

Modeling Compact Mergers detected by LIGO with Bivariate B-splines

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Hierarchical Bayesian Modeling

$$p(\vec{\Lambda}|\{\vec{d}_i\}) = \pi(\vec{\Lambda}) \prod_{i=1}^{N_{\text{obs}}} \frac{\frac{1}{S_i} \sum_{j=1}^{S_i} p_{\text{pop}}(j\vec{\theta}_i|\vec{\Lambda}) \frac{1}{\pi(\vec{\theta})}}{\int d\vec{\theta} p_{\text{det}}(\vec{\theta}) p_{\text{pop}}(\vec{\theta}|\vec{\Lambda})}$$

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Modeling compact mergers

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- Tracing out cosmological history

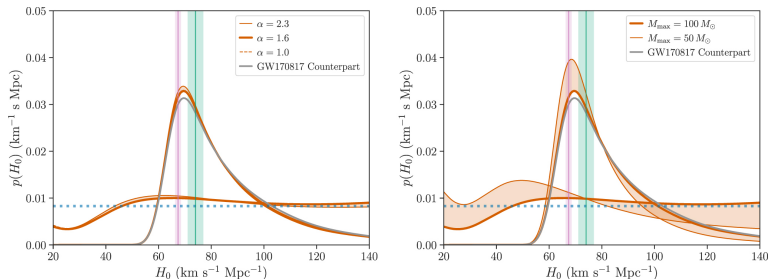


Figure: B.P. Abbott et al. 2021

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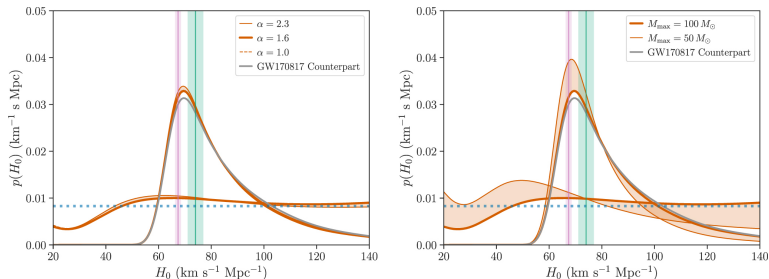


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- ...all of these rely on population models

Modeling compact mergers

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- *pros*: data-driven, introduces less bias, reveal substructures

Population models

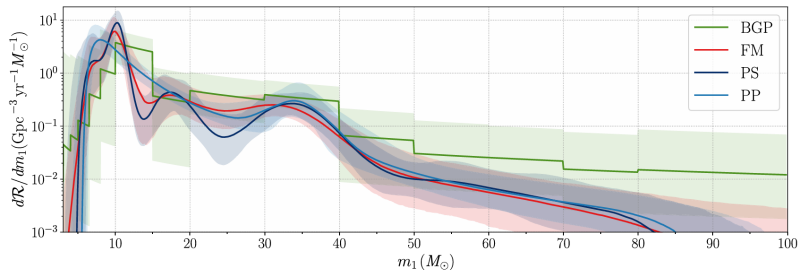


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Non-parametric model: B-splines



Figure: *Left*: parametric models. *Right*: non-parametric models.

