



Optical properties of silicon at low temperatures

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Outline

- motivation
- temperature dependence of the refractive index of silicon
 - measurement principle and experimental setup
 - acquired data
 - results
- outlook : absorption measurements

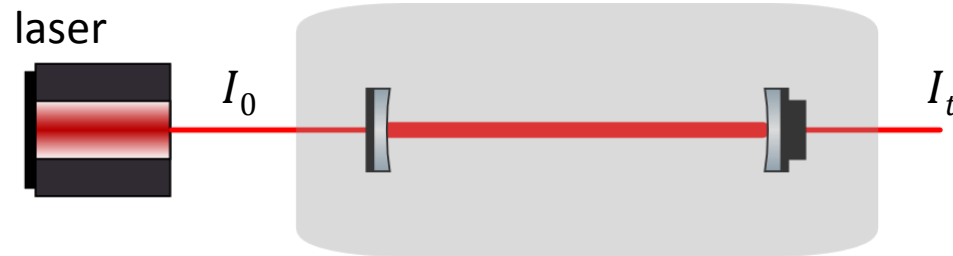


Motivation

- silicon is one material of interest for optical components for future detectors
- n , dn/dT for low temperatures are unknown, but needed for noise calculations
 - for example in the ET design study there are only extrapolated values for temperatures below 30 K
- for the design of the cryogenic parts of future detectors the optical absorption is needed (thermal equilibrium)



Fabry Perot Cavity (FPC)



- for a FPC the reflected light is:

$$I_r = \frac{I_0}{1 + \frac{\pi^2}{4F^2} \sin^2 \theta}, \quad \theta = \frac{2\pi n l}{\lambda}$$

l...cavity length, F ... finesse of the FPC, λ ...wavelength



Fabry Perot Cavity (FPC)



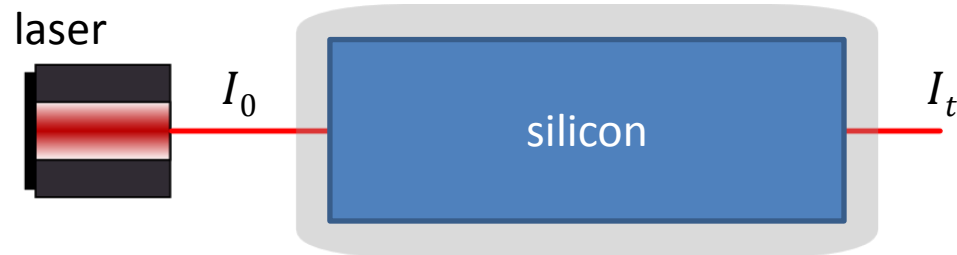
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$$I_r = \frac{I_0}{1 + \frac{\pi^2}{4F^2} \sin^2 \theta}, \quad \theta = \frac{2\pi n l}{\lambda}$$

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Fabry Perot Cavity (FPC)



- for a FPC the reflected light is:

