Stack-a-flare SGR burst search development

1.0.0

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Photo: Joe Becker



Use assumption of SGR burst similarity to improve search

Detection statement (as before)

Improved upper limits

Improved astrophysical context S5y1 SGR individual search should stimulate theorist interest

Perform the search with Stack-a-flare Initial Flare pipeline does heavy lifting Review should be quick



Parameter space choices for S5y1 SGR still apply

Ringdowns 1-3 kHz

Simulation frequencies: 1090, 1590, 2090, 2590 Hz 3 kHz upper bound: strange quark stars 1.5 kHz lower bound: lightweight star with stiff EOS [1] Simulation tau: 200 ms predicted range is 140-380 ms

[1] O. Benhar, V. Ferrari, and L. Gualtieri, Phys. Rev. D 70,124015 (2004)

WNB below 1 kHz

WNB injections to estimate upper limits; 11 ms and 100 ms durations Band-limited to detector's sensitive regions:

- 100 200 Hz (small band)
- 100 1000 Hz (large band)



SGR 1900+14 storm

BAT light curve with 100 us bins

Method for matching rising edge of bursts may be helpful

N = number of bursts to include
How to pick burst set to stack?
How to weight each burst?
How to align bursts?
Are there 2 distinct burst types here?
Theorist involvement could help





Method 1: T-Stack



Pros:

Greater potential sensitivity. Stacking *amplitude*

expect ~N^1/2 amplitude sensitivity dependence in WN

Cons:

Either precise timing, or expensive time shift combinatorics, is needed Sensitive to relative sign between detectors



Method 2: P-Stack



- 1. Time-align to signals & apply Flare N times
- 2. Add up resulting Power matrices

Pros:

Suited for stochastic (WNB) SGR search Less timing precision needed

Cons:

Less sensitive. Stacking *power*

expect ~N^1/4 amplitude sensitivity dependence in WN





Characterize:

sensitivity dependence on N sensitivity dependence on $\sigma_{\Delta T}$

Shift component injections by small times **ΔT** simulates timing imprecision subsample shifting

Two simulation designs tell consistent story Rough simulated SGR 1900+14 storm 18 identical ringdowns into **σ=1** WN **N** evenly spaced injections (RDs or WNBs)





RD Stack-of-N



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WNB Stack-of-N



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Ringdowns: 1090 Hz (left) 2590 Hz (right) P-Stack (red) is insensitive to time shifts (so long as signal TF pixels overlap) T-Stack is sensitive. Crossover points are given

hrss50 story looks very similar



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Precise trigger timing means smaller on-source region

[-1000,1000] second background region as before (cat2 DQ) estimate $\mu(f)$, $\sigma(f)$ used by Flare pipeline estimate local false alarm rate (FAR) background data will need to be stacked

Follow up on-source analysis events with significant FAR as before

Loudest event upper limits as before



Timing precision makes T-Stack unfeasible

P-Stack requires:

Propagation of all light crossing times to detectors

Light curves could yield precise burst times

Weight each burst according to antenna pattern before stacking

How to combine bursts All bursts from all SGR sources? All bursts from particular SGR sources? Brightest N bursts? Weight each burst according to fluence? Theorist input required



Future of Stack-a-flare



P-Stack

storm unmodeled search, isolated bursts

T-Stack

Requires timing precision of ~50 us or better to be worth it RD search only

Concentrate first on storm

T-Stack may be possible

P-Stack S5 closed box results soon (stay tuned)

Time-to-publish should be small, thanks to Flare review

People:

S5y1 search reviewers (Kipp, Ben) have agreed to review Max Factourovich, Columbia grad student

Experience from storm analysis will aid isolated bursts search