

Candidate Beam Tube Materials Vacuum Properties

Carlo Scarcia (w/ data contribution from Aiman H. Al-Allaq and James Fedchak) 2025/09/30, BTW3 workshop

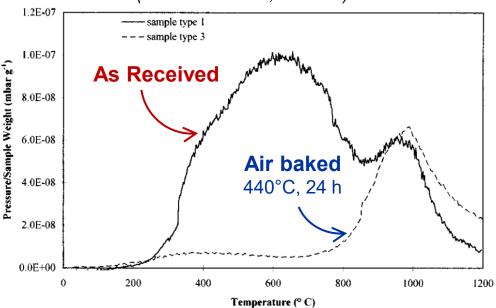
2G-GWD beam pipe material: 304L air-baked

Air bakeout (L. Petermann, 1965 FR1405264)

→ H₂ depletion + diffusion barrier w/ Fe-rich passivation layer

	LIGO ^[1]	Virgo ^[2]
Applied on	Coils	Tubes
Tube thickness [mm]	3.23	4
Air bake recipe	454°C, 36 h	410°C, 72 h
In situ bakeout	160°C, ~3 weeks	150°C, ~1 weeks
In-field specific outgassing rates [mbar I s ⁻¹ cm ⁻²]	6.1 - 8.4 × 10 ⁻¹⁴ (@ 23°C)	≈ 1×10 ⁻¹⁴ (@ 20°C)

Thermal desorption measurements on 304L specimens (30 x 10 x 5 mm, 5 K/min) [3]





[1] Weiss R., Residual Gas in the LIGO Beam Tubes: Science, Arts and Recipes, AVS 50, 2003

[2] Pasqualetti A., Virgo Vacuum System, INFN-LASA, 13/06/2024

[3] Bernardini et al., Air bake-out to reduce hydrogen outgassing from stainless steel, JVST A, 16, 188–193 (1998)



Scalability problem from 2G-GWD to 3G-GWD



Raw material & air-bakeout cost

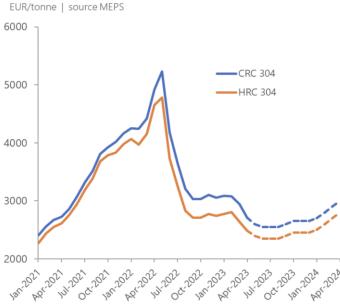


Production time



New requirements
vs
vacuum system cost

Europe Coil Price Forecast



VIRGO production rate (air-bakeout limited): 20 m/day

ET (120 km) production time: ~16 years (365 days production)

Residual gas noise contribution^[6,7]: ((x~10 better than LIGO/VIRGO)

ET: <10⁻²⁵ Hz^{-1/2}

CE: <5 ×10⁻²⁶ Hz^{-1/2}

Forecasted budget for beampipe vacuum (from 2G)^[6,7,8]:

ET: ~440 M€

CE: ~730 M\$

- [6] Matthew Evans et al. Cosmic Explorer Horizon Study, 2021
- [7] ET Science Team, Einstein gravitational wave Telescope conceptual design study, 2011;
- [8] ET collaboration, Einstein Telescope preliminary cost book, 2020



Cost reduction guidelines from previous workshops^[4,5]



Less expensive beam pipe materials

Mild steels & Ferritic stainless steels

Literature review: only four articles reporting on UHV studies on mild steels and ferritic stainless steels.

C. D. Park et al., J. Vac. Sci. Technol. A 26, 1166 (2008).

C. D. Park et al., J. Vac. Sci. Technol. A 34, 021601 (2015).

S. Kato et al., J. Vac. Soc. J. 55, 160–163 (2012).

J. Kamiya et al., Vacuum 98, 12–17 (2013).

Results limited to water outgassing at room temperature and the total outgassing rate after bakeout



Lower bakeout temperatures (<150°C)

Low binding energy coatings (Fe₃O₄, a-Si:H)





The landscape of labs and measurements









H ₂ O outgassing rate (Room temperature)	Throughput	Throughput	Throughput	-
Outgassing rates after bakeout	Coupled accumulation-throughput (H _{2,} CH _{4,} CO and CO ₂)	Rate of rise (H ₂)	Rate of rise (H ₂)	Rate of rise (H ₂)
Other measurements	TPD XPS Water binding energy	-	Water binding energy	XPS
Materials	Mild steel, Ferritic StS. Austenitic StS.	Mild steel Austenitic StS.	Mild steel Austenitic StS.	Mild steel
Coatings	a-Si:H	_	Fe ₃ O ₄	Fe ₃ O ₄

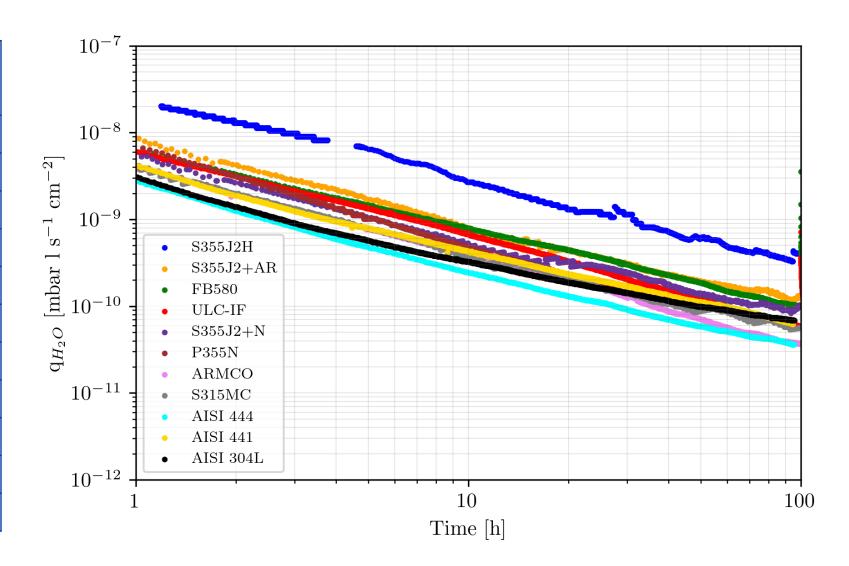
Reminder: The cleaning method/procedures were different!