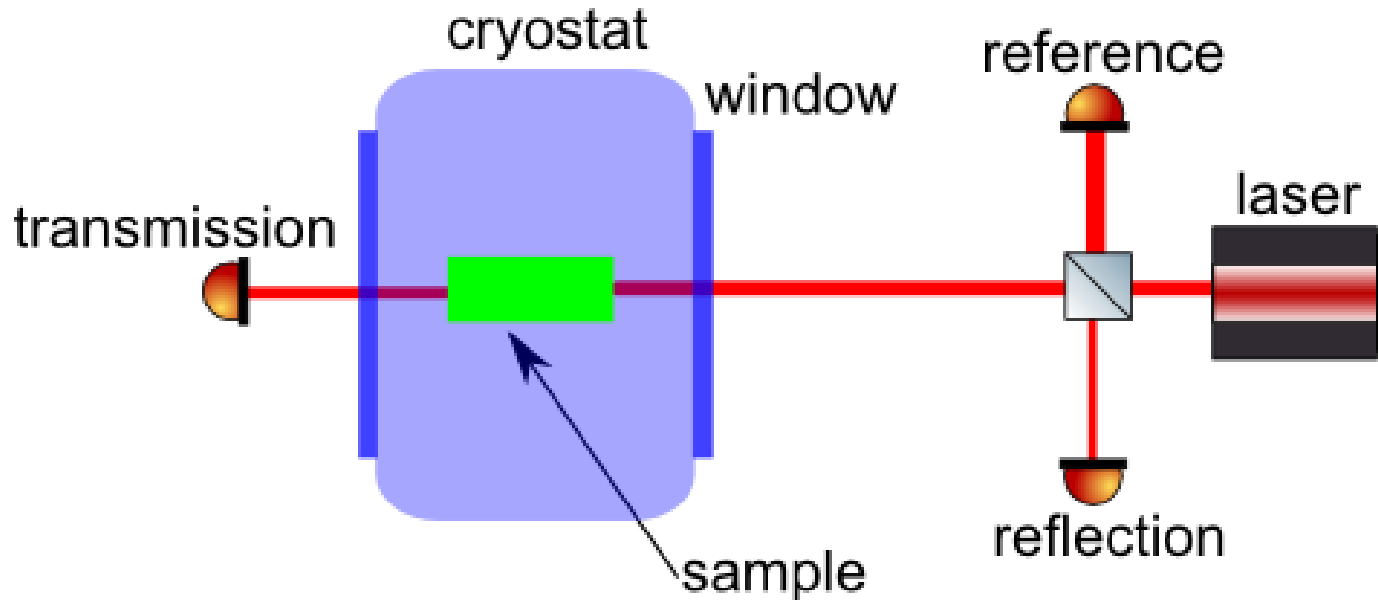
$$I_r = \frac{I_0}{1 + \frac{\pi^2}{4F^2} \sin^2 \theta}, \quad \theta = \frac{2\pi n l}{\lambda}$$

