



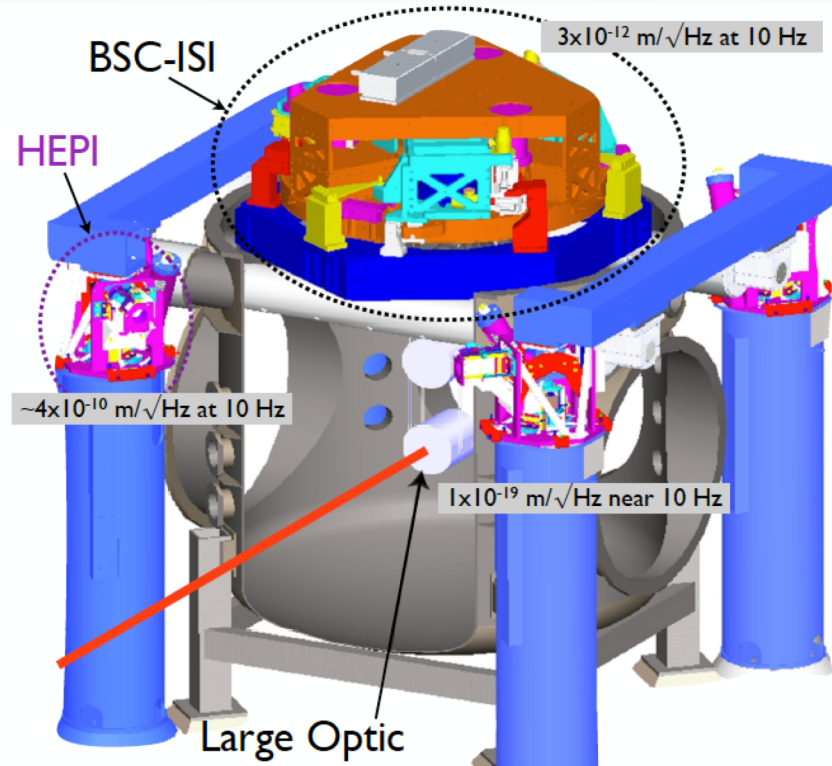
***Environmental noise
lessons from LIGO for
CE site selection***

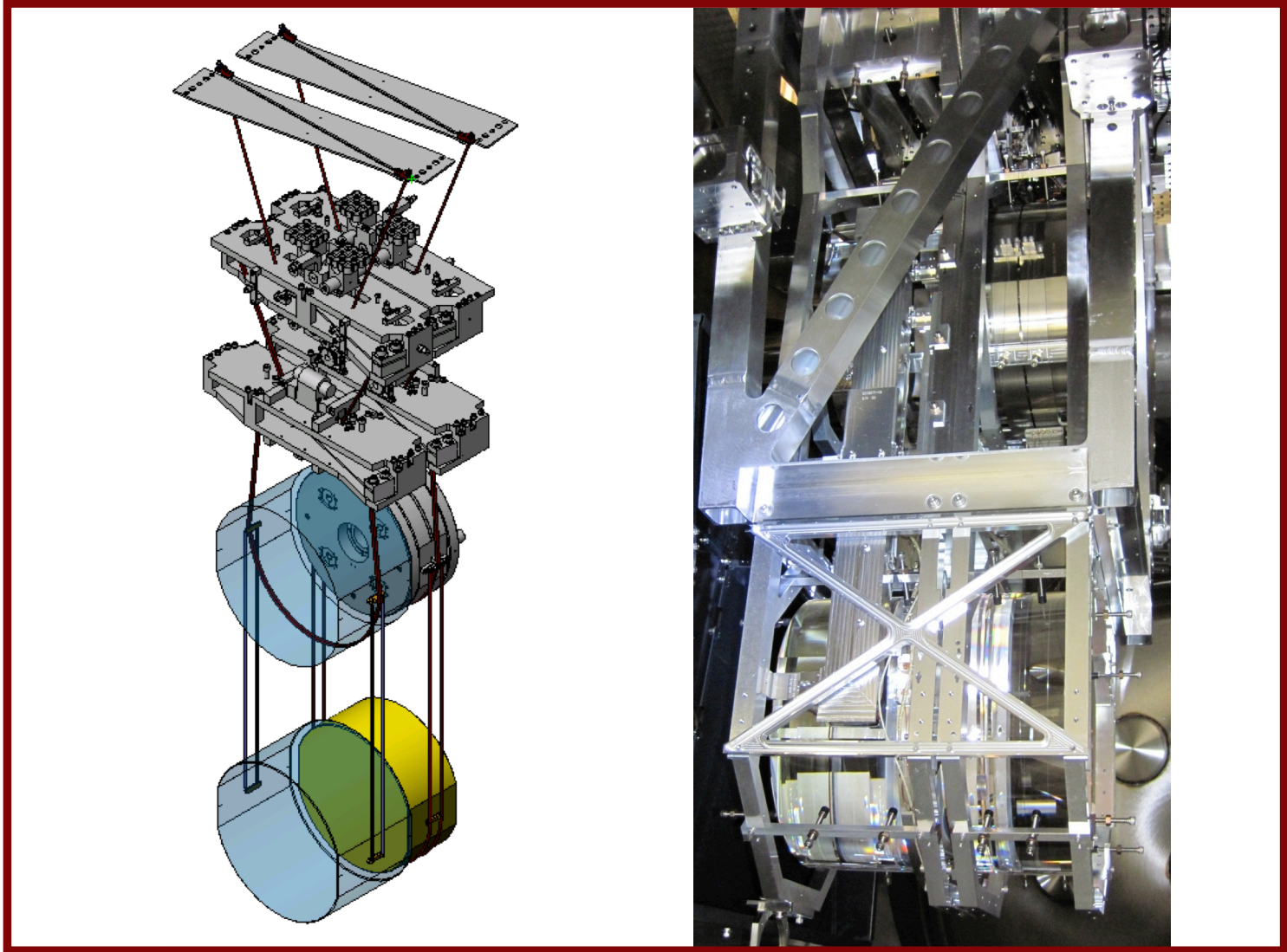


The two things we had to do to detect gravitational waves

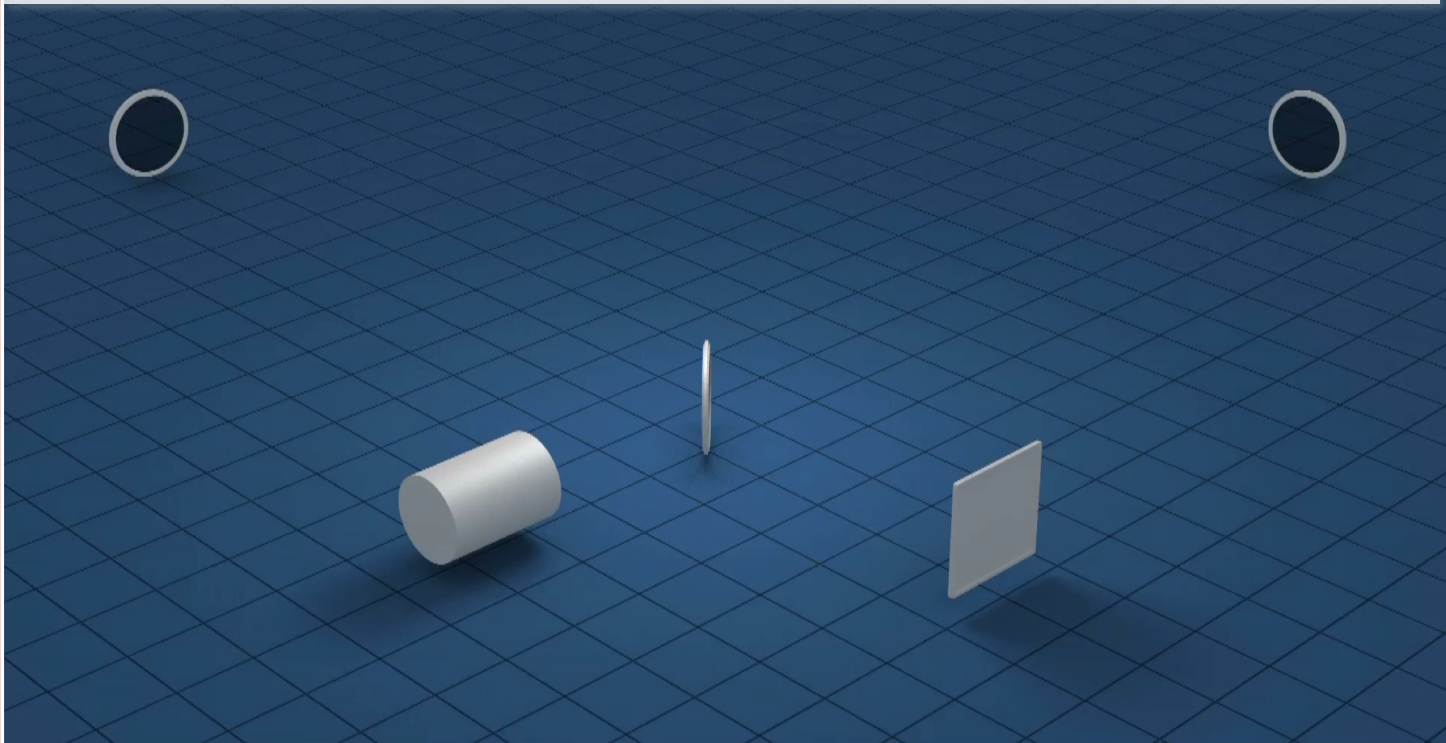
- 1. Quiet the motion of the test masses so that they are moving less than about $1e-19$ m, so that the motion produced by gravitational waves is not swamped by other motion**
- 2. Measure the tiny relative motion between test masses**

