# BHD Mode Matching 

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## 1 AS Path

Table 1: Telescope design for AS path.

| Component | AS1 | AS2 | AS3 | AS4 |
| :--- | :---: | :---: | :---: | :---: |
| Position $z[\mathrm{~m}]$ | 0.7192 | 1.2597 | 1.8658 | 2.5822 |
| RoC $[\mathrm{m}]$ | 2.8 | Inf | -2.0 | 0.60 |




Figure 1: Beam profile for the AS path.


Figure 2: Left: cumulative histogram of mode overlap. Right: histogram of gouy phase separation between the two actuator mirrors. In both plots, we use grey (olive) traces to represent the tangential (sagittal) plane. The solid-vertical lines are for the nominal values and the dotted ones are the $(10,50,90)$ percentiles of the distribution. When generating the distribution, we have assumed a Gaussian error on the location $z$ of each component $\sigma_{z}=5 \mathrm{~mm}$, and a fractional error on the $\operatorname{RoC} \sigma_{\mathrm{RoC}}=1 \%$.


Figure 3: Scattering plots of the mode-matching (MM) vs. errors for the tangential plane. We color-coded the plots according to the MM. The purple dots corresponds to those with $\mathrm{MM}>0.99$, the olive triangles with $0.98<\mathrm{MM} \leq 0.99$, and grey crosses with $\mathrm{MM} \leq 0.98$. The MM is most sensitive to the AS1 RoC error, and can be compensated by moving the distance between AS1-AS3.


Figure 4: Similar to Fig. 3 but for the sagittal plane.


Figure 5: Assuming $\pm 1 \%$ error on the AS1 RoC, the MM as a function of the AS1-AS3 distance adjustment.

## 2 LO Path

Table 2: Telescope design for LO path.

| Component | LO1 | LO2 | LO3 | LO4 |
| :--- | :---: | :---: | :---: | :---: |
| Position $z[\mathrm{~m}]$ | 0.4027 | 2.9835 | 4.5955 | 4.8880 |
| RoC $[\mathrm{m}]$ | 6 | 1000 | 0.75 | -0.45 |




Figure 6: Beam profile for the LO path. The actuators needs to be LO1 and LO2 in this case.


Figure 7: Similar to Fig. 2 but for LO path.


Figure 8: Similar to Fig. 3 but for the LO path, tangential plane. The MM is most sensitive to the LO3 RoC error and it can be corrected by moving the LO3-LO4 distance.


Figure 9: Similar to Fig. 8 but for the sagittal plane.


Figure 10: Assuming $\pm 1 \%$ error on the LO 3 RoC , the MM as a function of the LO3-LO4 distance adjustment.

