

Searching for Cosmic Ray Induced Noise at LHO

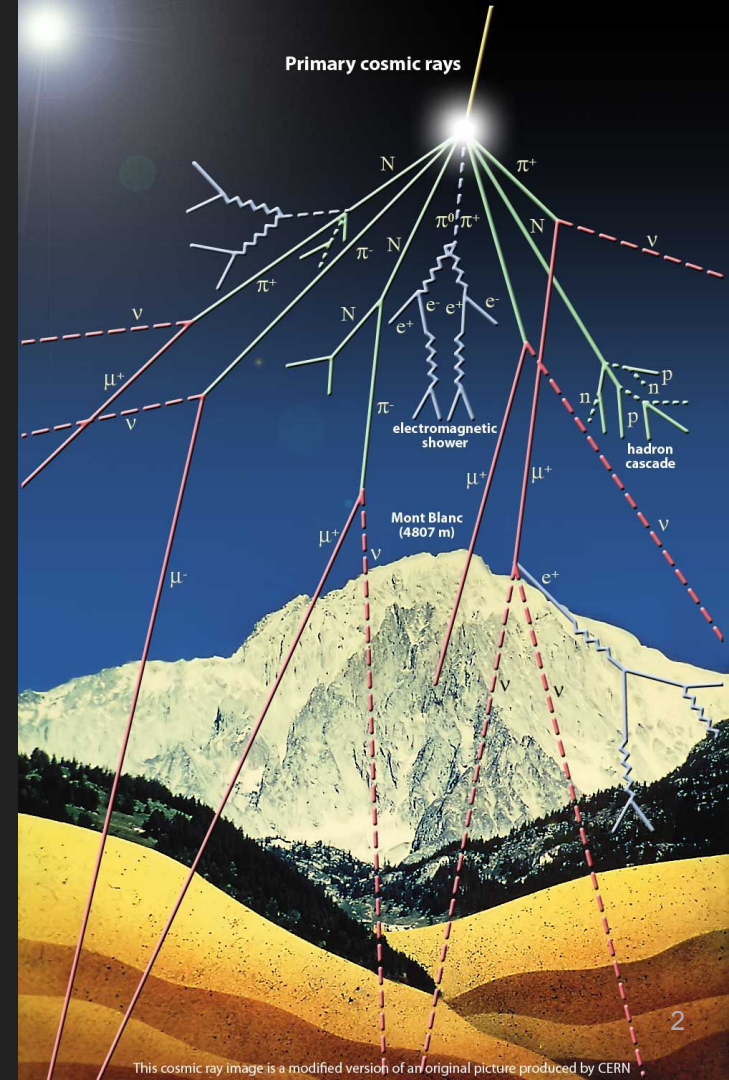
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GWANW 2024



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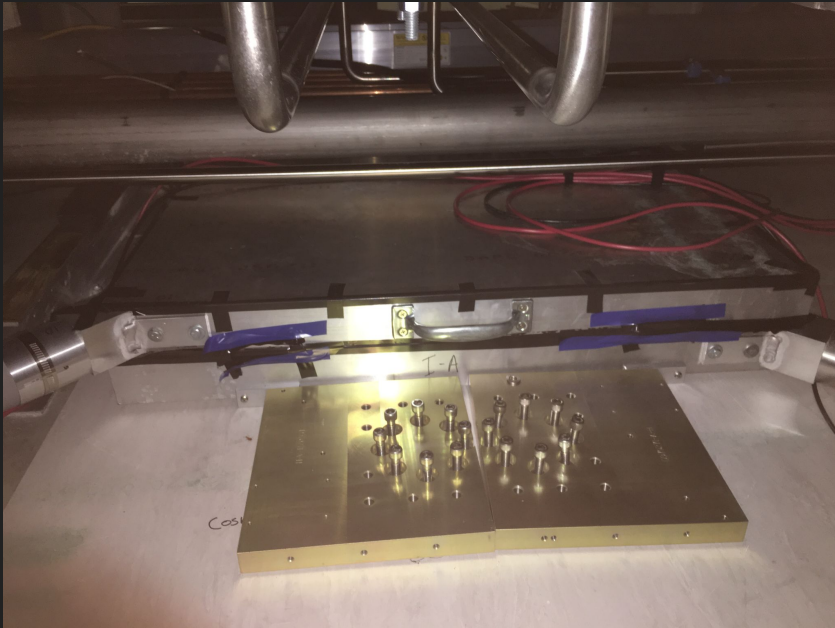
Cosmic Rays

