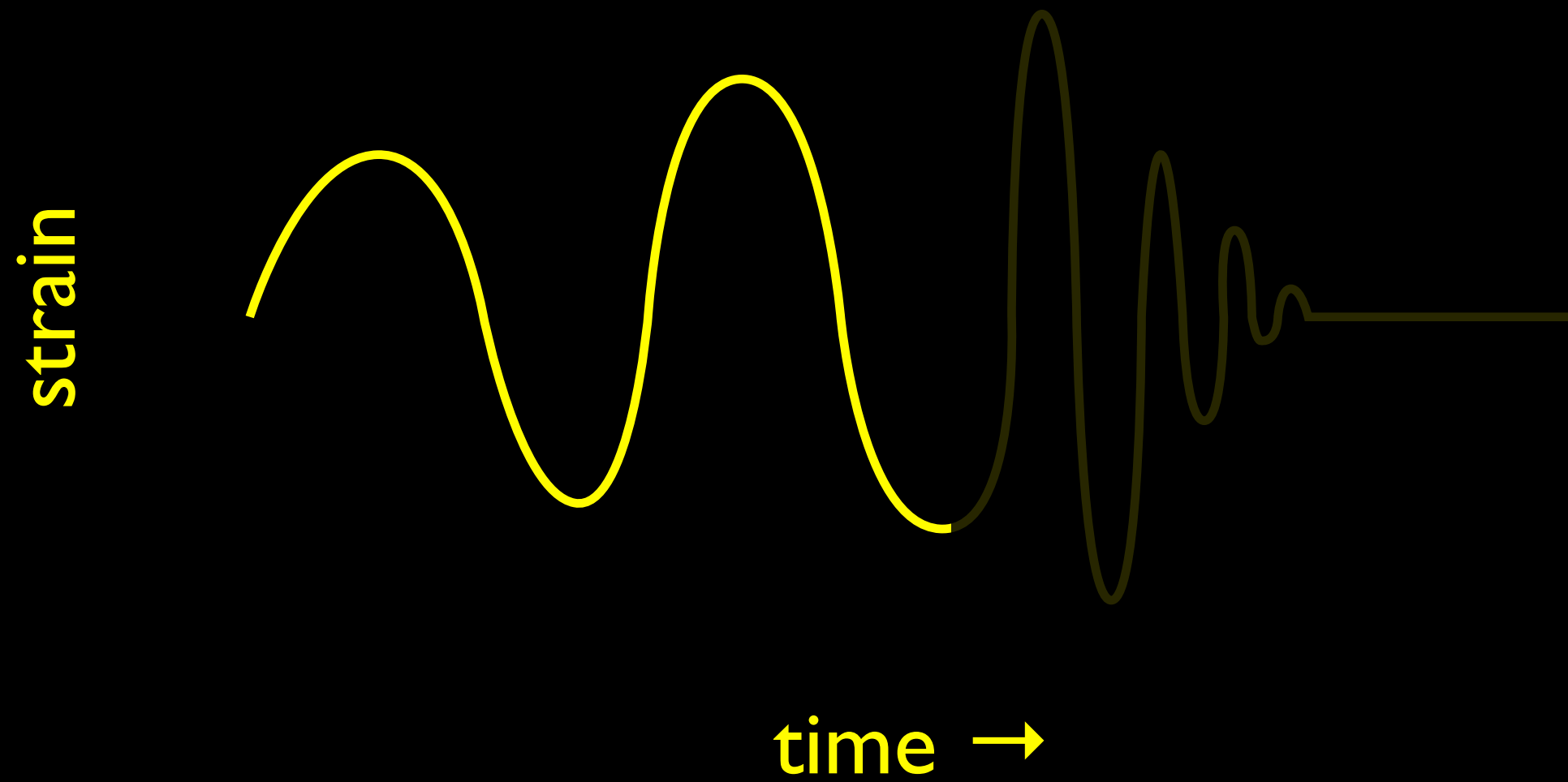
A horizontal white arrow pointing to the left, with the labels 'Inspiral', 'Merger', and 'Ringdown' positioned above it. Two small white dots are placed on the arrow, one under 'Merger' and one under 'Ringdown'.

strain



time →

Inspiral



Inspiral Search

- Matched filter is optimal:
 - determines: $\Pr\{ \text{Inspiral} \mid \text{Data} \}$
- Search over unknown params:
 $\{ t, \varphi, m_1, m_2, \dots \}$
- Search over $\{ t, \varphi \}$ is easy
- Use a bank to search over $\{ m_1, m_2 \}$

Inspiral Search

- Restricted 2 pN stationary phase templates
 - High mass?
 - Spin?
- χ^2 (chi-squared) waveform quality veto
- Upper limit depends on efficiency estimate

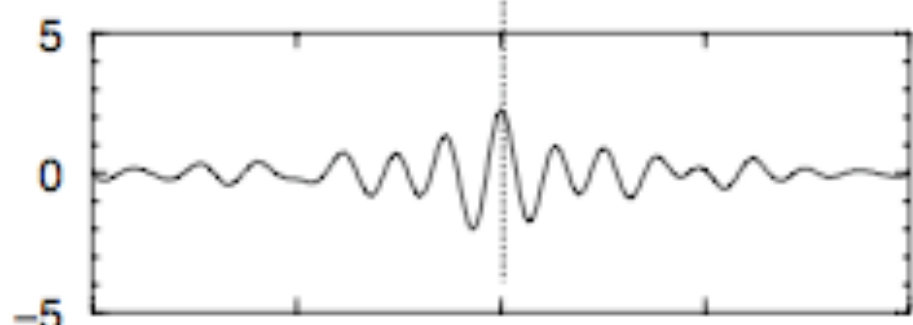
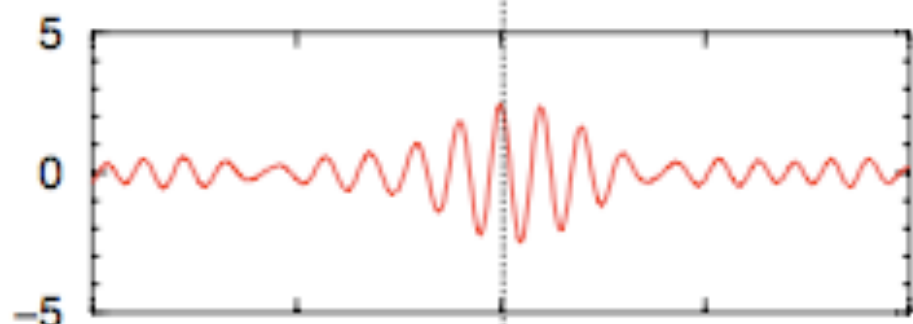
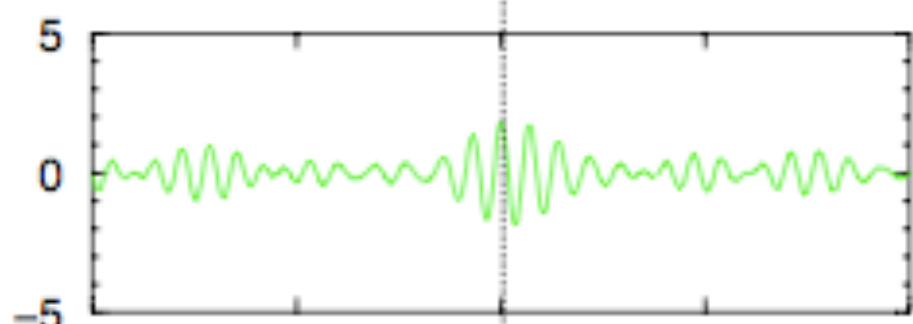
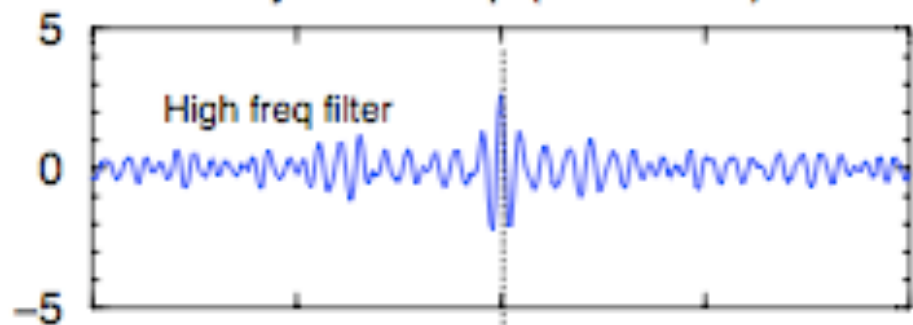


CONSISTENCY

IT'S ONLY A VIRTUE IF YOU'RE NOT A SCREWUP.

