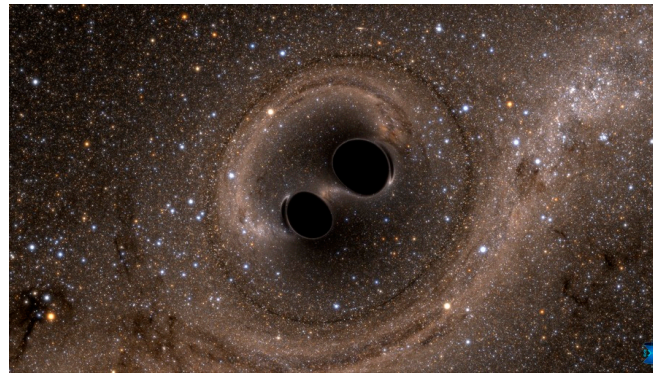




General Relativity

The Basics



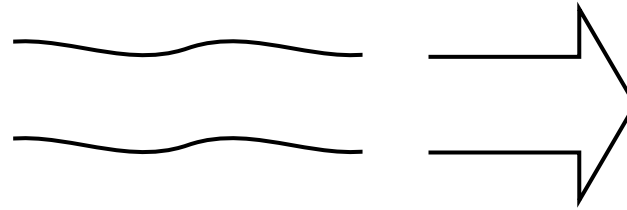
Credit: SXS Lensing; <http://www.black-holes.org/>

The Laser Interferometer Gravitational-wave Observatory: a Caltech/MIT collaboration supported by the National Science Foundation

Gregory Mendell, LIGO Hanford Observatory

LIGO-G1801389

Einstein Wondered:



Can we
catch light?



*Photo: Albert Einstein at the first
Solvay Conference, 1911; Public
Domain*



Niels Bohr



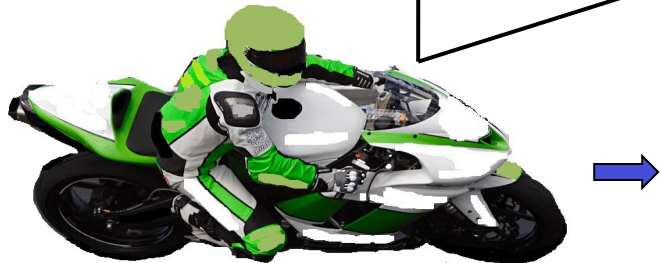
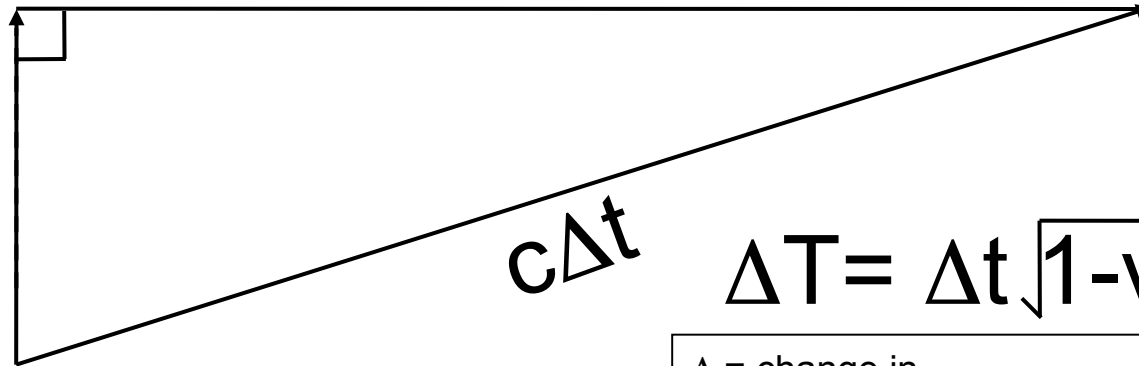
Albert Einstein



Time Dilation



$$\Delta x = v \Delta t$$



$$\Delta T = \Delta t \sqrt{1 - v^2/c^2}$$

Δ = change in

T = time measured by motorcycle riders

t = time measured by observer at "rest"

v = speed of motorcycles

c = speed of light

Start

Warning: thought experiment only; do not try this at home.
Motorcycle: http://en.wikipedia.org/wiki/Motorcycle_racing



The Pythagorean Theorem Of Spacetime

$$c^2\Delta T^2 + v^2\Delta t^2 = c^2\Delta t^2$$

$$c^2\Delta T^2 = c^2\Delta t^2 - v^2\Delta t^2$$

$$c^2\Delta T^2 = c^2\Delta t^2 - \Delta x^2$$

$$c = 1 \text{ light-year/year}$$

$$\Delta T^2 = \Delta t^2 - \Delta x^2$$

(Usually known as the spacetime interval)



