

LIGO Laboratory / LIGO Scientific Collaboration

LIGO-E1400310-v1

LIGO

24 Jun 2014

aLIGO Availability Estimate

Dennis Coyne

Distribution of this document: LIGO Scientific Collaboration

This is an internal working note of the LIGO Laboratory.

California Institute of Technology LIGO Project Massachusetts Institute of Technology LIGO Project

LIGO Hanford Observatory

LIGO Livingston Observatory

http://www.ligo.caltech.edu/

1 Introduction

This is a preliminary estimate of the availability (or uptime) of the Advanced LIGO (aLIGO) interferometer. Updates to this availability estimate will be made as the system FMEA and reliability assessment is completed, and as more operational experience is gained.

2 Applicable Documents

<u>T080160</u> -v1	D. Hoak, B. Bland, "Downtime Accounting in S5", Jul 2008
<u>T1300519</u> -v3	D. Coyne, "Sparing Analysis",
<u>T010075</u> -v2	P. Fritschel, "Advanced LIGO Systems Design"

3 Scope

The scope of this availability estimate is a single aLIGO interferometer (not an estimate of coincidence operation, yet).

This estimate does not (yet) address the impact of non-interferometer, observatory systems on the system availability. For example facility systems such as HVAC and vacuum equipment. The experience from iLIGO/eLIGO is that these systems do not significantly impact the interferometer uptime (see for example T080160).

4 **Previous Experience**

Some of the experience for Initial and Enhanced LIGO (iLIGO and eLIGO), in particular the S5 and S6 science runs are potentially relevant to anticipated aLIGO operations. Good downtime statistics have been assembled for the S5 run (T080160), but appear to be mostly lacking for the S6 run. (It is worth noting that LIGO operations intends to implement better operational metrics tracking that have been in place in the past.)

5 Terminology

The terminology for the categories of the causes for interferometer downtime (unavailability) used here are very similar to those used in T080160.

5.1 Uptime

Uptime is time the interferometer is in science mode. Science segments of any length are counted.

5.2 Calibration

Calibration is time spent performing calibration studies. These are done periodically throughout the science runs. The same percentage time for iLIGO/eLIGO should apply to aLIGO.

5.3 Hardware Failures

Hardware failures result in downtime due to needed repair or replacement of equipment. Examples are broken seismometers and electronics modules. The repair time includes diagnosis and isolation of the faulty equipment and the subsequent time to repair or replace and verify function (including reboots to computer equipment if appropriate), as well as the time required to bring the interferometer system back to low noise operation again.

Also included in this category is power grid outages.

5.4 Software Failures

Software failures result in downtime due to faults in the software code. Examples are memory leaks and inadvertent triggers of software watchdogs (i.e. when not appropriate due to environmental conditions or equipment faults). The repair time includes diagnosis and re-start/re-boot of processes/computers as appropriate, as well as the time required to bring the interferometer system back to low noise operation again. The time required to isolate of the software fault, or define more appropriate parameter values (e.g. watchdog trigger levels), and then revise and regression test a new version of the code is assumed to be done off-line.

5.5 Maintenance

Maintenance is downtime for planned detector maintenance, including scheduled repairs for failures which do not require prompt repair or replacement. During S5 & S6, four hours were scheduled every Tuesday for invasive, periodic maintenance tasks, such as liquid nitrogen deliveries, grounds-keeping, HEPI pump filter replacement, etc. This period is also used for scheduled detector maintenance such as PSL pump diode replacement, HWS SLED replacement, OptLev laser replacement, etc.

5.6 Commissioning

Commissioning is downtime for planned detector improvements. There were a few extended commissioning breaks at both sites during S5 and S6; these lasted for about a week and occurred once or twice a year. In addition, about 25 hours were allocated every month in S5 for discretionary studies.

5.7 Transition to Low Noise Operation State

In the S5 downtime accounting document (T080160), the "scripts" category covered time spent running the "up" and "down" software scripts that take the instruments from lock-acquisition mode to the low noise detection mode. This required about ten minutes, each time the interferometer lost lock. The Automation System, Guardian, employed for aLIGO serves a similar function, but is more robust and should be more efficient, once the system is fully commissioned.

5.8 Wind

The category "wind" includes downtime due to high winds. At Hanford, for S5 and S6 this is typically wind over 25-30 miles per hour.