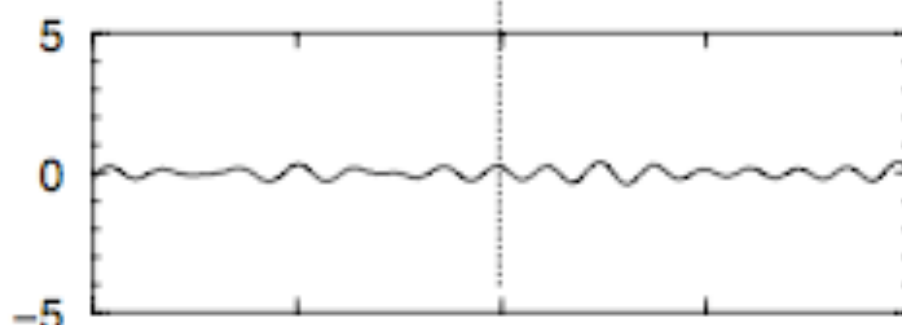
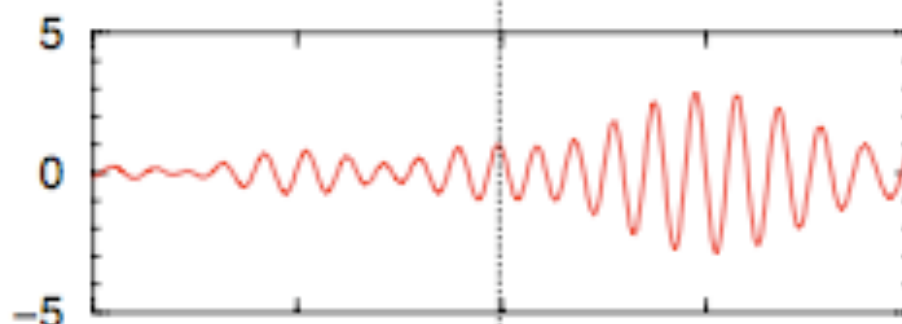
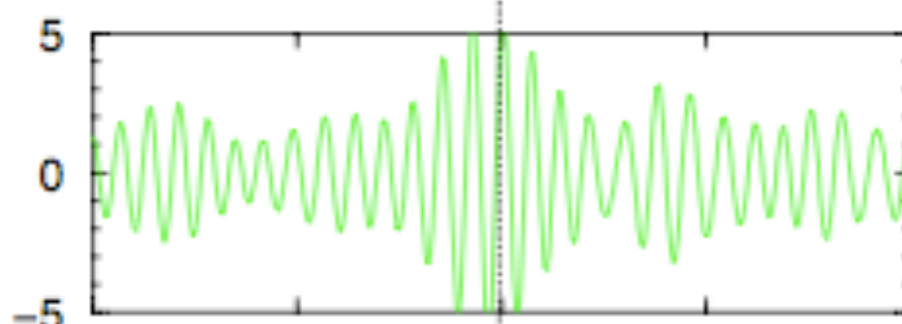
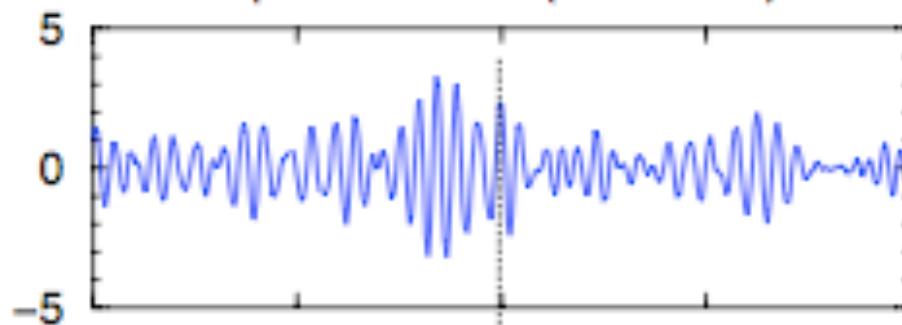
χ^2 Veto

Injected Chirp (SNR = 9.2)



$t = t_0$

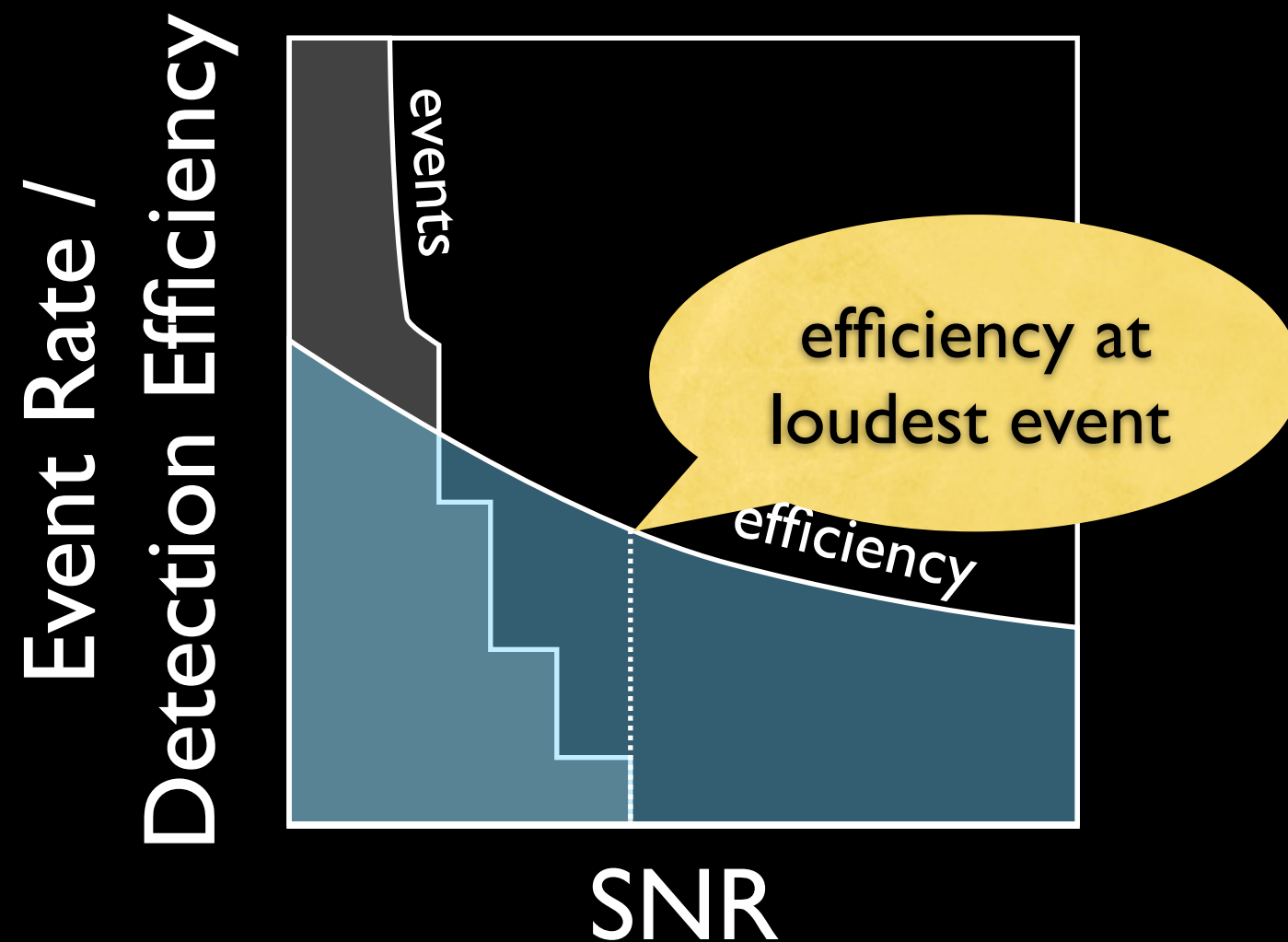
Spurious Event (SNR = 8.7)



$t = t_0$

Allen (GRASP)