Example

$c = 1$ light-year/year

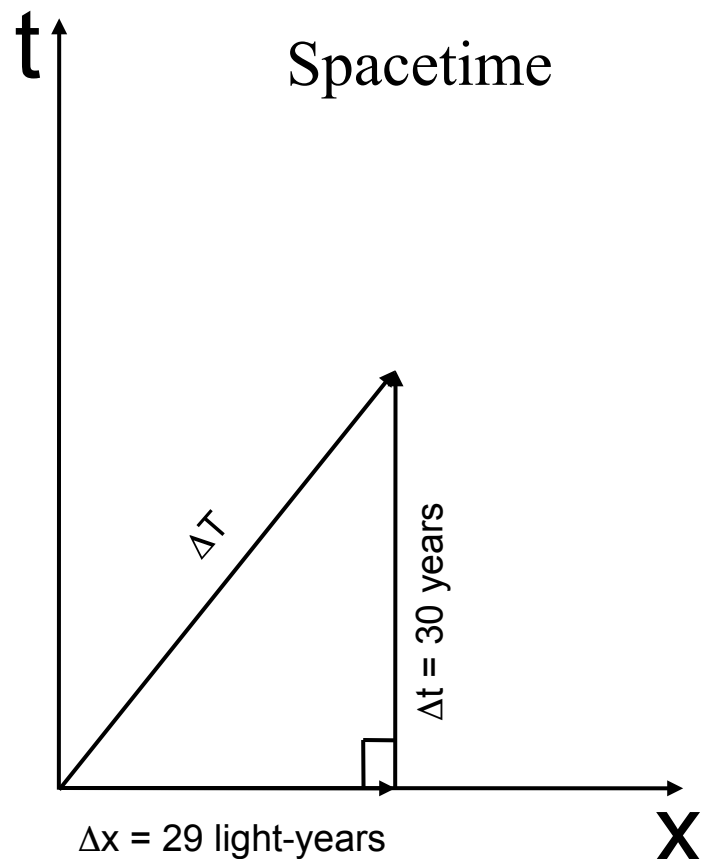
$$\Delta T^2 = \Delta t^2 - \Delta x^2$$

$\Delta t = 30$ years; $\Delta x = 29$ lt-yrs.

$v = 96.7\%$ the speed of light

$\Delta T^2 = 30^2 - 29^2 = 59 \text{ yrs}^2$

$\Delta T = 7.7$ years





Einstein's Happiest Thought

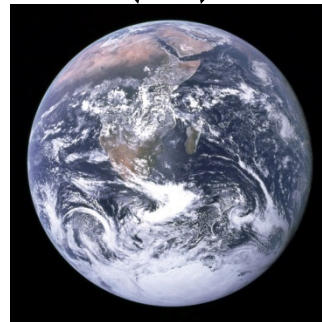


Photo: NASA

Gravity
disappears
when you
free fall!

http://en.wikipedia.org/wiki/Leaning_Tower_of_Pisa

Warning: thought experiment only; do not try this at home.



Einstein's General Theory of Relativity

$\Delta \rightarrow d =$ infinitesimal change

$$dT^2 = g_{tt}dt^2 + g_{xx}dx^2$$

$$dT^2 = g_{\mu\nu}dx^\mu dx^\nu$$

In GR the components of a 4x4 symmetric matrix called the metric tensor define the curvature of spacetime.

$$G_{\mu\nu} = \frac{8\pi G}{c^4} T_{\mu\nu}$$

$$G_{\mu\nu} = R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R$$

$$R_{\mu\nu} = R^\alpha{}_{\mu\alpha\nu}; R = g^{\mu\nu} R_{\mu\nu}$$

$$R^\alpha{}_{\mu\beta\nu} = \partial_\beta \Gamma^\alpha{}_{\mu\nu} - \partial_\nu \Gamma^\alpha{}_{\mu\beta} + \Gamma^\alpha{}_{\beta\gamma} \Gamma^\gamma{}_{\mu\nu} - \Gamma^\alpha{}_{\gamma\nu} \Gamma^\gamma{}_{\mu\beta}$$

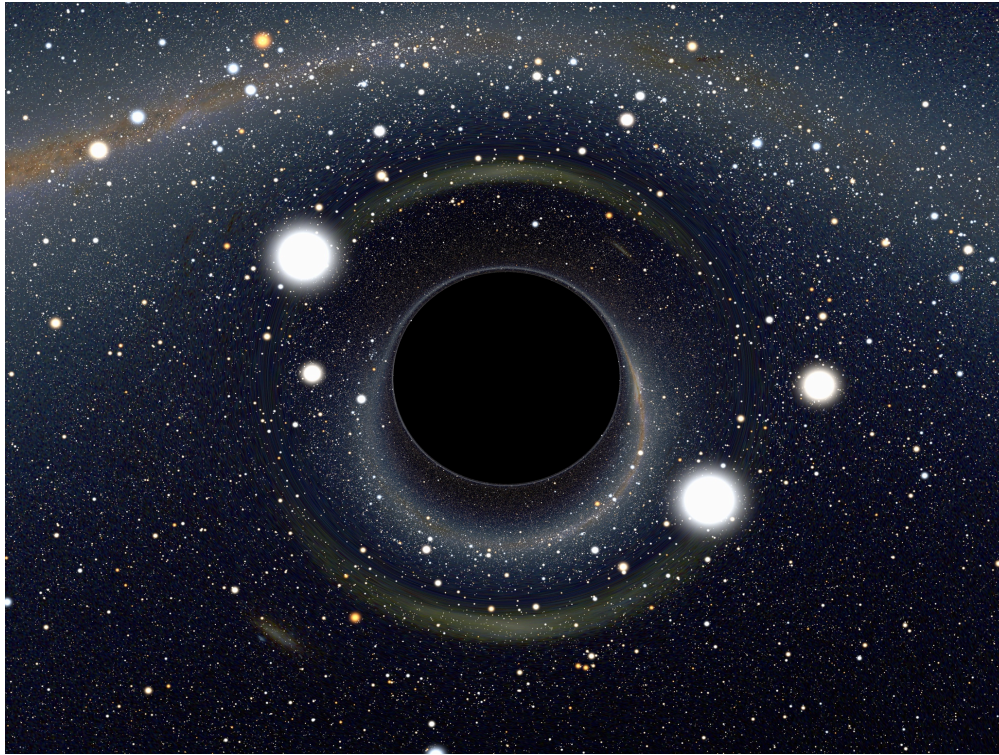
$$\Gamma^\alpha{}_{\mu\nu} = \frac{1}{2} g^{\alpha\beta} (\partial_\nu g_{\mu\beta} + \partial_\mu g_{\beta\nu} - \partial_\beta g_{\mu\nu})$$