In S5 and S6 at Livingston, the seismic pre-isolation system is very sensitive to the low-frequency ground motion caused by wind. Typically, Livingston could not operate in science mode when the

wind was gusting more than 15mph. (At Livingston, this category also includes downtime due to Storms.)

5.9 Seismic

For Hanford, the Seismic category includes any source of ground motion that prevents the interferometer from normal operations. Mostly, this downtime is due to earthquakes. The seismic environment is more complicated at Livingston, and this category is expanded into different sources:

Microseism is the term used to describe the low-frequency (0.1-0.35 Hz) ground motion caused by ocean waves striking the continental shelf. This is worst in the winter months, due to storms in the Atlantic.

Earthquakes produce large amplitude, low frequency ground motion (0.03-0.1 Hz), although higher frequencies can be driven depending on the earthquake's size and location. Typically, a 6.0-magnitude earthquake on the western Pacific Rim (Russia, Japan, the Philippines, or Indonesia) kept the Livingston interferometer out of lock for one or two hours in S5 and S6. Events like these happen about once every two days.

Trains are a phenomenon unique to Livingston. Railroad tracks pass by the site about 1.75 miles south of the ETMY end station, and freight trains move through Livingston three or four times a day. The largest trains occur regularly around 2am and around 7am, and typically take the interferometer out of low-noise operations for about half an hour.

6 aLIGO Availability Estimate

6.1 Uptime

The estimated uptime for aLIGO is 79%, as indicated in the Table and Figure below. Each of the factors that are part of this estimate are discussed in the sections below.

LIGO-E1400310-v1



		S5 H1	S5 L1 *	S5 L1 **	aLIGO				
Category		(T080160)	(T080160)	last 2 months	Estimation	aLIGO comments			
Uptime		76.9%	70.9%	76.1%	78.6%				
Calibration		0.5%	0.5%	0.5%	0.5%	same as for S5			
Hardware Failures					5.1%	Rate for aLIGO is based on failure rates (from T1300519, Sparing Analysis) for			
		3 0%	3.9%	2.7%		prompt (unscheduled) attention			
		5.976				and LLO power outage rate for last year with recovery WAG			
Software	Failures				0.5%	WAG; should be low after commissioning			
Maintena	nce	6.2%			8.1%	corresponds to 4 hr/wk, plus one 3 wk vent&pump period per yr			
			8.3%	6.1%		Note that for aLIGO the ISC in-vacuum output electronics/detectors are in			
						vacuum volumes separate from the main vacuum volume			
Commissi	oning	4.3%			0.0%	projection to steady-state operations; no commissioning			
Transitior	n to Low Noise	2.0%	2.8%	2.2%	1.1%				
Operation	n State					15 min every 24 hr. with Guardian automation system			
(Up/Dow	n Scripts)					(less rms motion and better automation)			
Wind		2.0%	1.7%	0.2%	2.0%	assume that we don't improve on sensitivity to wind-induced ground tilf			
						senstivity			
	microseism		2.4%			should have better isolation from earthquakes and microseism			
Seismic	trains	2.2%	2.4%	2.3%	2.2%	assume better for L1, but same as H1 S5			
	earthquakes		4.2%	6.5%					
Unknown		2.0%	2.7%	3.4%	2.0%	same as H1 S5			
		100.0%	100.0%	100.0%	100.0%				

* renormalized to remove the exceptional time on L1 for SEC construction and ITMy repair due to SEC construction

** calibration time assumed to be part of the unknown/other category. Unusually high earthquake activity in this period, but no microseism induced downtime reported.

6.2 Hardware Failures

The impact of availability due to hardware failures was estimated by taking the list of Line Replaceable/Repairable Units (LRUs) that was provided by the subsystem technical experts for each subsystem (and documented in T1300519, Sparing Analysis) and then projecting out the following three categories of failures:

- Repair/replacement of in-vacuum failed equipment [assigned to a once per year in-vacuum incursion event with a duration of 1 week duration, plus a 2 week pump-down period]
- Repair/replacement of in-air equipment which require prompt attention (i.e. failures which impact availability or performance, for which we have no credible early detection/forewarning of the failure)
- Repair/replacement of in-air equipment which can be scheduled (i.e. for which we have an early detection/forewarning of the failure) [assigned to be addressed in the weekly maintenance period]

For each of the failures, an estimate of the time required to diagnose, isolation, repair or replace and confirm function/operation has been made.