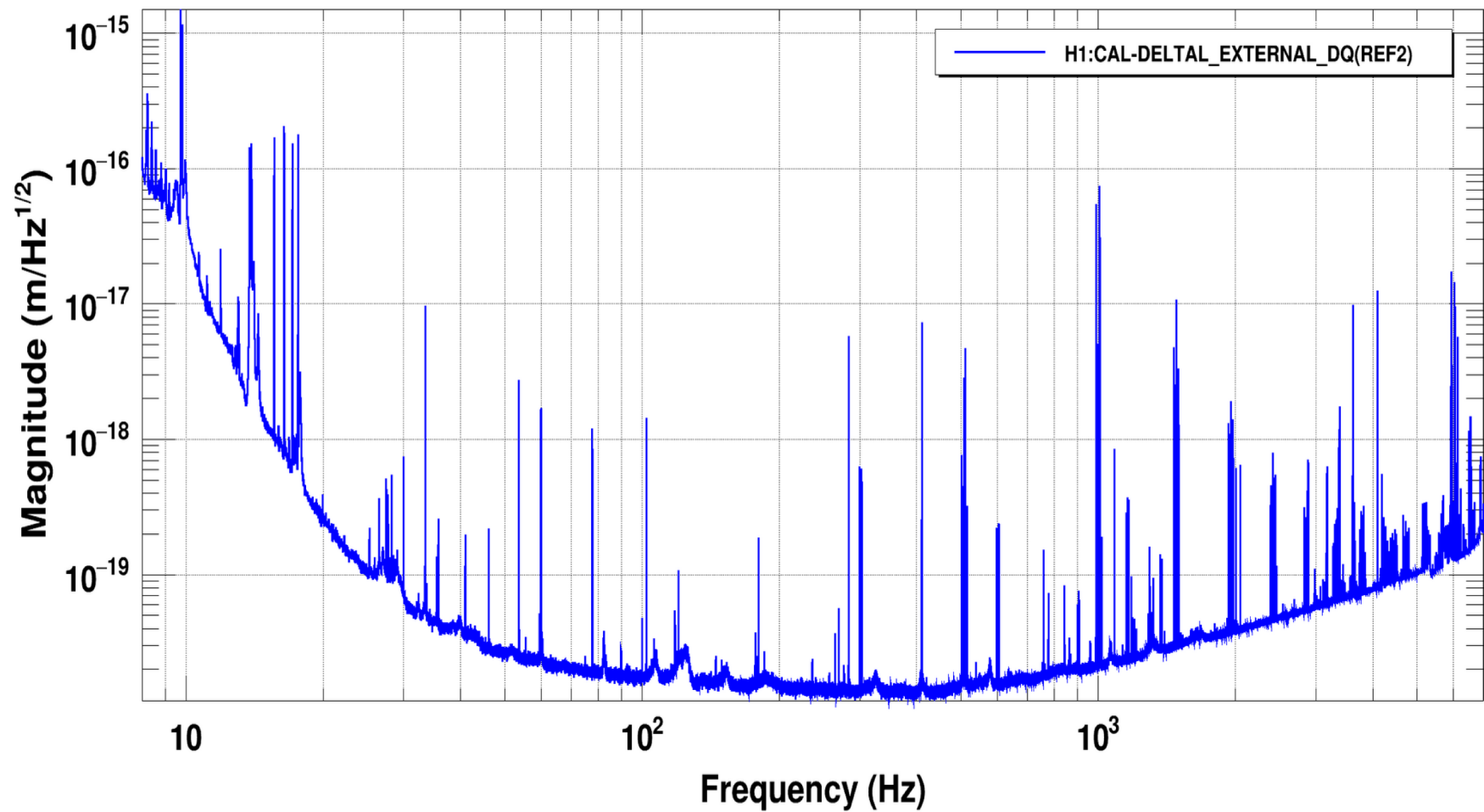
1) Quiet the motion of the mirrors to about $1e-19$ m at certain frequencies





***2) Measure the tiny distances:
interferometers are the best at
measuring small differences in distance***





T0=11/04/2024 01:30:00

Avg=100/Bin=2L

BW=0.0117098⁺

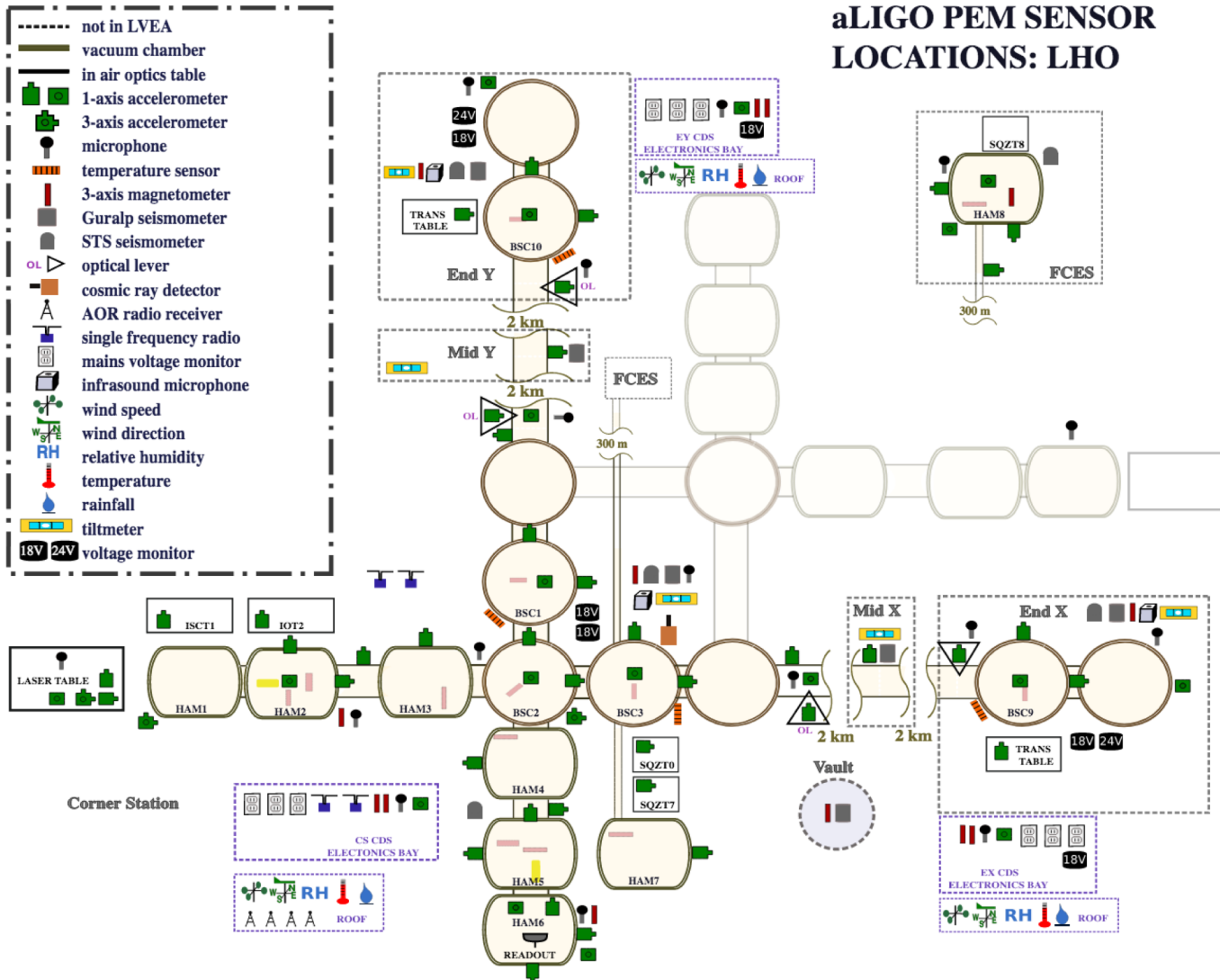
Seismic

Frequency band	Dominant Noise Source	How it can hurt GW detectors	Coupling
1/day	Earth tides	Changes distance between buildings	Stronger actuators are noisier. Higher velocity between buildings can make main laser light “mimic” GWs (scattered light noise).
Around 0.1 Hz	Ocean waves: microseismic peak	Same as above and also some relative motion within buildings.	Same as above: scattered light noise is one of the worst non-fundamental noise sources.
0.1 – 10 Hz and above	Wind	Shakes buildings, tilts ground	More motion gets through isolation, moving optics, and scattered light noise, also direct Newtonian coupling to TMs
1-30 Hz	Anthropogenic noise Trains, trucks, cars, OHVs, construction equipment, wind farms, dams	Motion within a building	Moving test masses and other optics, and scattered light noise, also direct Newtonian coupling to TMs
30 Hz to 7000 Hz	On site sources, HVAC, motors, people	Same as above	Same as above