Water outgassing rate at room temperature

Throughput method

	Sample	q _{10h} (x 10 ⁻¹⁰) [mbar l s ⁻¹ cm ⁻²]
	S355J2H	27
	S355J2+AR	8.1
	FB580	7.9
Mild steel	ULC-IF	6.7
Ĭ	S355J2+N	4.9
	P355N	4.7
	ARMCO	3.9
S	S315MC	3.7
Ferritic StS	AISI 444	2.4
erriti	AISI 441	4.3
ц	304L	3.3

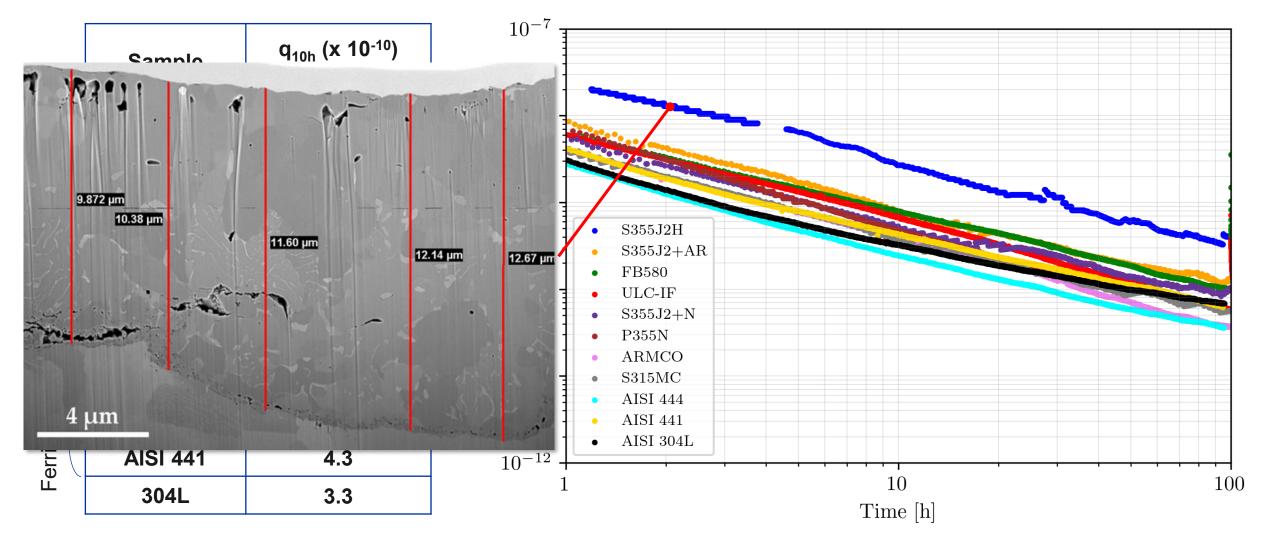


Background system removed.



Water outgassing rate at room temperature

Throughput method



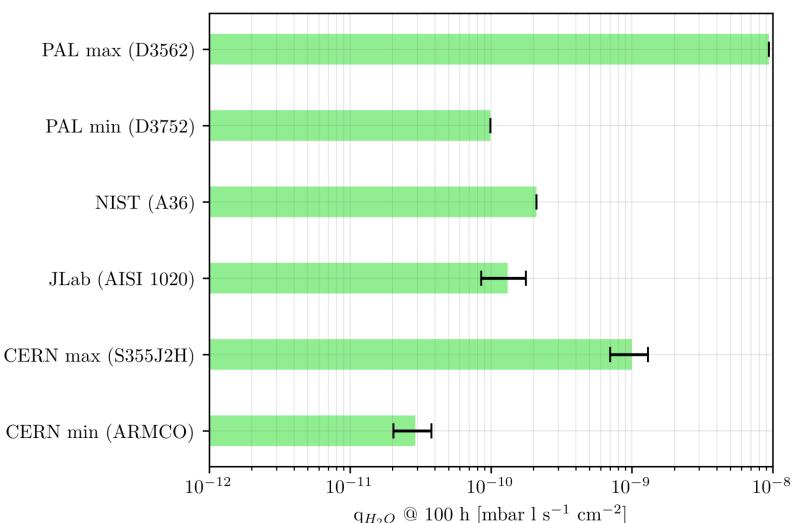
Background system removed.



Water outgassing rate at room temperature

Throughput method (100h pumping at RT)

	Tmeas. [°C]	Pre-treatment
CERN ^[a] (min, max)	21±2	None
Jefferson Lab [unpublished]	25	Acid etched
NIST ^[b]	24 (avg.)	(Machined?)
PAL ^[c] (min)	24±1	Machined + baked 150°C, 24 h + vented for 5h w/ N2
PAL ^[c] (max)	24±1	baked 150°C, 24 h + vented for 5h w/ N2



[[]c] Park. C. et al., J. Vac. Sci. Technol. A 34, 021601 (2016)



[[]a] Scarcia C. et al., J. Vac. Sci. Technol. B 42, 054202 (2024)

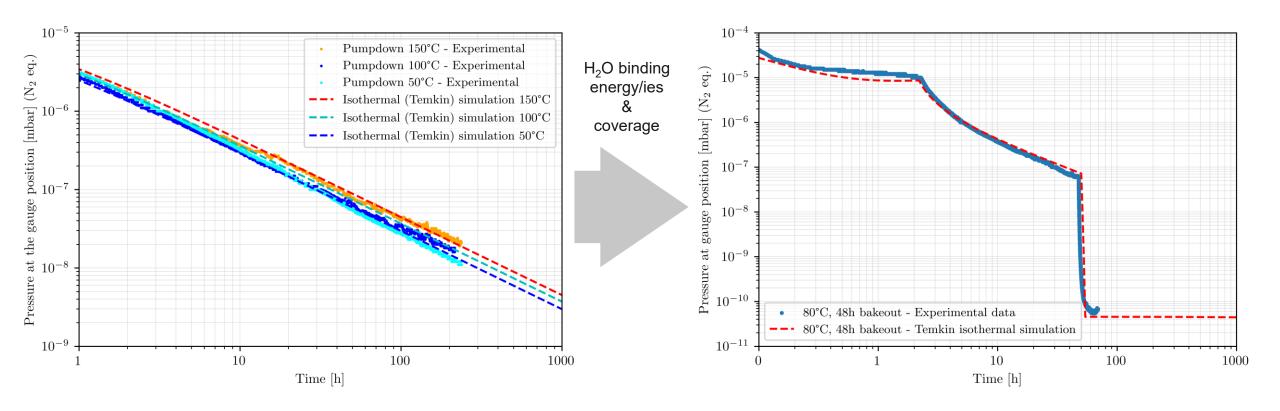
[[]b] Fedchak J., Beampipes for Gravitational Wave Telescopes 2023, CERN, 27-29 March 2023

Water vapour binding energy

Throughput method

Objective: extraction of water vapour binding energy/ies and coverages (distribution) to simulate bakeouts.