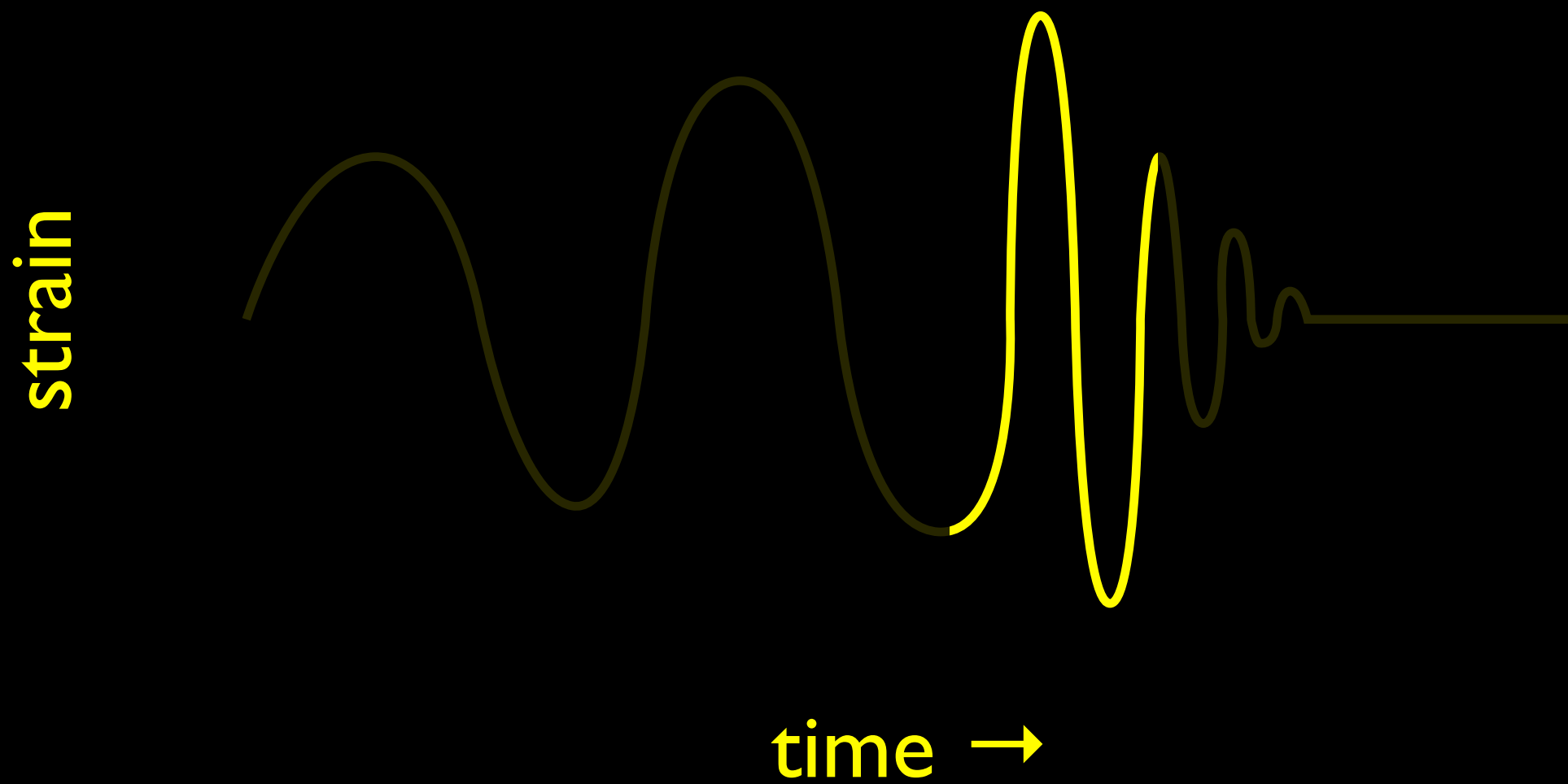
Detection Efficiency

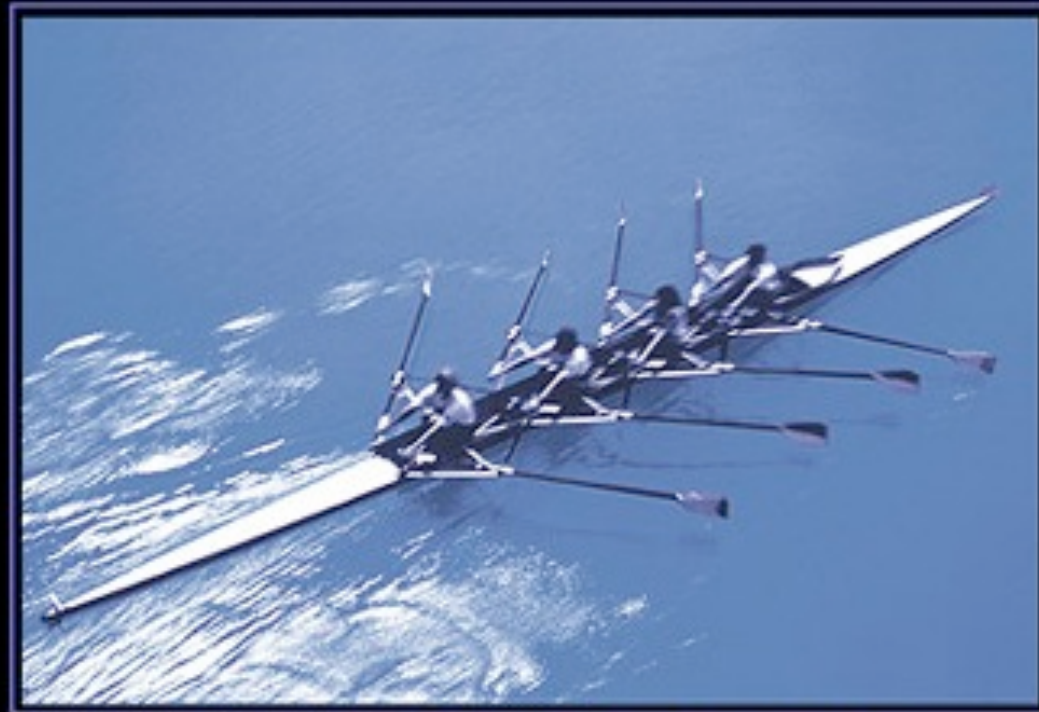


Issues

- Does our bank cover real signals?
- Does our veto veto real signals?
- Do we correctly compute efficiency?
- Can we do better?

Merger





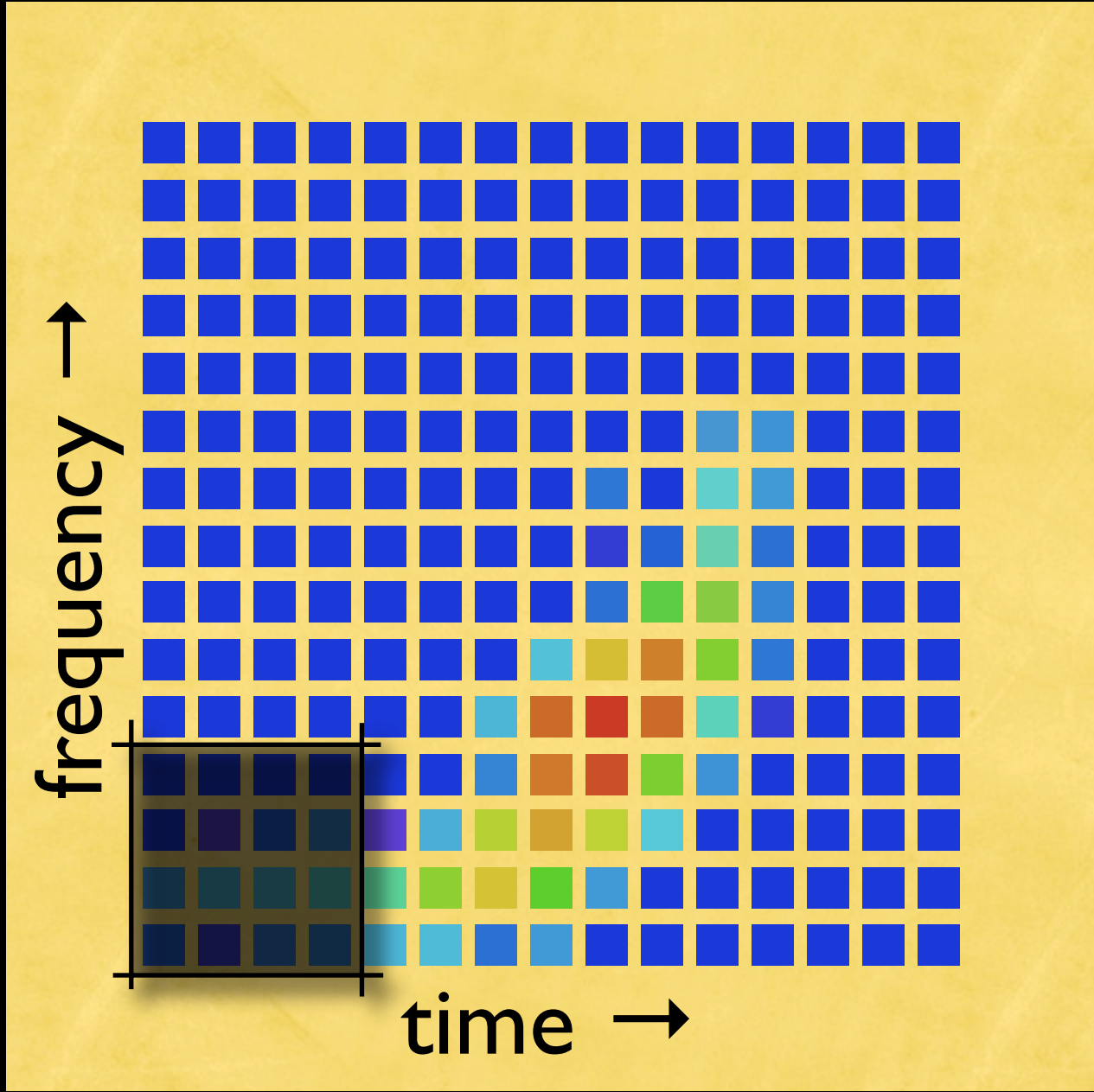
IGNORANCE

IT'S AMAZING HOW MUCH EASIER IT IS FOR A TEAM TO WORK TOGETHER
WHEN NO ONE HAS ANY IDEA WHERE THEY'RE GOING.