Einstein's Field Equations

$$\frac{dx^\alpha}{dT} = U^\alpha; \quad U_\alpha = g_{\alpha\beta} U^\beta \quad U = 4\text{-Vel.}; T = \text{Proper Time}$$

$$\frac{dU_\alpha}{dT} = \frac{1}{2} \partial_\alpha g_{\beta\gamma} U^\beta U^\gamma \quad \text{Geodesic Equation}$$

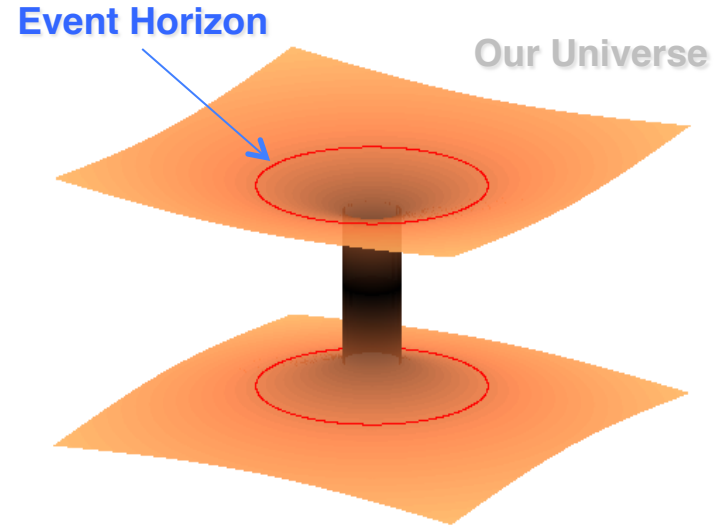
Black Holes are:



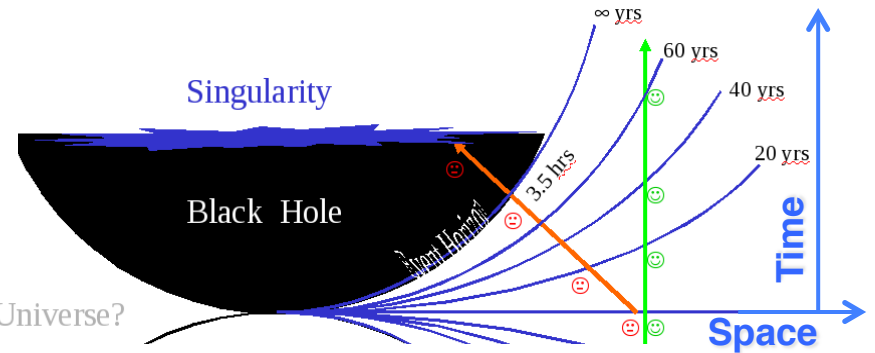
Credit: Alain Riazuelo, IAP/UPMC/CNRS

1. Black

Escape Velocity = Speed of Light



2. Holes in Space



3. Space & Time Warps

Schwarzschild 1916; Einstein & Rosen 1935; many others in the 1960s



Schwarzschild Black Hole

$$c^2 dT^2 = \left(1 - \frac{2GM}{rc^2}\right) c^2 dt^2 - \frac{1}{\left(1 - \frac{2GM}{rc^2}\right)} dr^2 - r^2 d\theta^2 - r^2 \sin^2 \theta d\phi^2$$

$$c^2 dT^2 = \left(1 - \frac{v_{esc}^2}{c^2}\right) c^2 dt^2 - \frac{1}{\left(1 - \frac{v_{esc}^2}{c^2}\right)} dr^2 - r^2 d\theta^2 - r^2 \sin^2 \theta d\phi^2$$



Karl Schwarzschild

$$v_{esc} = \sqrt{\frac{2GM}{r}}$$

• Escape Velocity

$$R_s = \frac{2GM}{c^2}$$

• Schwarzschild Radius

<u>Object</u>	<u>Schwarzschild Radius</u>
You	1 thousand, million, million, millionth the thickness of a human hair
Earth	1 cm (size of marble)
Sun	3 km (2 miles)



Gravitational Time Dilation



$$\Delta T = \sqrt{1 - \frac{2GM}{rc^2}} \Delta t$$

Gravity
slows time
down!