Example from a mild steel chamber (P355N)



Background removed. Pumping speed corrected to gauge position.



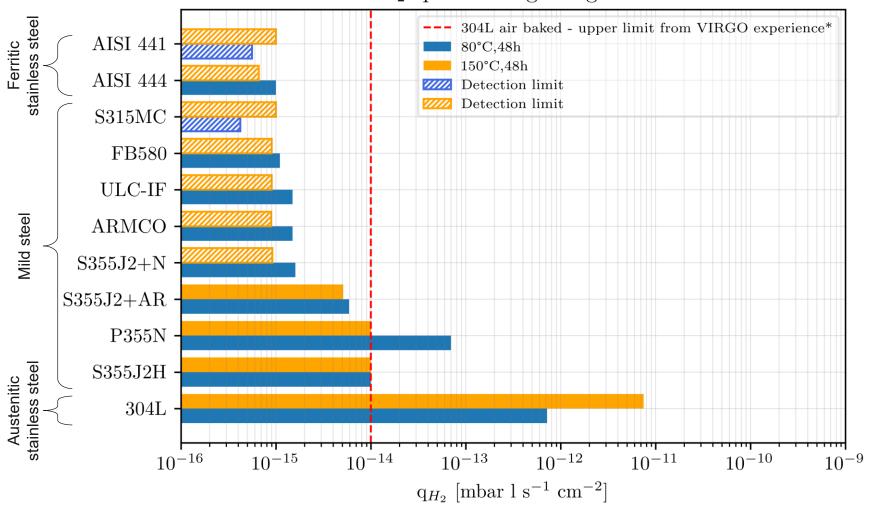
H₂ outgassing rates after bakeout

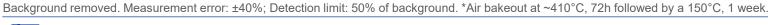
Throughput-accumulation method

H₂ specific outgassing rate

Two bakeout temperature sets:

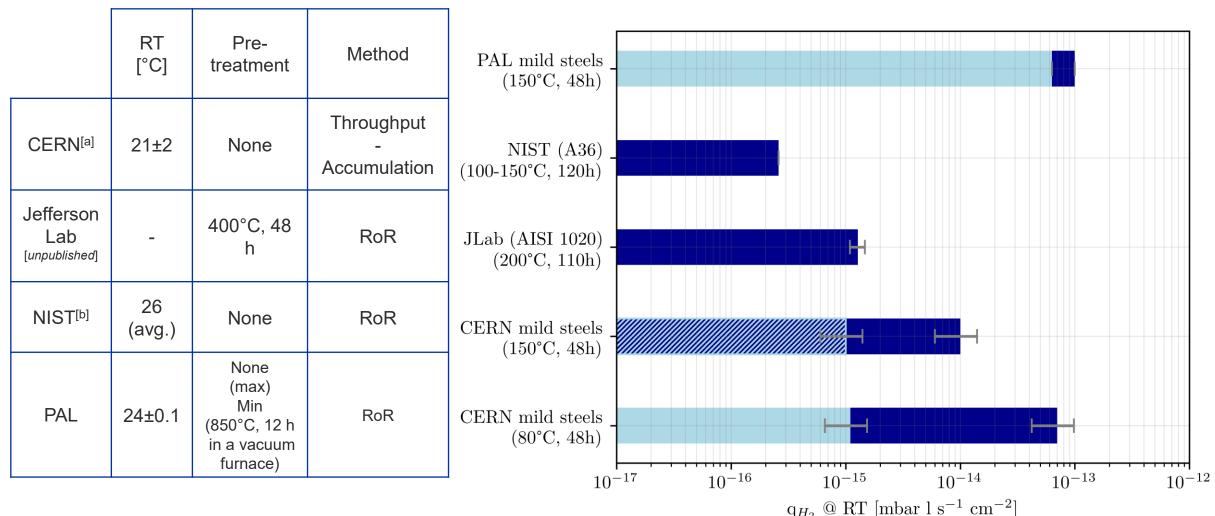
- 80°C: "Economical" alternative
- 150°C: Reference used in the current generation of GWD







H₂ outgassing rate after bakeout



[[]a] Scarcia C. et al., J. Vac. Sci. Technol. B 42, 054202 (2024)

[[]c] Park. C. et al., J. Vac. Sci. Technol. A 34, 021601 (2016)



[[]b] Fedchak J., Beampipes for Gravitational Wave Telescopes 2023, CERN, 27-29 March 2023

H₂ content

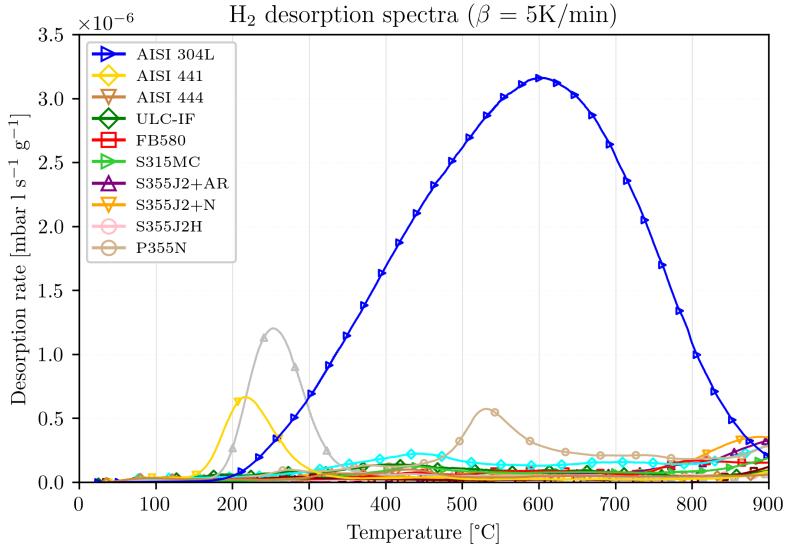
Temperature Programmed Desorption (TPD)