Unknown Burst Search

Unknown Burst Search

- Short burst — unknown signal shape
- Excess power marginalizes over unknown shape — optimal for known duration, band
- Just returns power in required time interval / frequency band
- Tricky bit: how to tie into preceding signal?





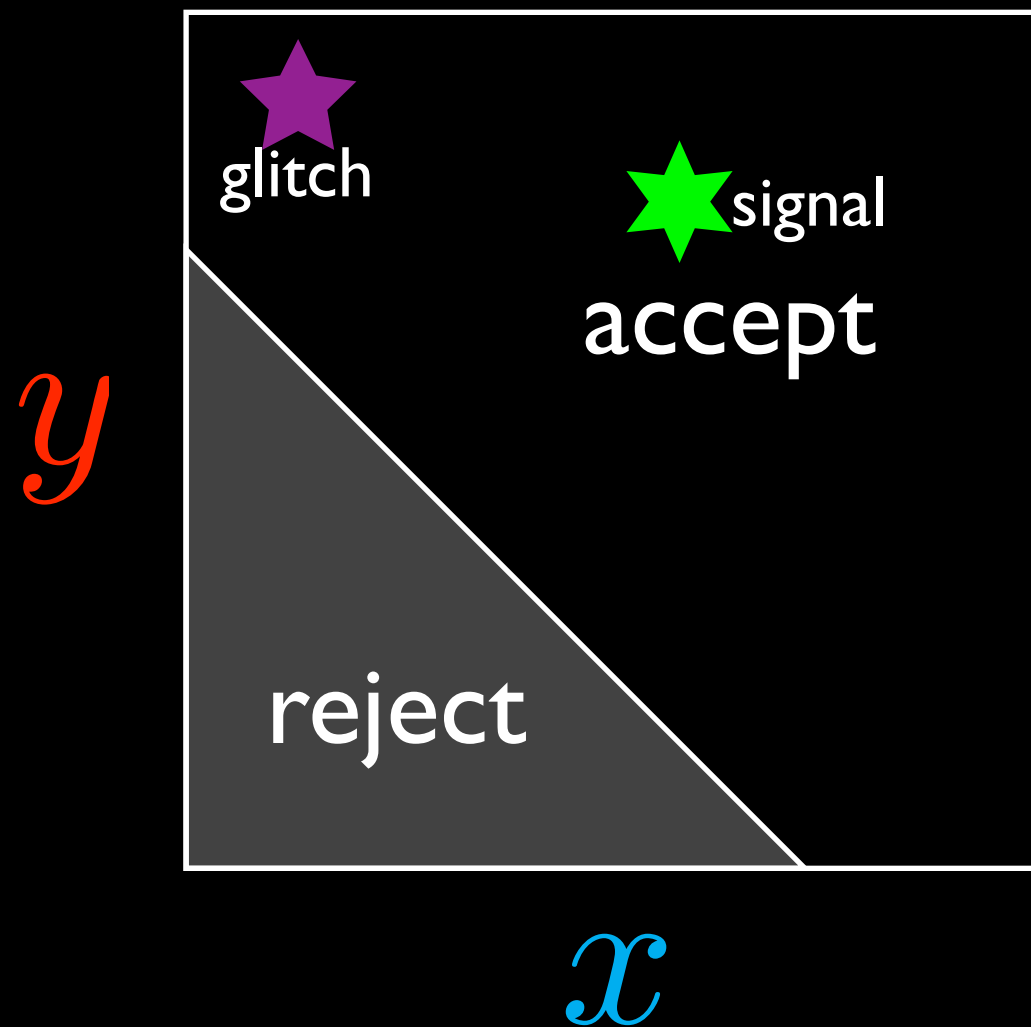
POWER

POWER CORRUPTS. ABSOLUTE POWER CORRUPTS ABSOLUTELY.
BUT IT ROCKS ABSOLUTELY, TOO.