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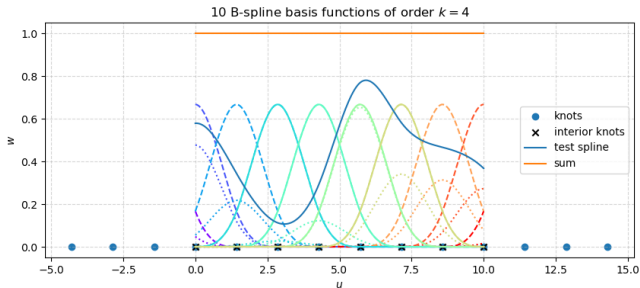
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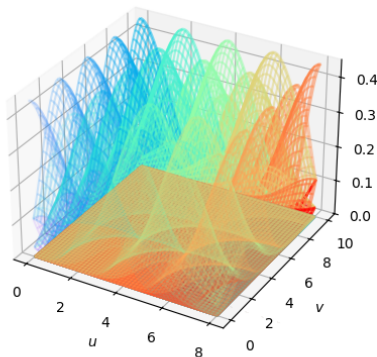
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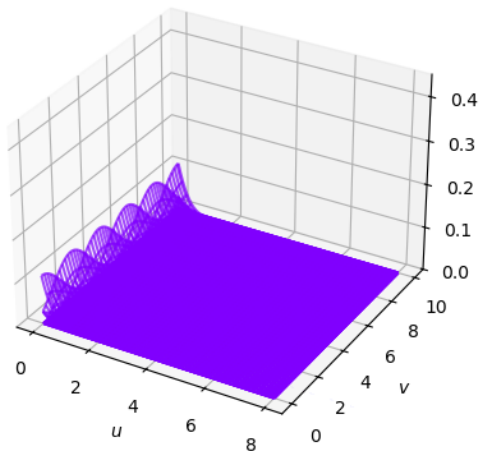
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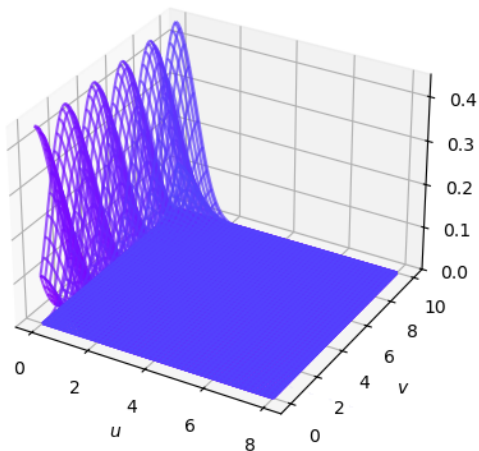
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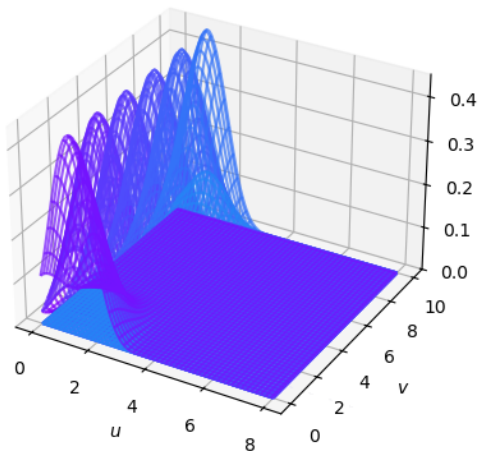
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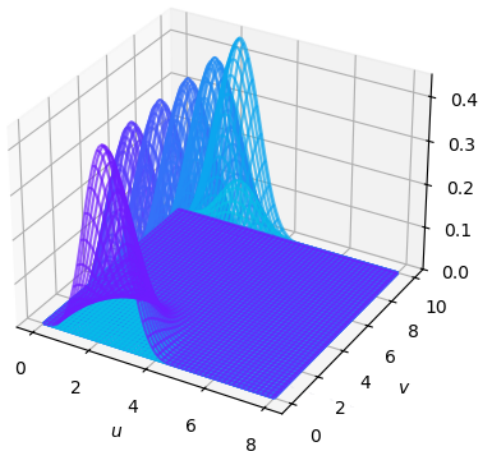
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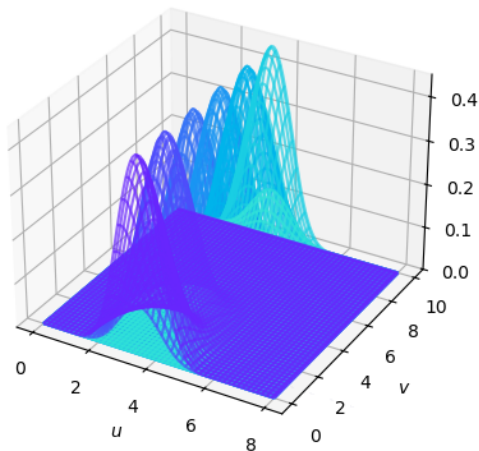
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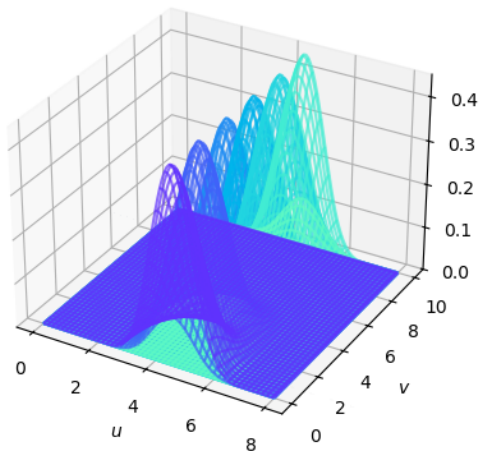