- Cosmic ray showers could potentially be a source of transient noise in LIGO detectors by:
 1. Transferring momentum to mirrors
 2. Heating the mirrors, creating vibrations
 3. Electromagnetic perturbation or depositing charge on the test mass
- We searched for effects of cosmic rays in LIGO Hanford data

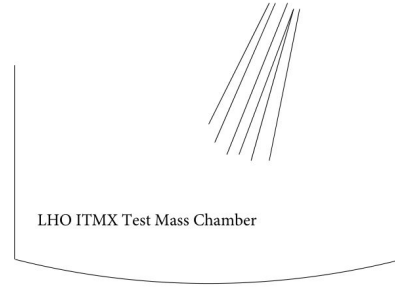


LHO Cosmic Ray Detector

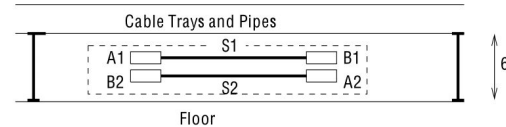
Cosmic ray detector under LHO ITMX



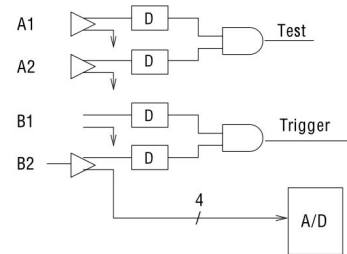
Detector Setup



- S1, S2 : 31"x31"x1" plastic scintillator
- A1, A2 : 10stage PMT (low gain)
- B1, B2 : 11stage PMT (high gain)
- D : Discriminator/comparator
- A/D : CDS voltage sampling
follows Q/V amp with ~10 us shaping



PMT = "Photomultiplier tube"

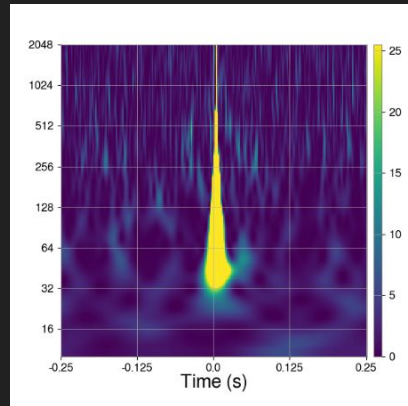


Cosmic Ray
electronics diagram
(as of O4a)

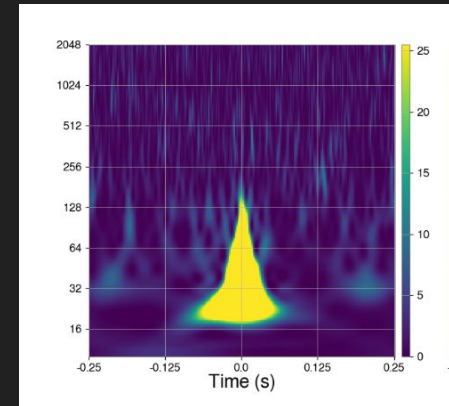
Glitches Considered

- Looking for something that could potentially be an effect of a cosmic ray shower impulse
- We consider glitches that are short in duration and of unknown origin (numbers correspond to amount reported from GravitySpy in Sept-Dec 2023):
 - Blips (661)
 - Low-Frequency Blips (330)
 - Repeating Blips (90)
 - Tomtes (1063)