Types of Power

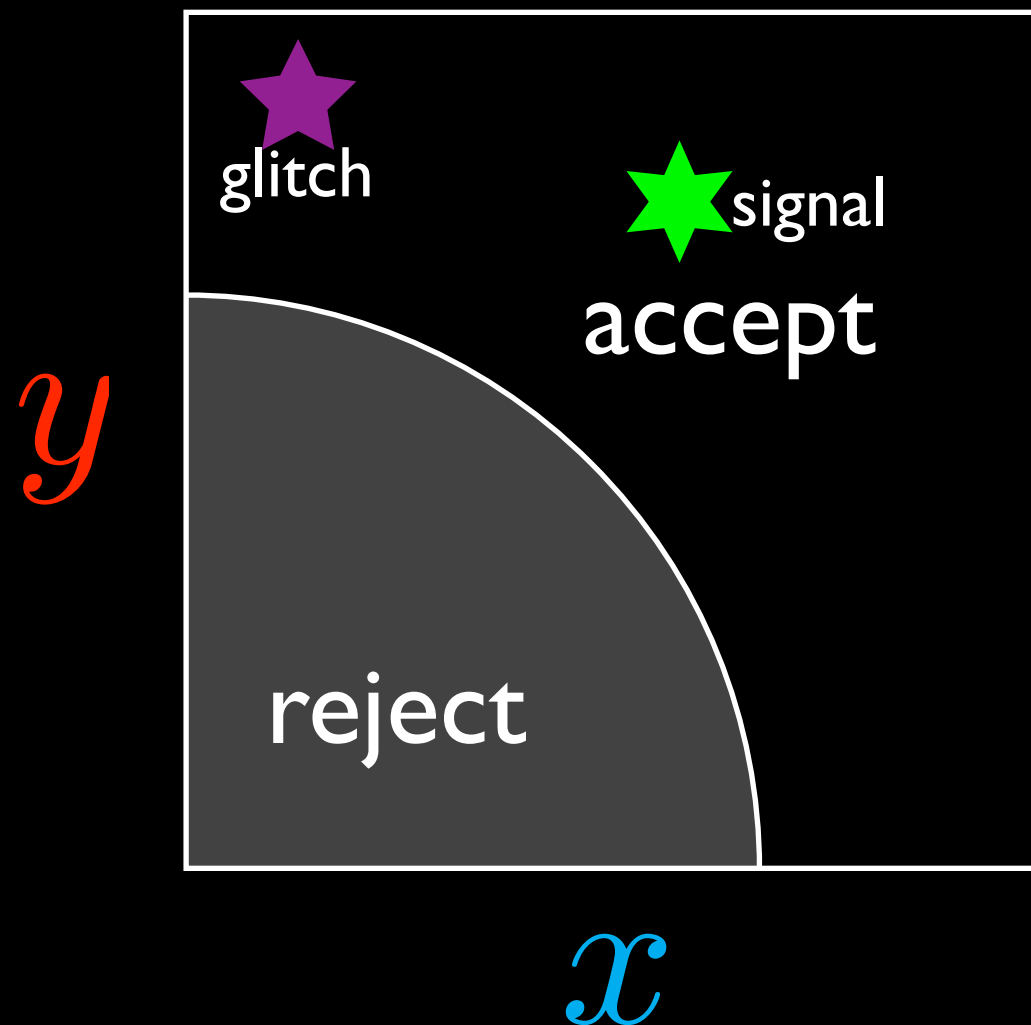
Coherent

$$(x + y)^2 = x^2 + 2xy + y^2 > \text{threshold}$$



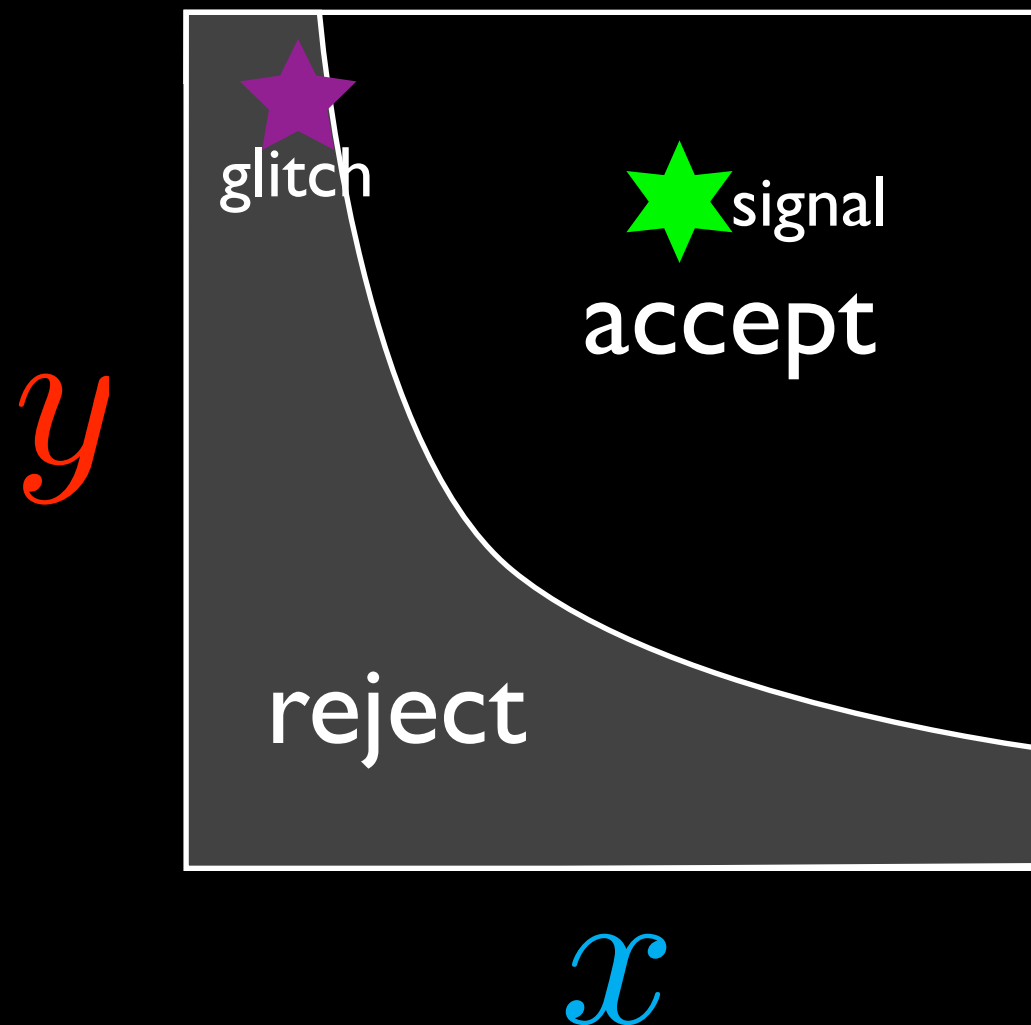
Incoherent

$$(x + y)^2 = x^2 + \cancel{2xy} + y^2 > \text{threshold}$$



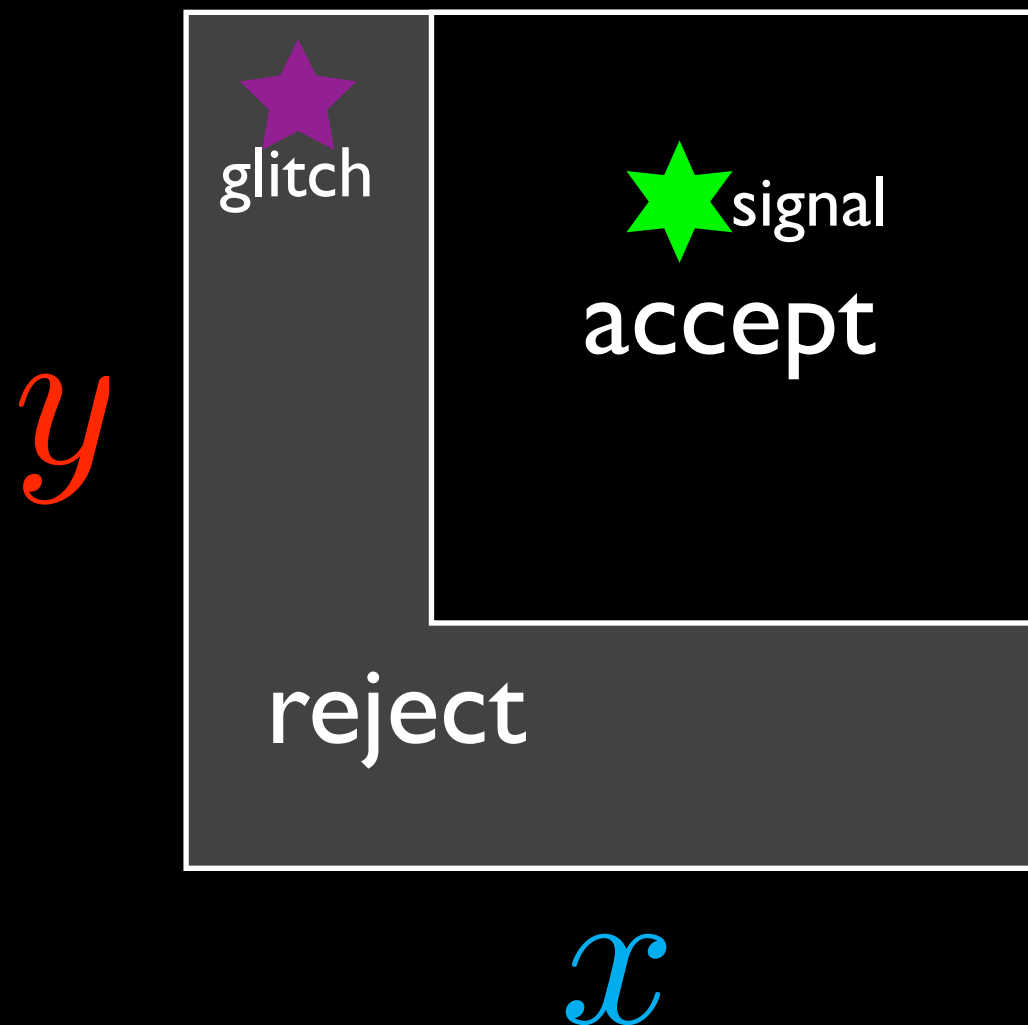
Correlated

$$(x + y)^2 = \cancel{x^2} + 2xy + \cancel{y^2} > \text{threshold}$$



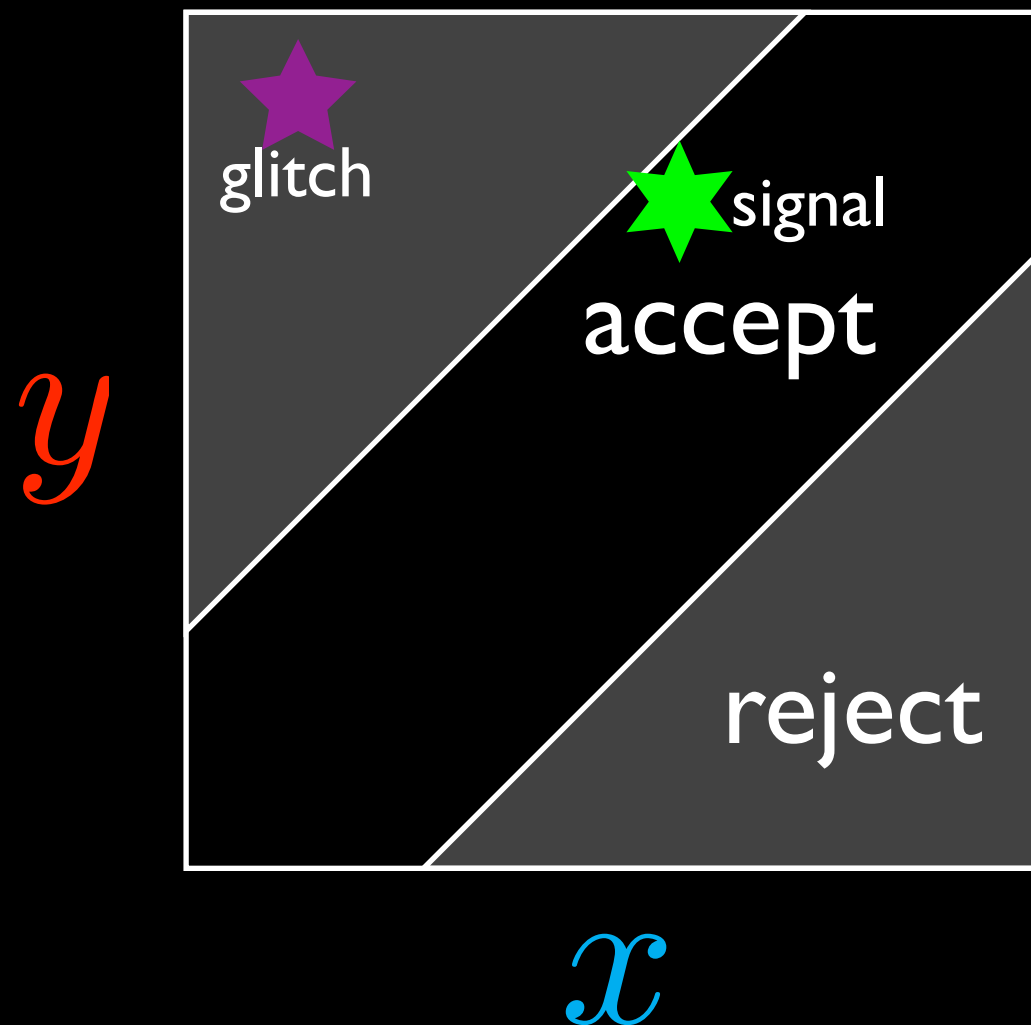
Coincidence

$$\min(x^2, y^2) > \text{threshold}$$



Null Veto

$$(x - y)^2 = x^2 - 2xy + y^2 < \text{threshold}$$



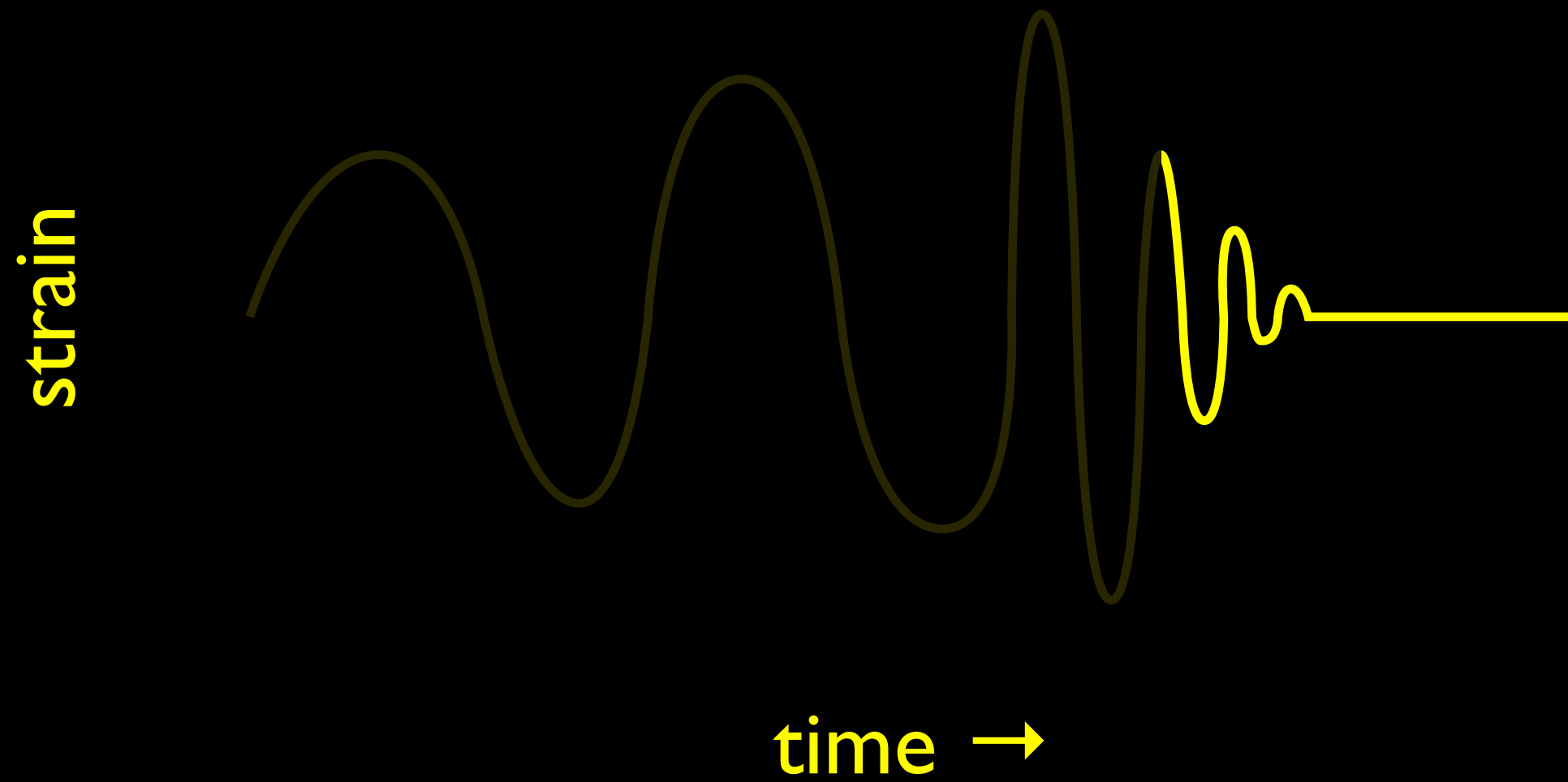
Excess Power Searches

- TFClusters (incoherent)
- WaveBurst (incoherent)