Photo:http://en.wikipedia.org/wiki/Leaning_Tower_of_Pisa

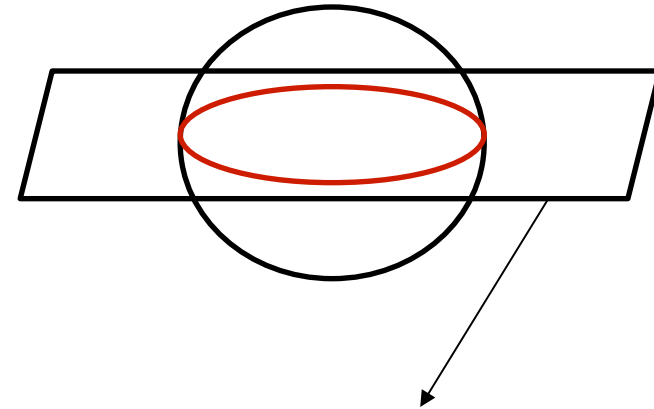
Clock_Photos:http://en.wikipedia.org/wiki/Cuckoo_clock



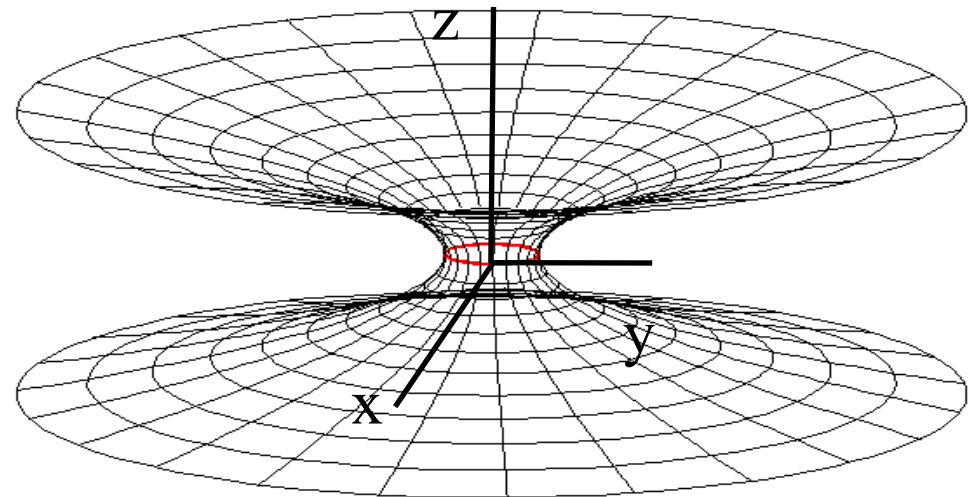
Black Hole Embedding Diagram

Schwarzschild solution to General Relativity for $t = \text{constant}$, $\theta = \pi/2$:

$$ds^2 = \frac{1}{\left(1 - \frac{2GM}{rc^2}\right)} dr^2 + r^2 d\varphi^2$$



The Schwarzschild Wormhole or Einstein-Rosen Bridge
(Flamm 1916, *Physikalische Zeitschrift*. XVII: 448; Einstein & Rosen 1935, *Phys. Rev.* 48 73; Misner & Wheeler 1957, *Ann. Phys.* 2: 525)

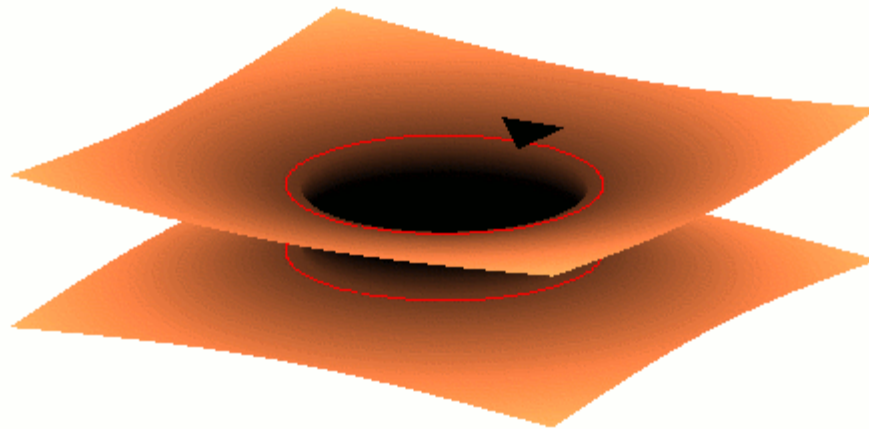




Non-traversable Wormhole

Our Universe

TIME INSIDE = 0.0 sec. ROCKET TIME = 0.0 sec. TIME OUTSIDE = 100.3 sec.