An estimate of the power outage rate was obtained by looking through the electronic logs for the period 6/1/2013 through 6/1/2014. Even short power "glitches" lasting a few seconds causes many of the interferometer detector systems to shut down and require a careful sequence of re-booting. For L1 there were 16 power outage events in the last year, of which only 5 had significant duration (many minutes or more). For H1 there appear to have only been 4 events.

It is currently difficult to estimate the duration of the impact of a power outage, beyond the duration of the outage itself. Here I have assumed a minimum 4 hour impact (in steady-state operations, once all procedures for re-start are well established). With this assumption, based on the last year of L1 experience, the total power outage impact is 83 hours, or 0.95%. This motivates implementation of adding UPS systems to all interferometer detector systems (not just the mass storage room computers); This is an item on the potential liens on aLIGO project contingency.

LIGO-E1400310-v1

LIGO

			Frequency	Duration		
#	Subsys	Event	(#/yr)	(hrs each)	Downtime	Comments
						once/year vent, parallel activities;
1	All	In-vacuum maintenance	1	504	5.75%	1 wk vent plus 2 wk pumpdown
2	All	weekly maintenance period	52	4	2.37%	All deferred/planned maintenance
					8.13%	

			Frequency	Duration			
#	Subsys	Maintenance/Repair	(#/yr)	(hrs each)	Downtime	Event	Comments
							duration is set by vertex optics set
1	сос	Clean optics accumulated hydrocarbon film off of optics	1	8	0.09%	1	ETMs cleaned (if needed) in parallel
		electronics failures, incl. all of the I/O chassis and real-					
2	DAQ	time/DAQ computers used by aLIGO	154	1	1.76%	NA	individual, separate events; no early detection
		clean or replace degraded opto-electronics on the PSL/IO					clean or replace when convenient during weekly
3	10	optics table	8.19	4	0.37%	2	maintenance period
4	10	replace, align and test EOM when it fails	0.365	4	0.02%	NA	
5	10	replace, align and test slow & fast PDs when they fail	0.526	4	0.02%	NA	
6	10	replace, align and test HP beam dump when it fails	0.4	4	0.02%	NA	
							very pessimistic MTBF estimates for degradation of the
							HAM2/3 IO optics; should revise based on iLIGO/eLIGO
7	10	replace/clean, align and test degraded HAM2 & HAM3 optics	15	24	4.11%	1	experience
8	10	replace/clean, align and test degraded IO table optics	13.538	4	0.62%	2	2
							HAM1 & HAM6 can be vented & pumped down separately
							from the main vacuum and turned around in 1 day (3 shifts)
9	ISC	replace, align & test ISC in-vacuum electronics	0.04	24	0.01%	NA	if necessary
10	ISC	replace, align & test ISC in-air electronics	6.973	4	0.32%	NA	
		replace, align & test ISC in-vacuum optomechanics,					
11	ISC	picomotors, detectors	1.802	4	0.08%	NA	
		replace, align & test ISC in-air optomechanics, picomotors,					
12	ISC	detectors	8.635	4	0.39%	NA	
							Optical levers are fly-wheels & diagnostic equipment, not
13	OptLev	replace, align & test diode laser, optical fiber or Quad PD	6.526	8	0.60%	2	critical to operation after commissioning (no downtime hit)
		repairs of failures without early detection:					
14	PSL	FrontEnd: EOM, AOM, shutter, PDs, pump fibers, NPRO driver	0.7	4	0.03%	NA	
		scheduled repair/replacement of degraded components:					
15	PSL	FrontEnd: Faraday Isolator, Amplifier, Diode set, NPRO	1.814	4	0.08%	2	2
		repairs of failures without early detection:					
16	PSL	HPO: PD Amp, PD int, PD Iso, PD BP, lock PD, Heads	4.13	4	0.19%	NA	
		scheduled repair/replacement of degraded components:					
17	PSL	HPO: Diode set, etc.	4.904	4	0.22%	2	2
18	PSL	Diagnostic Bread Board (DBB)	NA	NA		NA	diagnostic function only
19	PSL	ISS, PMC, FSS	1.9	2	0.04%	NA	
							ISI can perform (albeit degraded) when a sensor fails due to
							redundancy (control algorithms/laws and software for re-
20	SEL	in-vacuum sensors	13 193	16	2 41%	1	configuration for graceful recovery are still in development)
20	521	SEL renairs of failures without early detection:	15.155	10	2.41/0	-	
21	SEL	prompt attention to in-air electronics	12 554	2	0.29%	NΔ	
22	SEI	SEL scheduled repair/replacement or maintenance	223 083	2	5.09%	2	
	52.		225.005	-	5.0570	-	no maintenance or credible expected failures other than
23	SLC	NA			0.00%		PDs but they are not essential to operation
-25	520	N/A			0.0070		1 bs, but they are not essential to operation
							Can sustain a single OSEM failure on a stage generally
24	SUS	in-vacuum OSEM renair/renlacement	1 839	4	0.08%	1	(some redundancy) so schedule replacement in yearly vent
	505	SEL renairs of failures without early detection:	1.055		0.0070	-	
25	SUS	prompt attention to in-air electronics	10 022	2	0.23%	NΔ	
-23	505		10.022		0.23/0		pessimistic MTBE for CO2 beam coating life
							note that RH glass former is not a credible failure & RH RTD
26	TCS	in-vacuum CO2 laser beam relay mirror coating failures	2 608	R	0 24%	1	is not essential
27	TCS	in-air TCS prompt repair	15,516	4	0.71%	NA	NOTE HWS IS NOT ESSENTIAL
28	TCS	in-air TCS scheduled reapir	3 827	4	0.71%		NOTE HWS IS NOT ESSENTIAL
29	TMS	in-vacuum	0.876	4	0.04%	1	
30	TMS	in-air, prompt	0.291	2	0.01%	NA	
- 35			0.201	2	0.01/0		based on LLO for the period 6/1/2013 thru 6/1/2014
31	Power	power outage	16	5 1875	0 95%	NA	time to recover is a WAG
		r · · · · · · · · · · · · · · · · ·	10	5.2075	19.20%		