optical path length
 geometric path length

l...cavity length, F ... finesse of the FPC, λ ...wavelength



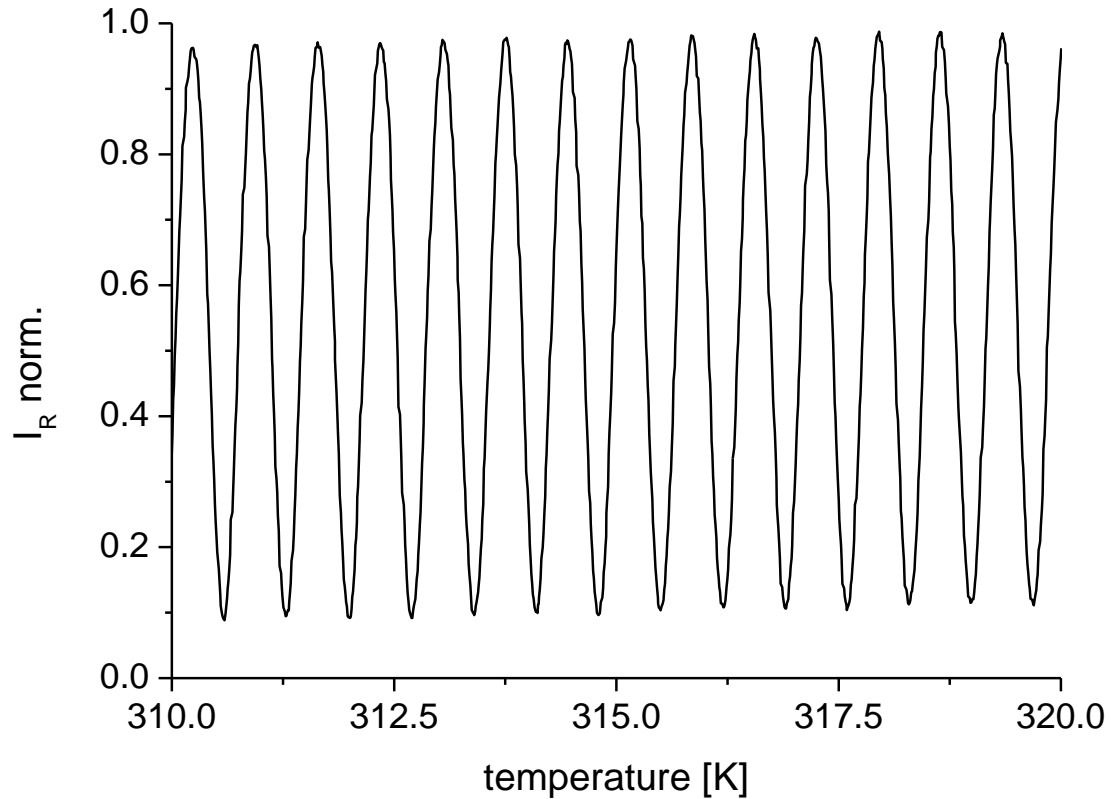
Setup for dn/dT measurement



- with the continuous helium flow cryostat it is possible to measure from <4 K up to 325 K
- sample thickness: 0.3 ... 14 mm



Data at room temperature



typical data from the measurement at RT



Extracting of dn/dT from exp data

- remember: $I_r = \frac{I_0}{1 + \pi^2 / 4F^2 \sin^2 \theta}$, $\theta = \frac{2\pi nl}{\lambda}$
- for the analysis θ is the term of interest.
- $\frac{\partial \theta}{\partial T} = \frac{2\pi}{\lambda} \left(l \frac{\partial n}{\partial T} + n \frac{\partial l}{\partial T} \right) = \frac{2\pi l}{\lambda} \left(\frac{\partial n}{\partial T} + n(T)\alpha(T) \right)$