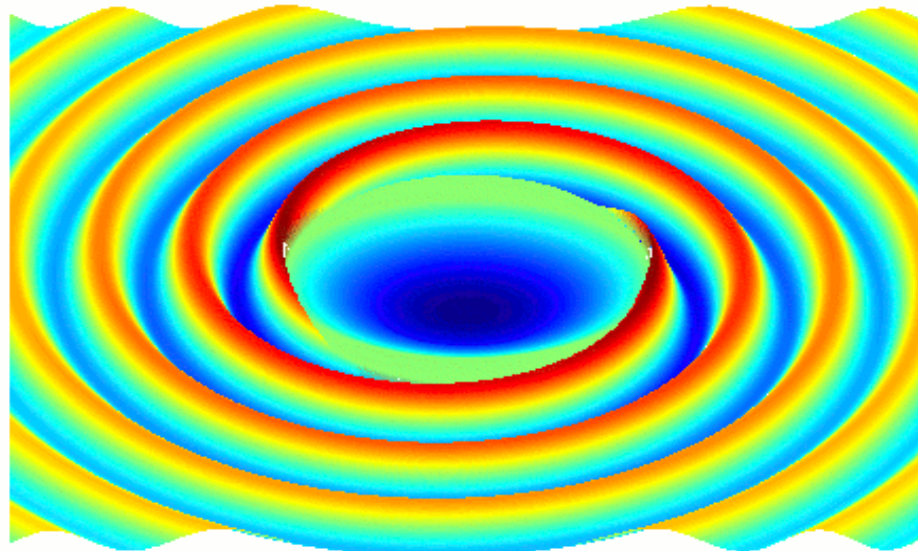


Another Universe



Gravitational Waves

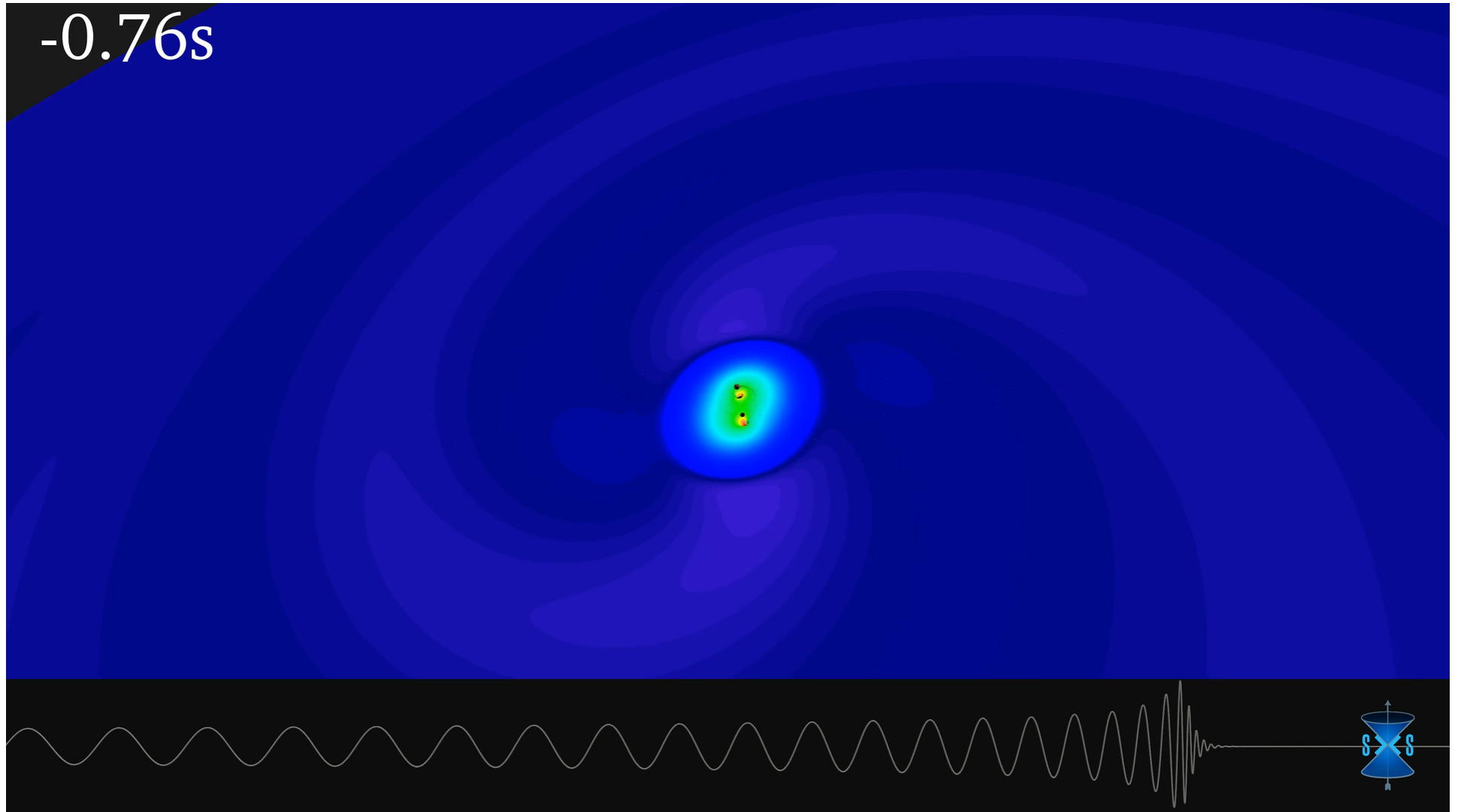
Gravitational waves are ripples in the fabric of spacetime, stirred up by the changing motions of matter and energy.