	Steel grade	H ₂ content [ppm at.]	Microstructure
ts (304L	80	Austenitic
S (AISI 441	1.7	Ferritic
Ferritic StS	AISI 444	1.3	Ferritic
Щ	ULC-IF	3.7	Ferritic
	FB580	2.8	Ferritic - Bainitic
	S315MC	2.7	Ferritic - Pearlitic
<u>a</u>	S355J2+AR	2.0	-
steel	S355J2+N	1.6	-
Mild	ARMCO	1.2	-
	S355J2H	7.8	-
	P355N	1.0	-

Concentration calculated from quantity of $\rm H_2$ (considered to be uniformly distributed) extracted with TPD (up to 850°C).

Graph normalized to sample weight.

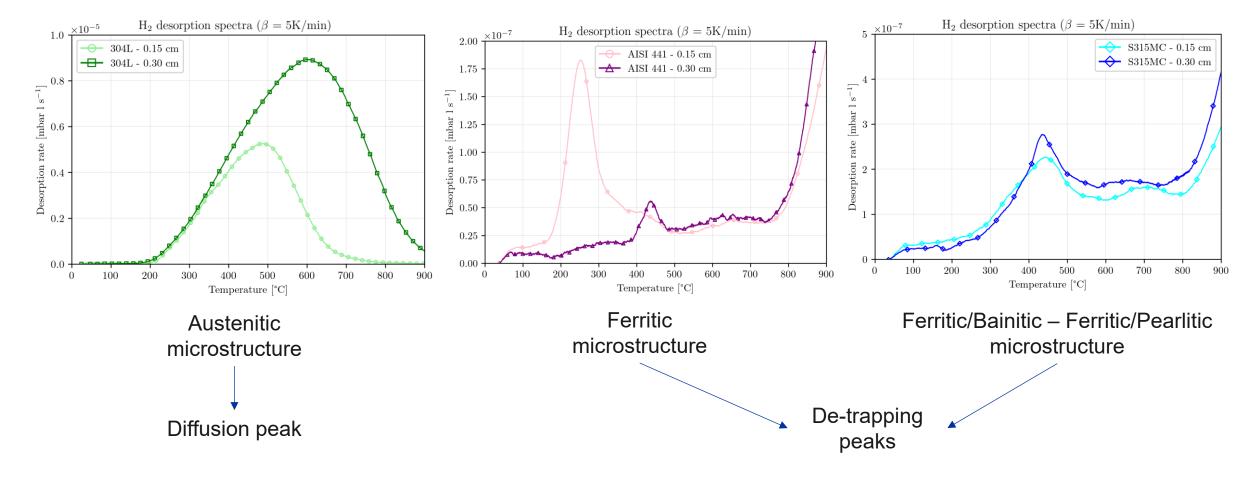
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H₂ thermal desorption: peaks origin

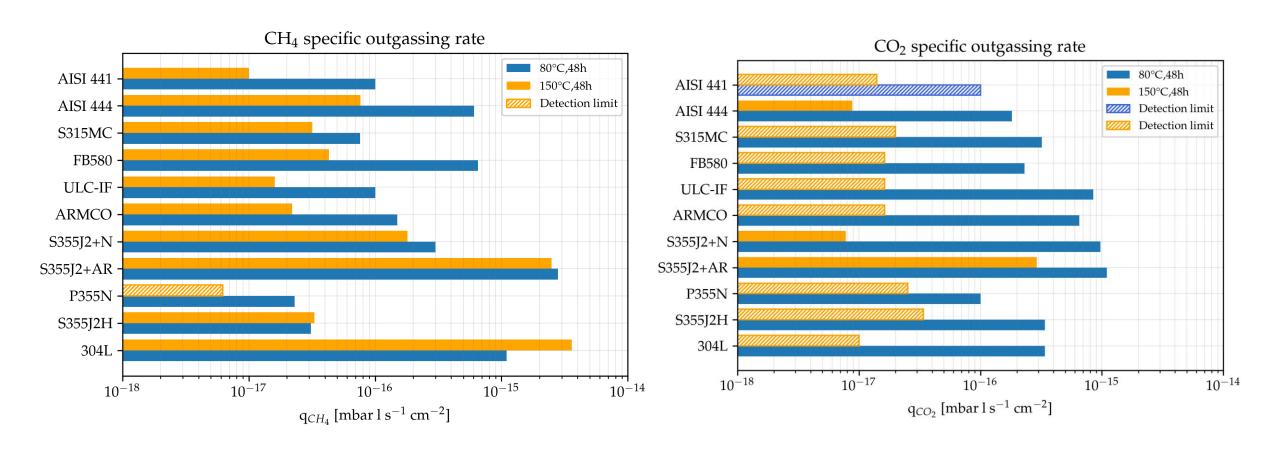
Temperature Programmed Desorption (TPD)