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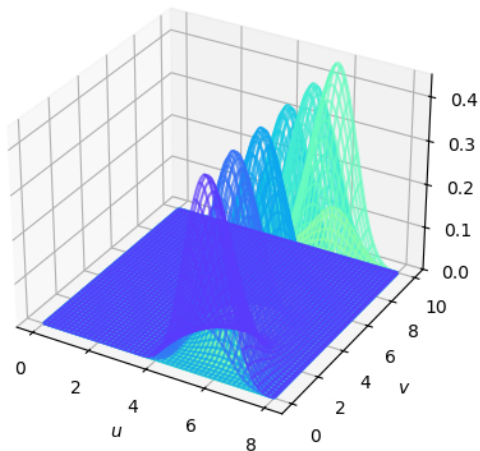
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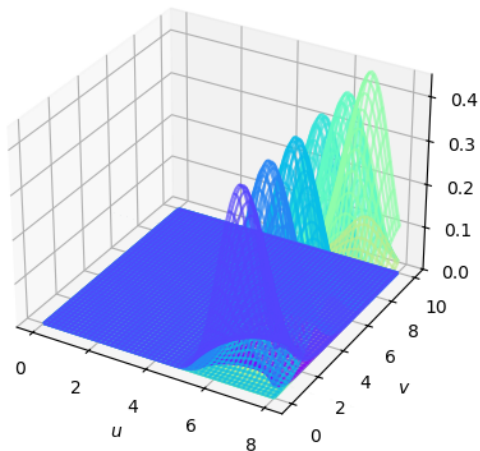
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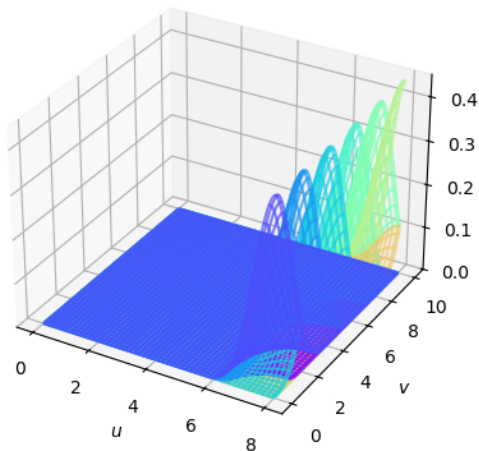
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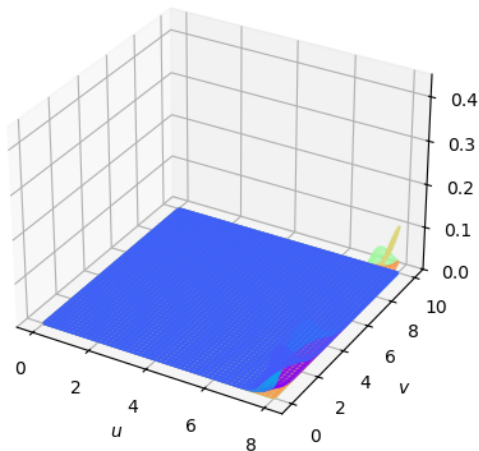
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$$\begin{aligned} p(a_1, a_2, \cos \theta_1, \cos \theta_2 | \alpha) &= \sum_{i,j}^{I,J} \alpha_{ij} B_{i,k}(a_1) B_{j,k}(\cos \theta_1) \\ &\times \sum_{\ell,m}^{L,M} \alpha_{\ell m} B_{\ell,k}(a_2) B_{m,k}(\cos \theta_2) \\ &= \mathcal{B}_1(a_1, \cos \theta_1 | \alpha) \mathcal{B}_2(a_2, \cos \theta_2 | \alpha). \end{aligned}$$