- Coherent WaveBurst (coherent)
- Corrpower (correlation)
- Q-pipeline (incoherent, null veto)
- X-pipeline (coherent)

Can We Do Better?

- More information on time-freq track
- Better guess of duration / freq-band
 - Tie into preceding inspiral waveform
- Sidebands?

Ringdown



Ringdown Search

- Very simple waveform — too simple!
- Matched filter available, but perhaps not much different from Excess Power
- Basically sees whatever is at 150 Hz
- How to tie into preceding waveform?

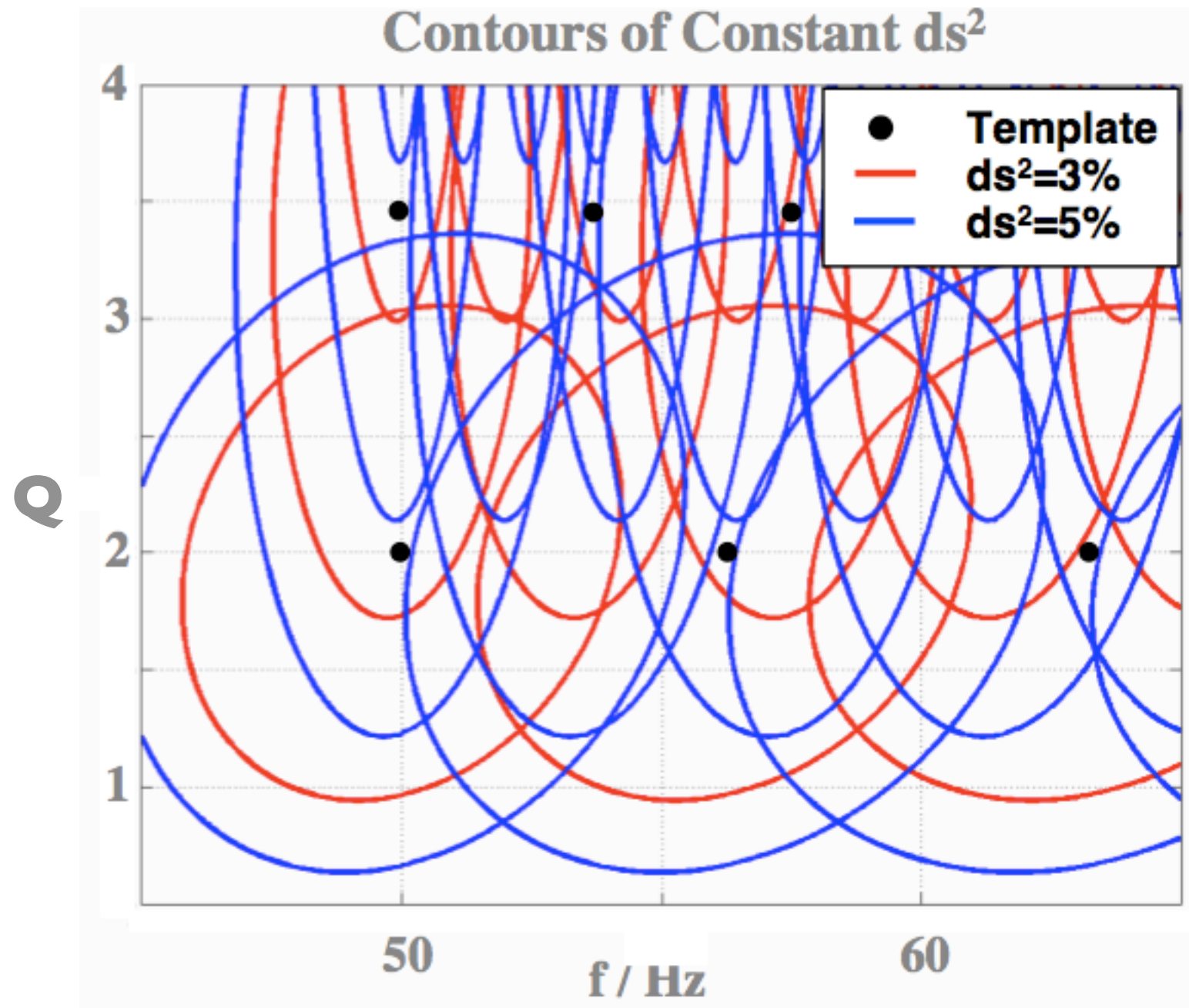


INDIVIDUALITY

ALWAYS REMEMBER THAT YOU ARE UNIQUE. JUST LIKE EVERYBODY ELSE.