CH₄ & CO₂ outgassing rate after bakeout

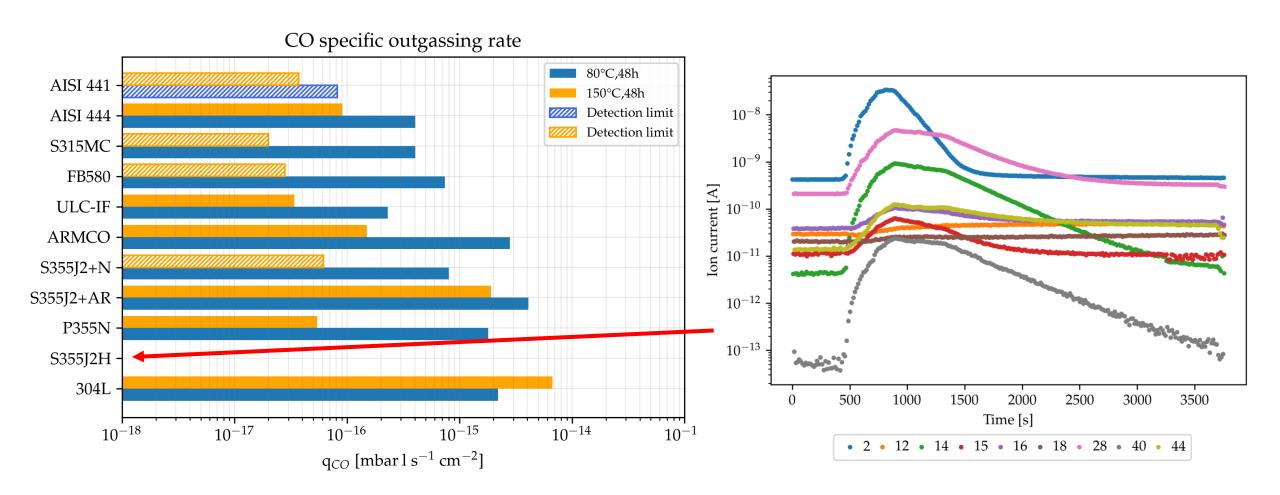
Throughput-accumulation method





CO outgassing rate after bakeout

Throughput-accumulation method





ET corrugated beampipe prototypes

System characteristics

DN400 - 2100 mm (t < 2 mm)

AISI 441 (FStS) **S315MC** (MS)

AISI 304L-VF* (AStS)

Pumping speeds scaled to ET dimensions

Objectives

Ultimate pressure after 80°C and 150°C bakeout

Verify water outgassing modelling

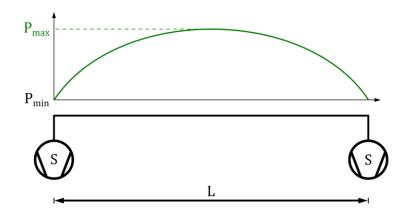
Test the effect of the increasing pumping speed during bakeout on ultimate pressure

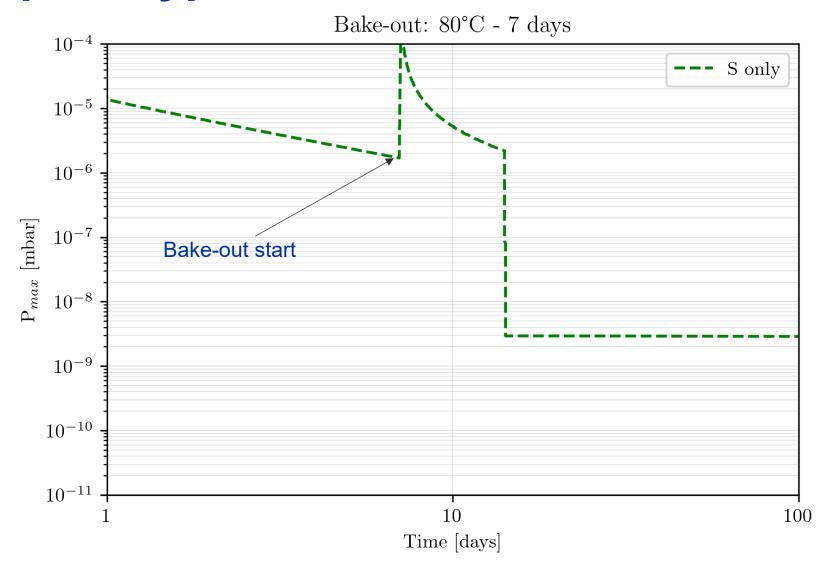


AISI 441 corrugated chamber.



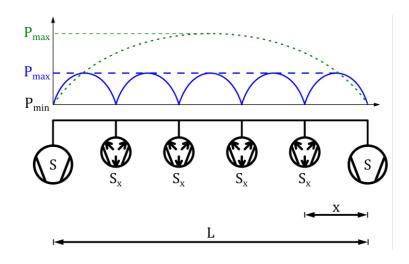
Modelling of effects of increased pumping speed during heating