Magnetic

Frequency band	Dominant Noise Source	How it can hurt GW detectors	Coupling
1–1000 Hz and up	Electric currents	magnetic coupling moves optics and other components	Actuation permanent magnets and magnetized components
1-1000 Hz and up but increasing with frequency	Electric currents	Induced currents add noise, mainly to analog signals	Electronic components and cables

aLIGO PEM SENSOR LOCATIONS: LHO





2 km

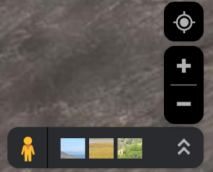
1.15 mi

LIGO Hanford Observatory
Recently viewed

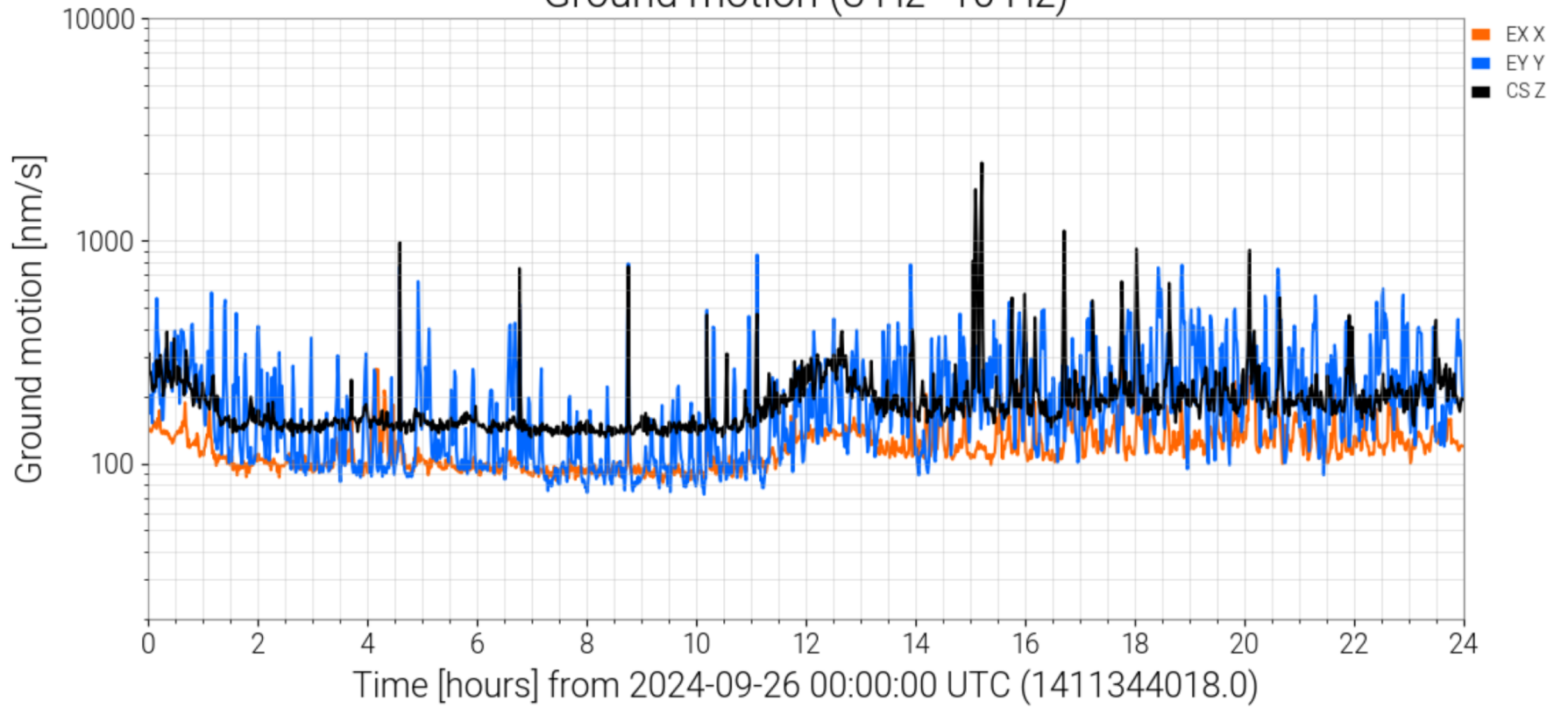
Energy Northwest

Geneva Junction

Measure distance ✕
Click on the map to add to your path
Total distance: 1.15 mi (1.84 km)

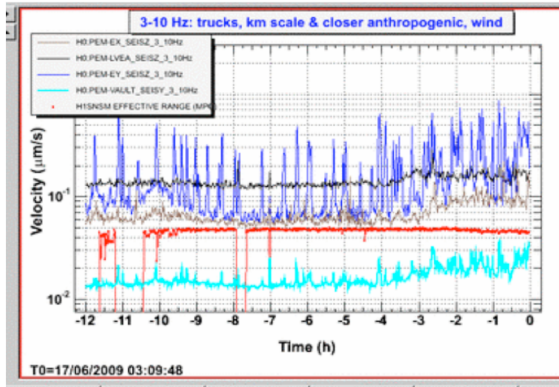


Ground motion (3 Hz--10 Hz)

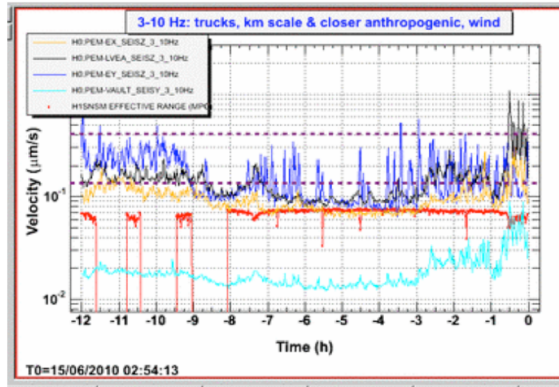


3-10 HZ PEAK SEISMIC NOISE DOWN BY >2 WITH RESURFACING OF HIGHWAY 240

Before



After



Percentiles from sorted 1 minute trend values (maxima) of H0:PEM-EY_SEISZ_3_10Hz for 150 days starting Nov. 2

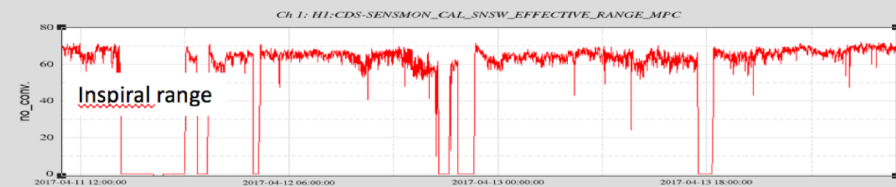
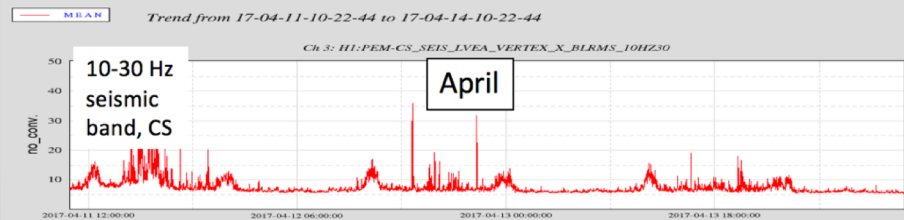
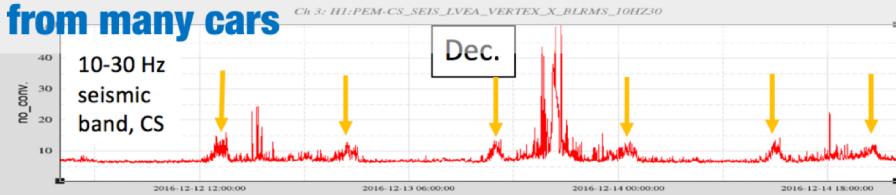
Percentile	Amplitude 2008-2009	Amplitude 2009-2010	After paving/ before paving
99.9	23982	8637	0.36
99	5560	2582	0.46
95	2630	1205	0.46
90	1633	769	0.47
75	575	309	0.54
50	251	152	0.61