Blip Glitch



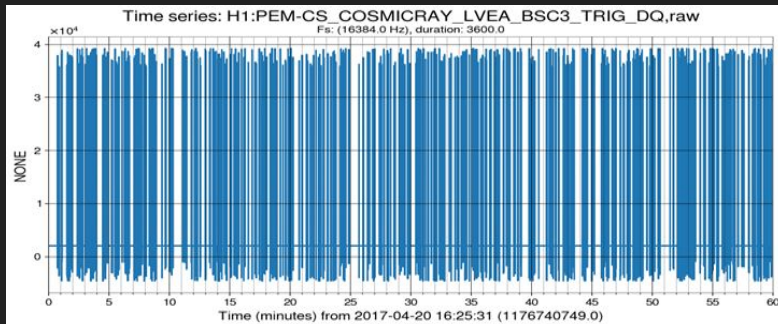
Tomte



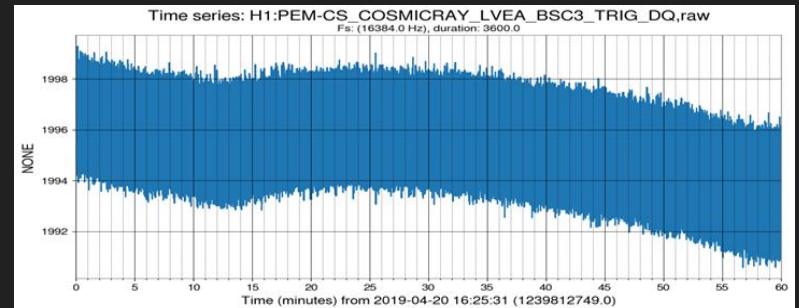
Previous Searches

- A search for correlations between blips and cosmic rays was done in O2 and O3. No connection was found.
- This search used data from the “cosmic ray trigger” channel but, cosmic ray rate dropped significantly from O2 to O3, likely due to an electronics issue

Cosmic Ray triggers in 1 hr of O2 data



No triggers in 1 hr of O3 data



See [DCC: LIGO-T2000261](#) or [DOI 10.1088/1361-6382/abfd85](#)

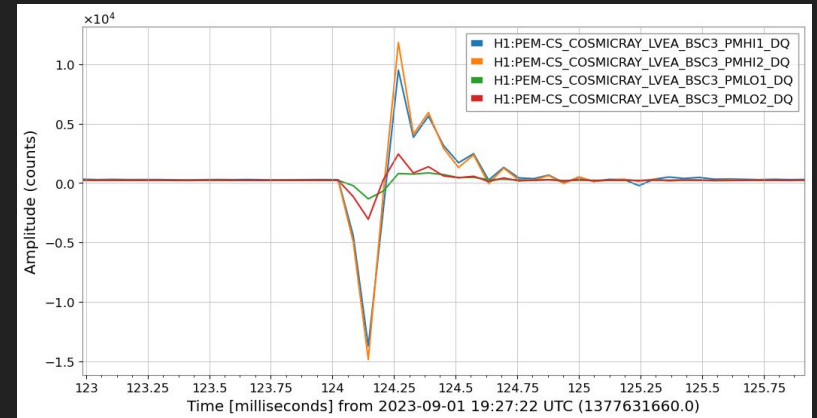
Electronics Upgrade

- The sampling needed to capture cosmic ray information is \sim MHz but LIGO data system is limited at \sim 16 kHz. Solved with voltage integrator scheme
- Now each photomultiplier tube has its own channel, giving both amplitude and timing information of cosmic ray events ([alog](#))