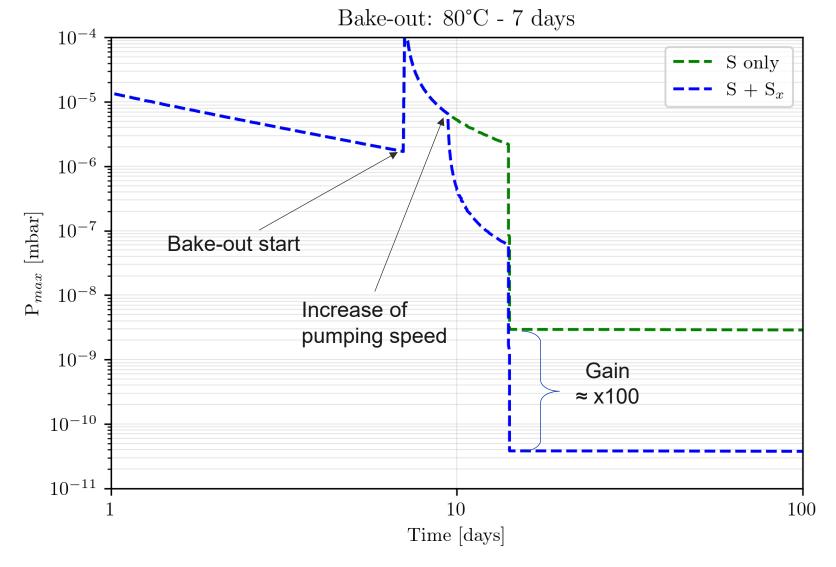




Modelling of effects of increased pumping speed during heating

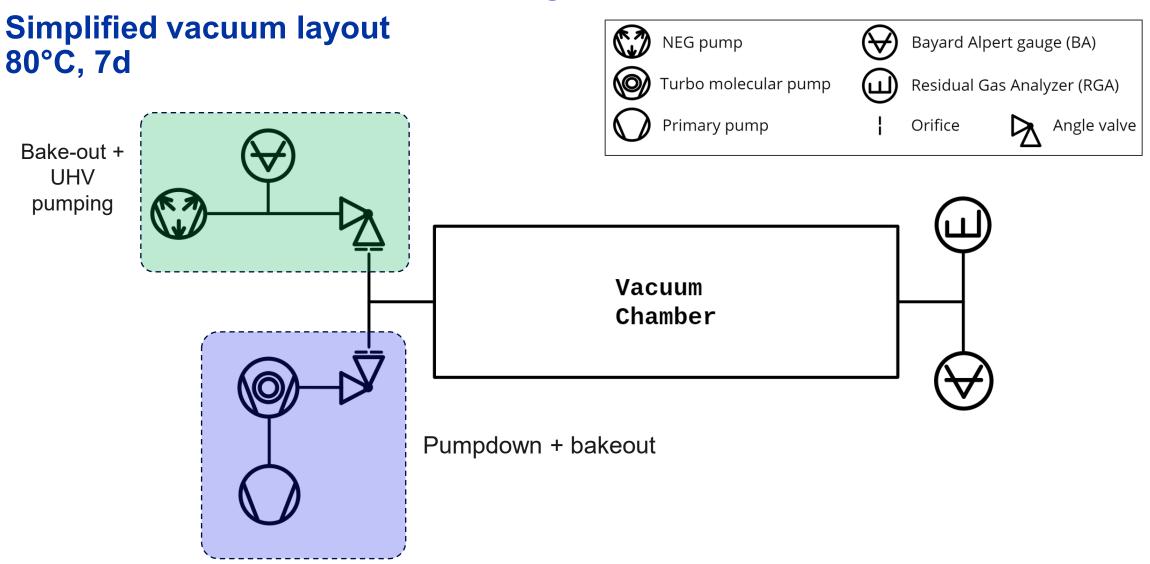


We could exploit the use of NEG pumps [SAES proposal, 2010]



Simulated layout: S = 1000 l/s, L = 2 km, $S_x = 1500 \text{ l/s}$ (H_2O), x = 50 m.

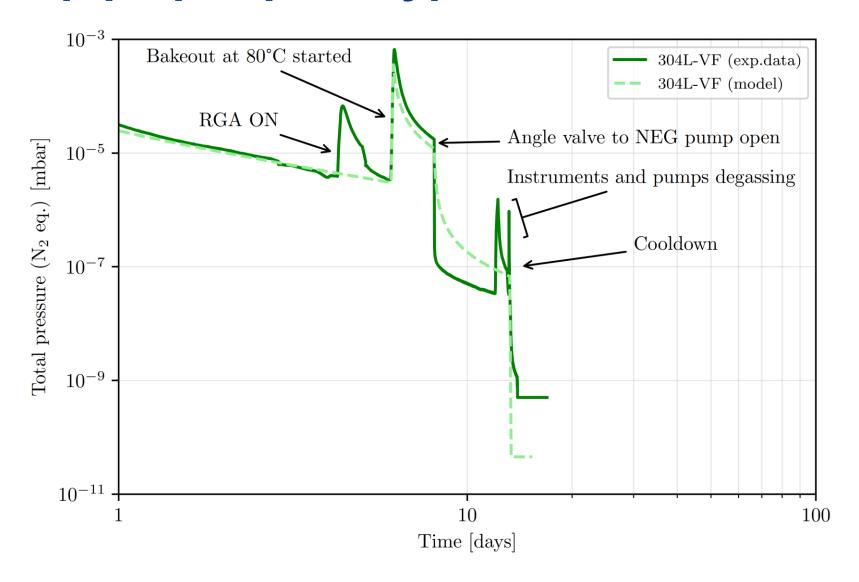




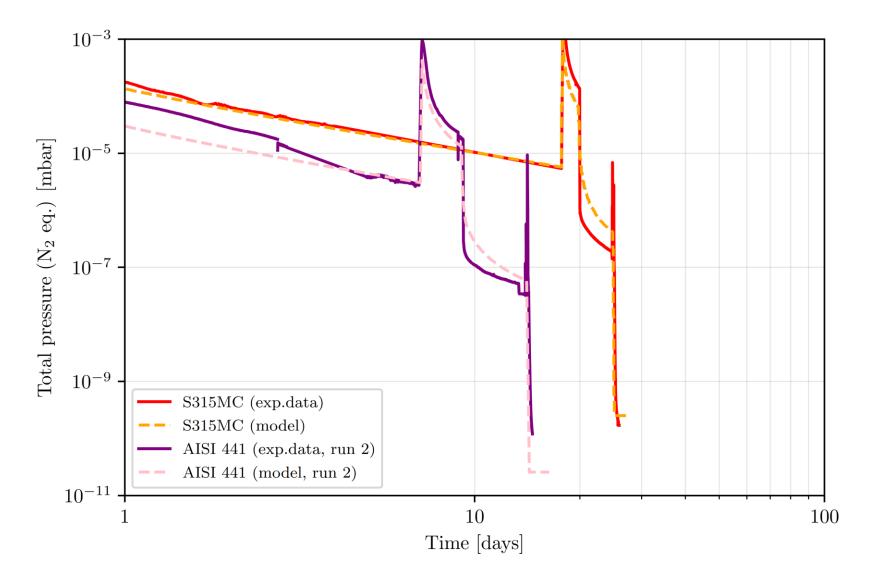


Simplified vacuum layout NEG pump Bayard Alpert gauge (BA) 150°C, 48h Turbo molecular pump Residual Gas Analyzer (RGA) Primary pump Orifice Angle valve Bake-out + **UHV** pumping **Vacuum** Chamber Pumpdown + bakeout