Binary Black Hole Merger Simulation

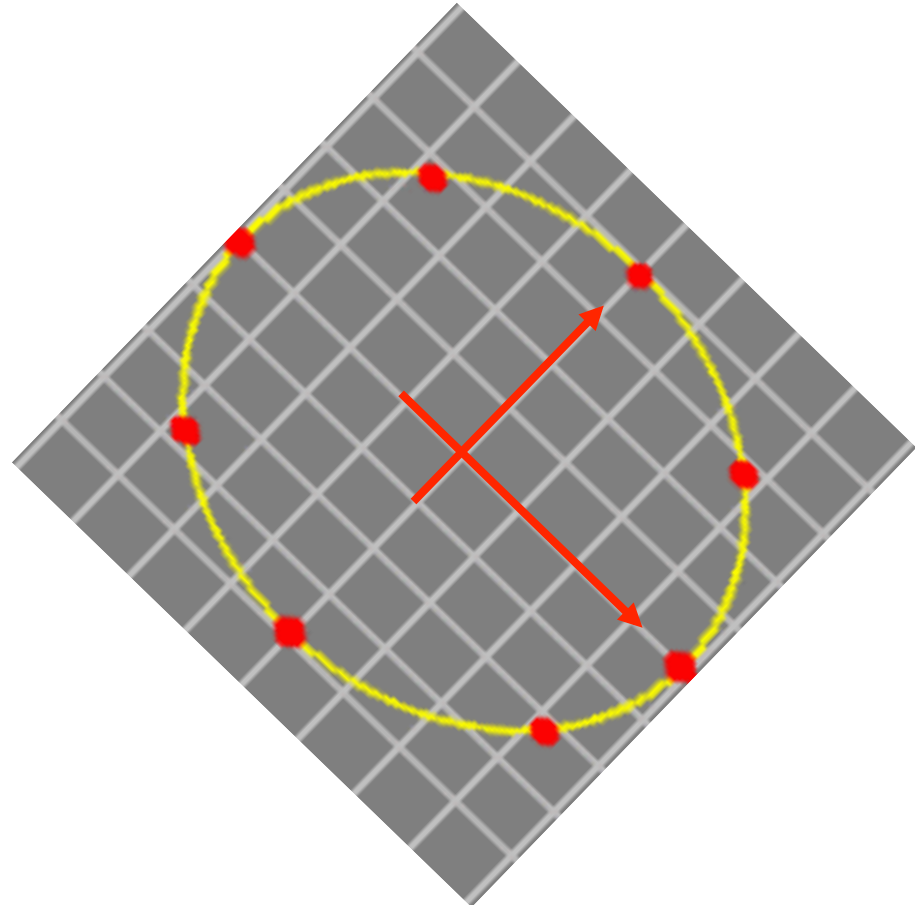
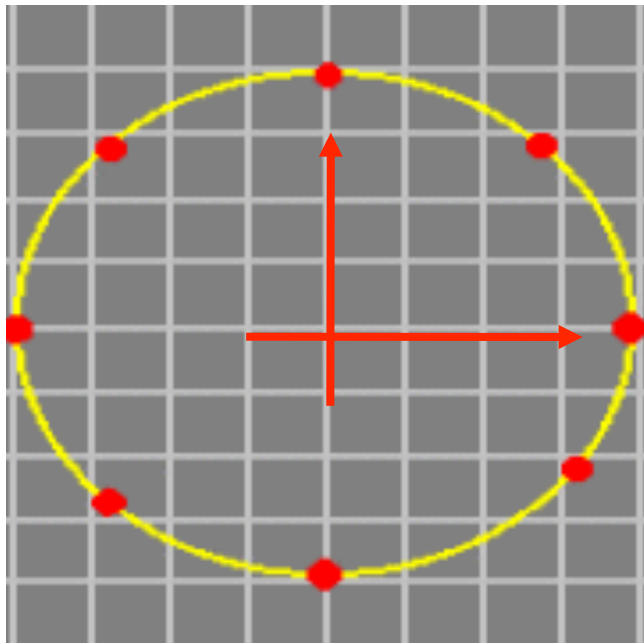
Credit: SXS Collaboration/Canadian Institute for Theoretical Astrophysics/SciNet; <http://www.black-holes.org/>