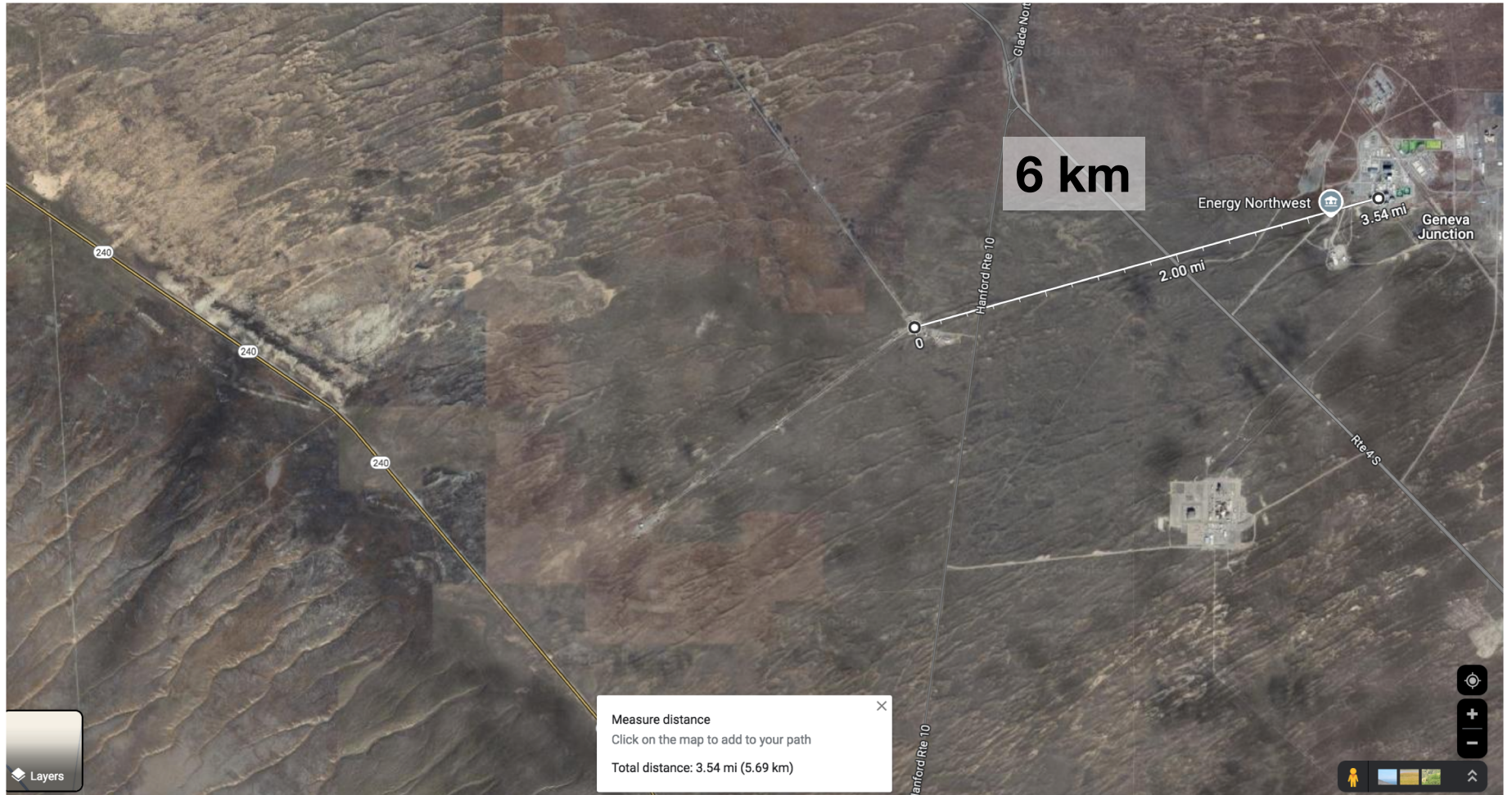
Maximum 10 Hz Newtonian noise estimates down by >2 for aLIGO

LHO noise from fire pump, HVAC, traffic

Work-day seismic peaks, and range dips, happen at rush hour and follow daylight savings, gaussian from many cars

B. Berger and other DetChar DQ shifters noticed seismic-associated events in DARM: trucks, rush hour traffic, HVAC, fire pump, all reduced range.

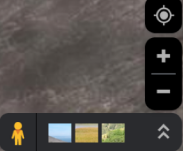




6 km

Measure distance
Click on the map to add to your path
Total distance: 3.54 mi (5.69 km)

Layers



Mobile seismometer suggested cooling towers.



Motor monitors and gear-teeth ratios allowed calculation of frequency for each of 36 fans: nearly all fell within our peak.

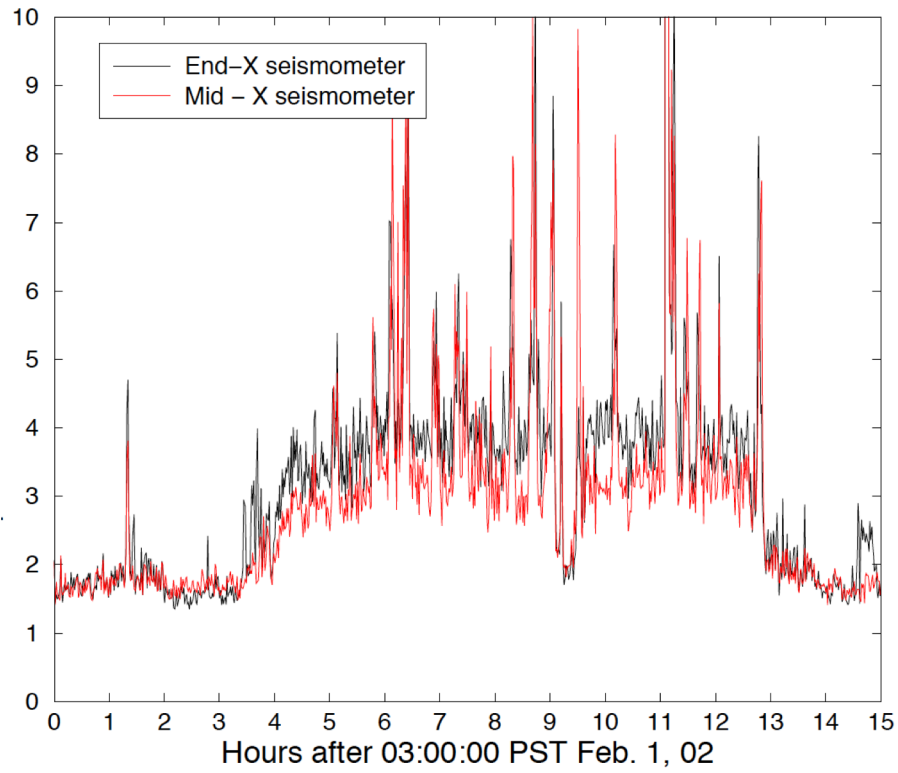


10 km

5.68 mi

Measure distance ✕
Click on the map to add to your path
Total distance: 5.68 mi (9.14 km)