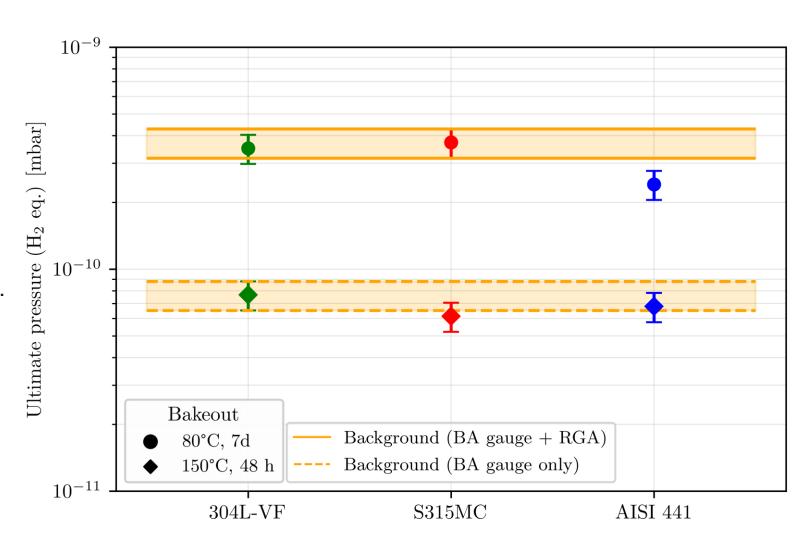








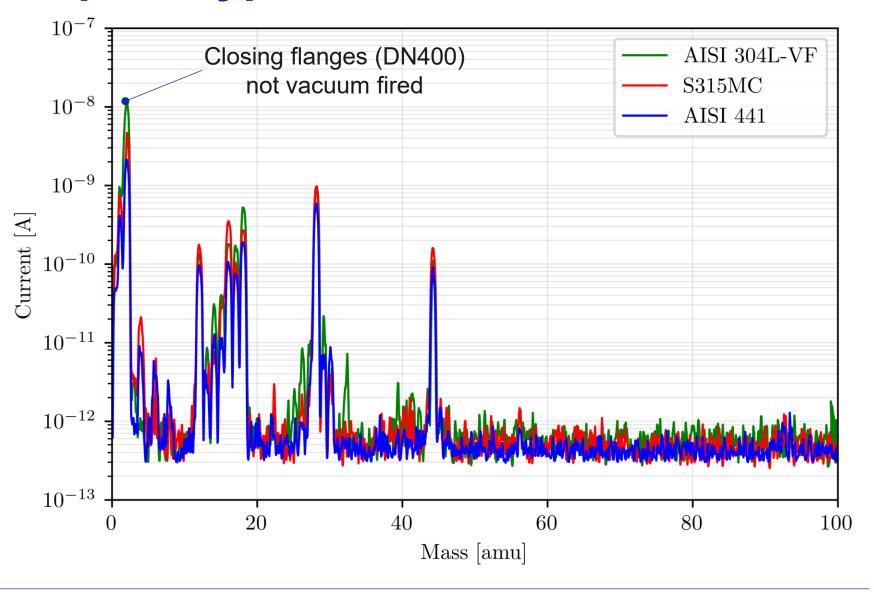
For all the chambers, **ultimate pressures are limited** by **instrumentation outgassing rates**.
(q_{H2} in line with sample measurements)



Orifice limited: 31 l s⁻¹ for H₂, beampipes area: 28800 cm², H₂ to N₂ conversion factor: 0.46, BA gauge error: 15%. BA: Bayard-Alpert, RGA: Residual Gas Analyser.



Gas composition measured at 20±3°C 48 h after the 80°C, 7 d bakeout.





Other prototypes







Implications of ferritic alloys on ET beampipe

Residual gas noise in GWD beam pipes

For a gas species, *i* [7]:

$$Strain(f) = \sum_{i} \sqrt{\frac{4 (2\pi \boldsymbol{\alpha_i})^2}{L^2 \boldsymbol{v_{0_i}}} \int_0^L \frac{\boldsymbol{\rho_i}(z)}{\omega(z)} e^{-\frac{2\pi f \omega(z)}{\boldsymbol{v_{0_i}}}} dz}$$

Where:

 $\rho(z)$ = gas density \Leftrightarrow pressure distribution

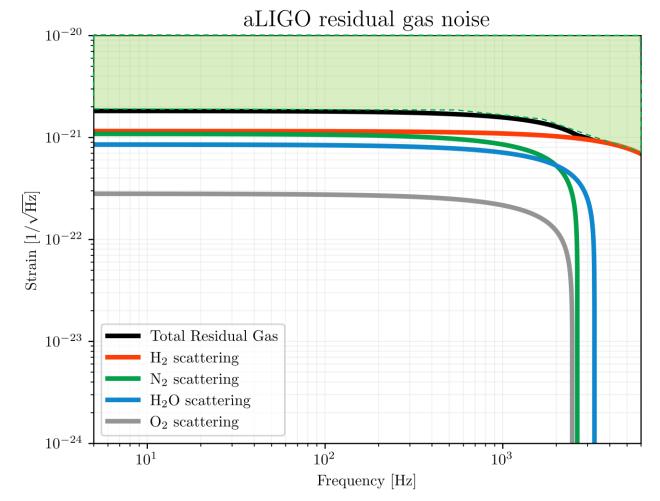
 α = gas optical polarizability

 v_0 = most probable thermal speed

f = frequency

 $\omega(z)$ = laser beam gaussian radius

L = interferometer arm length

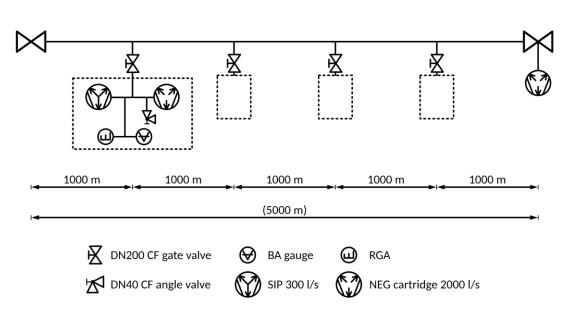


[7] M. E. Zucker and S. E. Whitcomb. "Measurement of optical path fluctuations due to residual gas in the LIGO 40-meter interferometer" (1994)



