

### **Interferometer Design and Noise Sources**

LIGO's discovery of gravitational waves ushered in a new era of multi-messenger astronomy, confirming Einstein's theory of general relativity.



### Wind at LIGO

LIGO employs both active and passive isolation to minimize seismic noise and to operate at high sensitivity. Horizontal tilt confuses LIGO's seismometers and can cause the system to execute spurious translations in response to low-frequency tilts, with the frequency response given by  $\frac{-g}{\omega^2}$ .<sup>3</sup> At LIGO Hanford, wind above 10 m/s causes a significant increase in tilt and occurs 15% of the time. To reduce problematic wind, LIGO has proposed building a fence around End Station X and End Station Y



### **Computational Fluid Dynamics (CFD) Modeling**

k-ε CFD models are reasonably reliable for characterizing turbulent flow with high Reynolds numbers. k-ɛ models solve for turbulent kinetic energy or the root mean square velocity fluctuations (k) and the rate of dissipation of k ( $\epsilon$ ).

Design flow domain in SolidWorks





Initialize solution and run computation

## Contact

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# Wind Proofing LIGO

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Table 1. Test fence and steady state model
results

Average Velocity	Test Fence (+/- .5 m/s)	Model
Freestream	4.36	5.83
Before Fence	3.65	3.05
After Fence	2.64	2.37
EX Roof	5.63	5.96

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