Mass ratio and effective spin

- GW observable χ_{eff}

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$$\begin{aligned}\chi_{\text{eff}} &= \frac{(m_1 \vec{a}_1 + m_2 \vec{a}_2) \cdot \hat{L}}{m_1 + m_2} \\ &= \frac{m_1 a_1 \cos \theta_1 + m_2 a_2 \cos \theta_2}{m_1 + m_2} \\ &= \frac{a_1 \cos \theta_1 + q a_2 \cos \theta_2}{1 + q},\end{aligned}$$

\hat{L} is a unit vector parallel to the orbital angular momentum of the binary system

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- Anti-correlations found in GWTC-3

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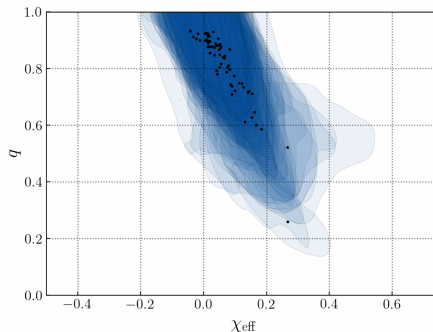


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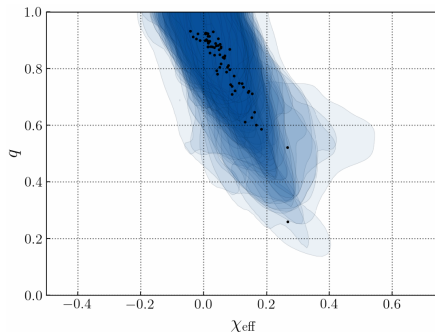


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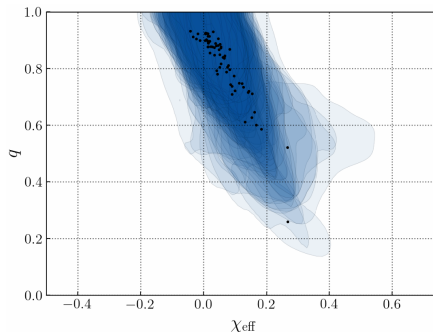


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- Break mass degeneracy between isolated and dynamical formation \rightarrow model as bivariate spline!
- Inform model building, disentangle formation channels

Backup

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and for $k > 1$,

$$B_{i,k} = \omega_{i,k} B_{i,k-1} + (1 - \omega_{i+1,k}) B_{i+1,k-1},$$

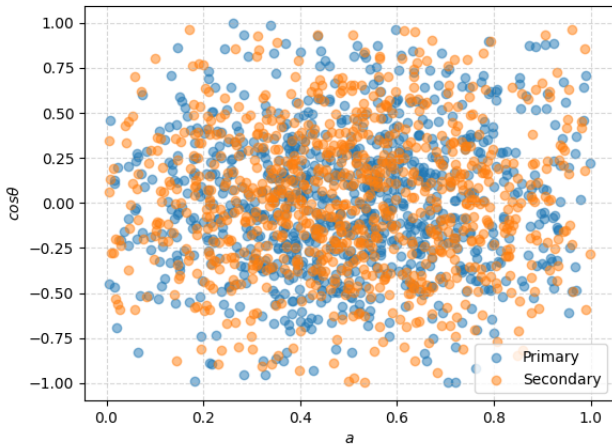
where

$$\omega_{i,k}(x) \equiv \frac{x - u_i}{u_{i+k-1} - u_i}$$

- Choosing priors

Current work

- Choosing priors



Choosing priors

Model	Degrees of freedom	Prior
1	100	$\mathcal{N}(0, 1)$
2	100	$\mathcal{N}(\alpha_s, 1)$
3	196	$\mathcal{N}(0, 1)$
4	196	$\mathcal{N}(\alpha_s, 1)$

Table: Choices for priors on the coefficients α . When a normal prior is centered on α_s , this means each α_{ij} is centered on a value α_s^{ij} . The values α_s are found by minimizing the least squares objective $S = \|y - B\alpha\|^2$, where B is the basis matrix containing the basis functions of the B-spline, and y contains the true values of the distribution.

Bivariate spline fit