Ringdown Morphology

Ring Template Bank



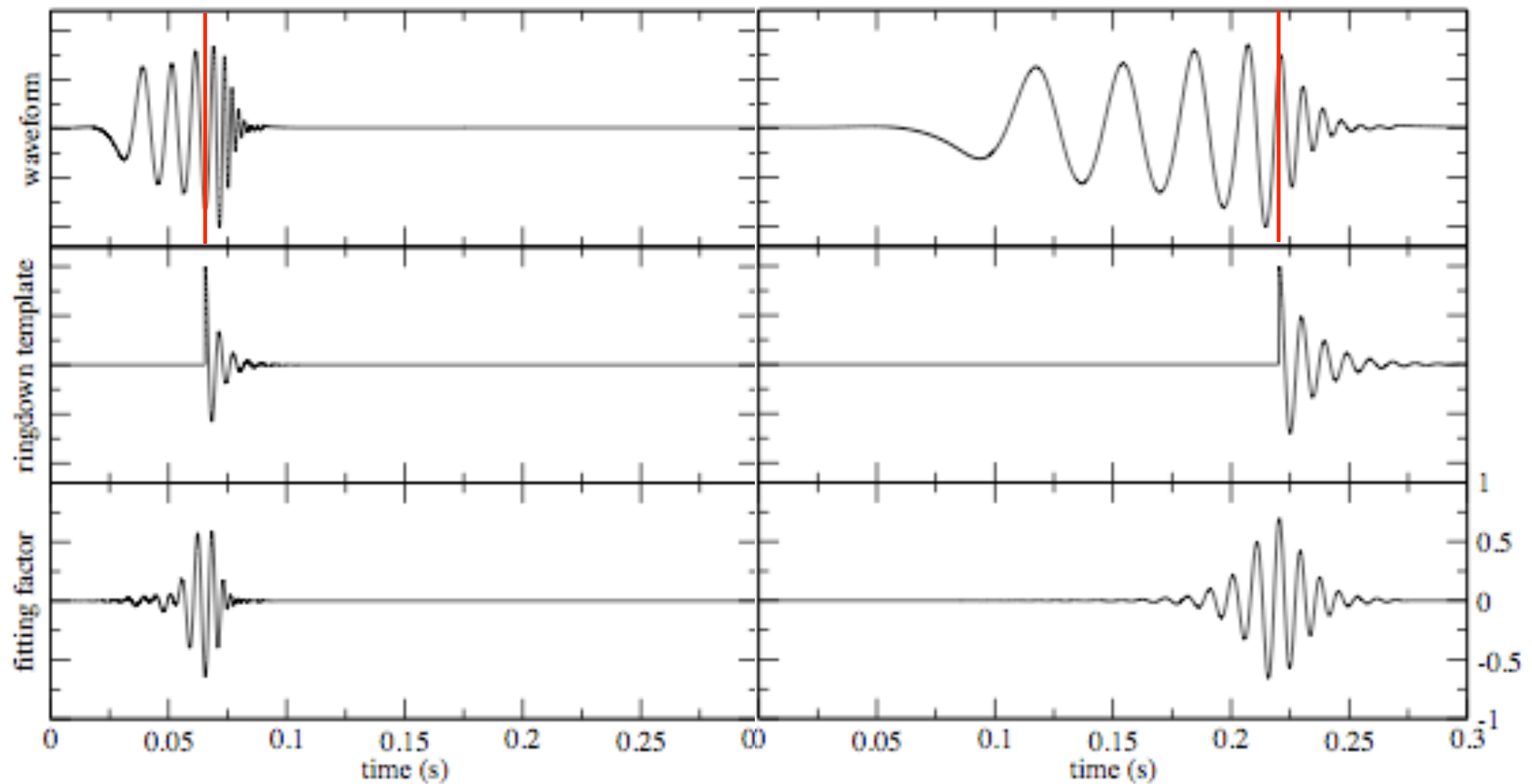
Goggin

Binary Black Hole Coalescence Waveforms

Baumgarte, Brady, Creighton, Lehner, Pretorius, DeVoe

$50 M_{\odot}$

$150 M_{\odot}$





HOPE

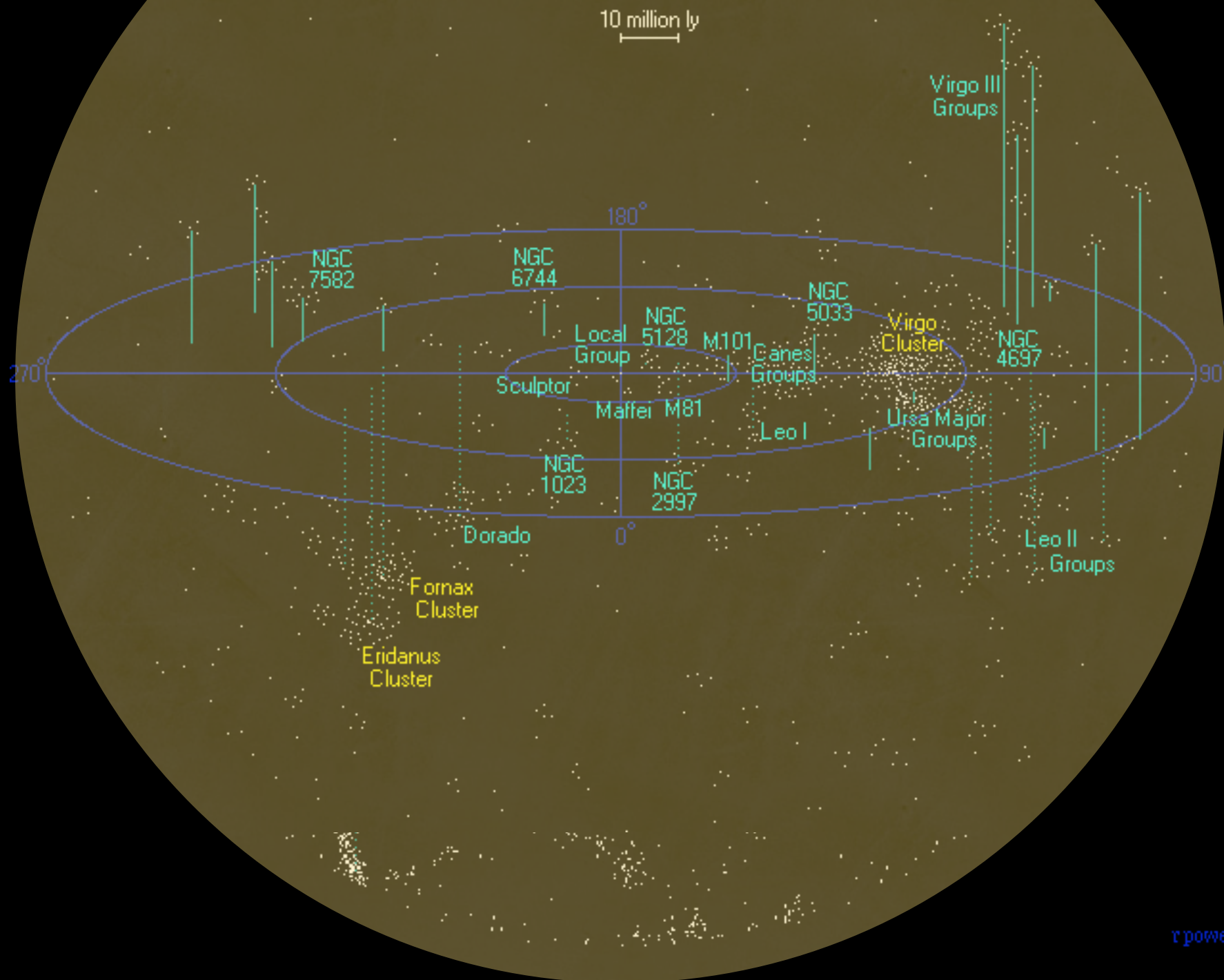
MAY NOT BE WARRANTED AT THIS POINT.

Outlook

LIGO Science Runs

- **S1**: 23 Aug – 9 Sep 2002 First Results!
- **S2**: 14 Feb – 14 Apr 2003
- **S3**: 31 Oct 2003 – 9 Jan 2004
- **S4**: 22 Feb – 23 Mar 2005
- **S5**: Nov 2005 – 1 Oct 2007 Design Sensitivity!
- Enhanced LIGO: 2009 – 2010 Factor of 2 better
- Advanced LIGO: 2013+ Factor of 10 better

Actual NGC (55)



Binary Neutron Star Range

Future...