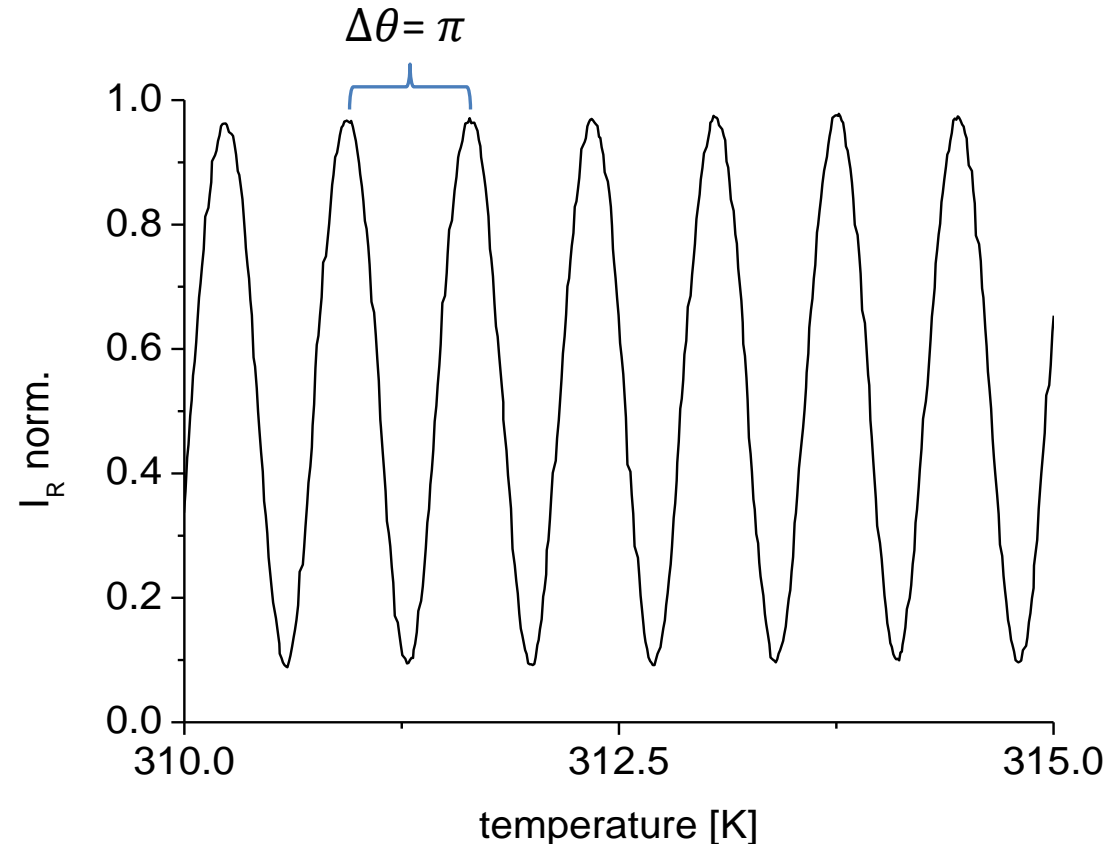


Extracting of dn/dT from exp data

interference maximum $\leftrightarrow \Delta\theta = \pi$

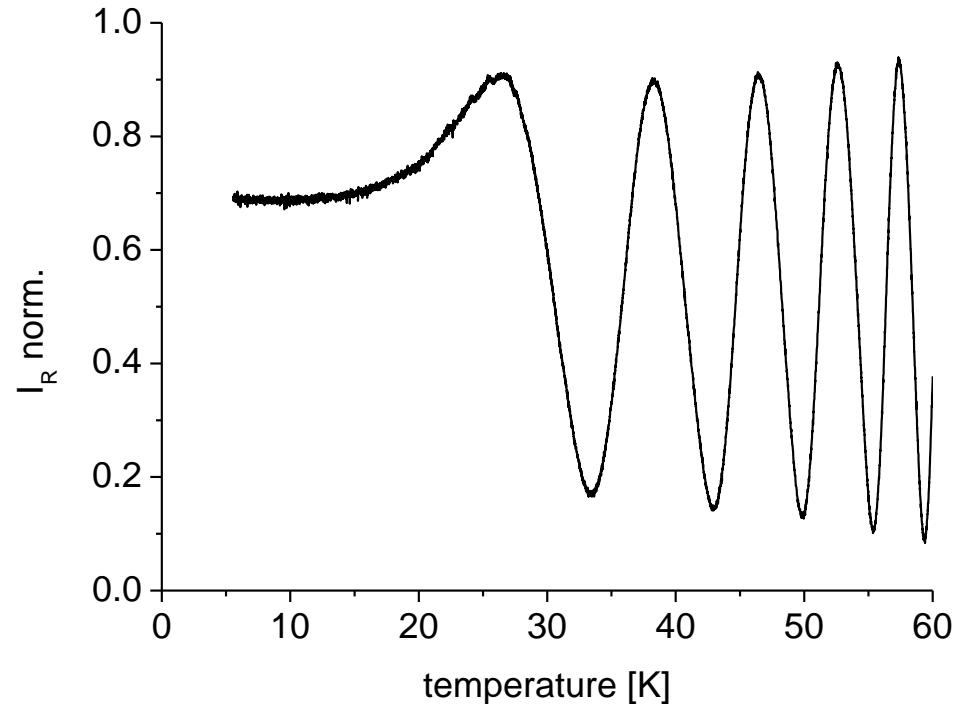
$$\rightarrow \frac{dn}{dT} = \underbrace{\frac{\lambda}{2l \Delta T}}_{\approx 1 \times 10^{-4} \text{ 1/K}} - \underbrace{n(T)\alpha(T)}_{\approx 1 \times 10^{-6} \text{ 1/K}}$$