LIGO-G1600258

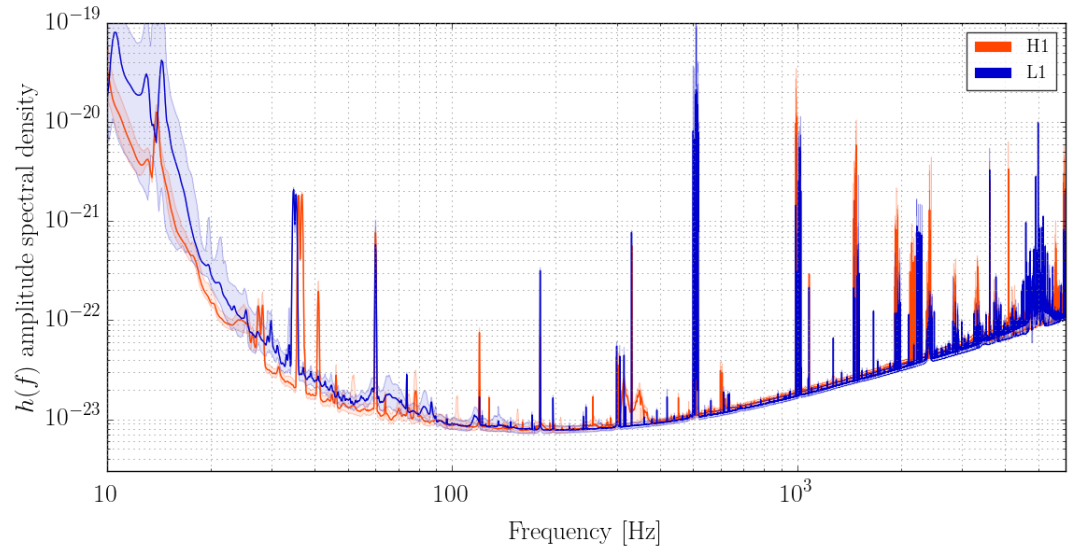
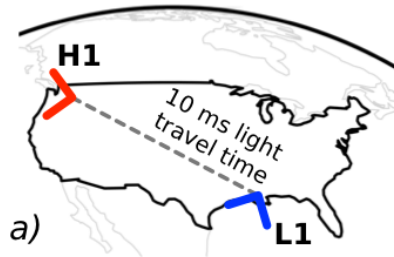
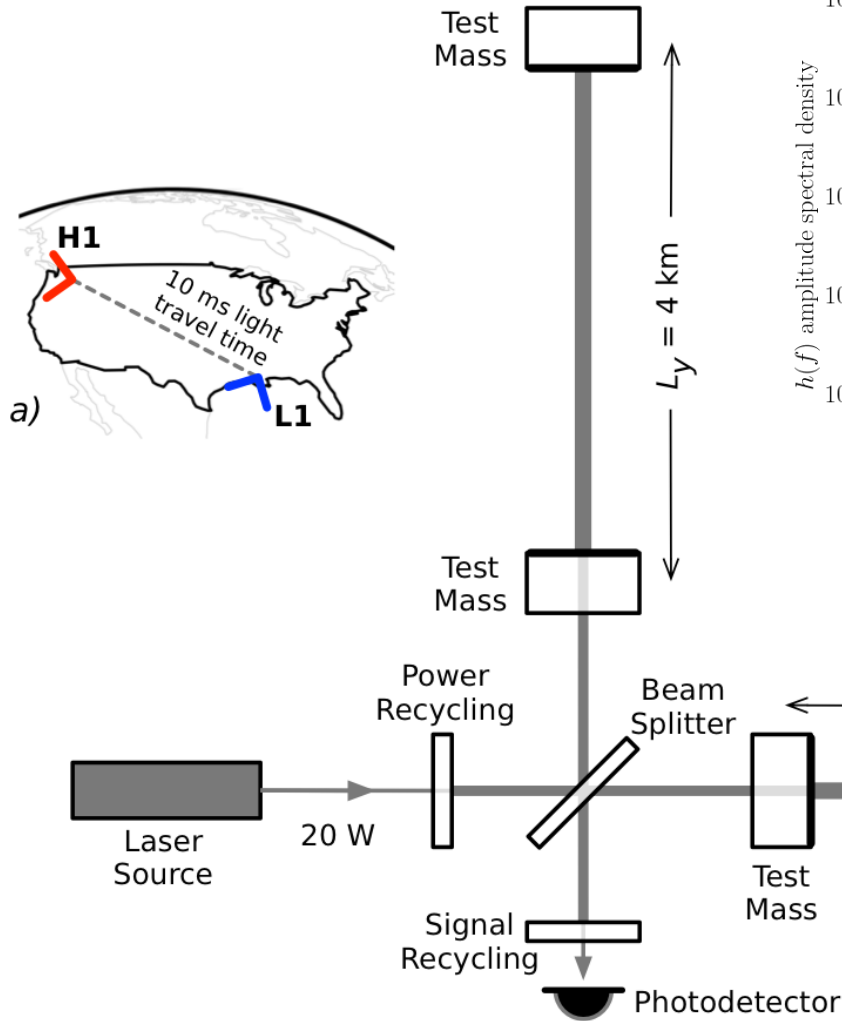


Gravitational Waves





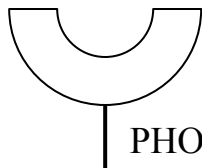
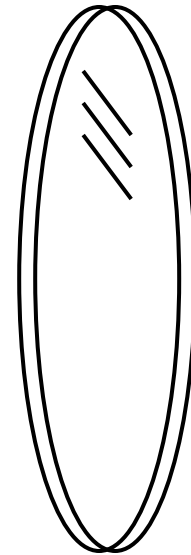
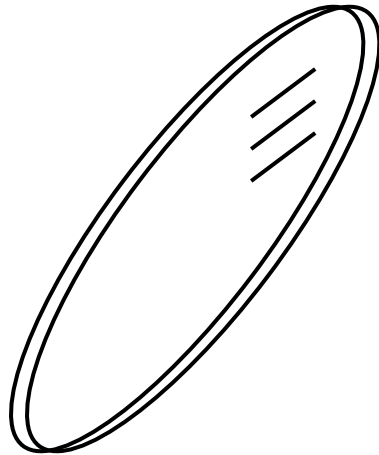
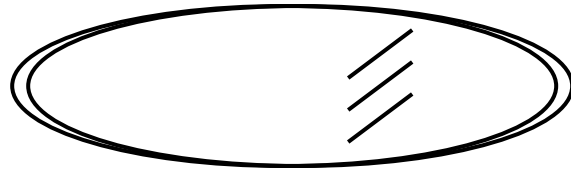
The LIGO Detectors