PMT in oscilloscope on nanosecond timescale

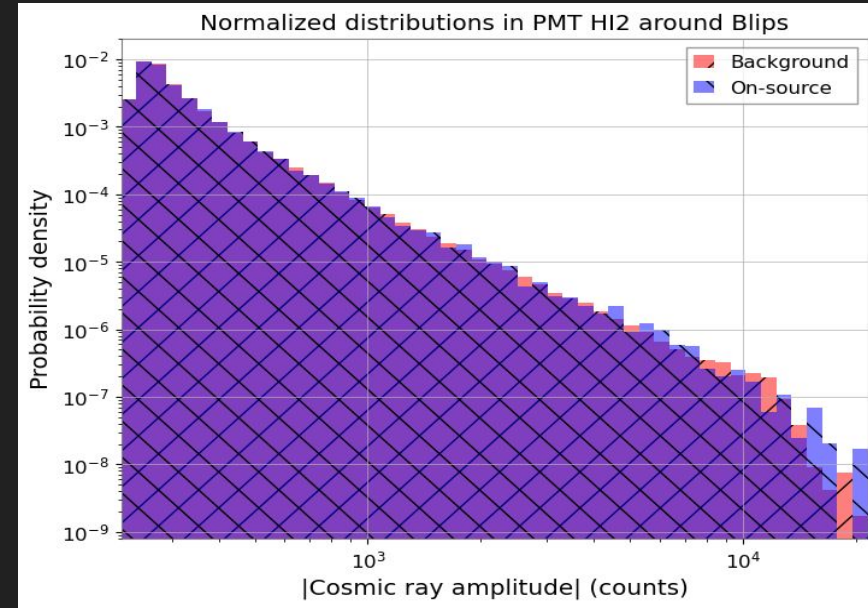


PMT channels on millisecond timescale



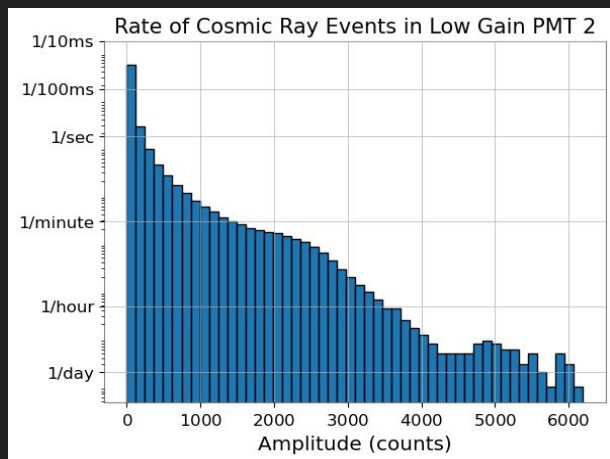
O4a results: Cosmic Ray Amplitude Near Glitches

- Look at amplitude of cosmic ray events within ± 1 second of glitch (on-source)
- For every glitch, take 20 seconds of cosmic ray amplitudes far away from glitches (Background)
- If cosmic ray showers were the source, we might expect to see more higher amplitude events in on-source
- Overlaying these normalized distributions, we see no significant difference

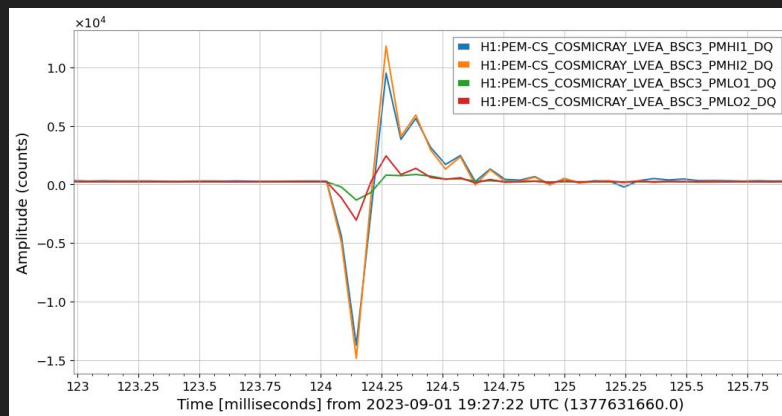


O4a results: temporal relation between cosmic rays and glitches

- We look at trends in cosmic ray data over a month to set some amplitude threshold (top 10% of of m-trend minimum events)
- We then require the event to be seen in multiple PMTs. This gives 1834 cosmic ray showers in the month
- This gives a set of the highest energy cosmic ray showers to compare to glitch times