$(l = 14 \text{ mm}, T = 300 \text{ K})$





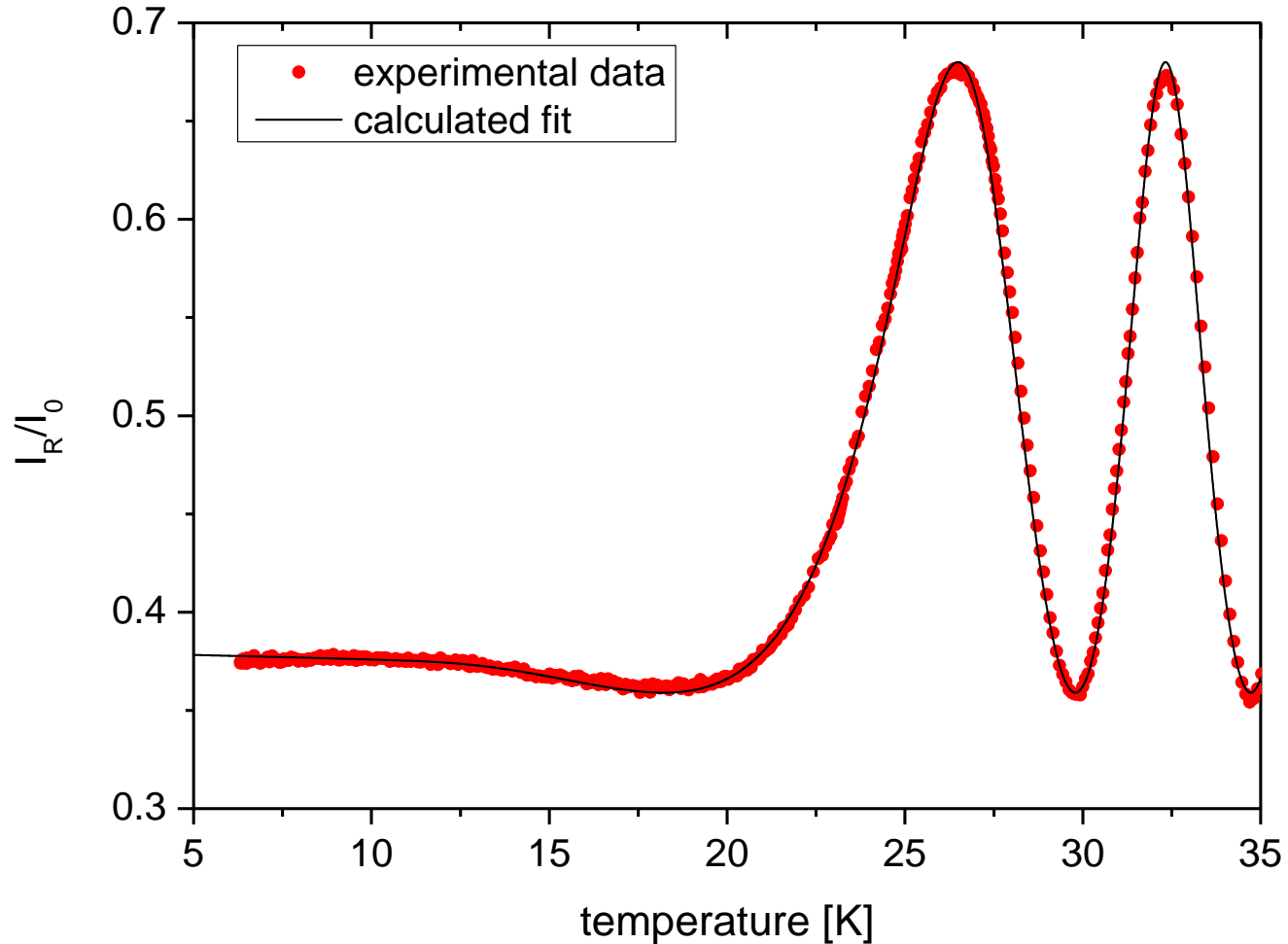
Data at low temperatures



- problem: only a small change of n at low temperatures \rightarrow only a few fringes
- $dn/dT \rightarrow 0$ for $T \rightarrow 0$ (3rd law of thermodynamics)