1-3 Hz band



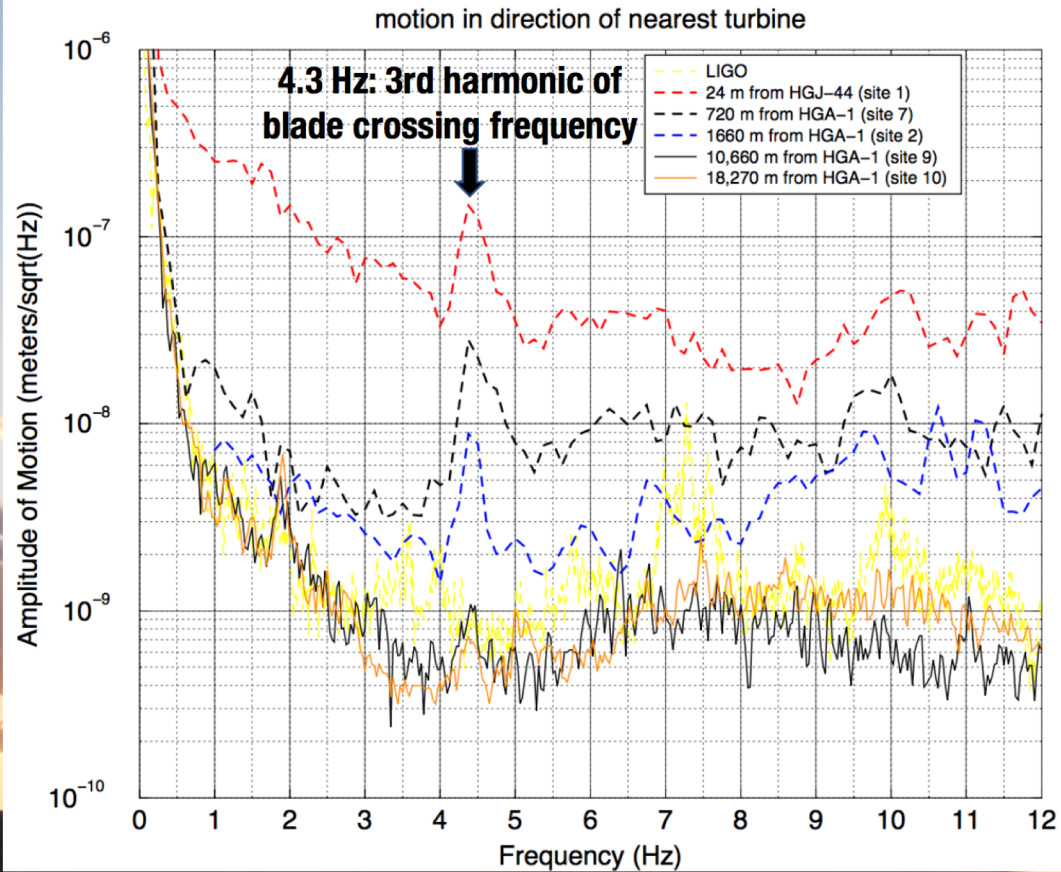
Traced to Vitrification Plant Project

10 km from X-end, up to 5000 workers expected during 5 year project



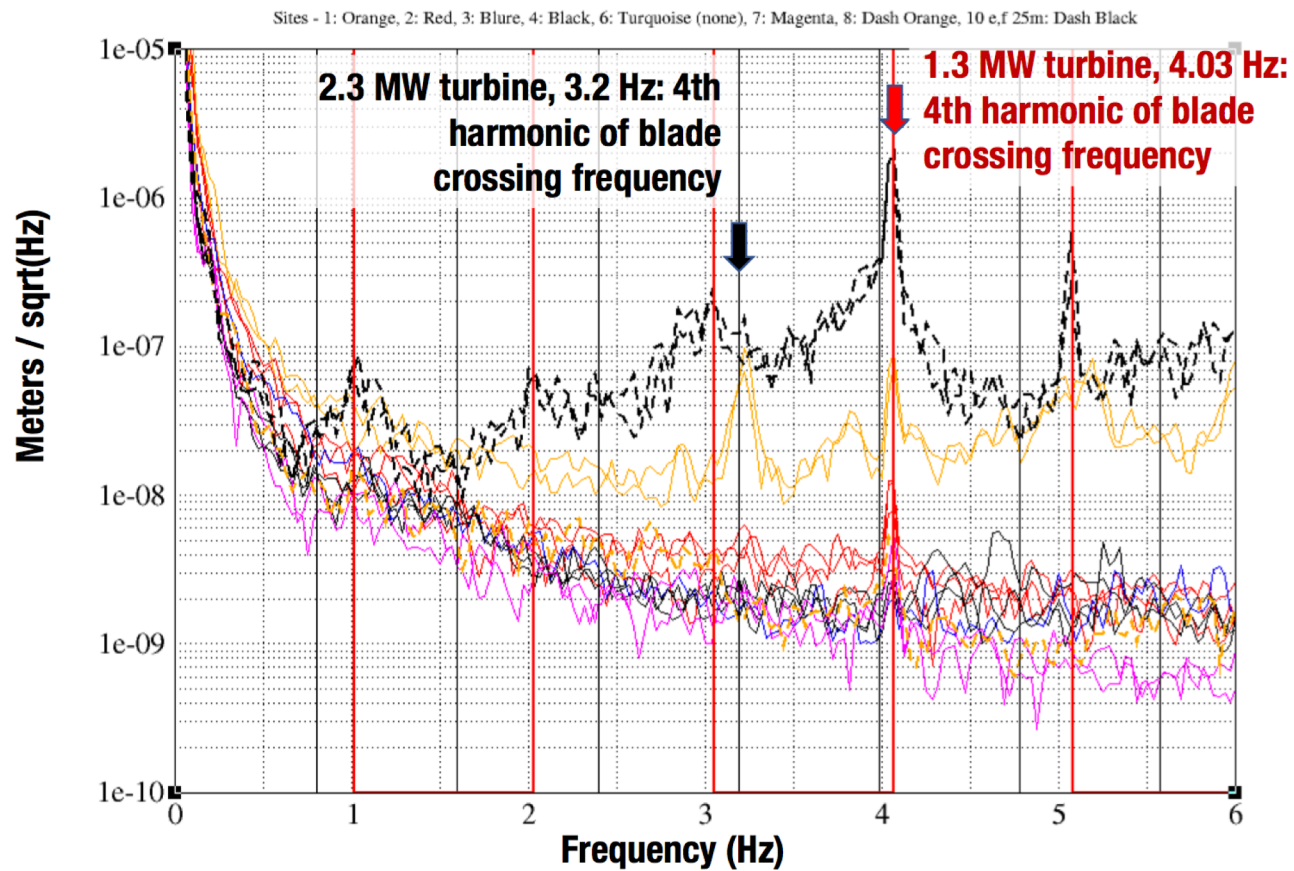
Sample of seismic measurements around State Line wind farm (State Line report, T020104-x0)

Horizontal Motion at Increasing Distances



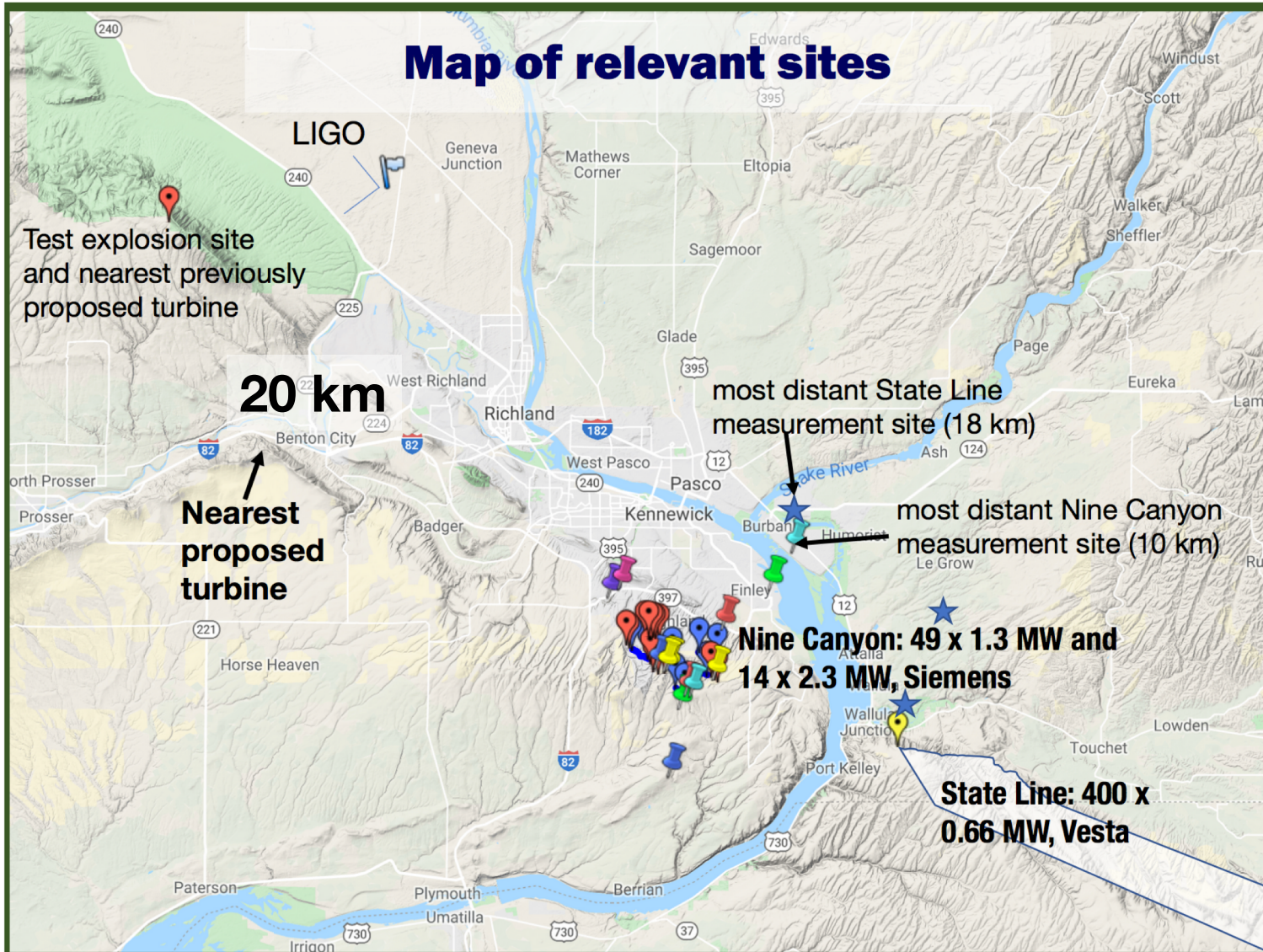
Sample of newer seismic measurements around Nine Canyon wind farm