6.97%event 1 - in-vacuum maintenance7.16%event 2 - weekly maintenance5.06%NA - prompt attention

6.3 Software Failures

The rate of software failures in the steady-state after commissioning should be rather low. However given the complexity of the system as well as the need to update/maintain the software as operating systems and hardware platforms are updated, it is not reasonable to assume that there is no impact. For this estimate a 0.5% impact has been assumed until better information, or the basis for a better estimate, is developed.

6.4 Maintenance

Just as for the S5 & S6 runs, a four hour maintenance period has been assumed weekly.

In order to accommodate the scheduled equipment maintenance/repair activities within this weekly 4-hour period, the observatory team would need to address ~3 repairs/tasks associated with the interferometer, on average, each week.

6.5 Commissioning

In the aLIGO availability estimate I have assumed that no commissioning is being performed in steady-state operation.

6.6 Transition to Low Noise Operation

The Automation System, Guardian, employed for aLIGO serves a similar function, but is more robust and should be more efficient, once the system is fully commissioned. The recent aLIGO lock acquisition for L1 using the Guardian takes about 10 minutes. Additional time will be required to bring the system to low noise operation as well. The goal is for lock acquisition lengths of > 40 hr. Here it has been assumed that the Guardian will require 15 minutes every 24 hr.

For the first aLIGO science run, the intent is to operate at 25W. When (eventually) aLIGO operates at high laser power, the time required to get to a low noise operational state, from a cold interferometer state, will be much longer. Even the interruption of high power operation for a short period will cause the interferometer optics to cool and require more time to re-establish the proper "hot" state. The availability estimate in this version does not take into account these considerations.

6.7 Wind

The aLIGO seismic isolation system (SEI) was designed to provide much better seismic isolation performance at low frequencies than we had in iLIGO and eLIGO. Integrated test experience with aLIGO indicates that the promised SEI performance has basically been delivered. However the performance is a trade-off with robustness to environmental disturbances. This trade-off, and optimization for the environmental disturbances, has not yet been completed, and may not until we're well into commissioning. There is evidence that wind at LHO causes significant ground tilt which the seismometer sensors sense as translation, and limits performance.

For this estimate, I have conservatively assumed that we don't improve on the sensitivity to windinduced ground tilt sensitivity observed in S5 for LHO. In fact we are likely to improve our robustness to wind as the SEI systems are fully optimized.

6.8 Seismic

As stated above, we fully expect (and have initial integrated test observations in support of this expectation) that we will have considerably better isolation from earthquakes and microseism disturbances. However we don't yet have sufficient data to make assertions on the percentage of downtime with the new seismic isolation systems. For this estimate, I have assumed that downtime due to seismic disturbances is the same as H1 in S5 (and therefore considerably better for L1 than experienced in S5).