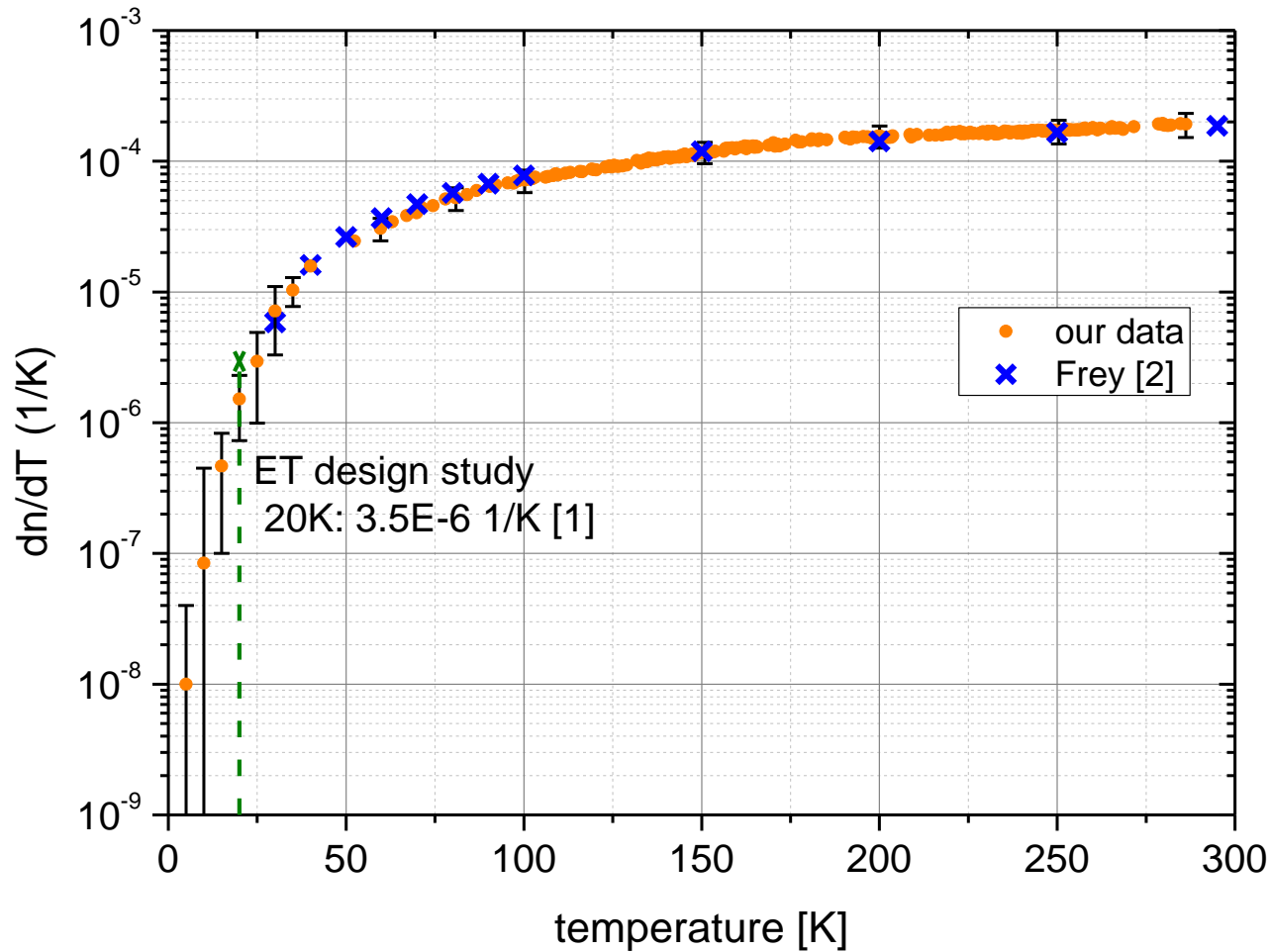


Extracting of dn/dT from exp data





Results for dn/dT

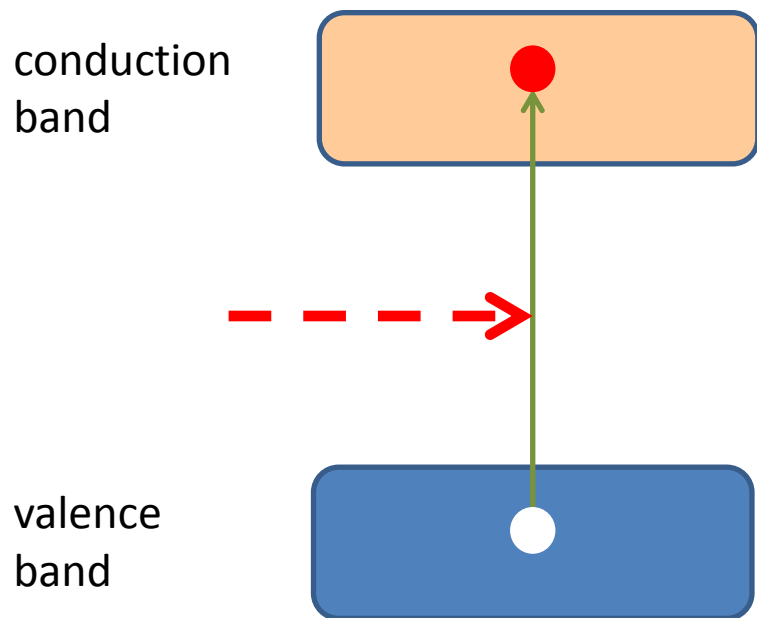


[1] The Einstein gravitational wave Telescope conceptual design study
M. Abernathy et al.