Implications of ferritic alloys on ET beampipe

Gas	Specific outgassing rate [mbar I s ⁻¹ cm ⁻²]	
	AISI 441 ^[8] (150°C, 48 h)	AISI 304L air baked ^[1] [<u>454°C, 36 h]</u> (<u>160°C, 504 h)</u>
H ₂	<1×10 ⁻¹⁵	6.1×10 ⁻¹⁴
CH ₄	<1×10 ⁻¹⁷	<1.3×10 ⁻¹⁸
H ₂ O	3×10 ⁻¹⁷ Simulated w/ isotherms for 150°C, 168 h ^[9]	<1.1×10 ⁻¹⁸
СО	<3.7×10 ⁻¹⁷	<2.7×10 ⁻¹⁸
CO ₂	<1.4×10 ⁻¹⁷	<1.13×10 ⁻¹⁵



UHV layout proposed for ET^[9]

^[9] Scarcia C., The Einstein Telescope beam pipe vacuum system: Exploring novel techniques and materials for a cost-effective design solution, Doctoral thesis, 2025

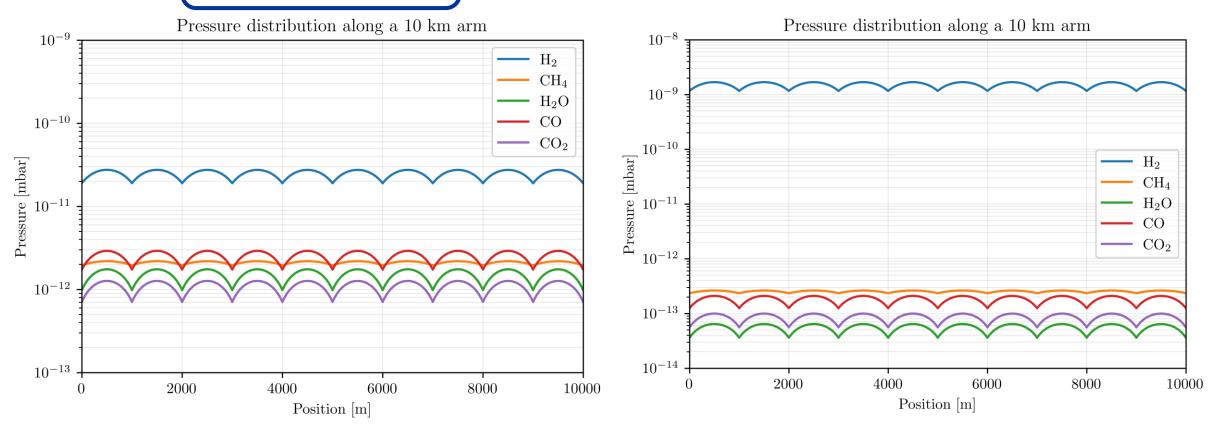


^[8] Scarcia C. et al., J. Vac. Sci. Technol. B 43, 044203 (2025)

Implications of ferritic alloys on ET beampipe

AISI 441^[8]
(150°C, 48 h)
Water contribution simulated after 150°C, 168 h bakeout

AISI 304L air baked^[1]
[454°C, 36 h]
(160°C, 504 h)

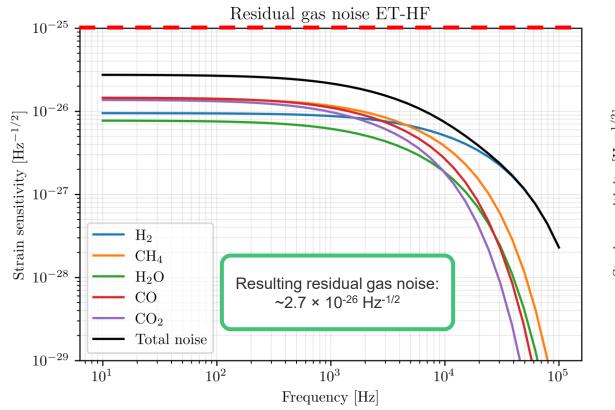


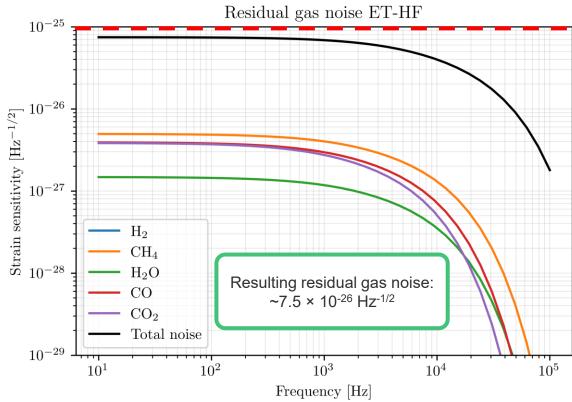


Implications of ferritic alloys on GWD beampipe

AISI 441^[8]
(150°C, 48 h)
Water contribution simulated after 150°C, 168 h bakeout

AISI 304L air baked^[1]
[454°C, 36 h]
(160°C, 504 h)





■ Upper limit from requirements: < 1 × 10⁻²⁵ Hz^{-1/2} (ET-0385A-24)

Optical parameters taken from ET-0385A-24. Gas parameters taken from [9]



Conclusions/Considerations

Studies of vacuum properties

- Mild steels and ferritic stainless steels are compatible with UHV applications.
 As received, their <u>bulk</u> H₂ outgassing rates outperform the austenitic stainless steels currently used in GWDs by several orders of magnitude.
 - Outgassing rates of ferritic alloys are more susceptible to surface quality rather than bulk.
- Isothermal models were developed to predict the water outgassing rate after bakeout.
- General agreement of results across the labs

Prototyping

Models and materials experimentally validated with GWD sector-scaled mockups.

Beam pipe vacuum system impact

- The experimental data and models developed can be applied to design the full vacuum commissioning sequence.
- Considering strict boundary conditions and applying a rationalised set of design criteria, and the new beam pipe materials
 we could potentially save up to 80% of the forecasted budget for the vacuum system's hardware
 (excluding beam pipe production).



A friend once said...

"The issue is not can one make the requirement but rather can one make them at significantly reduced cost."

Rainer Weiss, Beampipes for Gravitational Wave Telescopes 2023, CERN, 27-29 March 2023

Well Rai, we are getting there!



Thanks for your attention



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