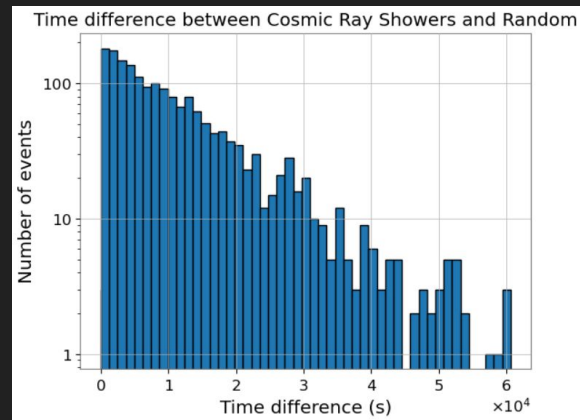
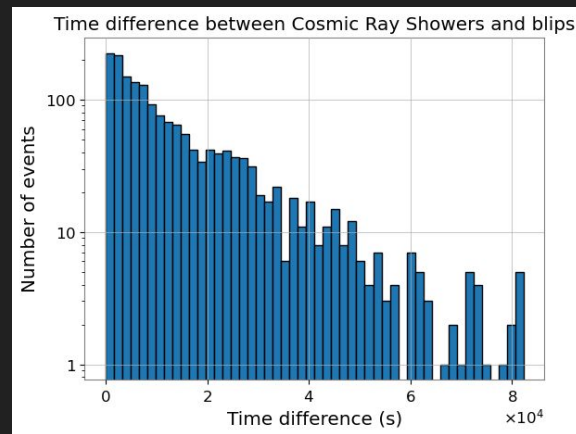


Coincident cosmic ray event in multiple PMTs



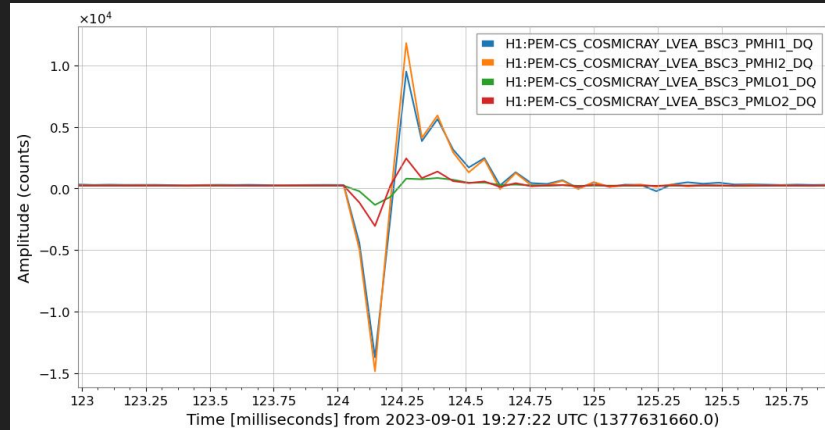
O4a results: Temporal Correlation Between Cosmic Rays and Glitches

- In case there is some sort of delay between cosmic rays and reaction, we investigate a temporal correlation
- Take time difference between large cosmic ray events and next-nearest glitch in time
- Compare to the time difference between cosmic ray showers and random times
- No correlation was found here