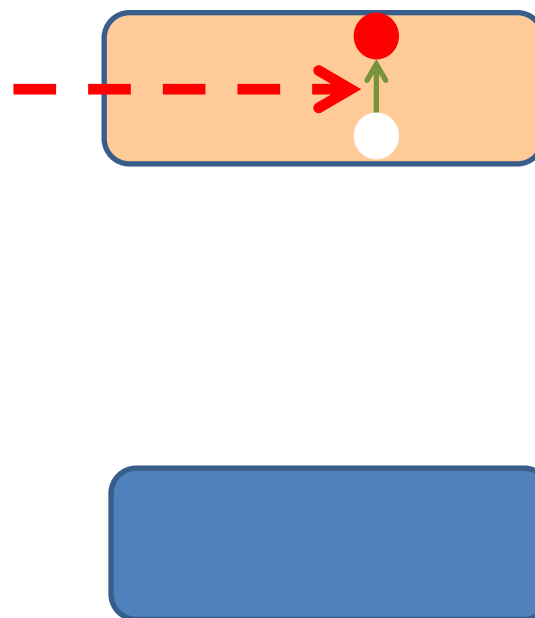
[2] **Temperature-dependent refractive index of silicon and germanium**
Bradley J. Frey et al.



Band-band and free carrier absorption



inter band absorption
(band-band abs.)