Horizontal good data



Map of relevant sites

20 km







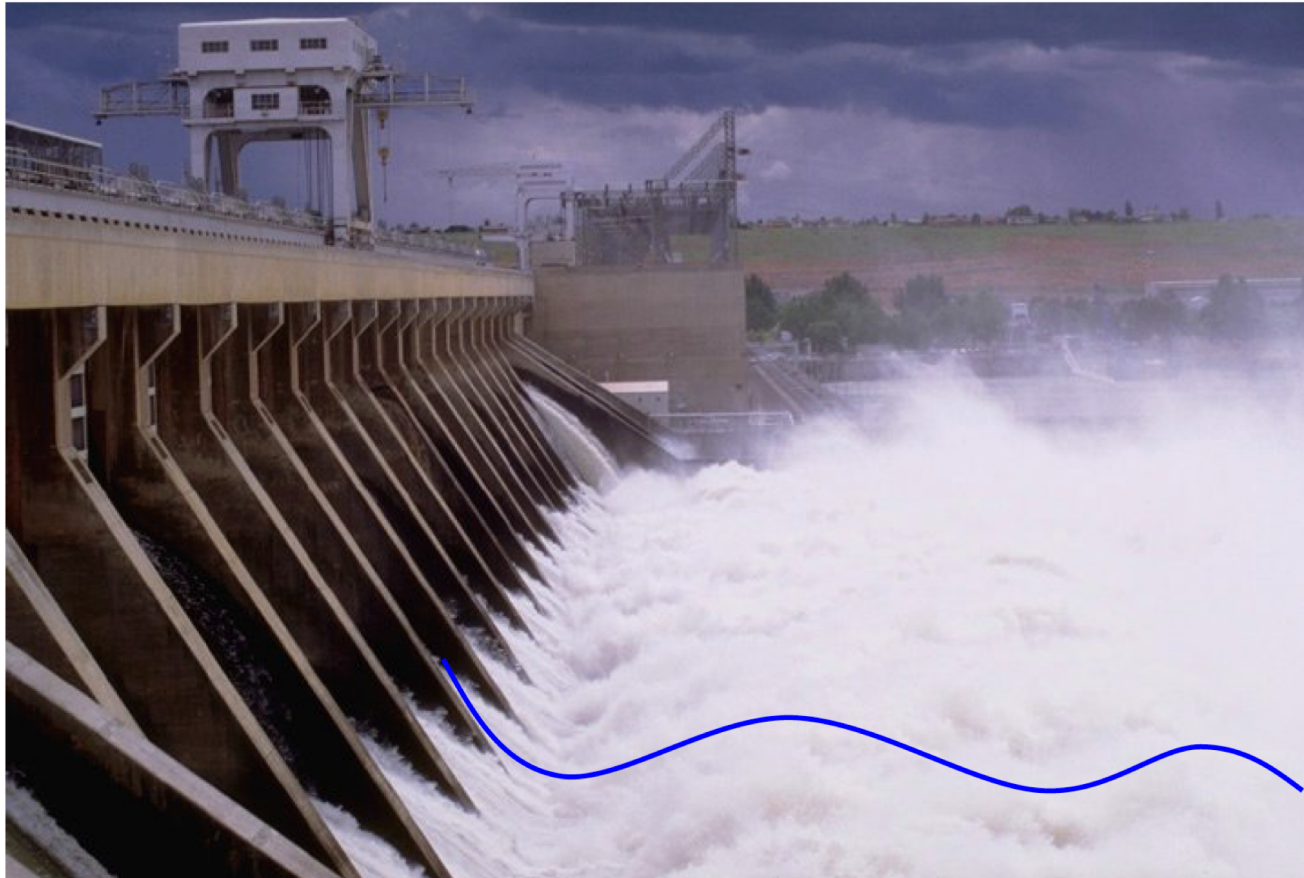
Measure distance
Click on the map to add to your path
Total distance: 36.34 mi (58.48 km)

60 km

20.00 mi

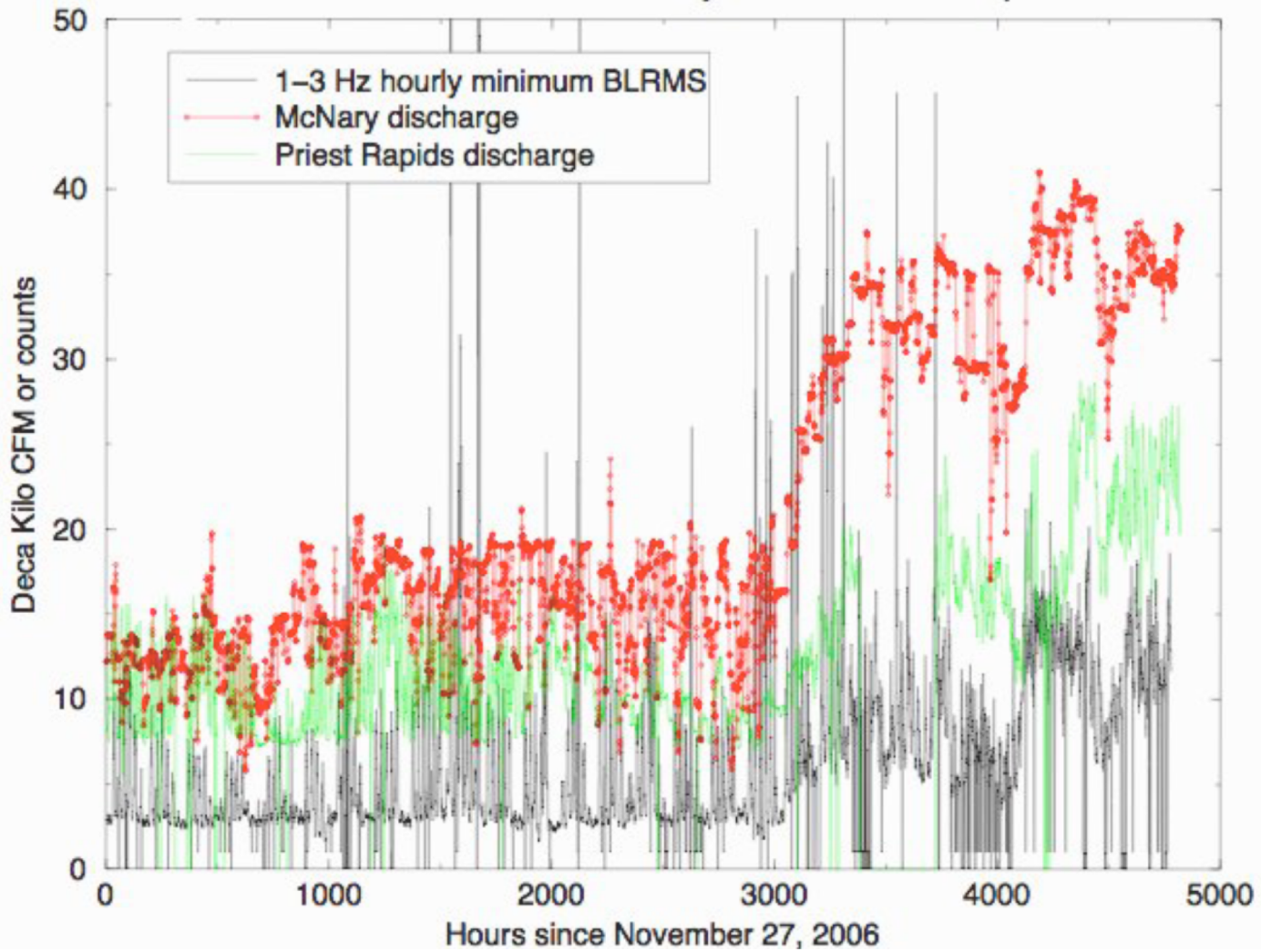
36.34 mi

**Dam affects inspiral range through upconversion of seismic signal at 1.2 Hz.
“Bounce” of water timed at 1 to 0.3 second.**