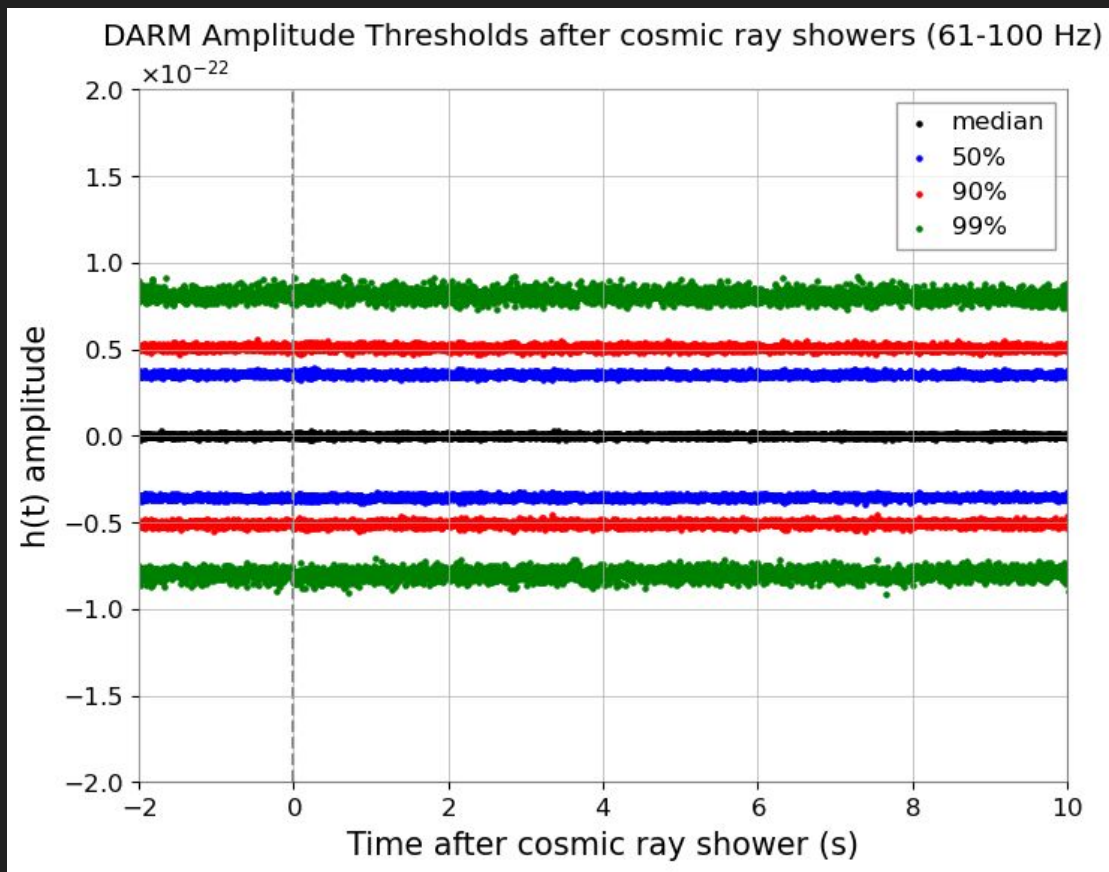
Looking at DARM

- Looking at large cosmic ray showers during observing time, can we see an impact in DARM (our gravitational wave channel)?
- Use cosmic rays as triggers to study the statistical behavior of DARM before and after a shower passes through



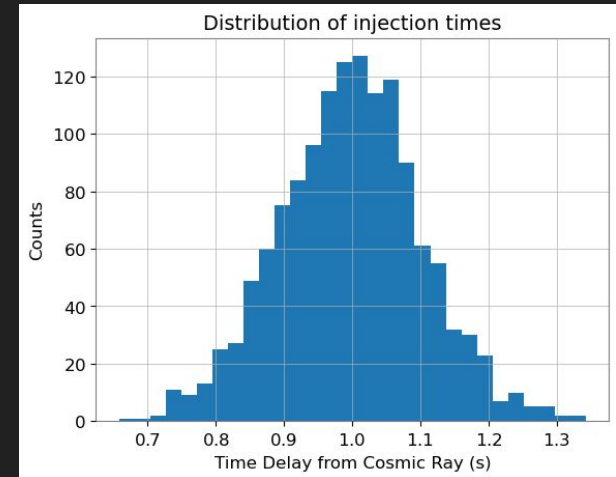
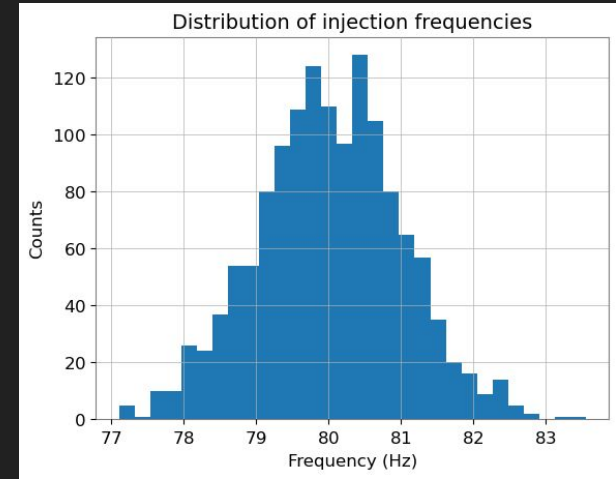
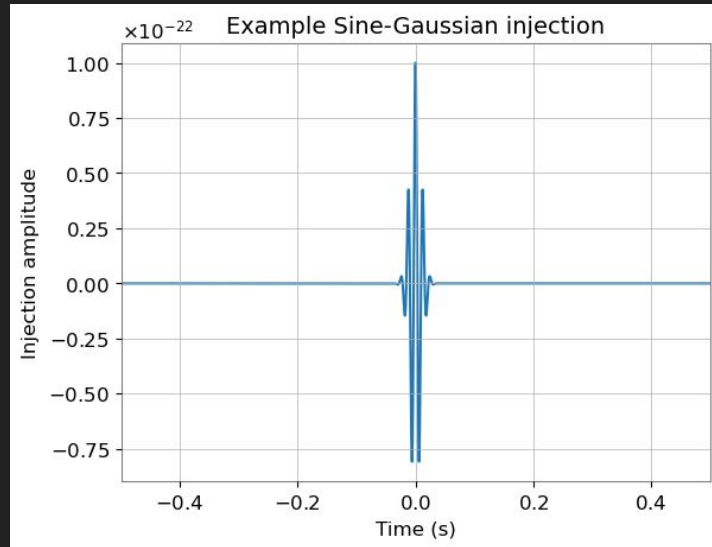
Looking at DARM