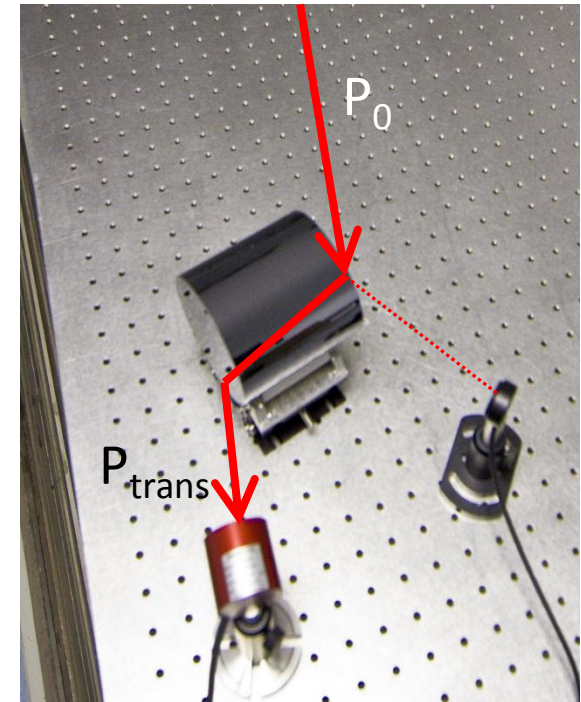
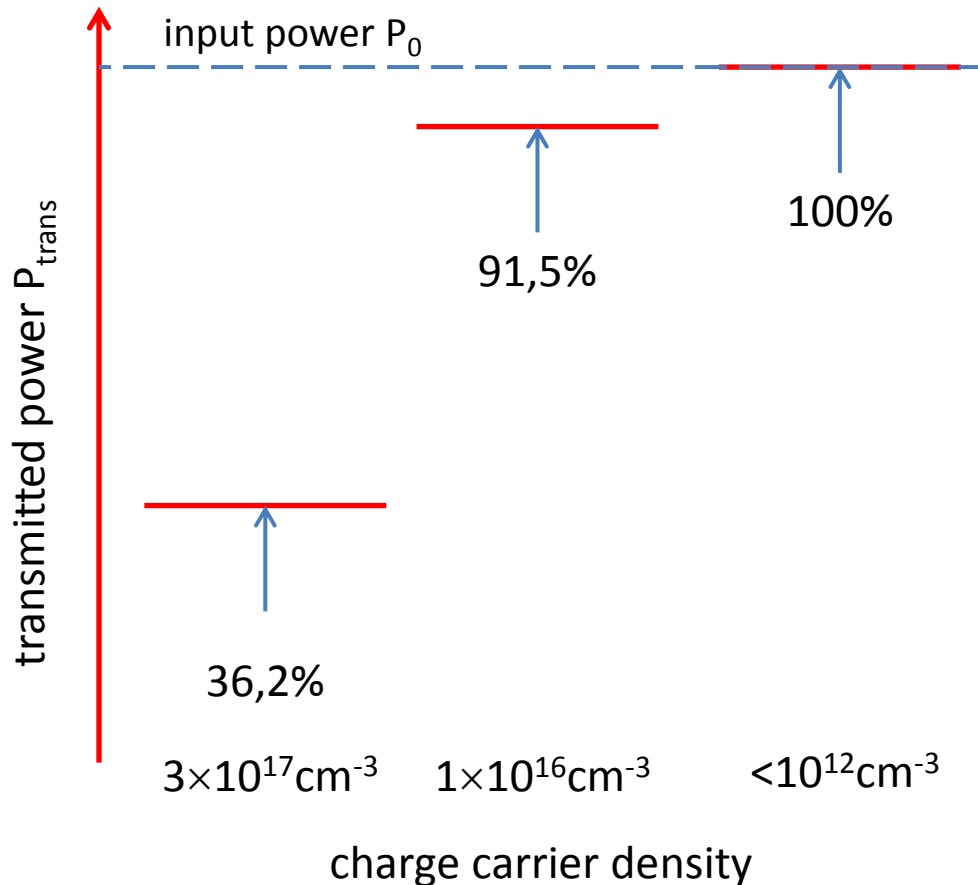


intra band absorption
(free carrier abs.)



Absorption measurement in brewster angle

- room temperature test for the absorption of different Si samples





Summary

- dn/dT was measured over a wide temperature range (5 ... 300 K)

$$\frac{dn}{dT} = 1.5 \times 10^{-6} \text{ 1/K @20K and } 8 \times 10^{-8} \text{ 1/K @10K}$$

- experimental values are in agreement with values assumed for the ET design study
- free carrier absorption is the dominant process in silicon @ 1550 nm and 300 K → high purity samples needed (free carriers freeze out at cryogenics)

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