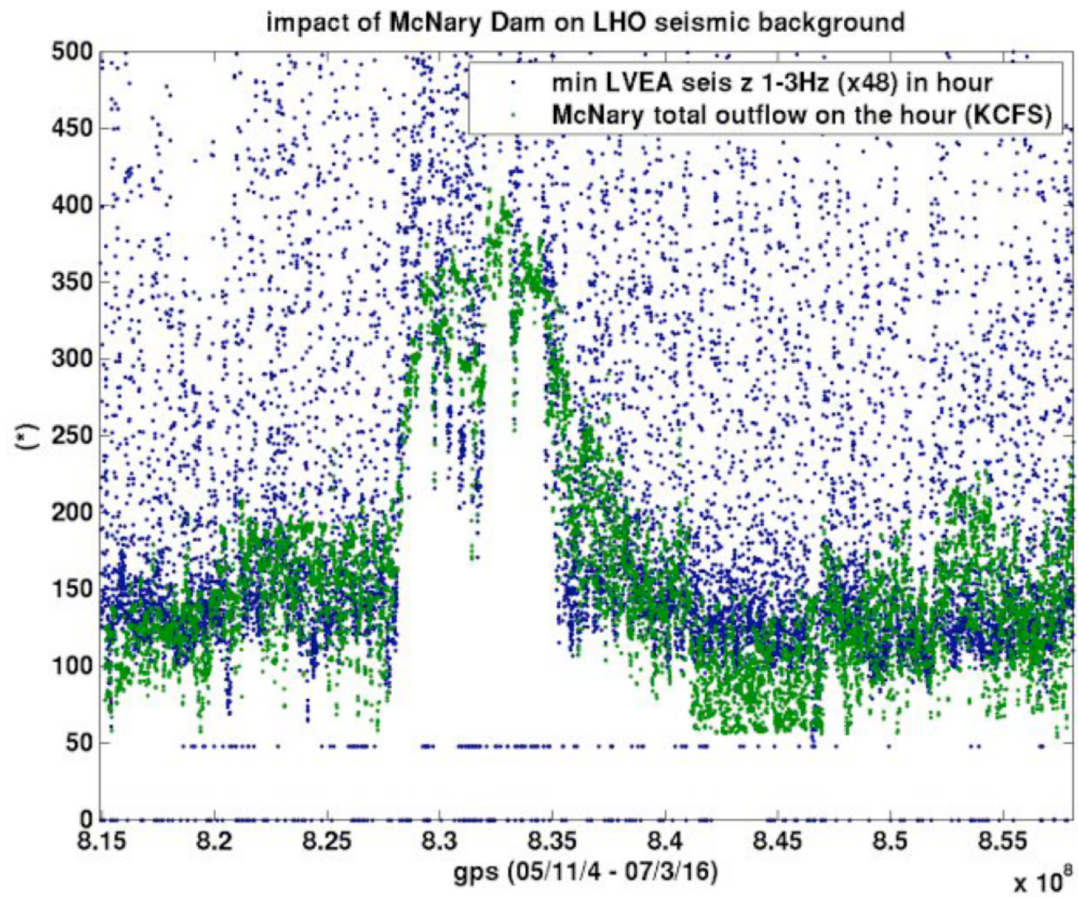


Dam Discharge and 1–3 Hz BLRMS hourly minimum

BLRMS minimum follows McNary better than Priest Rapids



Justins update showing entire run until present: dam has affected us for about 1/4 of the run so far



Transmission lines partly responsible for 60Hz peak in H1



Study of 31 site-wide magnetometer events

magnetic event no.	GPS Time	LocalTime	voltage event?	DARM KW Event?	BPA event (resolution mostly 2s)	Q-scan LVEA_V2 peak f (Hz)	Predicted DARM 70 Hz from mag field	Actual DARM 70 Hz
1	817736343	Dec 04, 2005 04:58:49 PST	Yes		Capacitors in, Benton substation	497		
2	817824114	Dec 05, 2005 05:21:40 PST	Yes		Capacitors in, Benton substation	592	6.00E-08	2.00E-07
3	817911868	Dec 06, 2005 05:44:14 PST	Yes		Capacitors in, Benton substation	631		
4	818073843	Dec 08, 2005 02:43:50 PST	Yes	H1	Capacitors in, Ashe substation	352		
5	818159114	Dec 09, 2005 02:25:00 PST	Yes		Capacitors in, Benton substation	691		
6	818426006	Dec 12, 2005 04:33:12 PST	Yes	H2	Trip, Marion north bus when switching reactor	69		
7	818507471	Dec 13, 2005 03:10:58 PST	Yes		Capacitors in, Benton substation	592	6.00E-08	2.00E-07
8	818511507	Dec 13, 2005 04:18:13 PST	Yes		Capacitors in, Richland substation	417		
9	818515503	Dec 13, 2005 05:24:49 PST	Yes		Capacitors in, Ashe substation	417	6.00E-09	2.00E-07
10	818684261	Dec 15, 2005 04:17:28 PST	Yes		Capacitors in, Ashe substation	417		
11	819002198	Dec 18, 2005 20:36:24 PST	Yes		Capacitors in, Ashe substation	389		
12	819036727	Dec 19, 2005 06:11:54 PST	Yes	H2	Capacitors in, Richland substation	352		
13	819160196	Dec 20, 2005 16:29:42 PST	Yes		Capacitors in, Richland substation	337		
14	819211052	Dec 21, 2005 06:37:18 PST	Yes		Capacitors in, Richland substation	367		
15	819257186	Dec 21, 2005 19:26:12 PST	Yes	H1 And H2	Fault on Ashe-Slatt 500 kV line	68	3.00E-06	4.00E-06
16	819384812	Dec 23, 2005 06:53:18 PST	Yes	H1 And H2	Fault on Slatt-Buckley 500 kV line	95	6.00E-07	6.00E-07
17	819386152	Dec 23, 2005 07:15:39 PST	Yes	H1 And H2	Fault on Benton-Othello 115 kV line	53		
18	819673745	Dec 26, 2005 15:08:52 PST	Yes	H1 And H2	None found	95		
19	820394439	Jan 03, 2006 23:20:25 PST	Yes	H1 And H2	Trip, Ice Harbor-Franklin 115 kV line	68	5.00E-07	3.00E-07
20	820595204	Jan 06, 2006 07:06:31 PST	Yes	H2	Trip, Vantage-Schultz 500 kV line	50		
21	821270678	Jan 14, 2006 02:44:24 PST	Yes	H1 And H2	Trip, Benton-Franklin 230 kV line	53	1.40E-06	2.00E-06
22	821459495	Jan 16, 2006 07:11:22 PST	Yes	H1	Capacitors in, Richland substation	350		
23	823009424	Feb 03, 2006 05:43:31 PST	Yes	H1 And H2	Fault on Lower Monumental-Ashe 500 kV line	53		
24	823355073	Feb 07, 2006 05:44:19 PST	Yes	H1	Capacitors in, Benton substation	592		
25	825082556	Feb 27, 2006 05:35:42 PST	Yes		Capacitors in, Benton substation	691	2.00E-08	6.00E-07
26	825170442	Feb 28, 2006 06:00:28 PST	Yes		Capacitors in, Benton substation	497		
27	826119810	Mar 11, 2006 05:43:16 PST	Yes		Capacitors in, Benton substation	592		
28	827089513	Mar 22, 2006 11:04:59 PST	No		None found	mag at about 10	9.00E-08	2.00E-07
29	827506076	Mar 27, 2006 06:47:42 PST	Yes		Capacitors in, Benton substation	757	3.00E-08	2.00E-07
30	827592239	Mar 28, 2006 06:43:45 PST	Yes	H1	Capacitors in, Benton substation	592		
31	828112266	Apr 03, 2006 08:10:53 PST	No		None found	mag at about 10		