- Combine DARM time series from 1375 loudest cosmic ray showers
- DARM noise should follow Gaussian noise
- At every point in time measure what amplitude contains X% of values
- If an effect is consistently present, you might expect this curve to widen



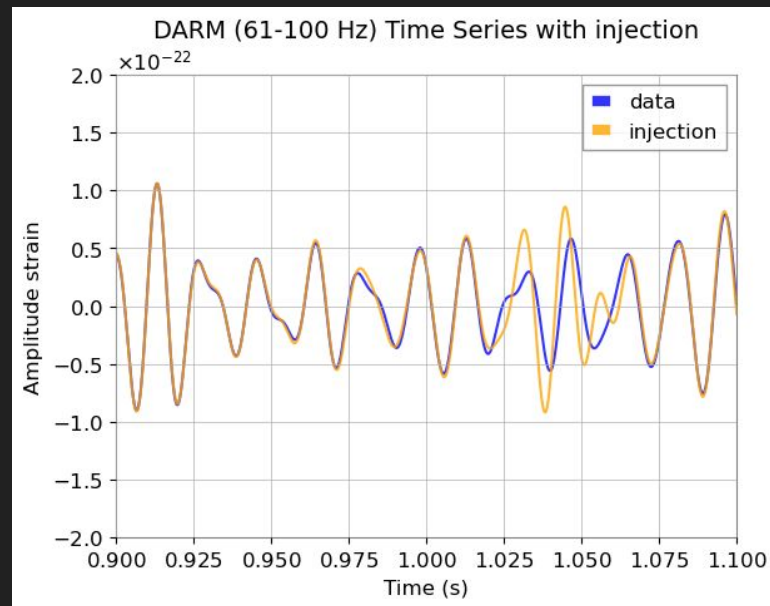
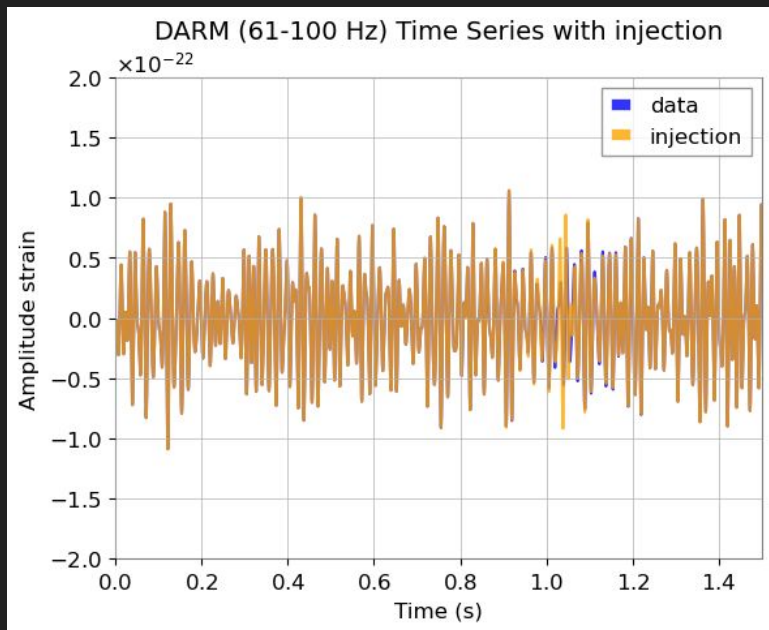
Injected Signal Test