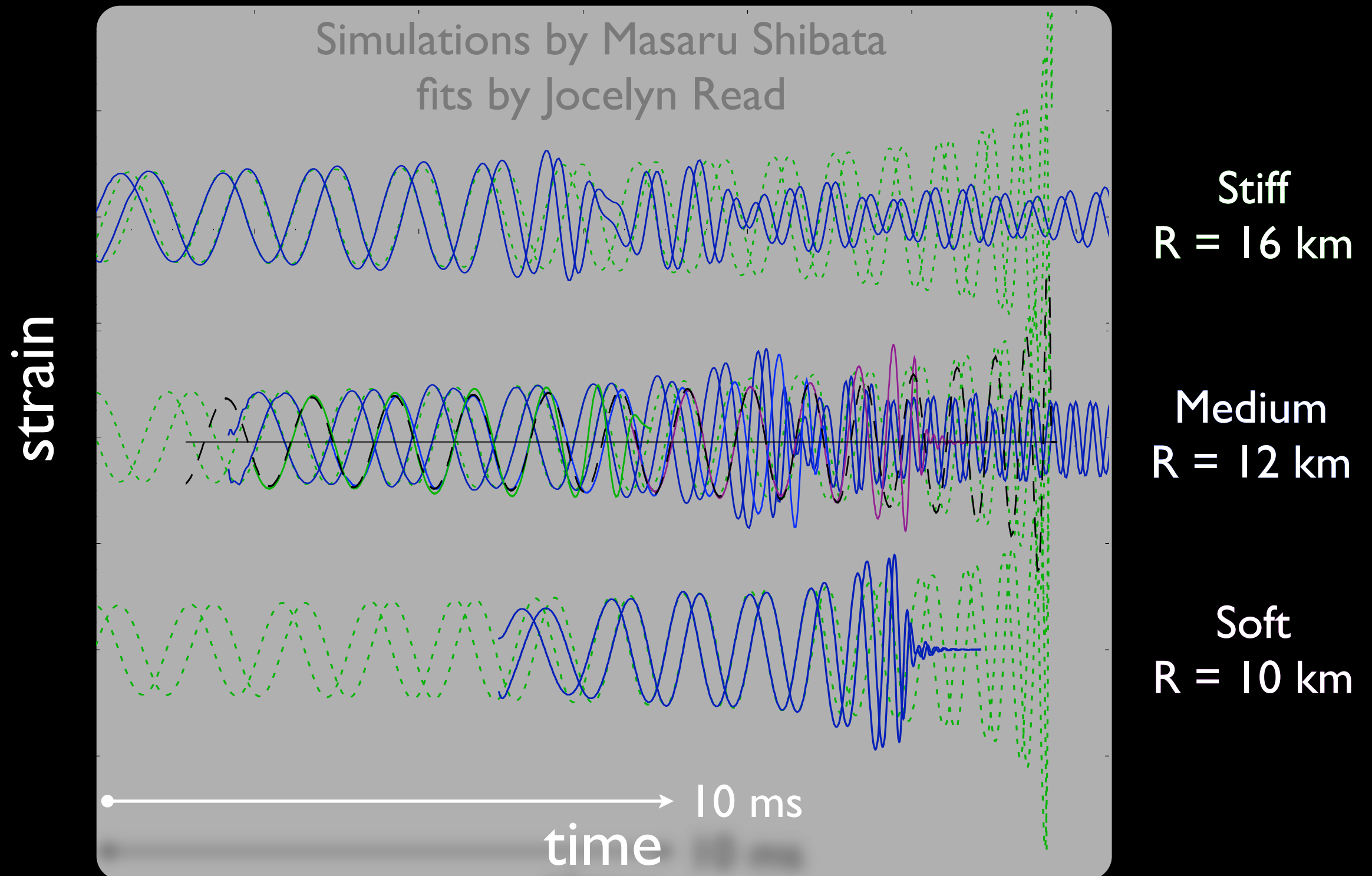
Parameter Estimation and
Astrophysics



INCOMPETENCE

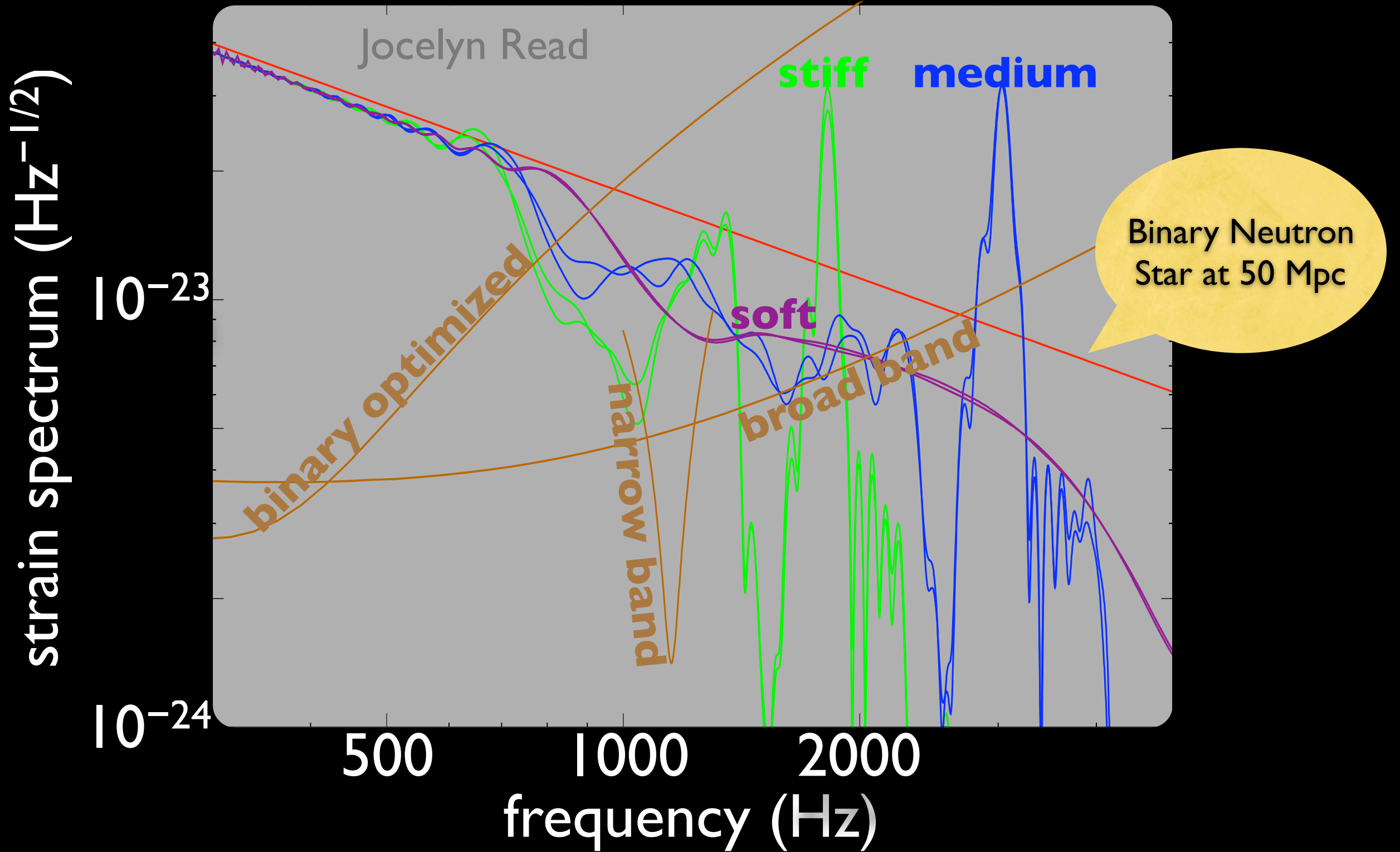
WHEN YOU EARNESTLY BELIEVE YOU CAN COMPENSATE
FOR A LACK OF SKILL BY DOUBLING YOUR EFFORTS,
THERE'S NO END TO WHAT YOU CAN'T DO.

Binary Neutron Star Coalescence Waveforms



Neutron Star Equation of State

Advanced LIGO Noise Curves



Conclusions

- Present: We want waveforms to
 - Improve/validate methods, hone vetoes
 - Determine efficiencies
- Future: We need waveforms to
 - Parameter estimation, astrophysics
 - Construct template banks?



MOTIVATION

IF A PRETTY POSTER AND A CUTE SAYING ARE ALL IT TAKES TO MOTIVATE YOU,
YOU PROBABLY HAVE A VERY EASY JOB. THE KIND ~~ROBOTS~~ WILL BE DOING SOON.
Computers

The End

Matched Filter Inner Product

$$\langle s|h \rangle = 4\Re \int_0^\infty \frac{\tilde{s}^*(f)\tilde{h}(f)}{S_h(f)} df$$

Fisher Matrix

$$\Gamma^{ij} = \left\langle \frac{\partial h}{\partial \lambda_i} \left| \frac{\partial h}{\partial \lambda_j} \right. \right\rangle$$

Measurement Accuracy

$$(\Delta \lambda_i)_{\text{rms}} = \sqrt{(\Gamma^{-1})_{ii}} \quad (\text{no sum})$$

One-Parameter Measurement Accuracy

$$(\Delta\lambda)_{\text{rms}} = \frac{|\lambda_1 - \lambda_2|}{\sqrt{\langle h_1 - h_2 | h_1 - h_2 \rangle}}$$

Fitting Factor

$$\text{FF} = \frac{\langle h_1 | h_2 \rangle}{\sqrt{\langle h_1 | h_1 \rangle \langle h_2 | h_2 \rangle}}$$