The waves are extremely weak by the time they reach Earth:
 $\Delta L/L \sim 10^{-21}$



The End



PHOTODIODE



Detector Response

$$g_{\mu\nu} dx^\mu dx^\nu = 0 \quad (\text{Light Travels On Null Geodesics})$$

$$c^2 dt^2 - \begin{pmatrix} dx & 0 & 0 \end{pmatrix} \begin{pmatrix} 1 + h_{xx} & h_{xy} & h_{xz} \\ h_{yx} & 1 + h_{yy} & h_{yz} \\ h_{zx} & h_{zy} & 1 + h_{zz} \end{pmatrix} \begin{pmatrix} dx \\ 0 \\ 0 \end{pmatrix} = 0$$

$$c^2 dt^2 = (1 + h_{xx}) dx^2$$

$$c \int_0^{\Delta t} dt = \int_0^L \sqrt{1 + h_{xx}} dx \cong \int_0^L \left(1 + \frac{1}{2} h_{xx} \right) dx$$

$$c\Delta t = L_x = L + \frac{L}{2} h_{xx}$$

$$\frac{\Delta L}{L} = \frac{1}{2} (h_{xx} - h_{yy}) = F_+(\theta, \phi, \psi) h_+(t) + F_\times(\theta, \phi, \psi) h_\times(t)$$



Summary



- Special Relativity, 1905: nothing can faster than light and the faster you go the slower time goes. Space and time become spacetime!
- Einstein's happiest thought, 1907: gravity disappears when you free fall.
- General Relativity, 1915: gravity is a warping of space and time.
- The Schwarzschild Solution, 1916: Black Holes!
- Gravitational Waves, 1916: the changing motion of matter and energy can produce ripples in the fabric of spacetime.

Photo: Albert Einstein at the first Solvay Conference, 1911; Public Domain



Would you believe you could fall into hole in completely empty space, a hole from which nothing can escape, not even light?



Gravitational Waves

Gravitational waves carry information about the spacetime around black holes & other sources.

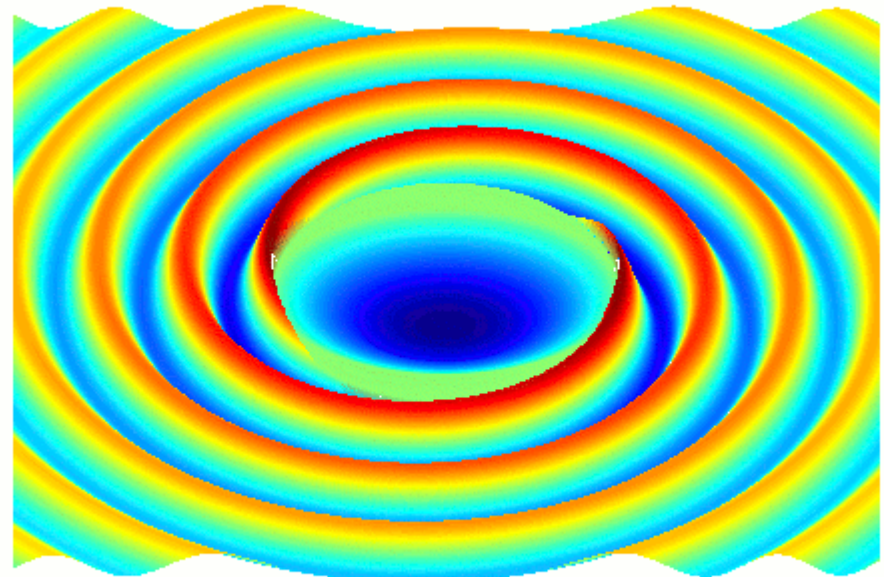
$$dT^2 = g_{\mu\nu} dx^\mu dx^\nu$$

$$g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}$$

$$\left(\nabla^2 - \frac{1}{c^2} \frac{\partial^2}{\partial t^2} \right) \bar{h}^{\mu\nu} = 0$$

$$h_{\hat{\theta}\hat{\theta}}^{TT}(\theta = \pi/2) \propto \frac{1}{r} \cos[2\pi f(t - r/c) + 2\phi]$$

$$h_{\mu\nu}^{TT} = \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & h_+ & h_\times & 0 \\ 0 & h_\times & -h_+ & 0 \\ 0 & 0 & 0 & 0 \end{pmatrix} e^{2\pi i f(t-z/c)}$$





Gravitational Waves

Gravitational waves are ripples in the fabric of spacetime, stirred up by the changing motions of matter and energy.

The waves are extremely weak by the times they reach Earth.