- As a test of this method, inject a sine-gaussian pulse into each time series of varying frequencies and at varying times



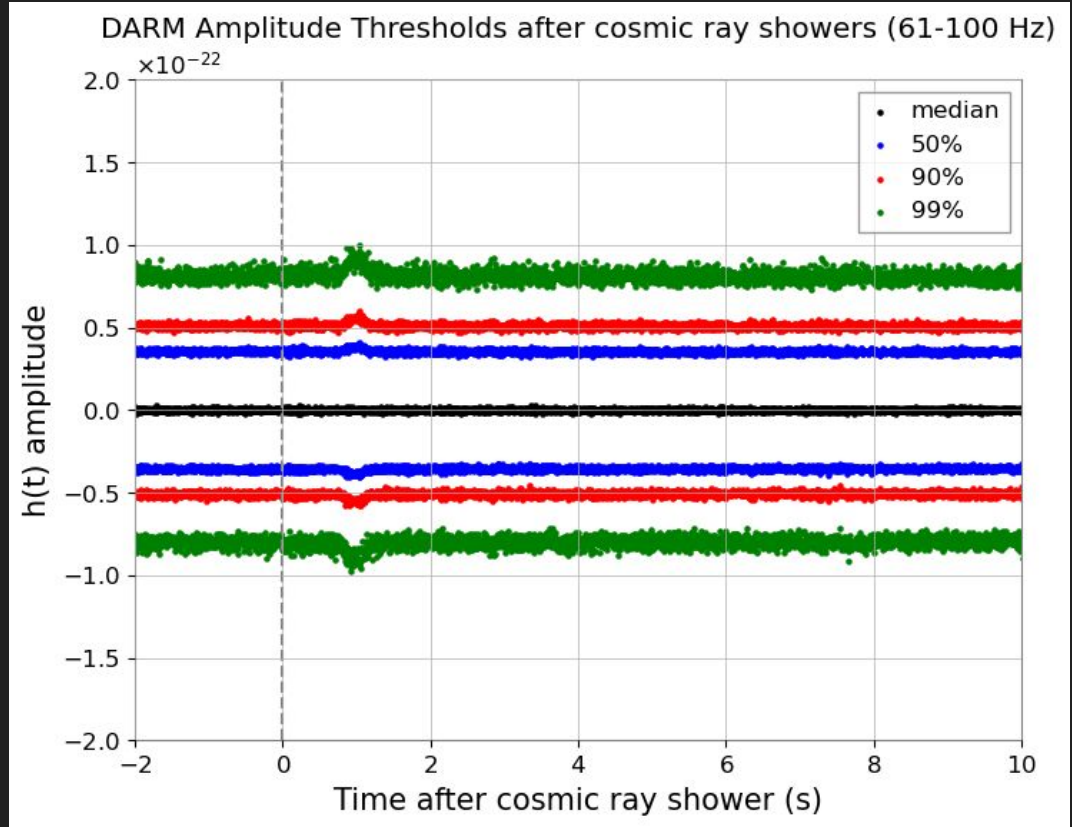
Injected Signal Test

- Example Time Series with an injection ($f = 78.6$ Hz, $t = 1.04$ s)



Injected Signal Test

- Over a large number of time series with a signal present, we start to see an effect emerge around 1 second



Conclusions and Future Work

- Seeing no connection between cosmic ray showers and blips/tomtes
- Seeing no effect in DARM so far. (Analysis still underway)
- Given our current sensitivity, we would like to say whether or not this will be a problem for third generation detectors