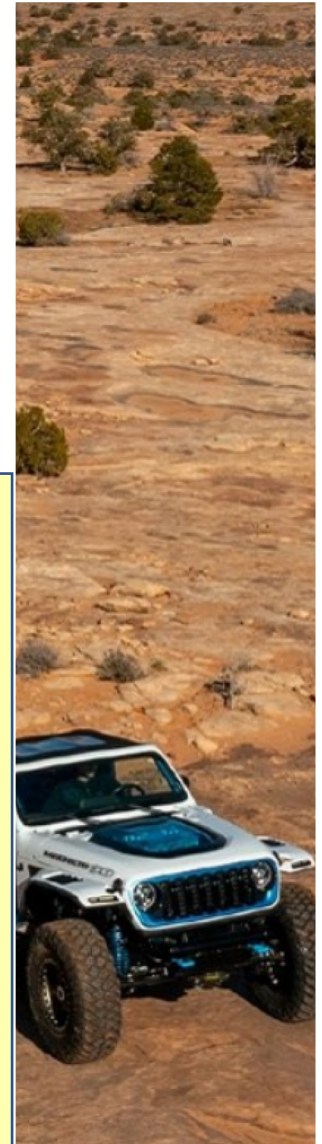
Collaboration with Mike Viles, BPA

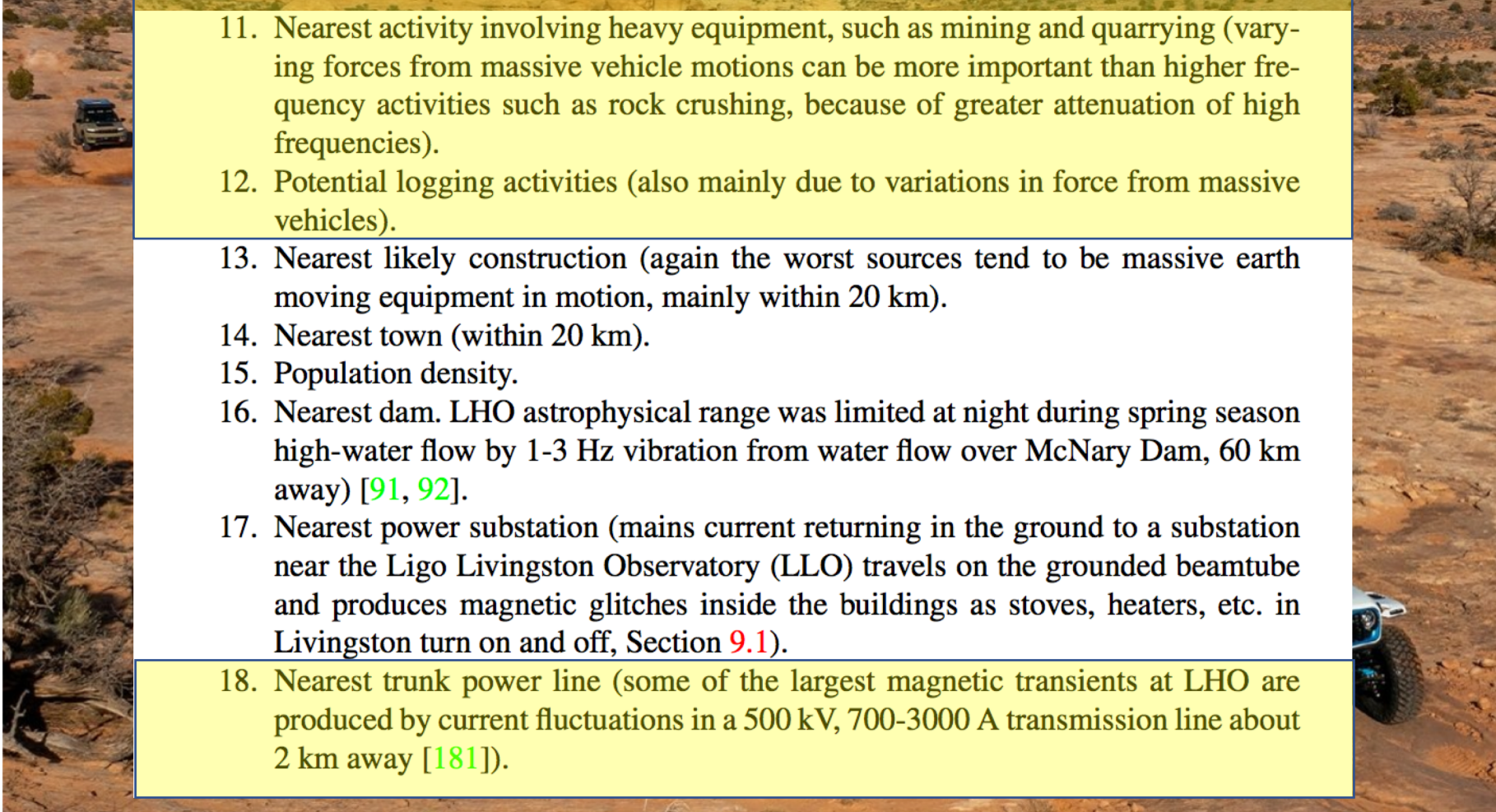
Suggested day-scale instrumented site evaluation

- Will not deal with earth tides, microseismic peak, wind driven seismicity that vary on week scales and greater
- Focus on local sources that are present daily but may not be captured by seismic maps and other external sources.
- Use imperfect but nimble setups so that we can measure at multiple sites, and alter our plans in order to investigate a signal.
- Focus on seismic and magnetic signals, which are the worst environmental noise for LIGO

Prior to actual site measurements, candidate sites can be evaluated using information available on the internet and Google Earth. This includes most of the considerations in the list below:

1. Geology.
2. Microseismic peak level (minimize relative motion of detector stations, Section 6.3).
3. Wind speed distribution (minimize wind greater than about 10 m/s).
4. Expected weather changes associated with climate change.
5. Nearest train tracks and density of train traffic on those tracks (trains can produce 1-3 Hz seismic noise, especially within 20 km, and, if electric, can produce magnetic noise transients below 10 Hz especially less than 10 km away, Section 9.4).
6. Nearest highway, and truck traffic density (especially within 10 km; frequency, given by axle spacing and speed, is generally in the 4-15 Hz band; transient vibrations from trucks can be an order of magnitude worse than from cars, and at lower frequencies due to greater axle spacing [180]).
7. Nearest bridges and viaducts (including those for trains), which may have structural resonances that are excited by vehicles (see also Section 6.4).
8. Conditions of nearest roads (the largest 4-15 Hz seismic signal at a LIGO Hanford Observatory (LHO) end station was reduced by a factor of 2 to 3 when highway 240, 2 km away, was resurfaced [98]).
9. Nearest gravel or dirt road and traffic density (gravel and dirt roads typically produce much more vibration noise per vehicle than other roads).
10. Off-highway recreational vehicle activity within 10 km (the signal from variations in force against the ground is typically in the 3-15 Hz band).



- 
11. Nearest activity involving heavy equipment, such as mining and quarrying (varying forces from massive vehicle motions can be more important than higher frequency activities such as rock crushing, because of greater attenuation of high frequencies).
 12. Potential logging activities (also mainly due to variations in force from massive vehicles).
 13. Nearest likely construction (again the worst sources tend to be massive earth moving equipment in motion, mainly within 20 km).
 14. Nearest town (within 20 km).
 15. Population density.
 16. Nearest dam. LHO astrophysical range was limited at night during spring season high-water flow by 1-3 Hz vibration from water flow over McNary Dam, 60 km away) [91, 92].
 17. Nearest power substation (mains current returning in the ground to a substation near the Ligo Livingston Observatory (LLO) travels on the grounded beamtube and produces magnetic glitches inside the buildings as stoves, heaters, etc. in Livingston turn on and off, Section 9.1).
 18. Nearest trunk power line (some of the largest magnetic transients at LHO are produced by current fluctuations in a 500 kV, 700-3000 A transmission line about 2 km away [181]).

19. Does it share an electrical grid with another GW detector (unlikely, but may be a source of correlation between detectors)?
20. Is it on a heavily used flight path of a nearby airport (does it line up with a runway within about 50 km)?
21. Proximity to military training areas (LHO experiences 1-12 Hz seismic and acoustic transients from tank and artillery practice at the Yakima Training Center, about 50 km away [182]).
22. Proximity to potential wind farms (infrasound coupling to the ground may be important, Section 6.4).
23. Proximity to other facilities with large rotating equipment (cooling fans at a nuclear power plant 6 km from LHO produce seismic signals that are at least an order of magnitude above background at about 2 Hz [93]).

This workshop's activities

- Identify truck transients from roads around the site
- Determine if the Ashe-Slatt, Ashe-Marion transmission lines are the main sources of magnetic fields and transients
- Try out a nimble system where we collect data and evaluate it “in the van”
- We will use a WebDAQ acquisition system that allows us to monitor spectra real-time
- We will use jupyter notebooks to display the data in different ways “in the van”